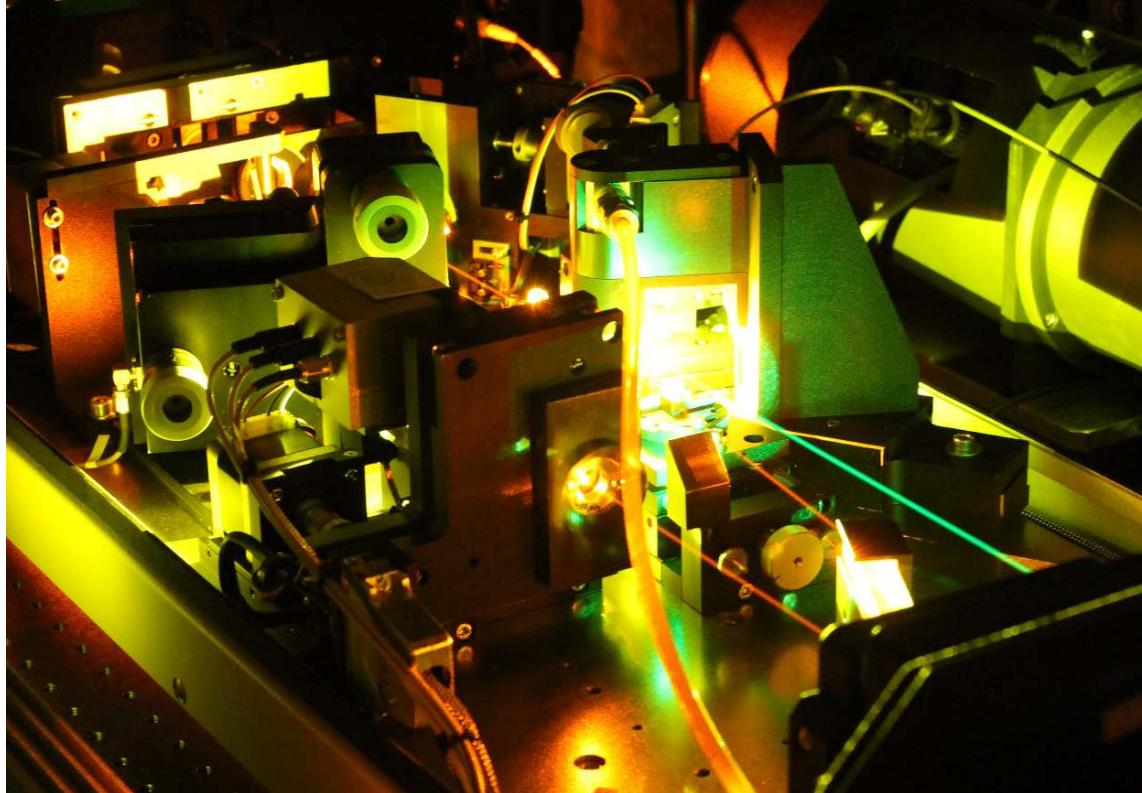


Recent laser spectroscopy results at NSCL/MSU

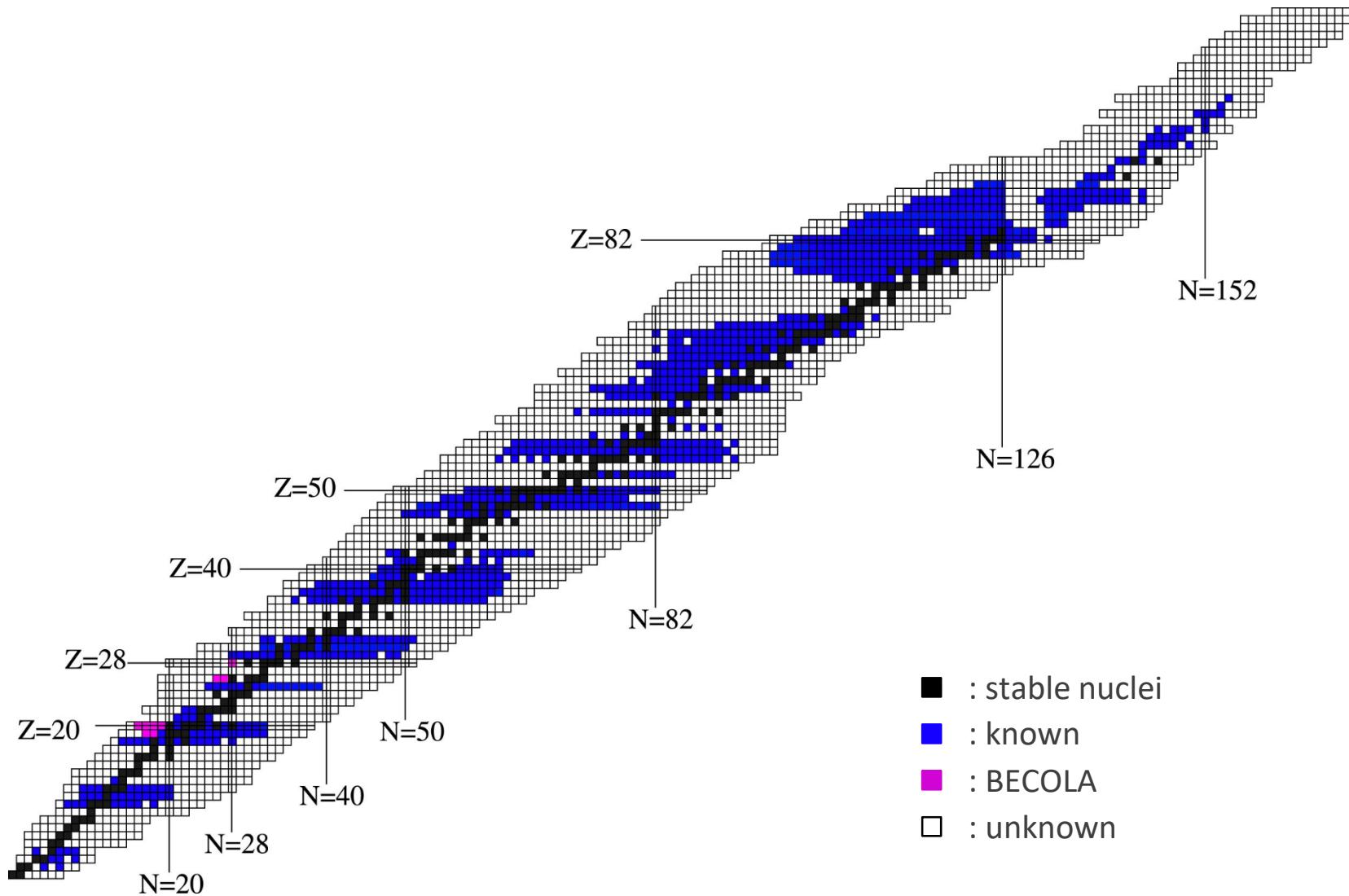


Kei Minamisono

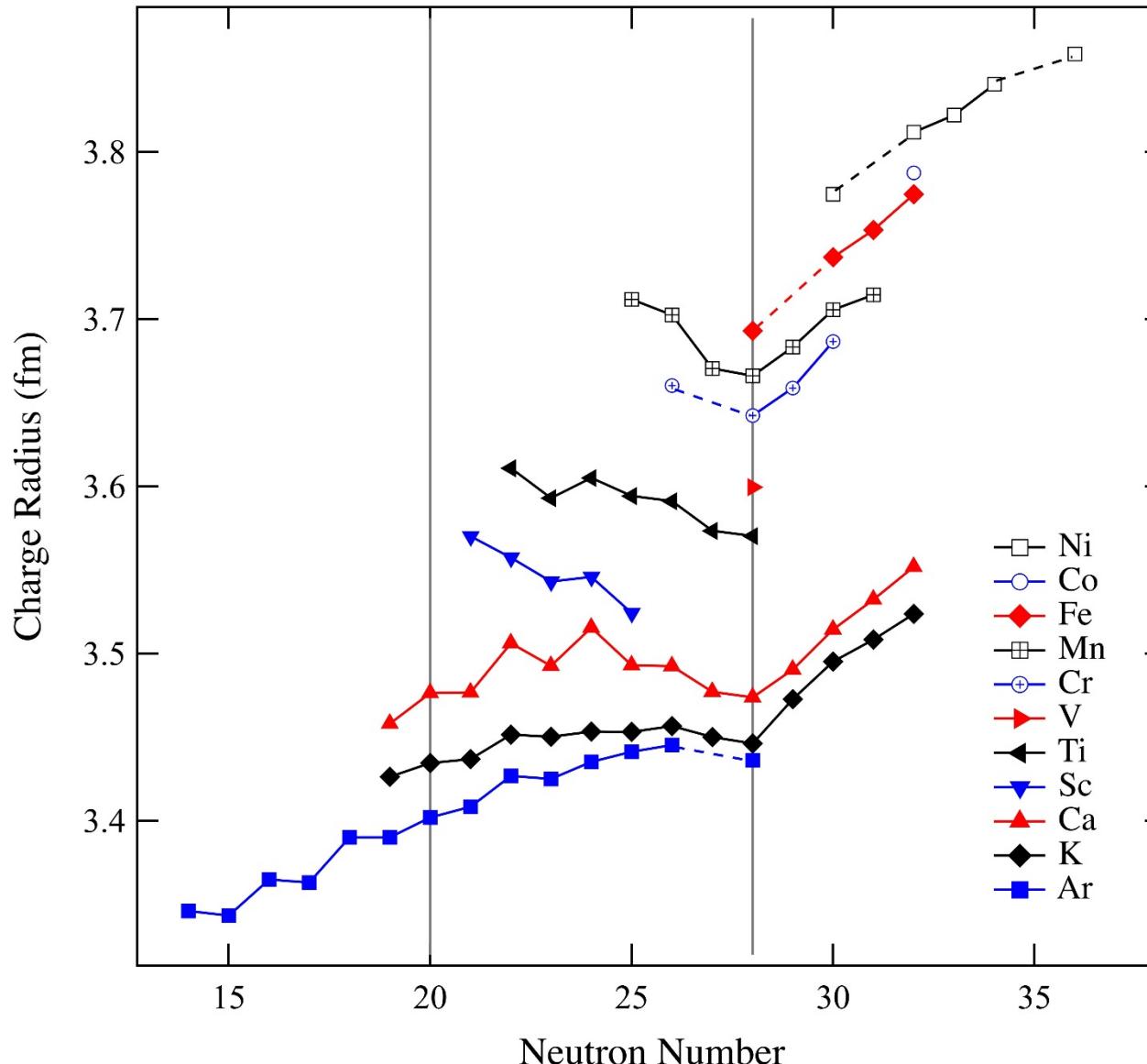
Plan of my talk

- Very brief introduction
- NSCL, and its transition to FRIB
- About BECOLA
- Some science cases
- Future prospects

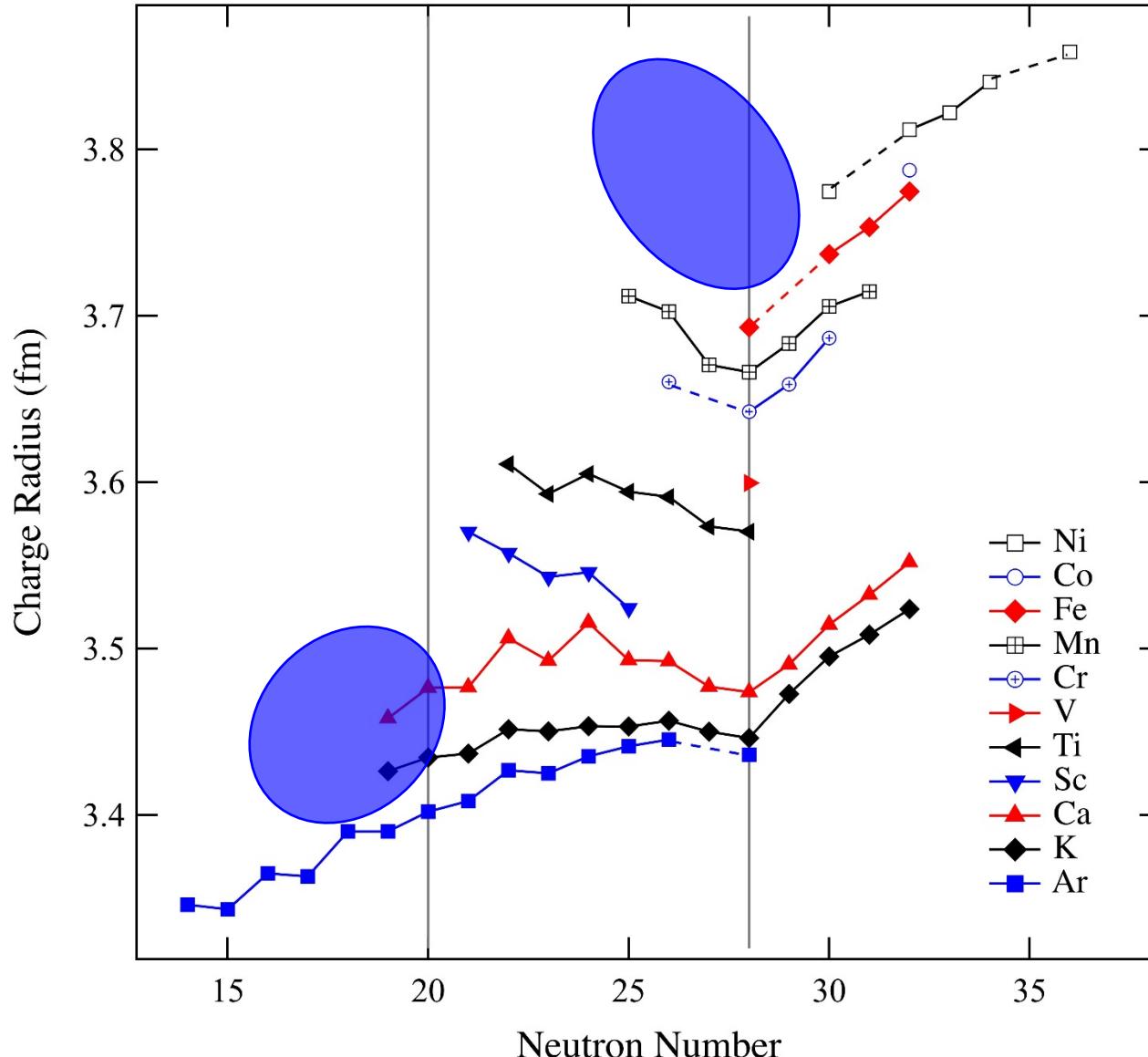
Laser spectroscopy experiments at BECOLA



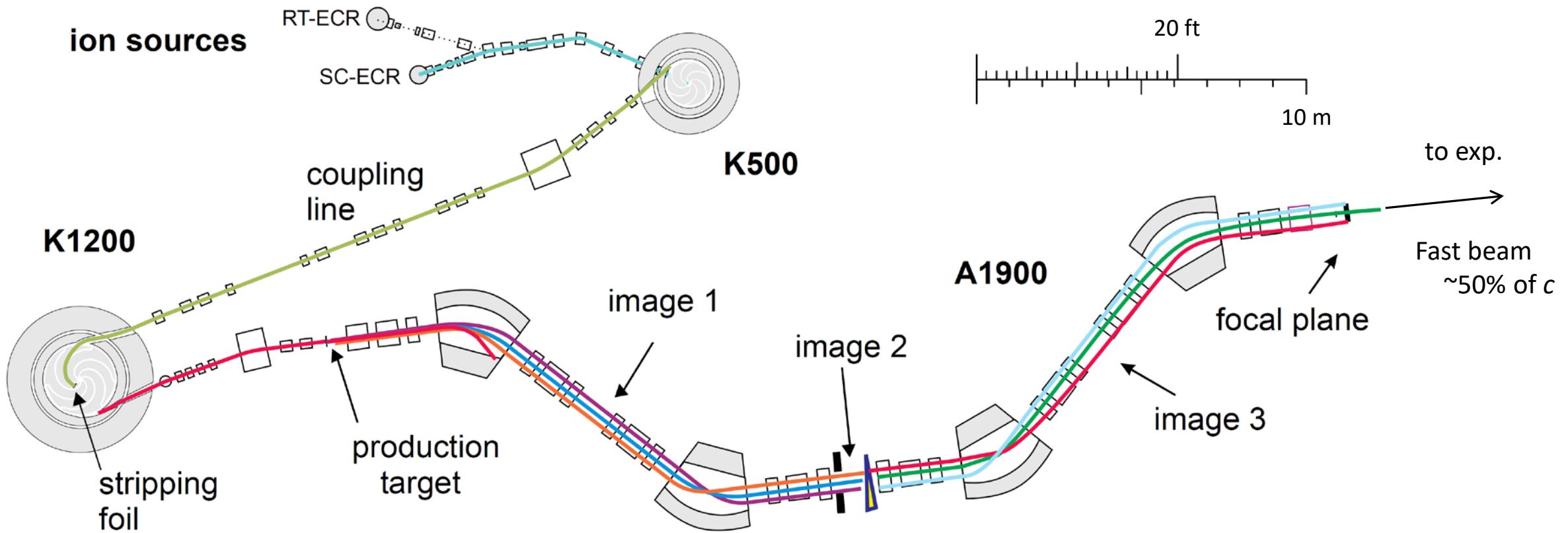
Recent interests: Ca & Ni at $N = 20$ & 28



Recent interests: Ca & Ni at $N = 20$ & 28



Coupled cyclotron facility at NSCL/MSU



Projectile-fragment reactions

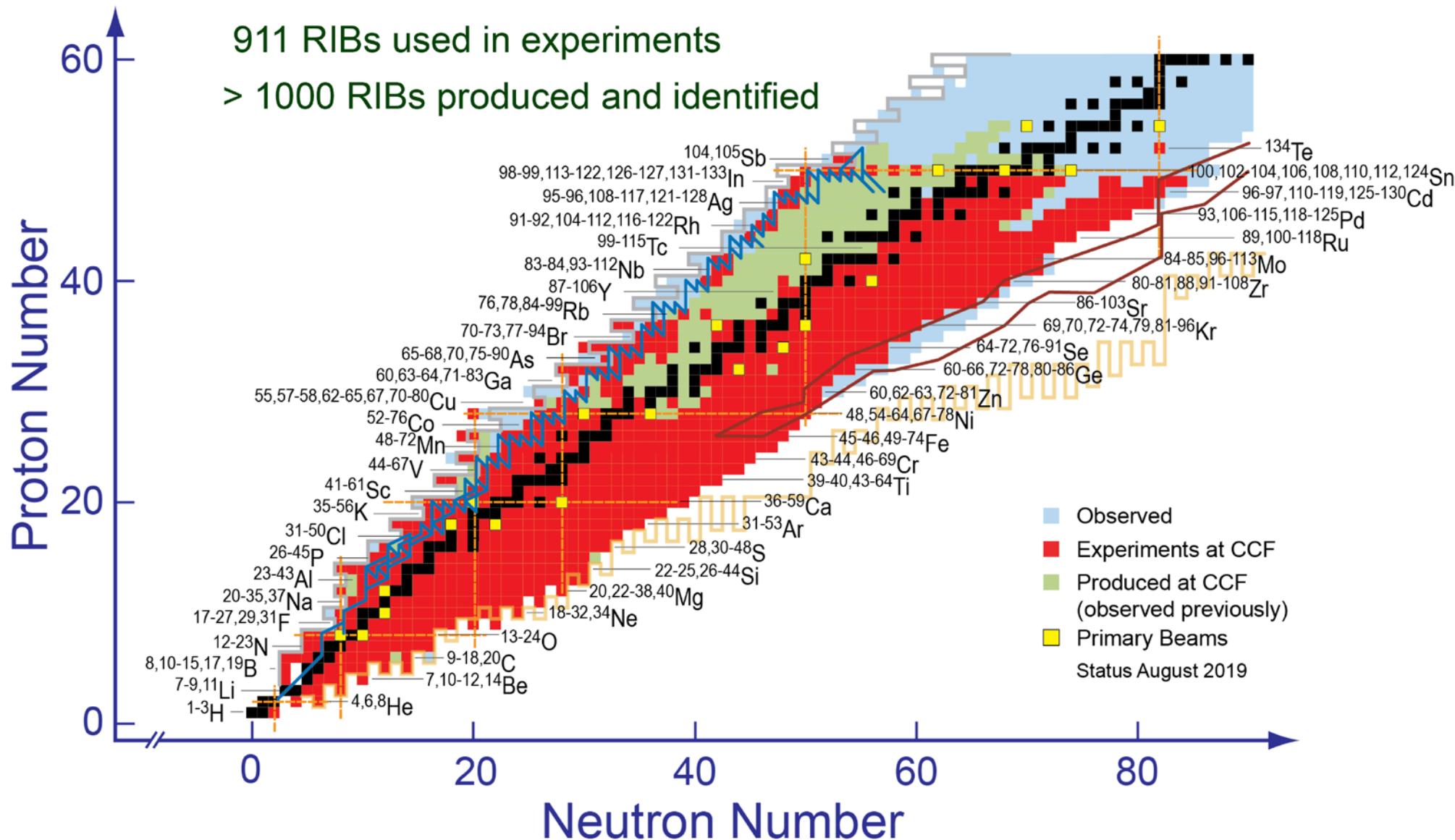
- Forward focusing, fast separation
- Produces nuclei lighter than primary beam (so far the heaviest is U)
- Chemistry free
- Complementary to e.g. ISOL mechanism

Good at

- Short lived isotopes
- Neutron-deficient side

Production at NSCL

911 RIBs used in experiments
> 1000 RIBs produced and identified

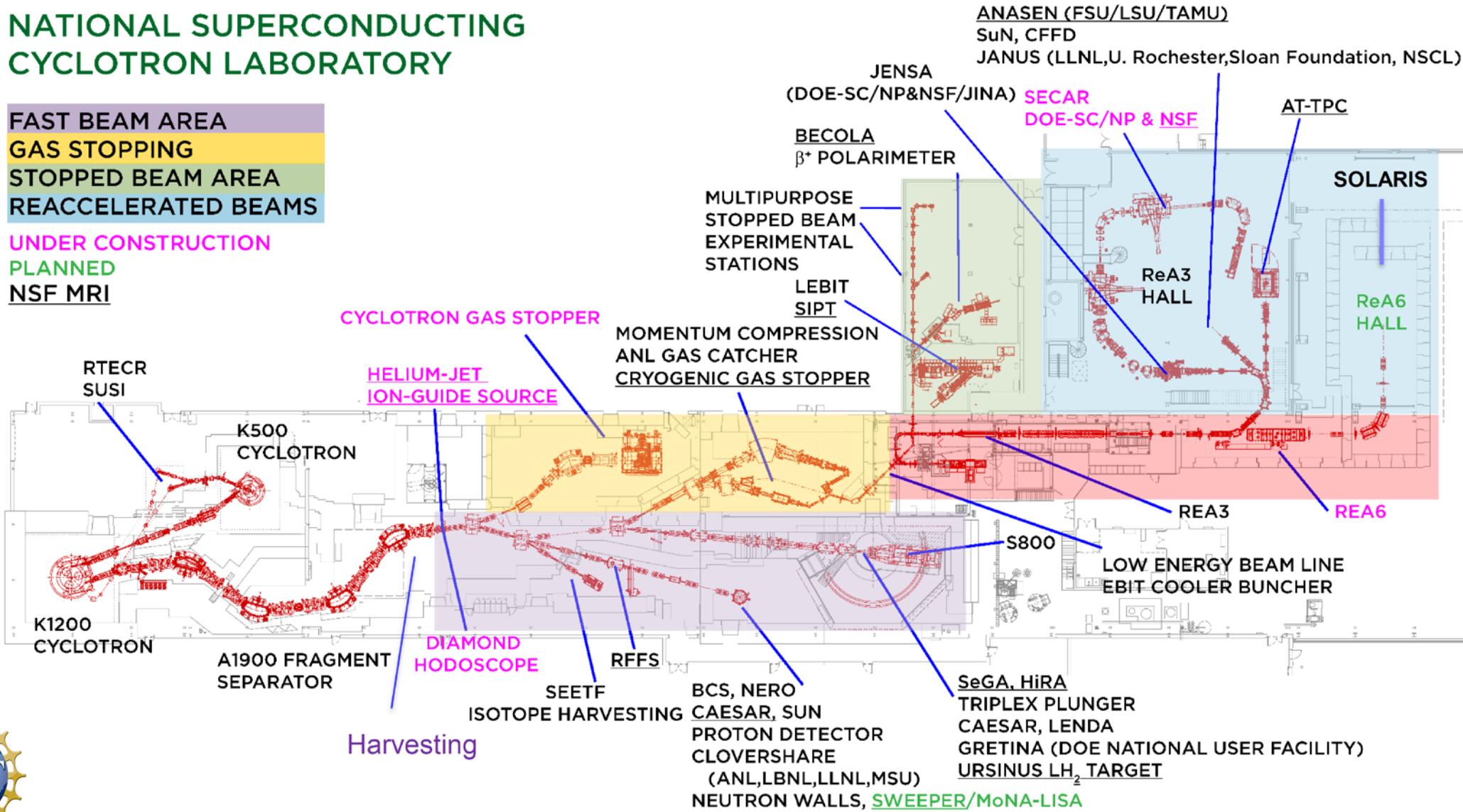


Current NSCL experimental layout

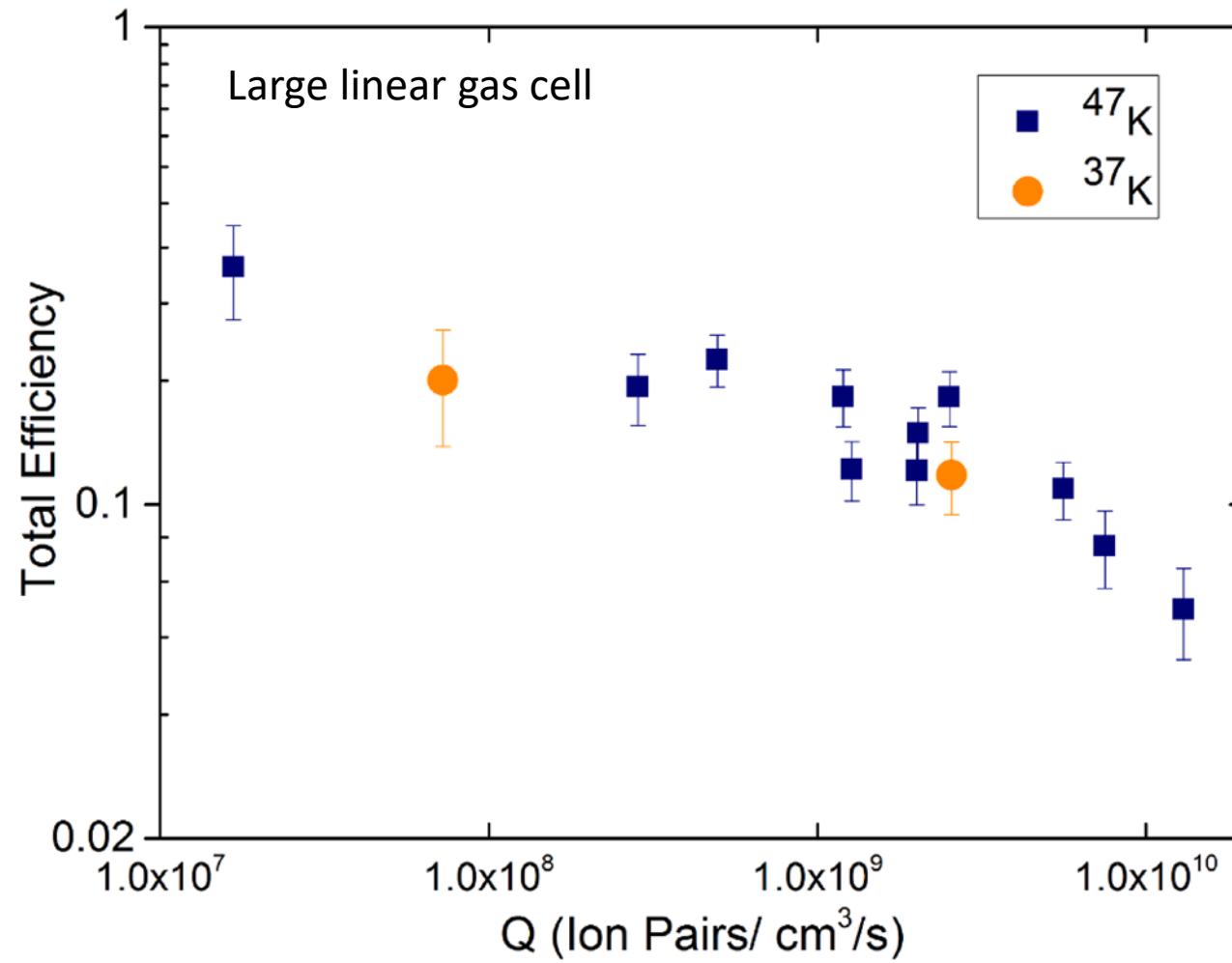
NATIONAL SUPERCONDUCTING CYCLOTRON LABORATORY

FAST BEAM AREA
GAS STOPPING
STOPPED BEAM AREA
REACCELERATED BEAMS

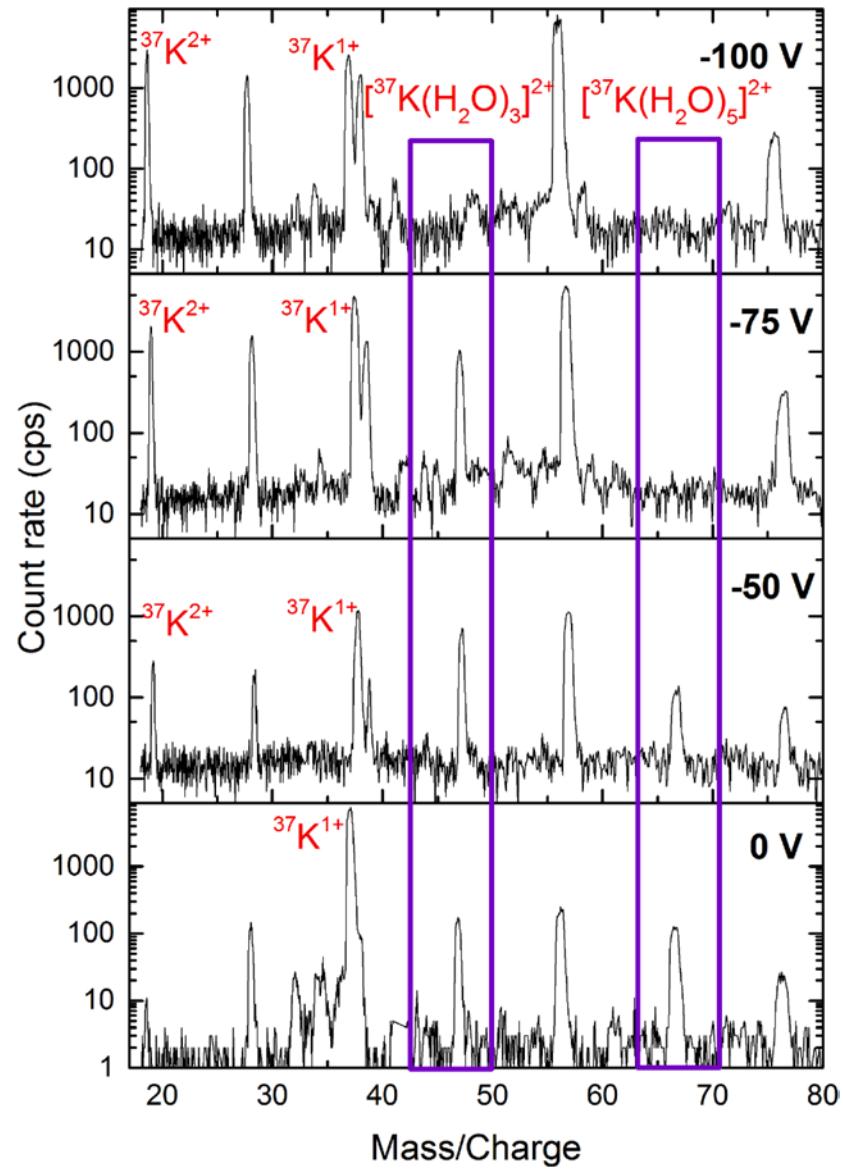
UNDER CONSTRUCTION
PLANNED
NSF MRI



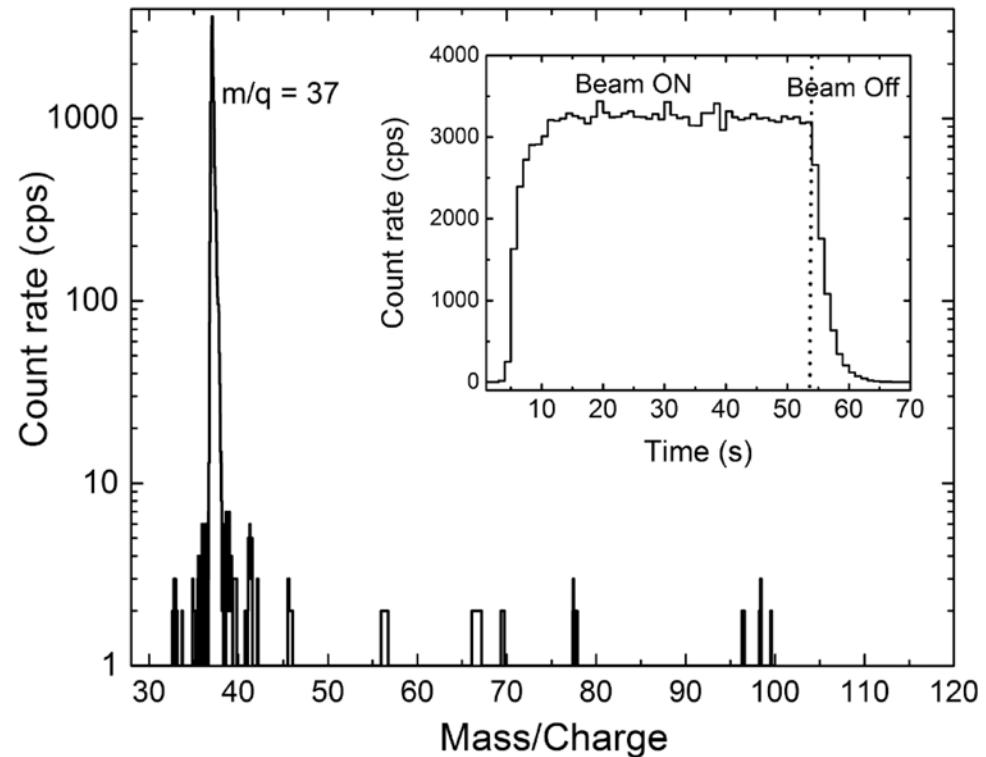
Gas stopping at NSCL



Gas stopping at NSCL

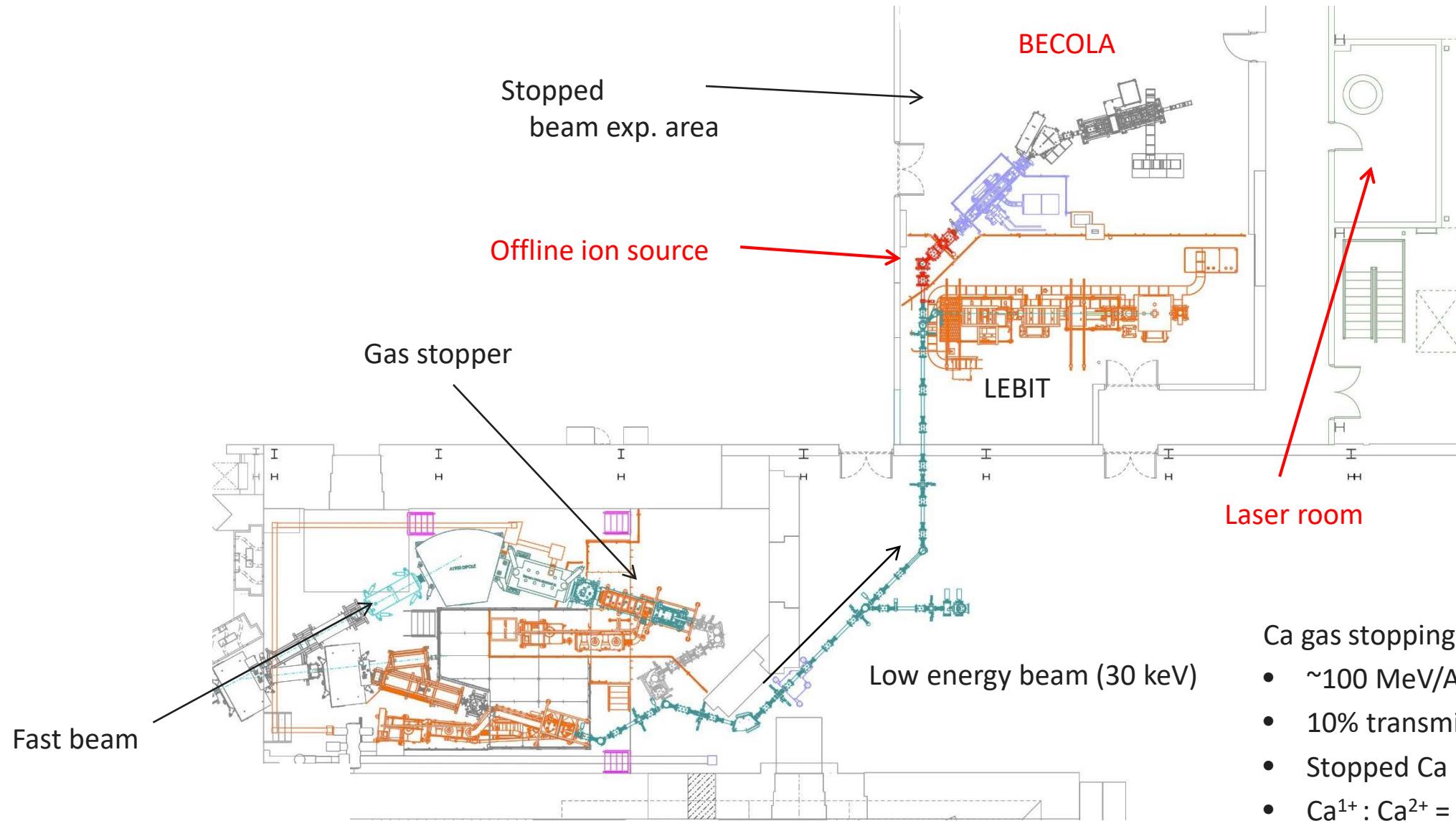


With “cleaner” cell condition



When the cell is “dirty”.

Gas stopping & ion beam transport to BECOLA



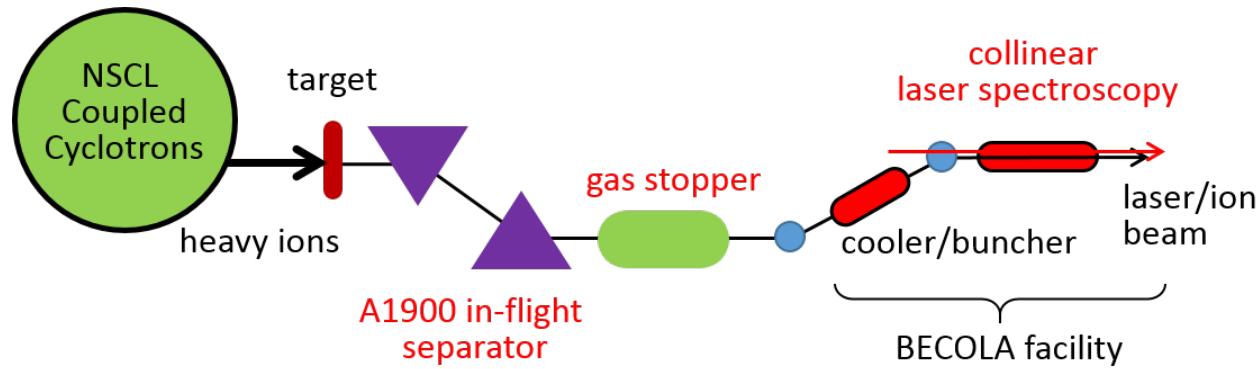
Ca gas stopping

- ~100 MeV/A injection
- 10% transmission efficiency
- Stopped Ca nH₂O attached
- Ca¹⁺ : Ca²⁺ = 1 : 1 after CID

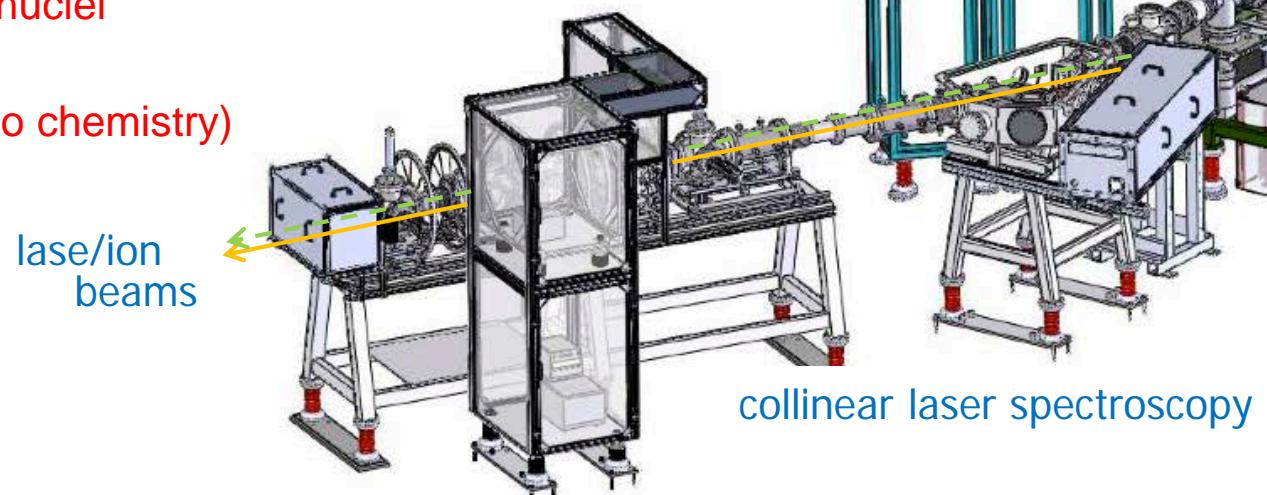
BECOLA facility @ NSCL/MSU

- Bunched beam collinear laser spectroscopy -

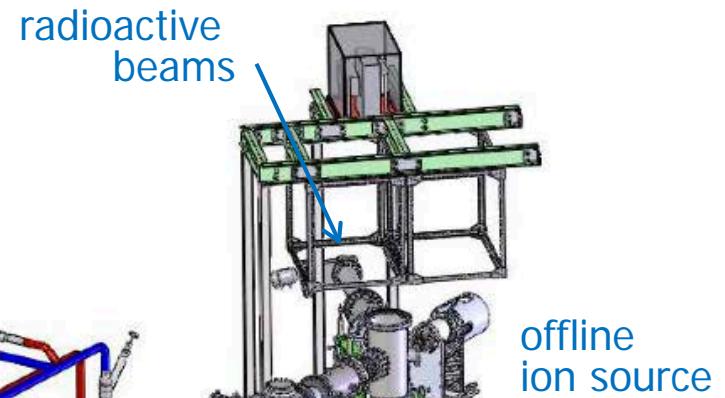
The production scheme complements existing capabilities of e.g. ISOL facilities.



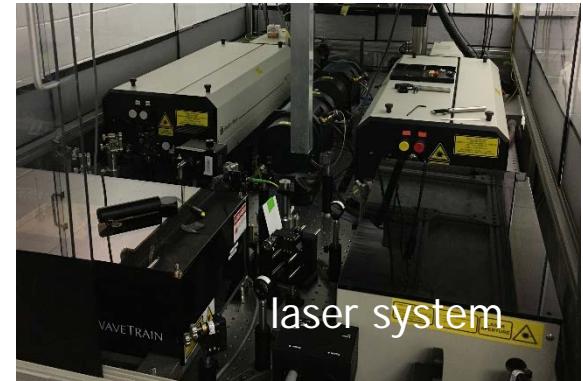
- Shorter lived isotopes (fast separation)
- Neutron-deficient nuclei
- Transition metals elements (no chemistry)



collinear laser spectroscopy



radioactive beams
offline ion source

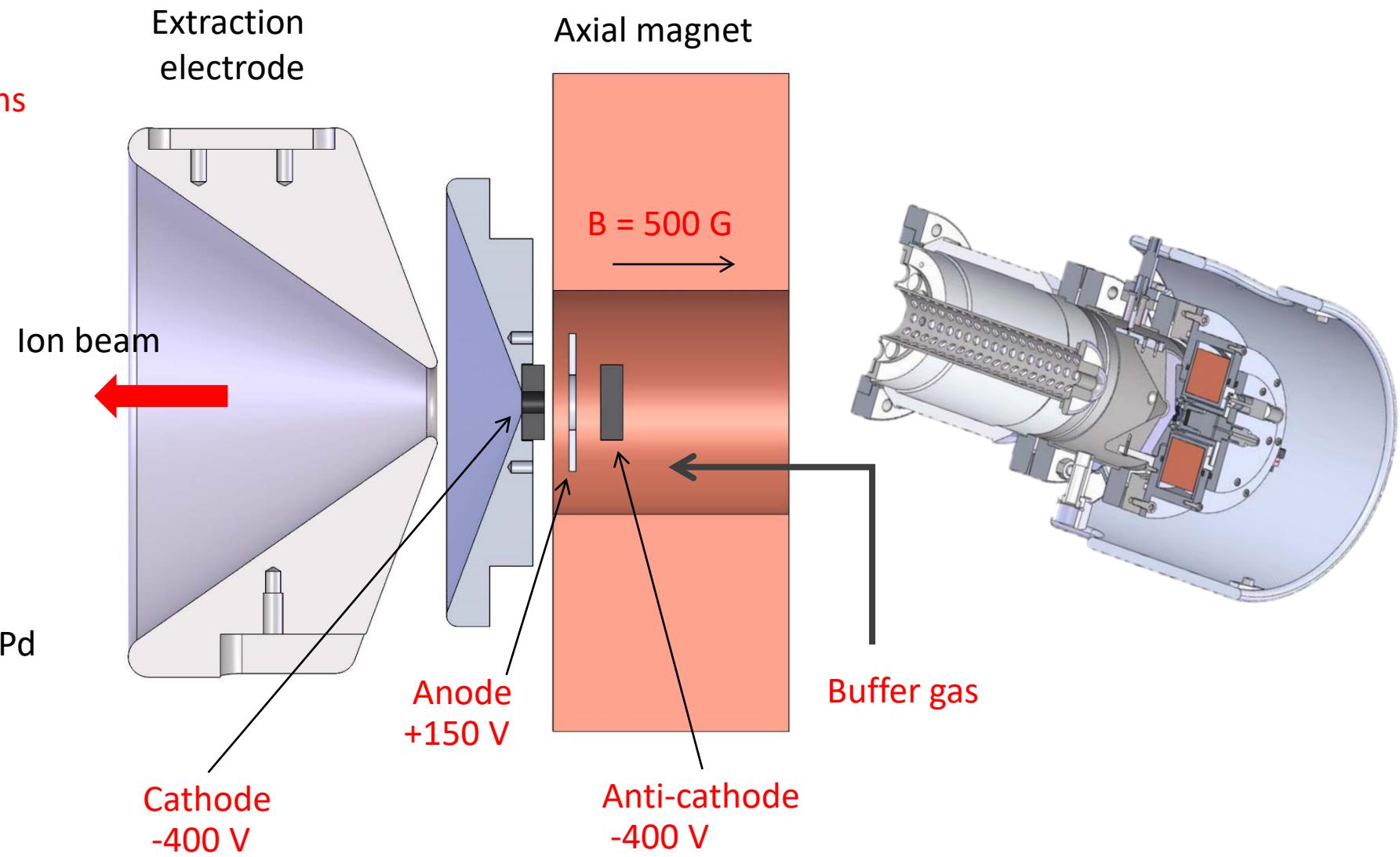


laser system

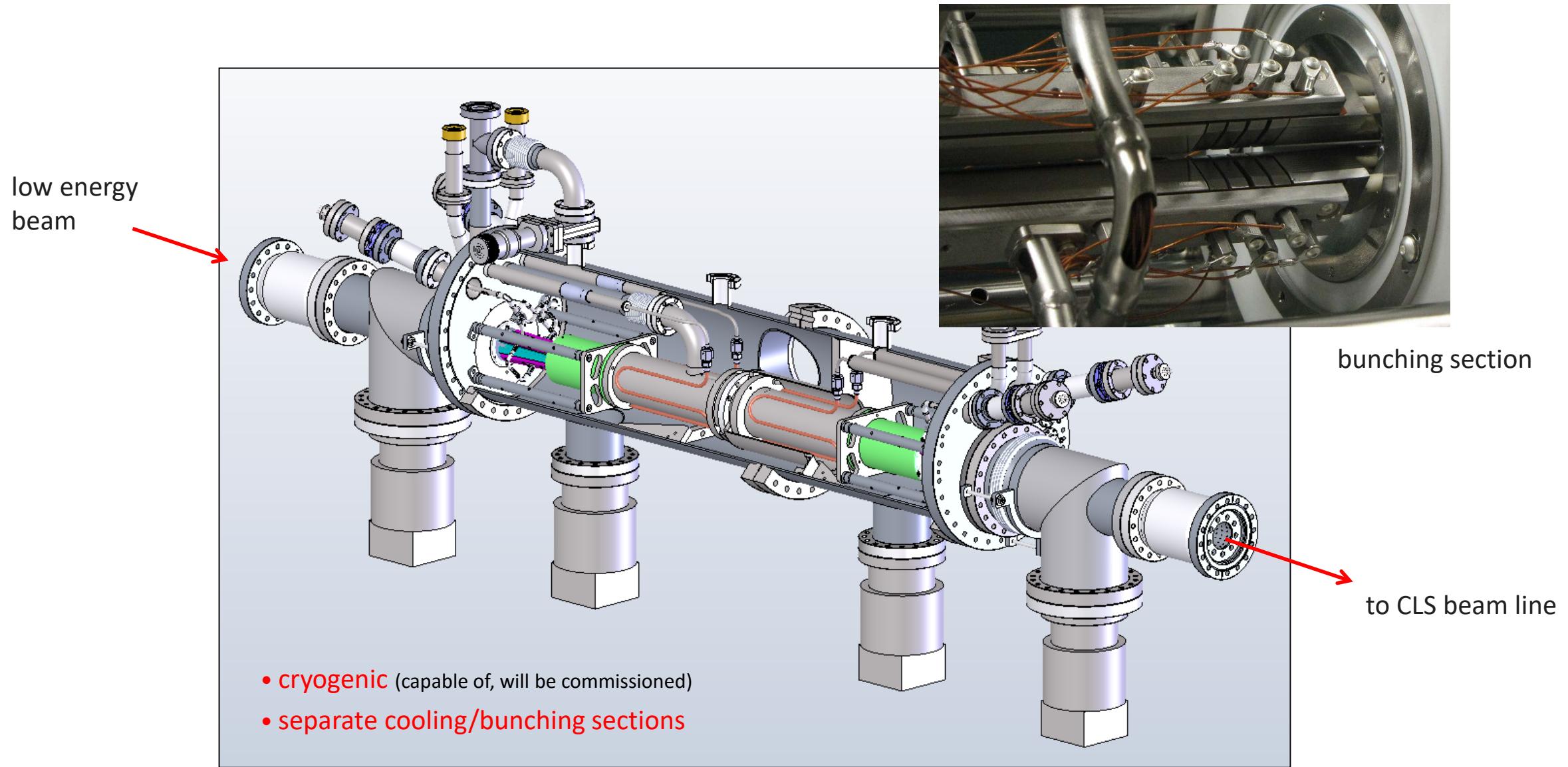
Penning ionization gauge (PIG) ion source

Critical for offline testing and online calibrations

- Plasma/discharge sputter source
- generates ions from cathodes buffer gas
- He, Ne, Ar, Ca, Sc, Mn, Fe, Ni, Zr, Sn, Pd

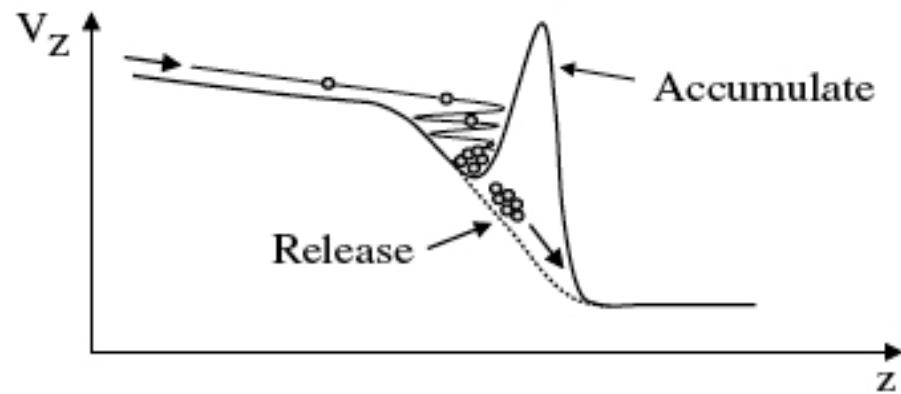
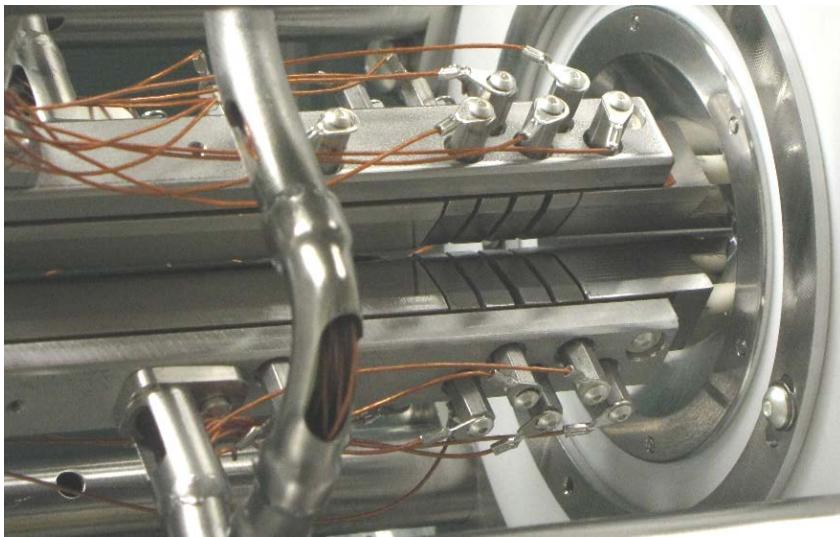


BECOLA cryogenic cooler/buncher

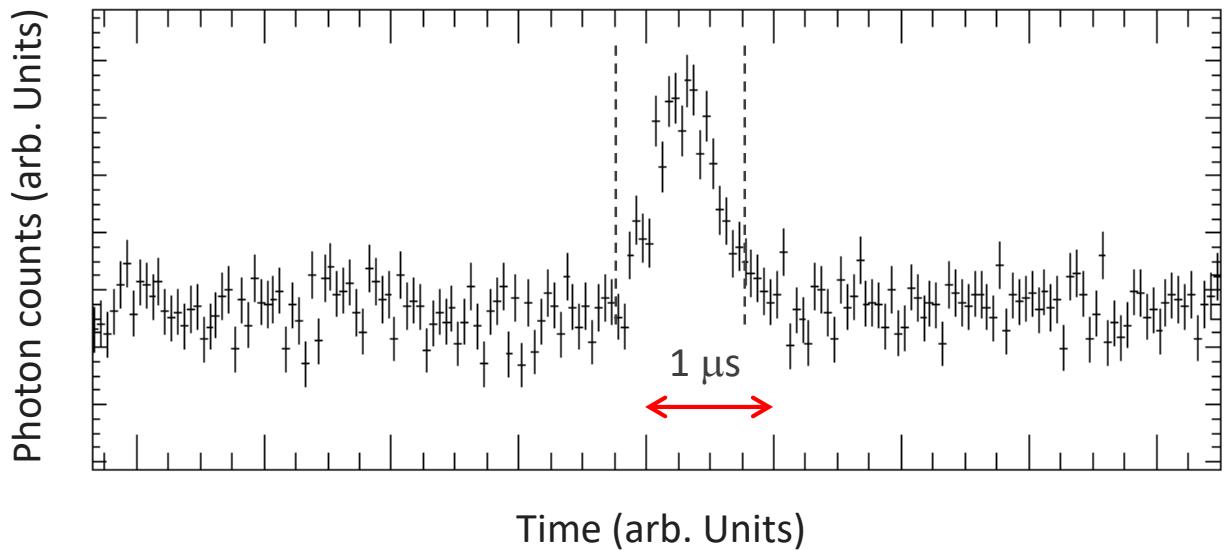


High sensitivity: bunched beam CLS

RFQ bunching section



Time spectrum of 30 keV ^{36}Ca ($T_{1/2} = 102$ ms)



Bunched beam CLS

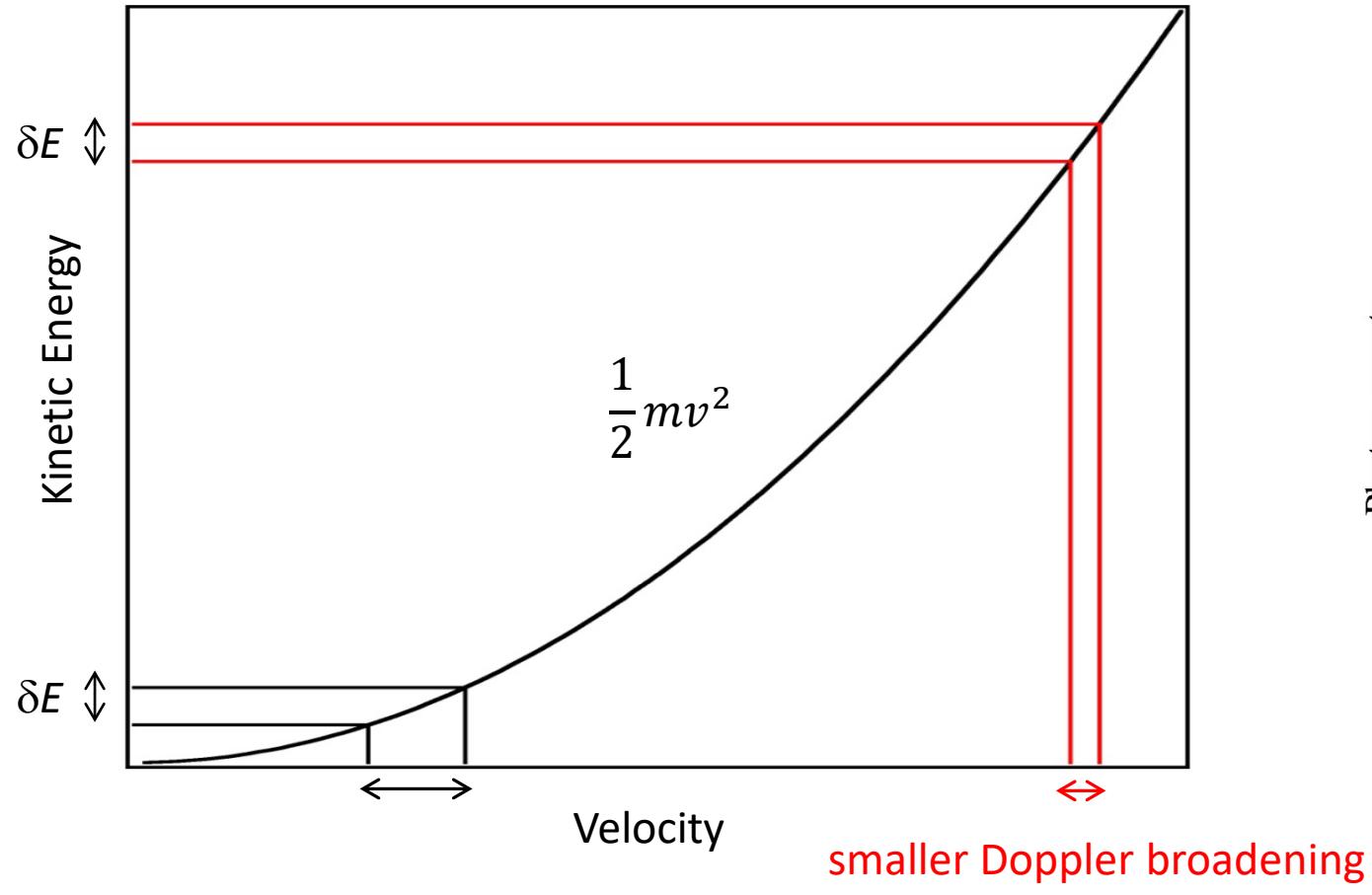
- Gate on ion-beam bunches
- No loss of signal, but suppresses background → high SN
- Suppression factor $\sim 10^6$ for 1 s bunch cycle
- Lifetime consideration for short-lived isotopes

B. R. Barquest et al., NIMA 866, 18 (2017);

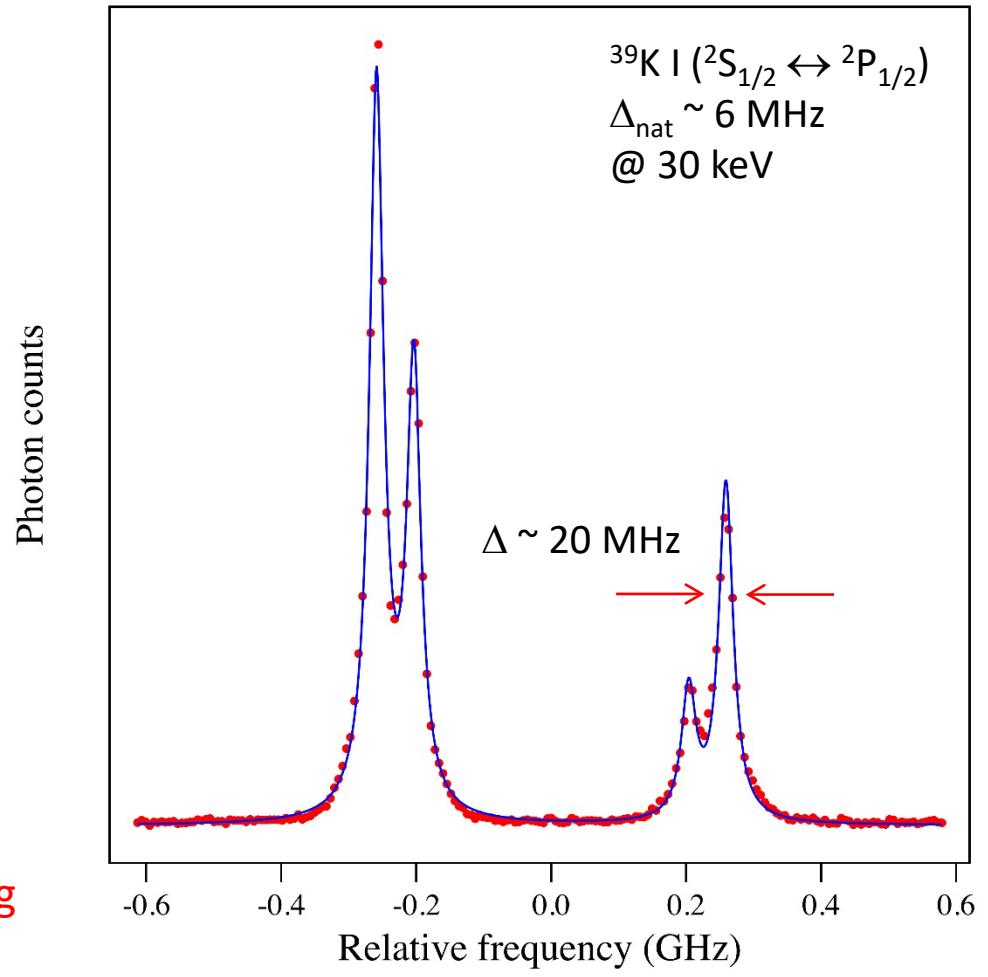
P. Campbell et al., PRL 89, 082501 (2002);

A. Nieminen et al., PRL 88, 094801 (2002).

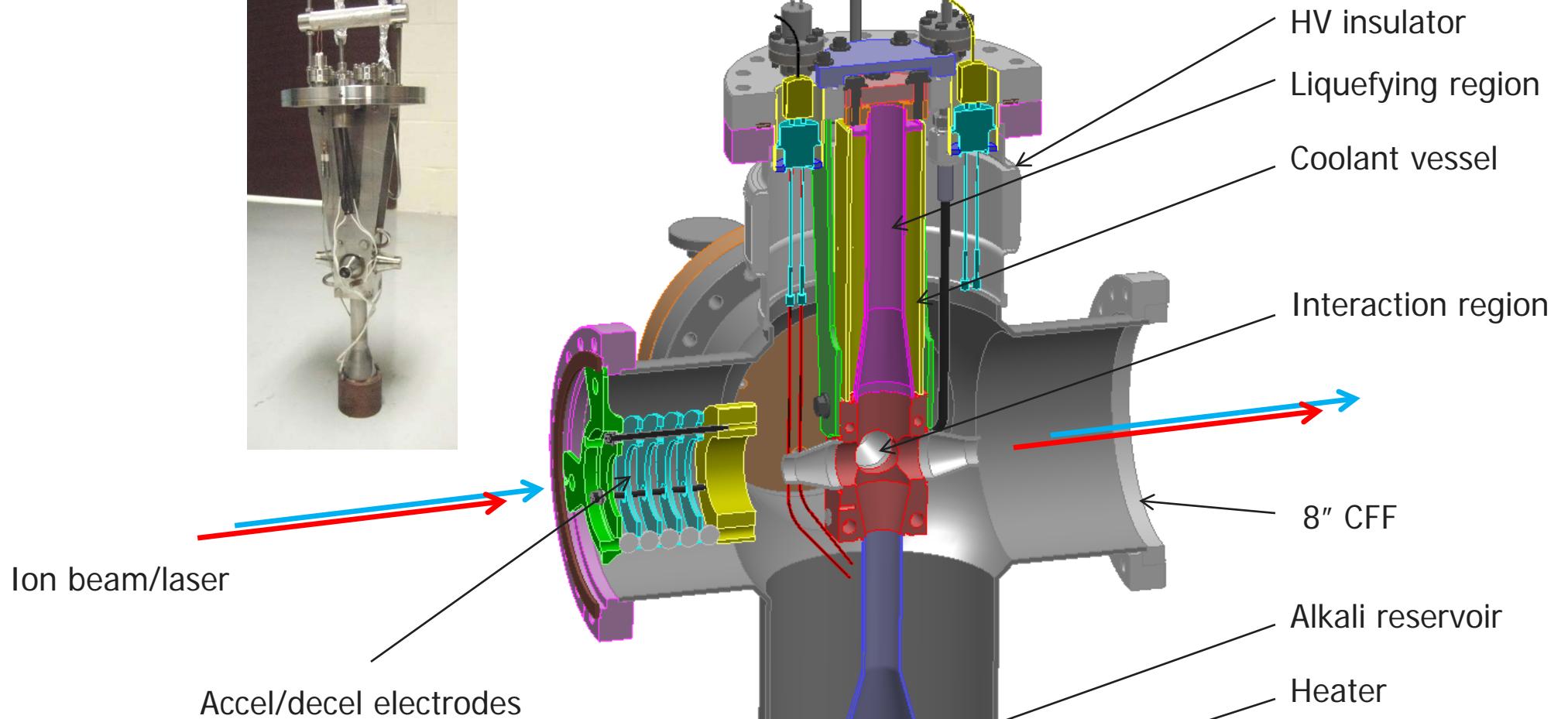
High resolution: kinematical compression (velocity bunching)



smaller Doppler broadening



CEC



A. Klose, K.M. et al., NIMA678, 114 (2012);

A. Klose, K.M. et al., PRA 88, 042701 (2013);

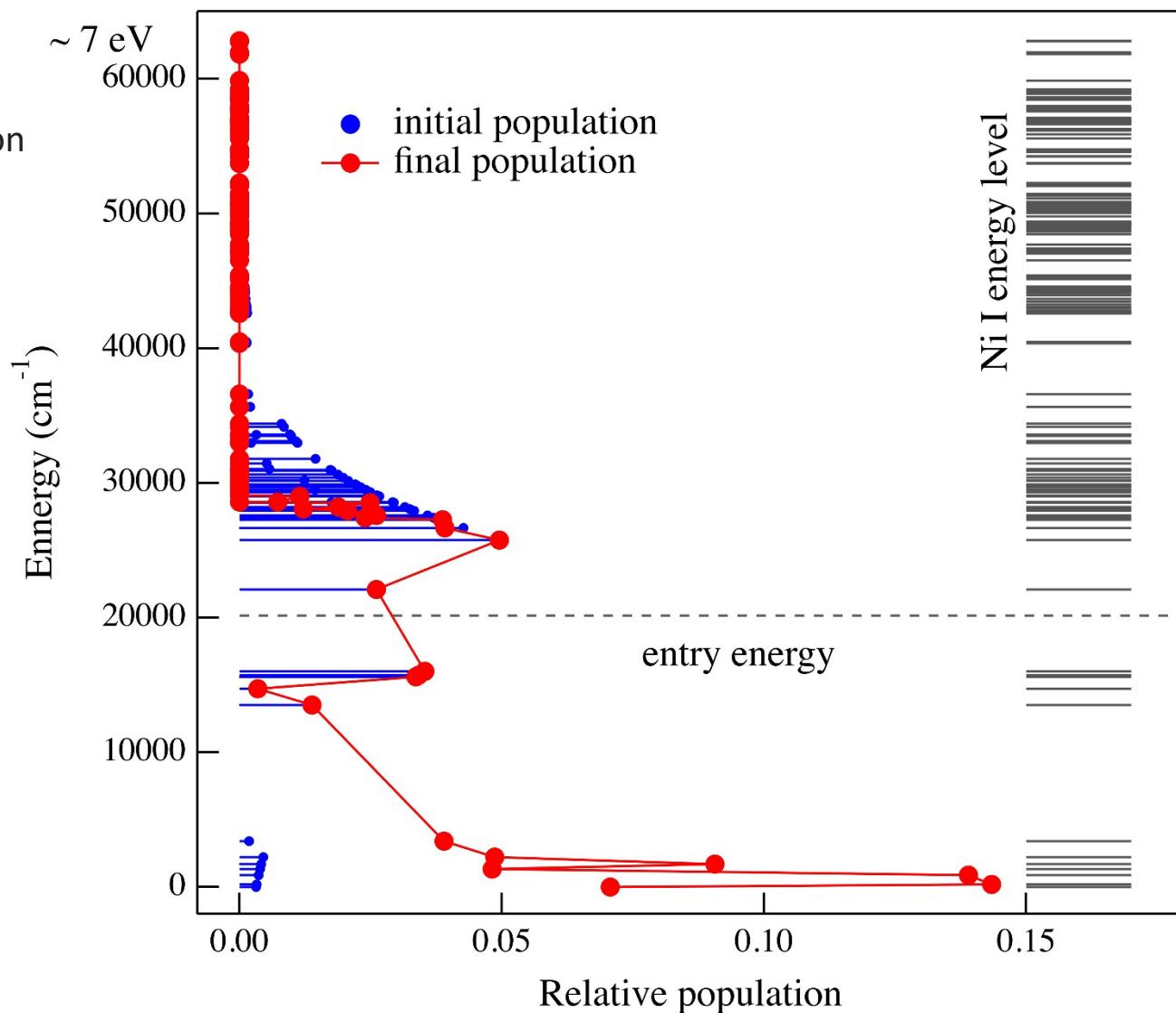
C. A. Ryder, K.M. et al., Spectrochim. Acta B 113, 16 (2015).

Designed by C. D. P. Levy at TRIUMF

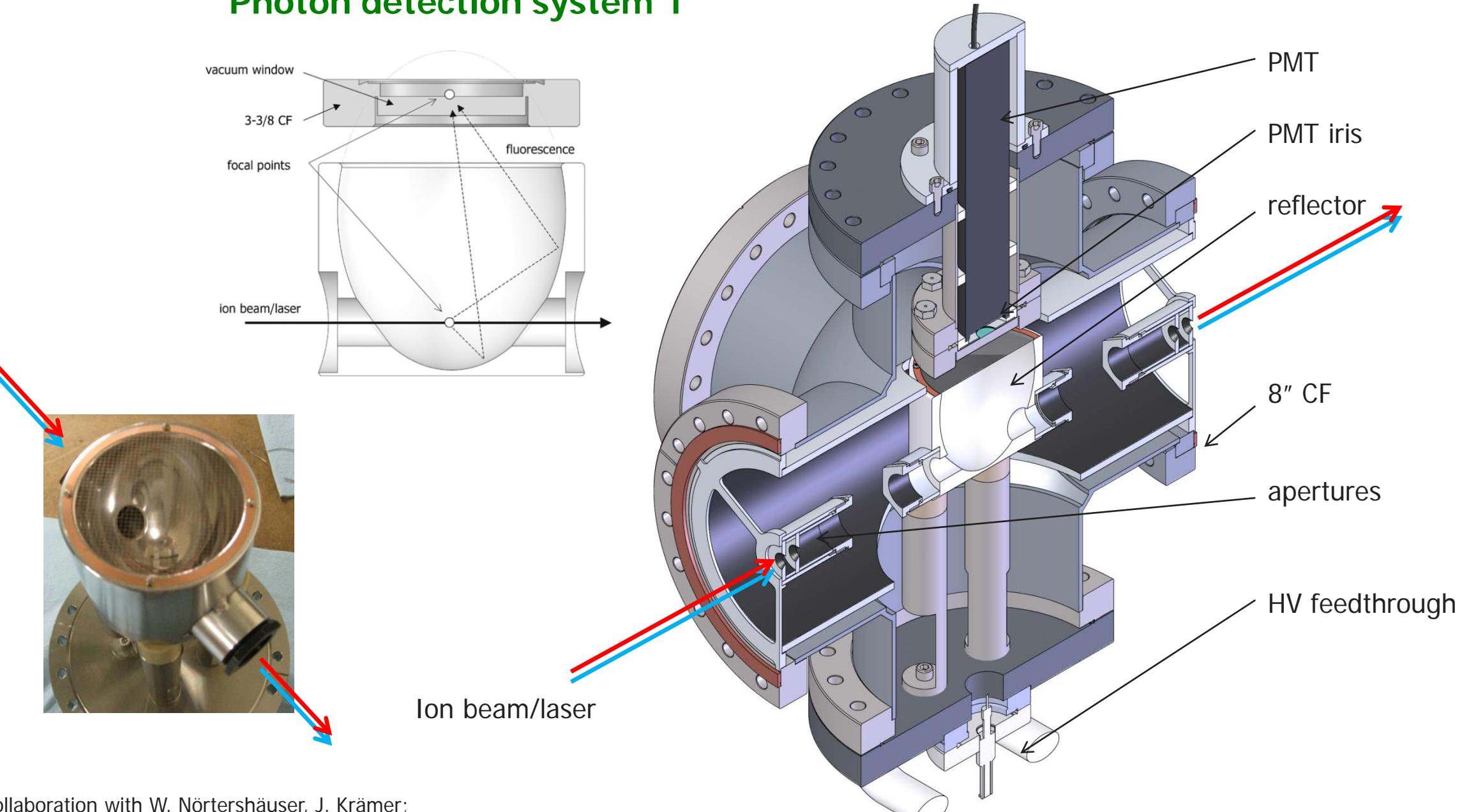
Atomic charge exchange

- charge exchange with Na vapor
 $\text{Ni}^+ + \text{Na} \rightarrow \text{Ni} + \text{Na}^+ + \Delta E$
 - High atomic energy-level density of Ni leads to severe fractionalization of population
- Small signal for laser spectroscopy
- Common issue for transition metal elements
 - In Ni experiment only 15% was probed

Simulation: 30 keV Ni^+ on Na vapor



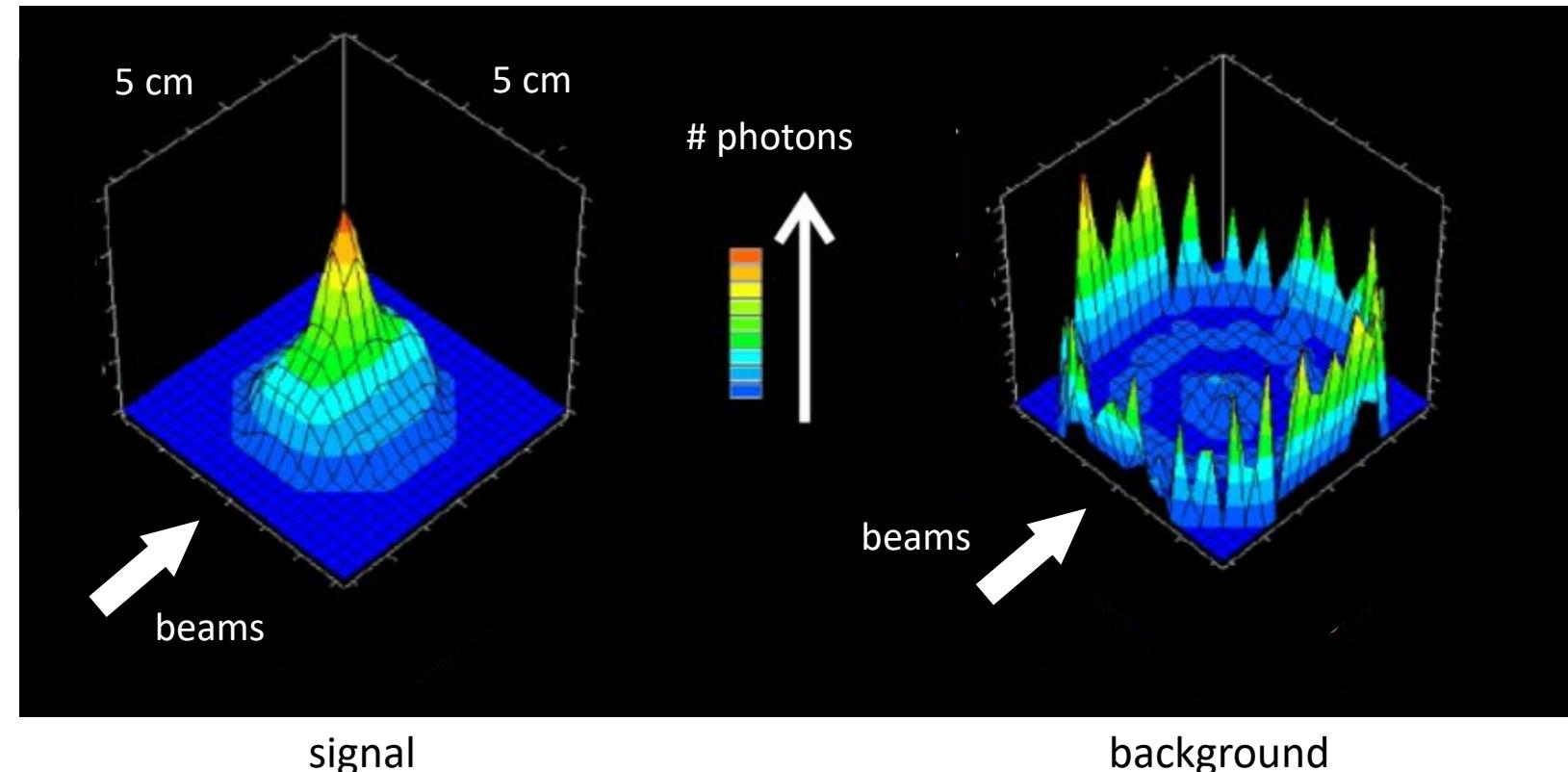
Photon detection system 1



In collaboration with W. Nörterhäuser, J. Krämer;
K. Minamisono et al, NIMA 709, 85 (2013).

Photon detection system 1: ellipsoidal reflector

Ellipsoidal reflector
simulation: at the detection plane

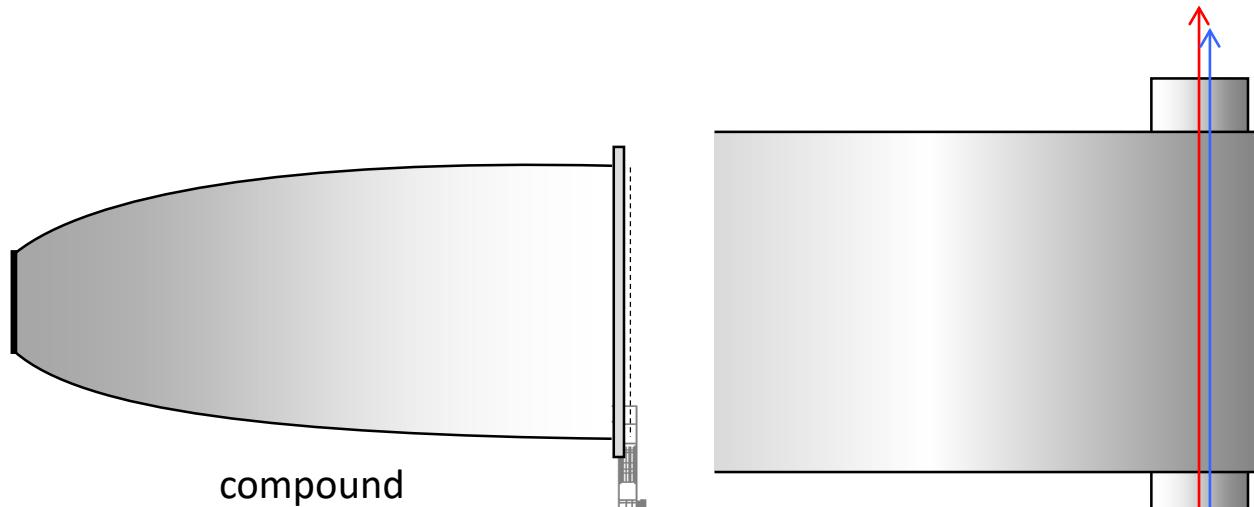


apertures

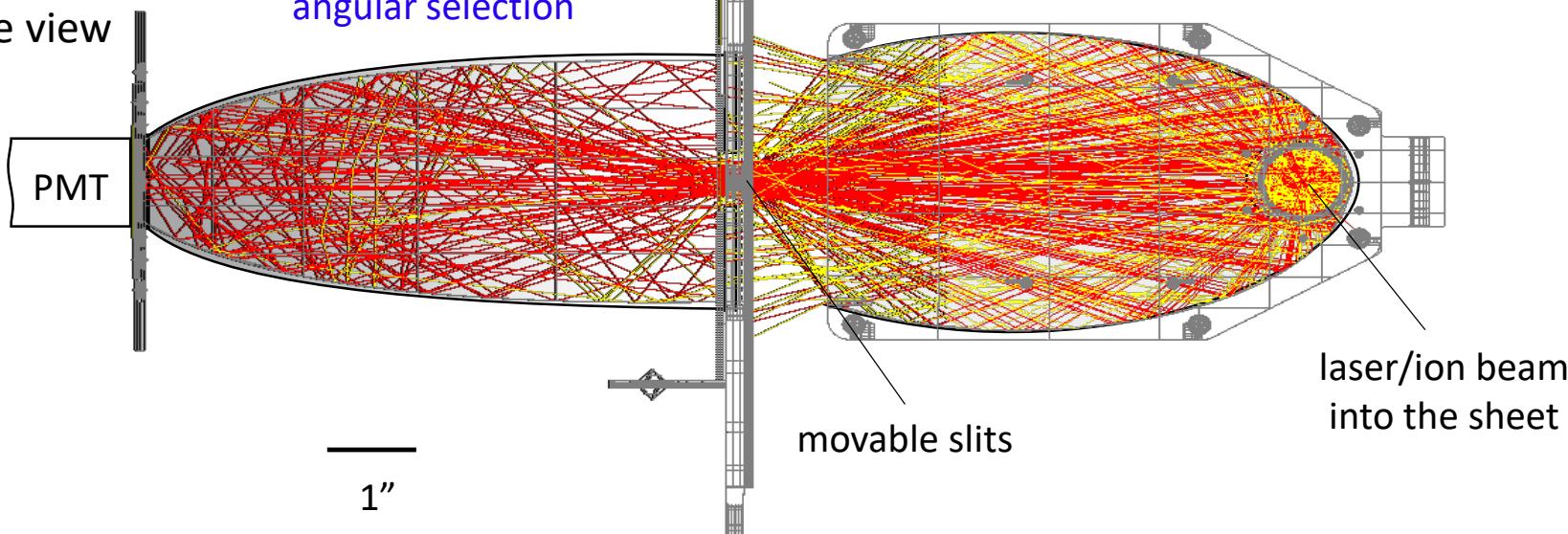
- Large solid angle
- Maximizing signal
- Relatively large background
- Wins S/N

Photon detection system 2

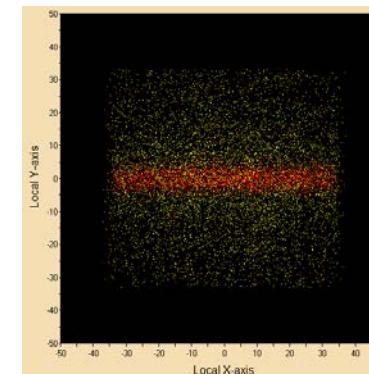
Top view



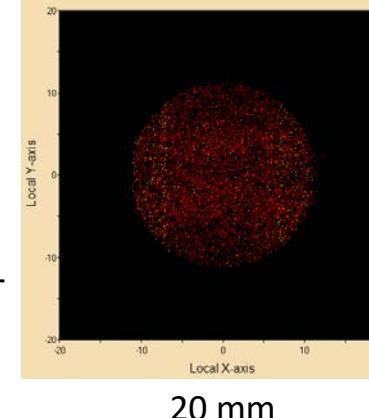
Side view



after ER

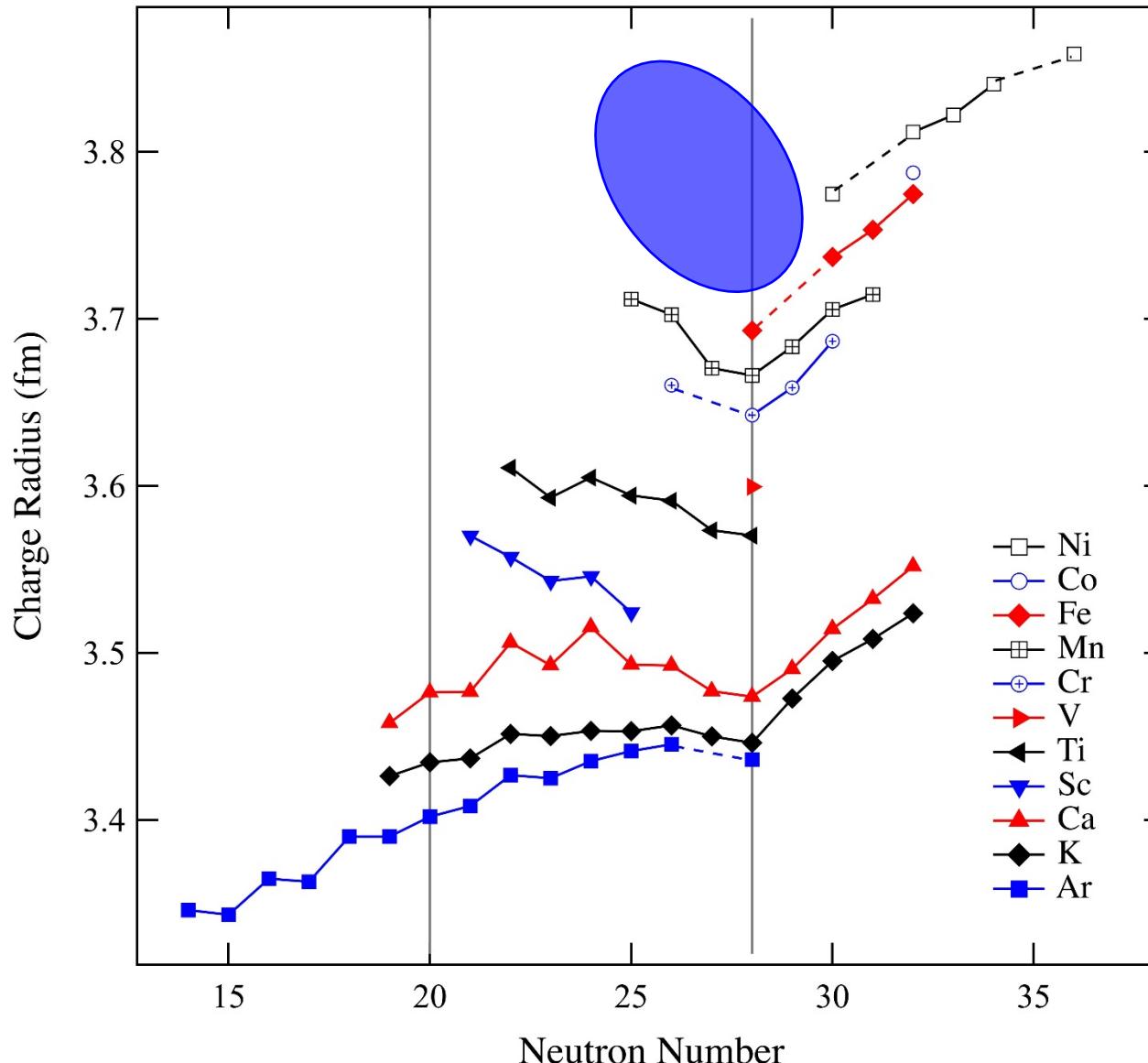


at PMT



- Filter background
- Minimizing background
- Relatively small signal
- Wins S/N

Current interests: Ca & Ni at $N = 20$ & 28

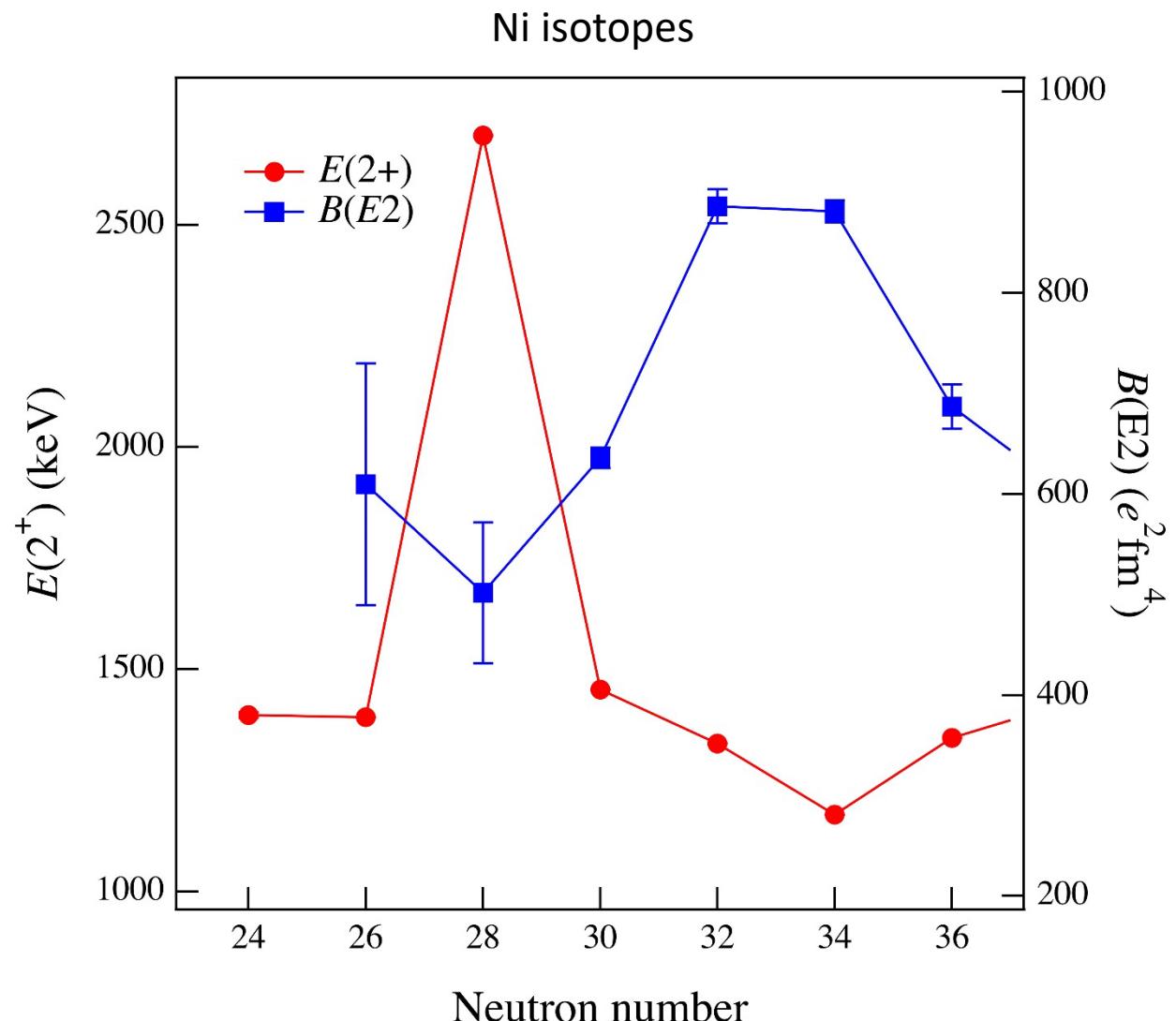


^{56}Ni nucleus is “soft”

- High $E(2^+_1)$, but no significant change in $B(E2)$
 - Canonical ^{56}Ni nucleon configuration: 50% in fp shell with FPD6
- Magnetic moment of $^{56}\text{Ni} \pm$ one nucleon nucleus ($^{55}\text{Ni}, ^{55}\text{Co}, ^{57}\text{Cu}, ^{57}\text{Ni}$)
 - Canonical ^{56}Ni nucleon configuration: 60% in fp shell with GXPF1



What about the charge radius?
size, shape and deformation



Isotope shift $^{52, 53}\text{Fe}$ & ^{56}Fe

$^5\text{D}_4 \leftrightarrow ^5\text{F}_5$ transition in Fe I at $26874.550 \text{ cm}^{-1}$

$$\delta\nu^{56,52} = -1834 \pm 3 \pm 3 \text{ (MHz)}$$

$$\delta\nu^{56,53} = -1249 \pm 4 \pm 3 \text{ (MHz)}$$

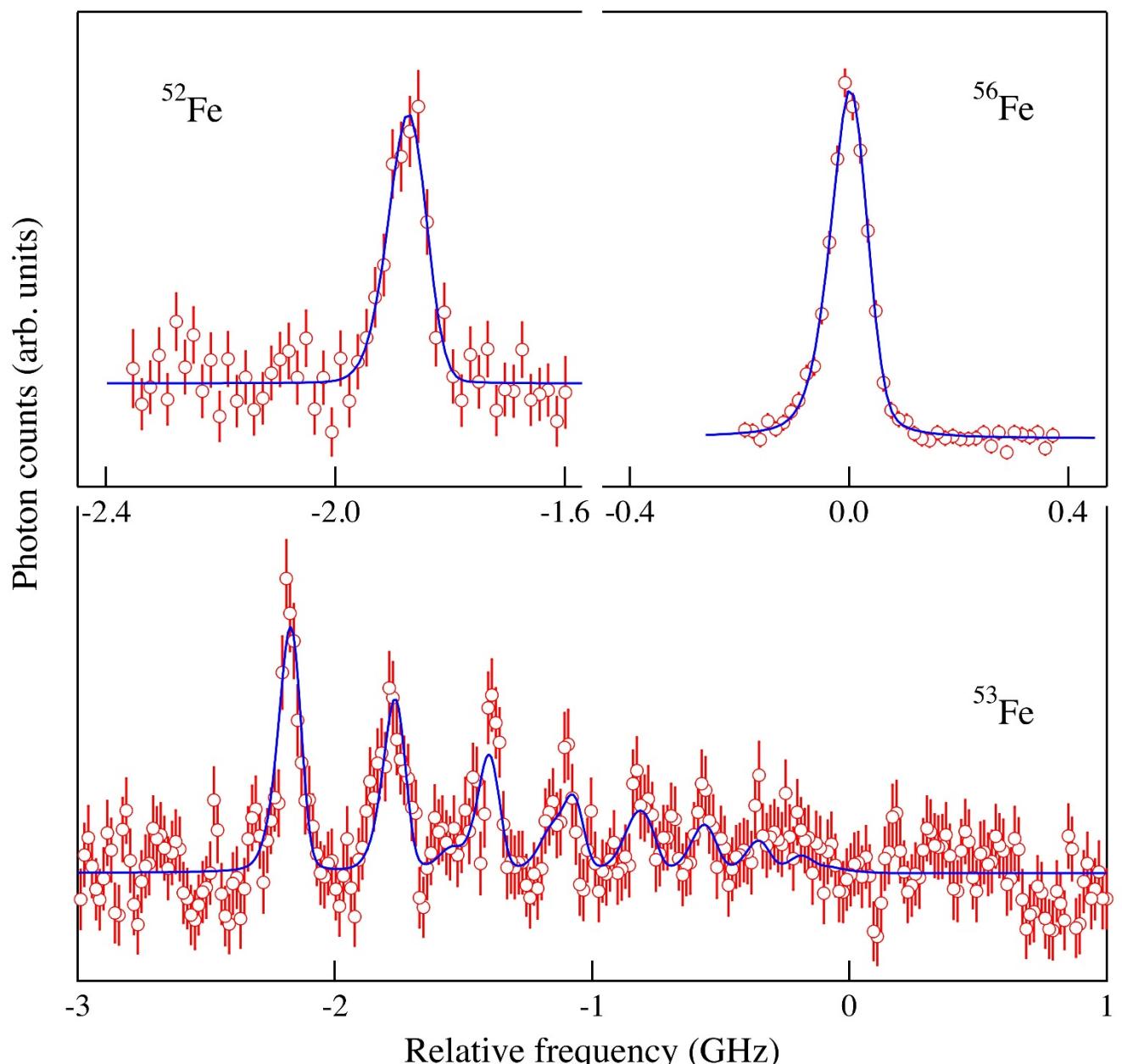
Theory + King plot

$$\begin{cases} F = -0.52 \pm 0.08 \text{ (GHz/fm}^2\text{)}: \text{theory} \\ k_{\text{SMS}} = 911 \pm 43 \text{ (GHz amu)} \end{cases}$$

$$\delta\nu^{A,A'} = \nu^{A'} - \nu^A = k \frac{M' - M}{M'M} + F \times \delta\langle r^2 \rangle^{A,A'}$$

$$\delta\langle r^2 \rangle^{56,52} = -0.034 \pm 0.013 \text{ (fm}^2\text{)}$$

$$\delta\langle r^2 \rangle^{56,53} = -0.218 \pm 0.013 \text{ (fm}^2\text{)}$$



Stable Fe-isotope shift: S. Krins et al., PRA 80, 062508 (2009);

Stable Fe $\delta\langle r^2 \rangle$: G. Fricke & K. Heilig, *Nuclear Charge Radii* (Springer, 2004);

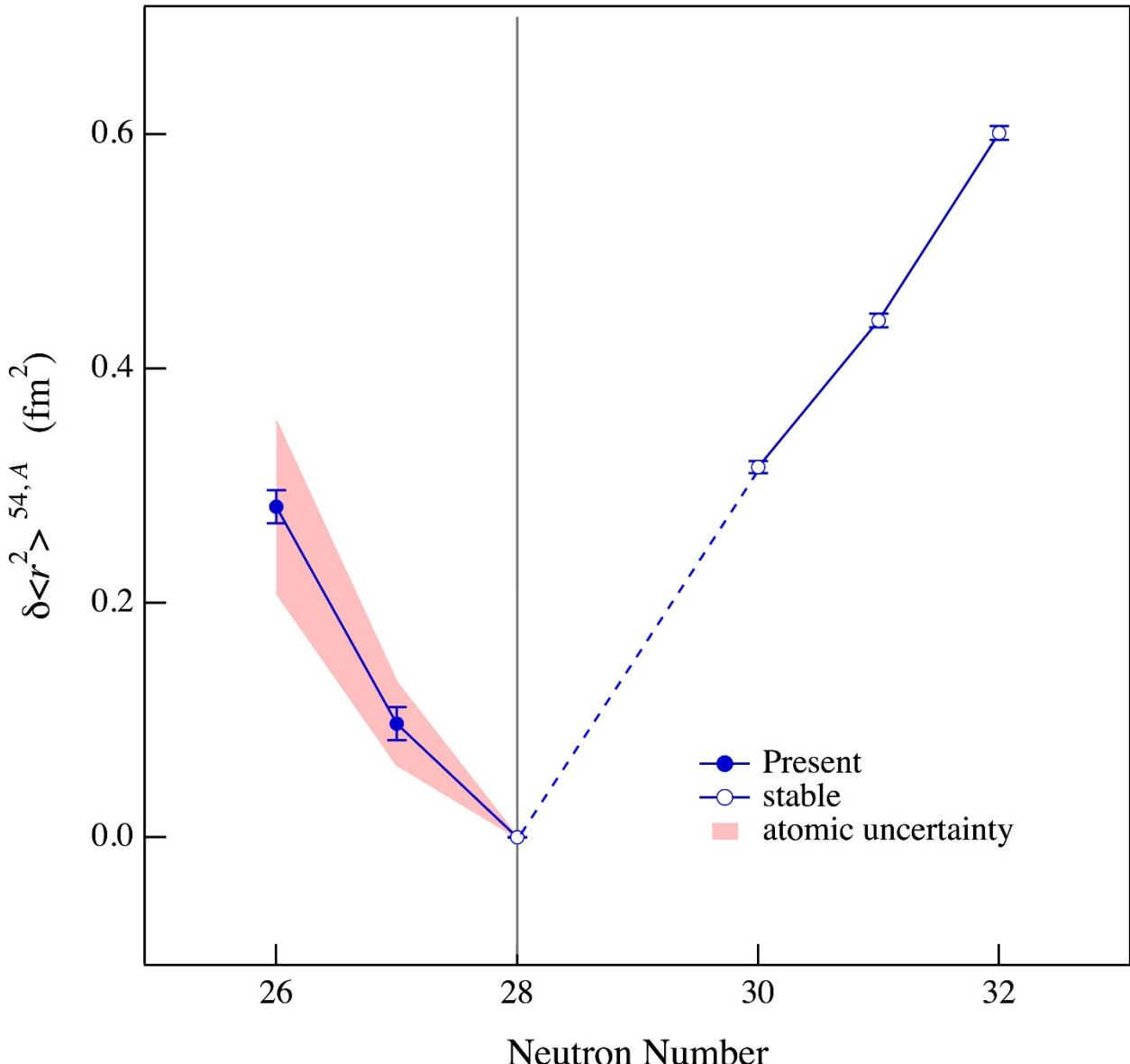
Field shift calculation: Multi-configuration Dirac-Fock (MCDF) method;

K. Minamisono et al., Phys. Rev. Lett. 117, 252501 (2016);

A. J. Miller, K. M. et al., Phys. Rev. C 96, 054314 (2017).

$\delta\langle r^2 \rangle^{54,A}$ of Fe isotopes

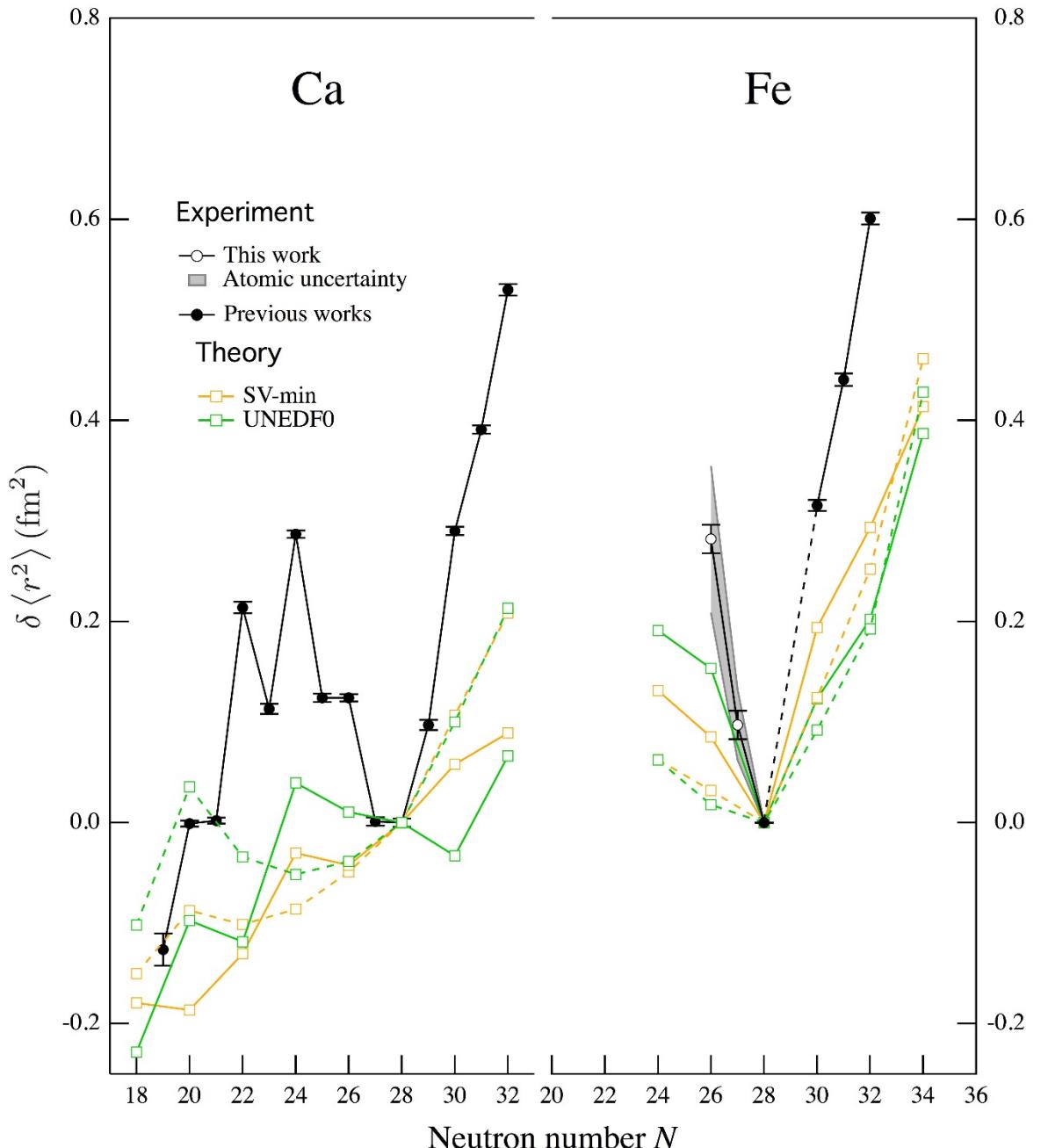
- Shell closure signature; “Kink” at $N = 28$
- Radii of ^{52}Fe and ^{56}Fe similar



DFT calculation for Ca & Fe

Skyrme EDF

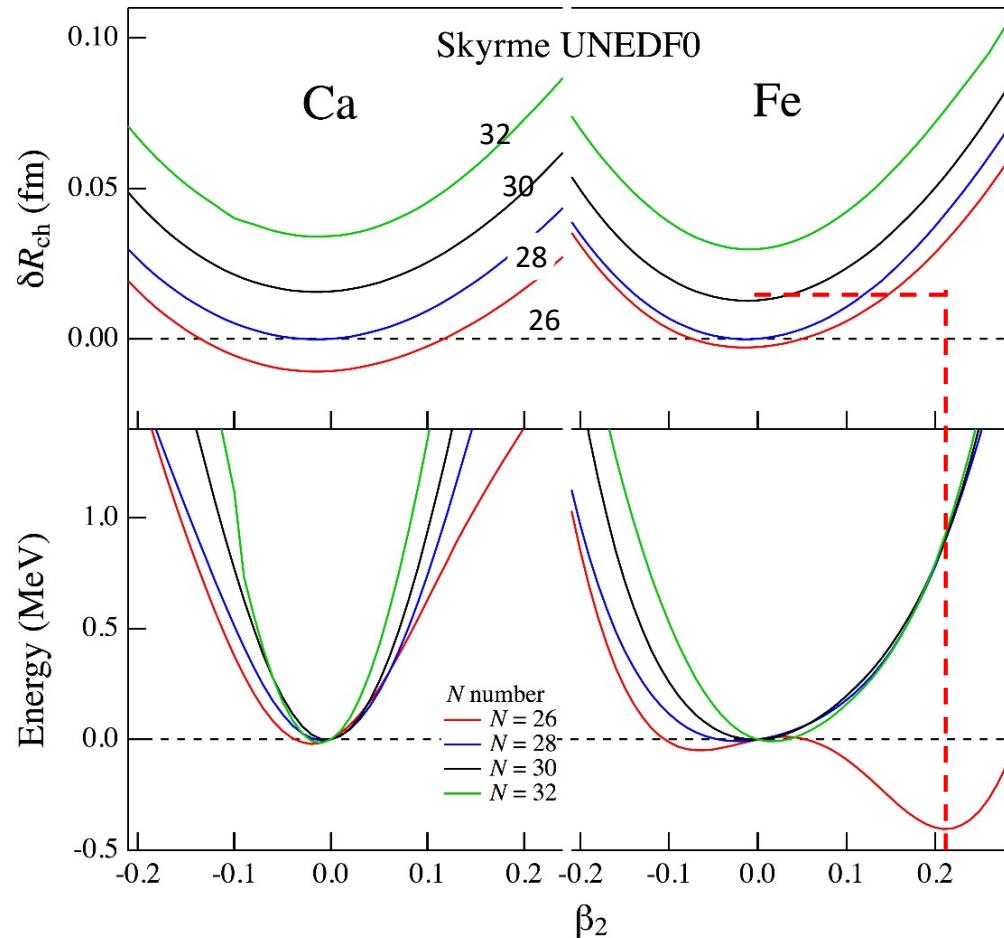
- Spherical shape (dashed line)
No good agreement
- Deformed shape (solid line)
No good for Ca
Kink better reproduced for Fe



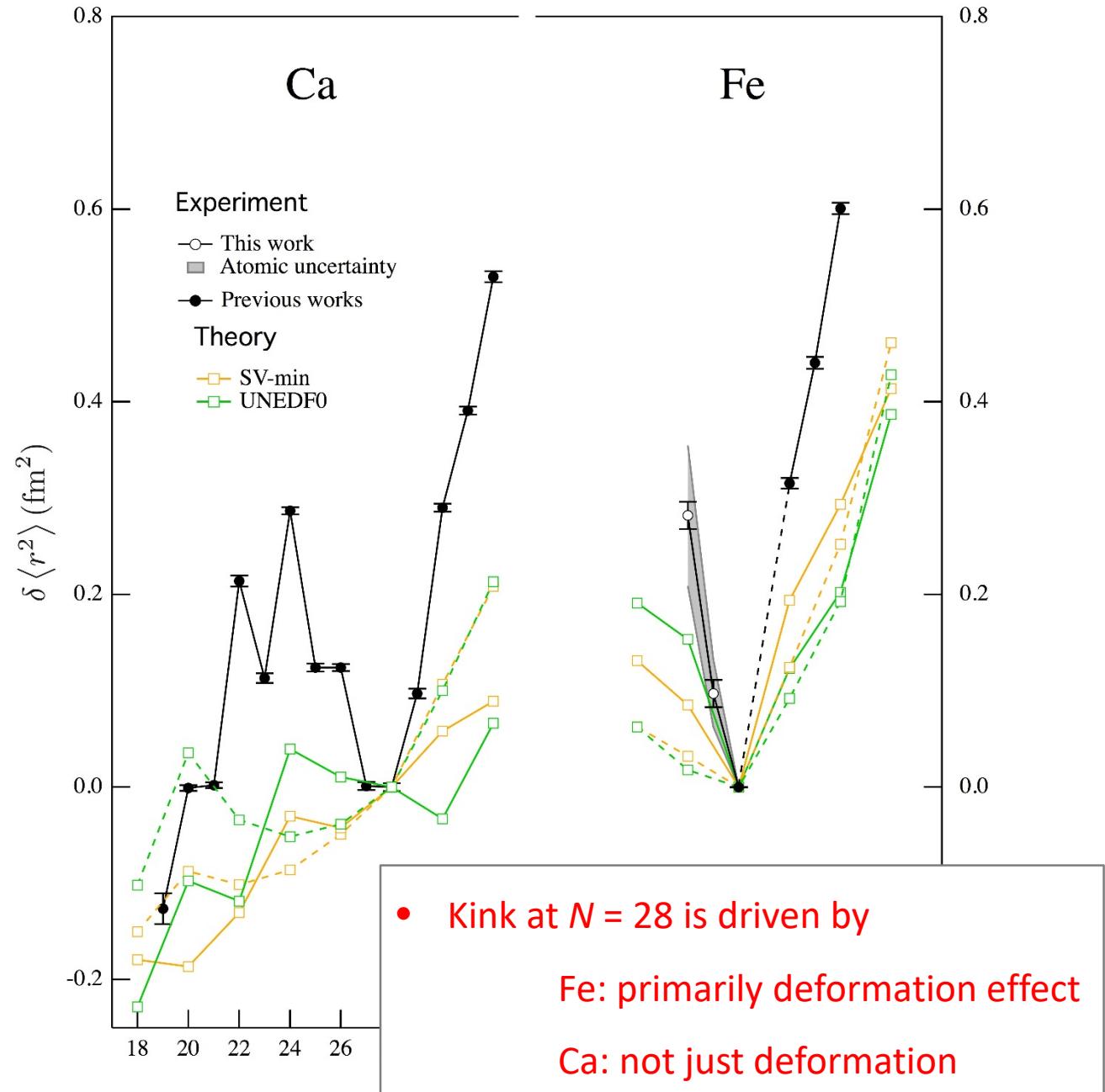
K. Minamisono et al., Phys. Rev. Lett. 117, 252501 (2016);
SV-min: P. Klöpfel et al., PRC79, 034310 (2009);
UNEDFO: M. Kortelainen et al., PRC82, 024313 (2010).

DFT calculation for Ca & Fe

Skyrme EDF



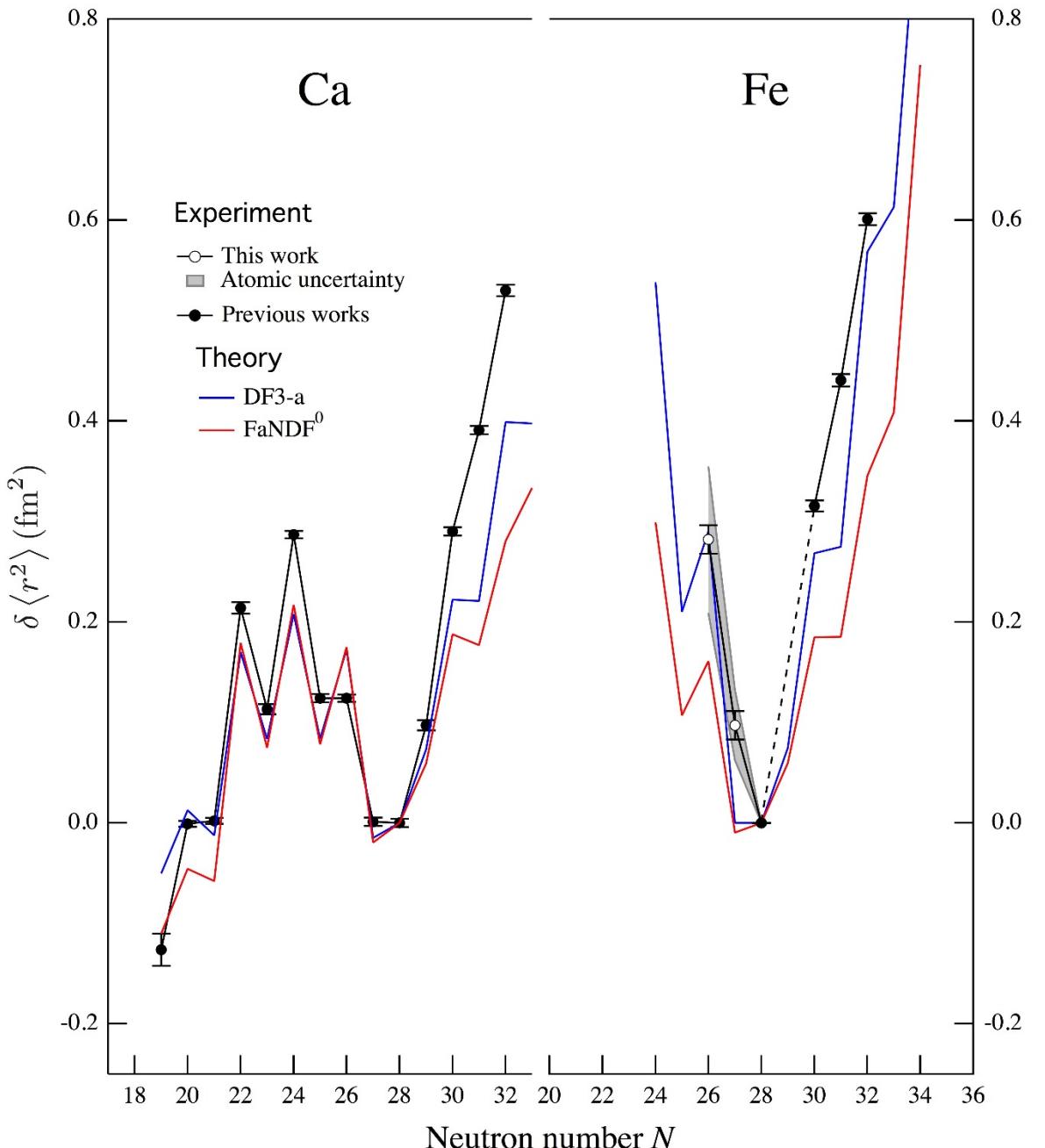
Consistent with $B(E2)$ of ^{52}Fe isotopes.
 K. L. Yurkewicz et al. PRC70, 034301 (2004).



DFT calculation for Ca and Fe: 2

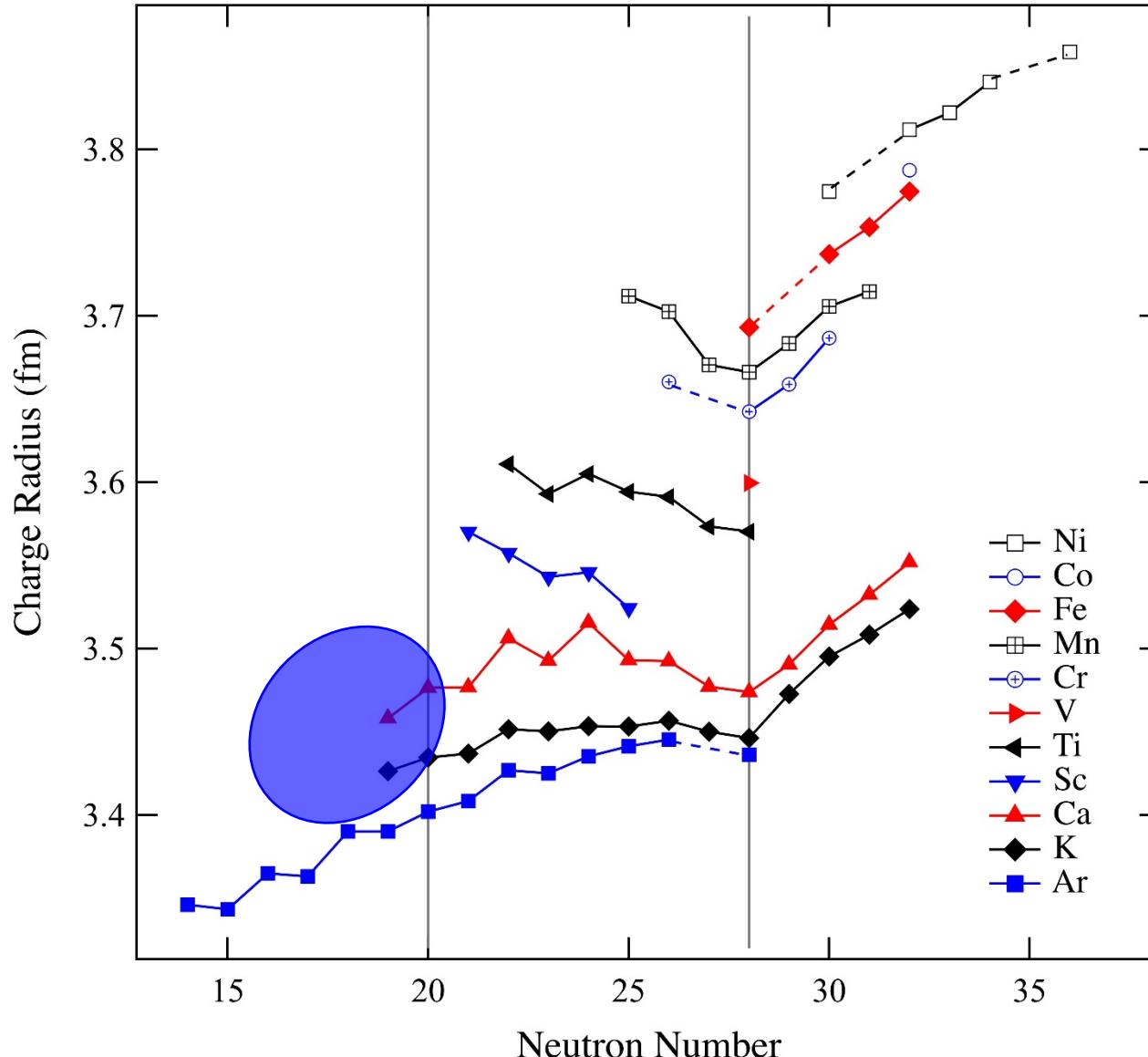
Fayans EDF

- a nuclear-density gradient dep. pairing term.
(coupling to surface vibrations)
- Parameters for the paring term fitted to S_n .
- Spherical shape
- **Very good overall agreement** though
 - underestimate R_{ch} of very heavy Ca
 - overestimate odd-even staggering for Fe



K. Minamisono et al., Phys. Rev. Lett. 117, 252501 (2016);
 DF3-a: S. V. Tolokonnikov & E. E. Saperstein, PAN73, 1684 (2010);
 FaNDF⁰: S. A. Fayans, JETP Lett. 68, 169 (1998).

Current interests: Ca & Ni at $N = 20$ & 28

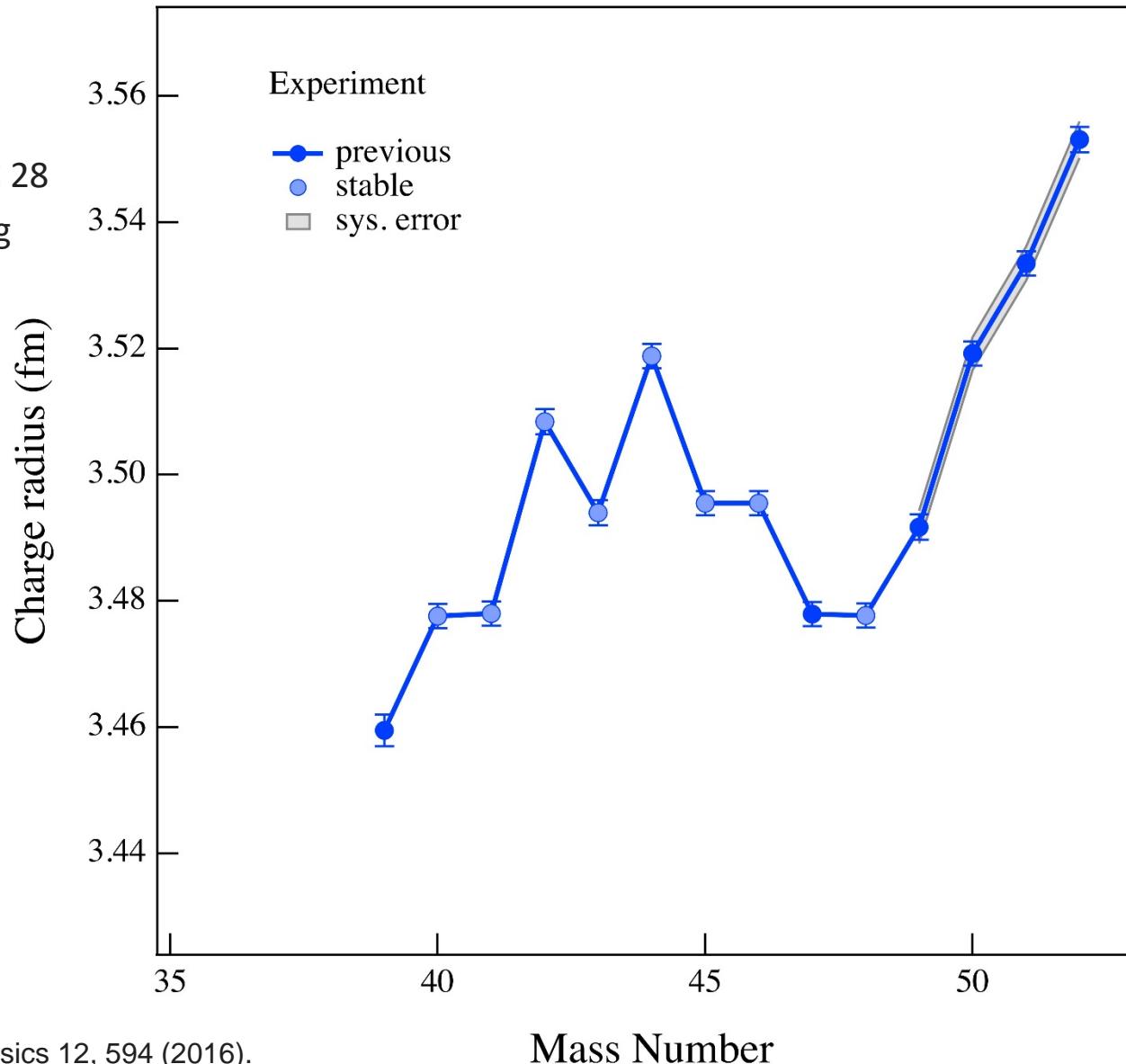


Example: charge radii of Ca isotopes

A “textbook” example

- $r(^{48}\text{Ca}) \approx r(^{40}\text{Ca})$
- Parabolic shape in $20 \leq N \leq 28$
- Strong odd-even staggering
- Steep increase $N \geq 28$
- No “kink” at $N = 20$

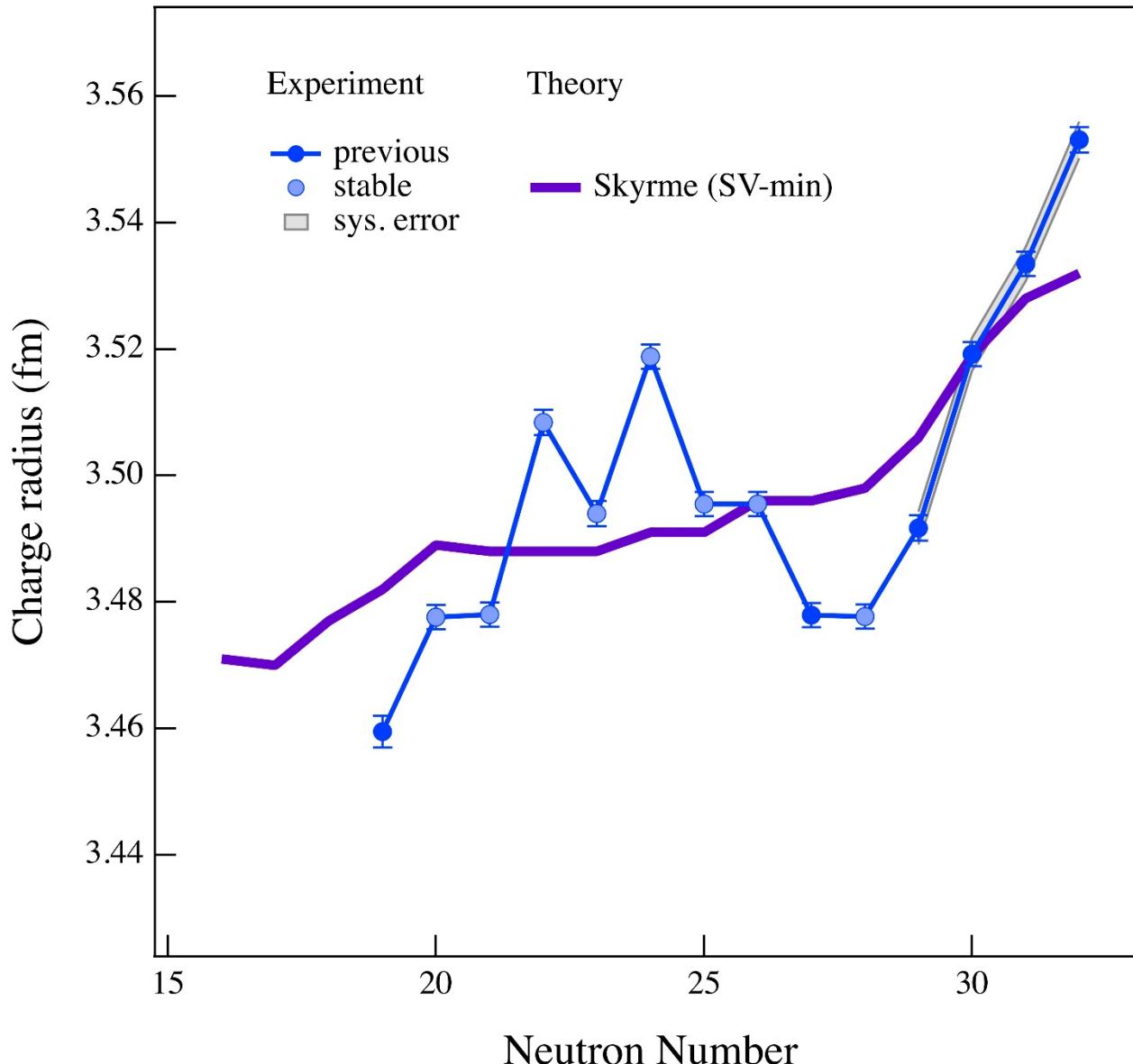
Ca radii show a very intricate pattern.



Data: R. E. Garcia Ruiz et al., Nature Physics 12, 594 (2016).

$R(^{40}\text{Ca})$: G. Fricke and K. Heilig, Nuclear charge radii (Springer Berlin, 2004).

Charge radii of Ca isotopes

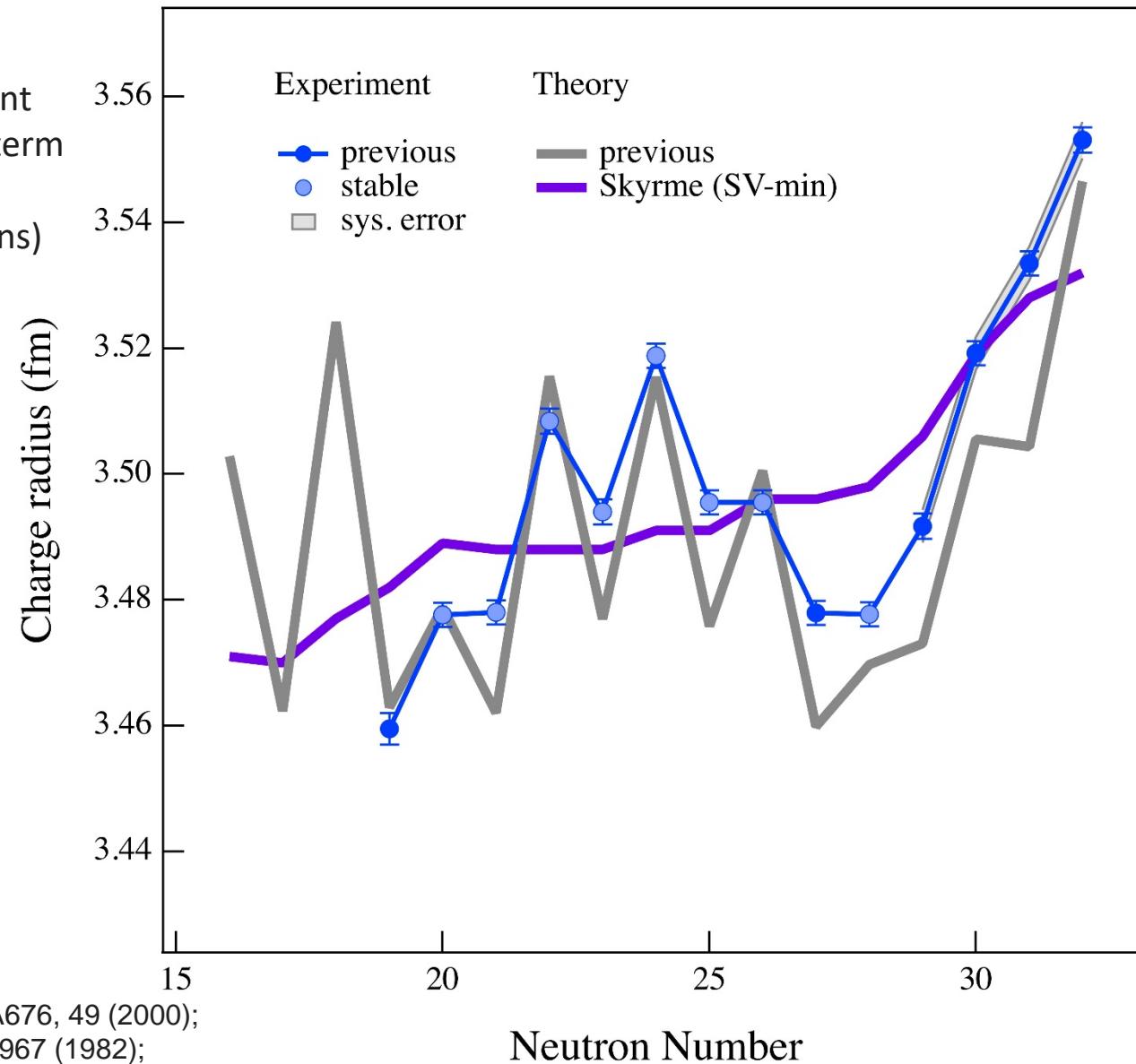


Typically DFT with Skyrme EDF does not reproduce details of the chain of Ca charge radii.

Charge radii of Ca isotopes

Fayans EDF

- Density gradient dependent pairing term
- Reflects surface phonons (vibrations)



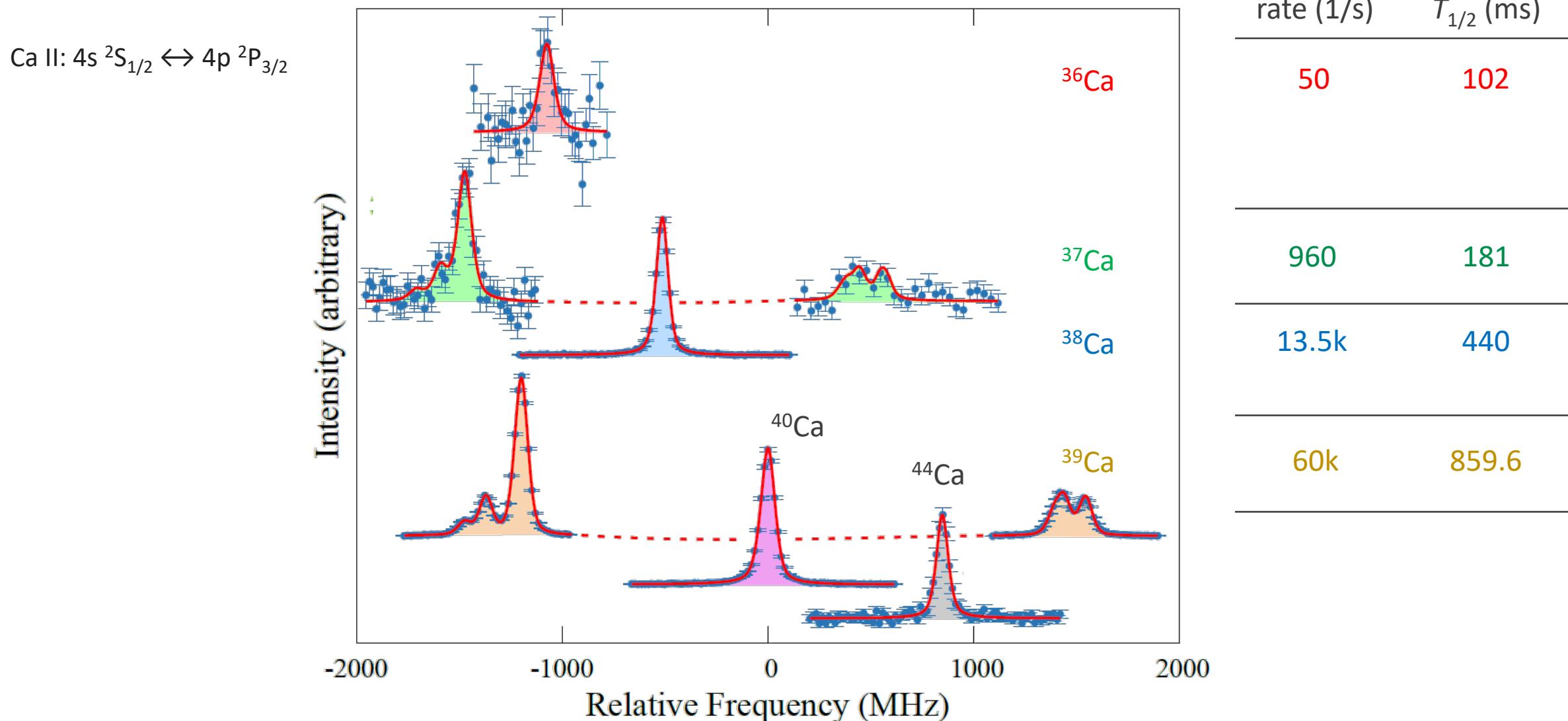
Fayans EDF does good job, but some discrepancy, and predicts very large radii for light Ca.

Theory: S. A. Fayans et al., Nucl. Phys. A676, 49 (2000);

V. A. Khodel et al., J. Phys. G 8, 967 (1982);

P. -G. Reinhard and W. Nazarewicz, PRC 95, 064328 (2017).

Isotope shift of HFS for neutron-deficient Ca



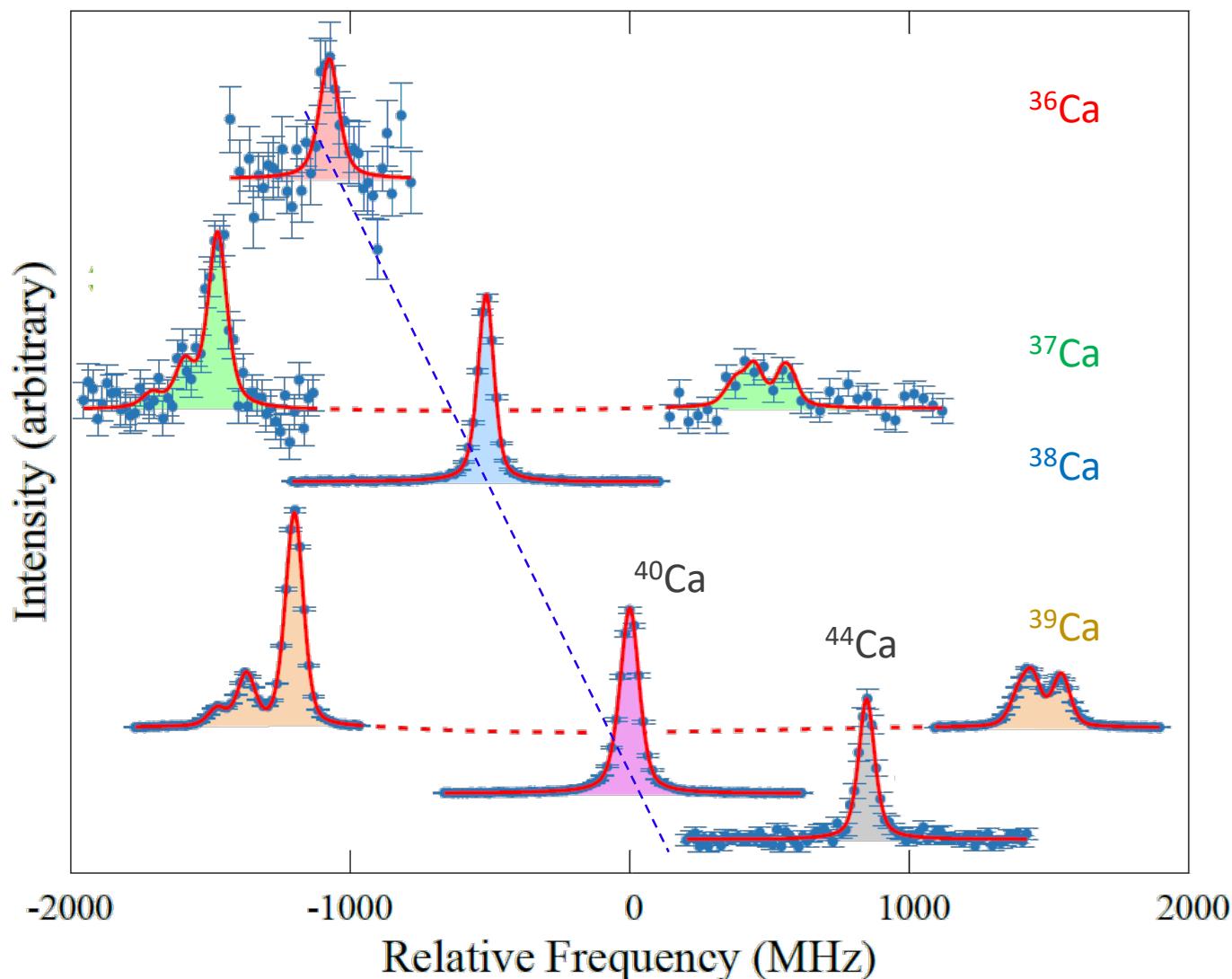
Isotope shift of HFS for neutron-deficient Ca

Ca II: $4s\ ^2S_{1/2} \leftrightarrow 4p\ ^2P_{3/2}$

$$\begin{aligned}\delta\nu^{A,A'} &= \nu^{A'} - \nu^A \\ &= k \frac{M' - M}{M'M} \\ &\quad + F \times \delta\langle r^2 \rangle^{A,A'}\end{aligned}$$

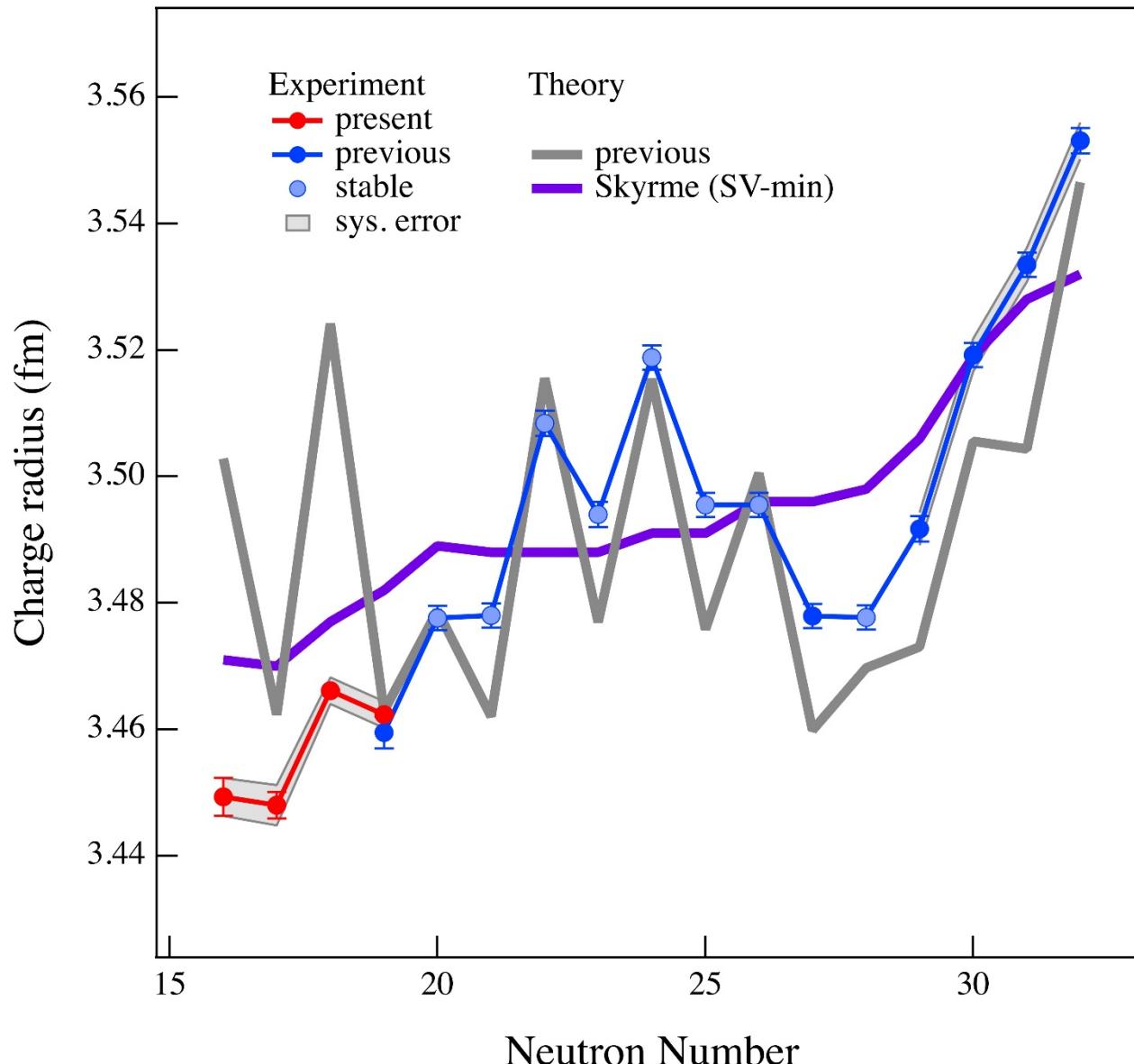
Atomic factors well determined:

$$\begin{cases} k = 409.35(42) \text{ GHz amu} \\ F = -284.7(82) \text{ MHz fm}^{-2} \end{cases}$$



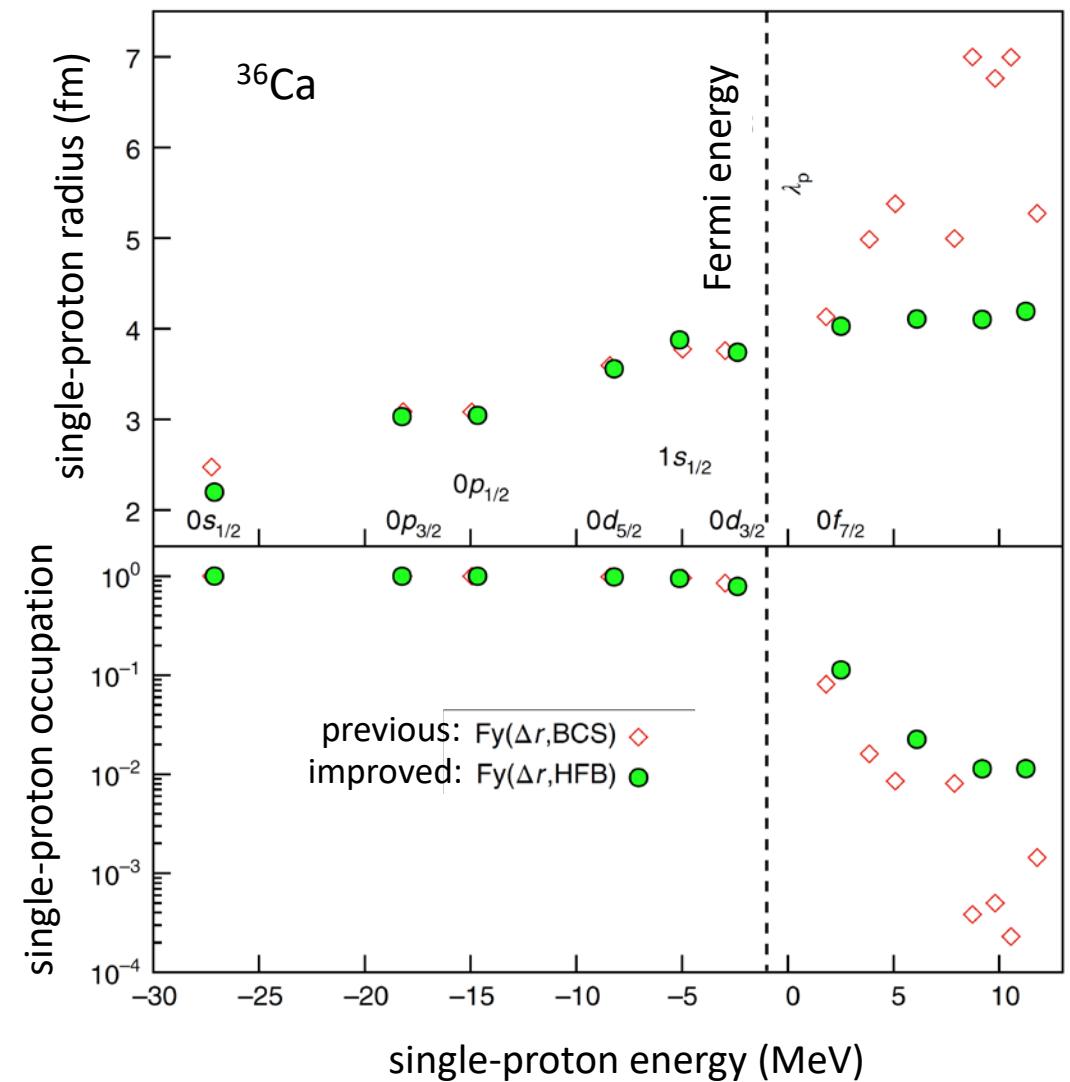
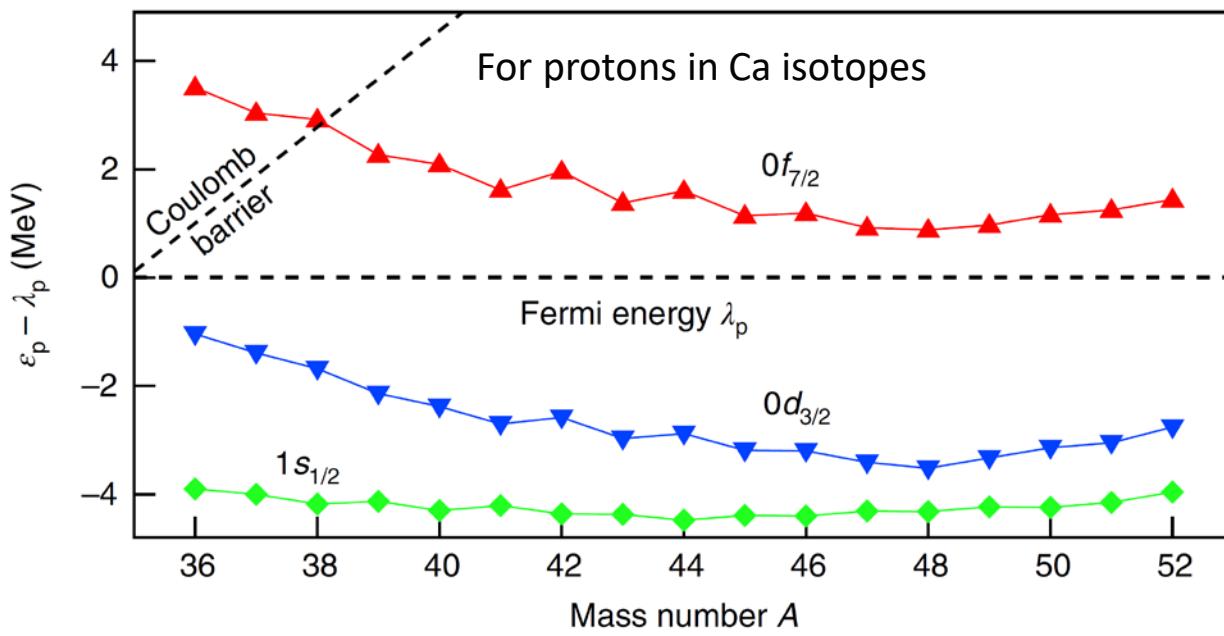
rate (1/s)	$T_{1/2}$ (ms)
50	102
960	181
13.5k	440
60k	859.6

Charge radii of Ca isotopes

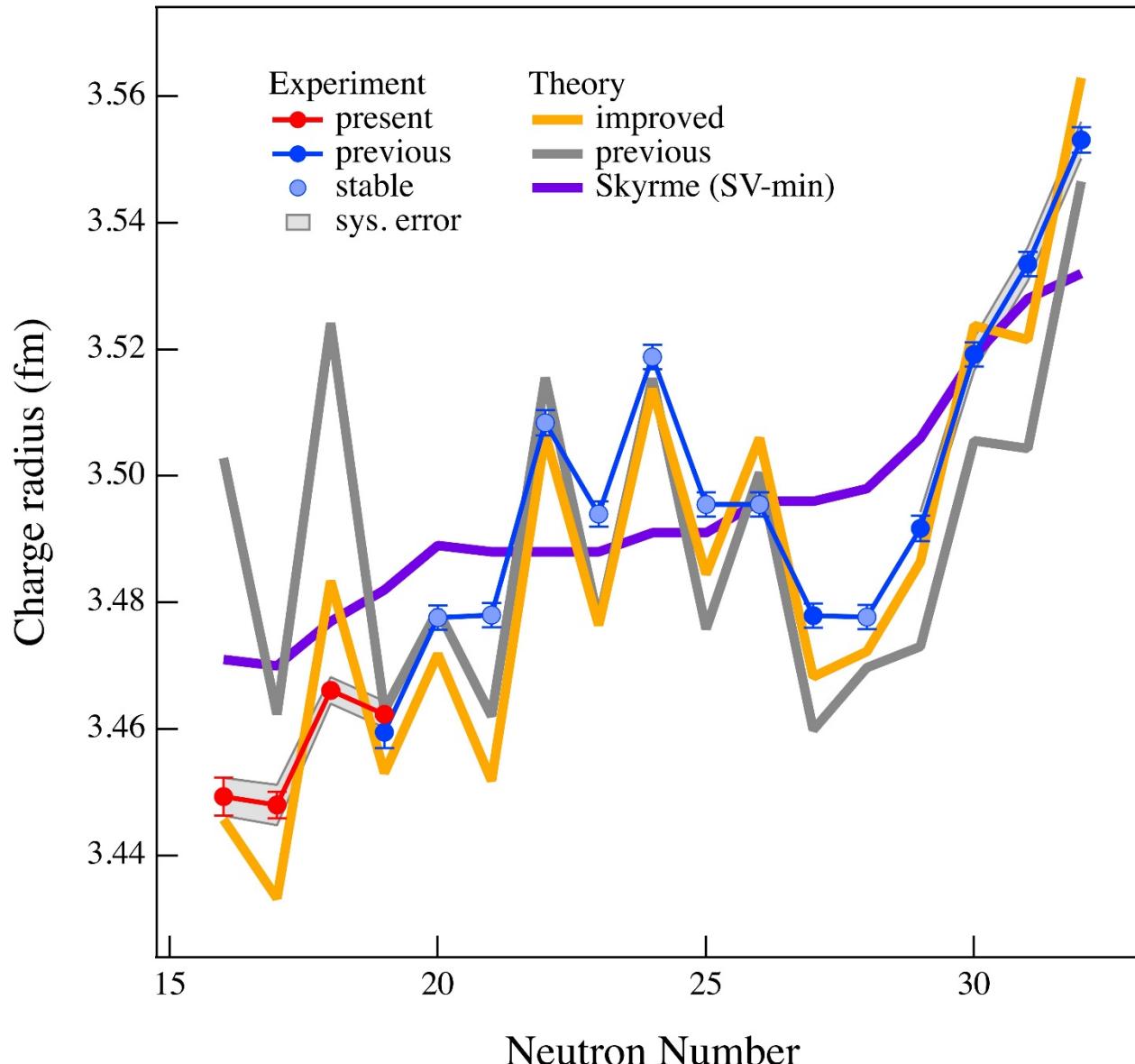


Weakly-bound protons

- Moving toward ^{36}Ca , the $0f_{7/2}$ state rises above the Coulomb barrier.
- Pairing interaction can scatter proton pairs into particle continuum (superfluidity).
- Improved theory avoids the nonphysical increase of radius in the continuum



Charge radii of Ca isotopes



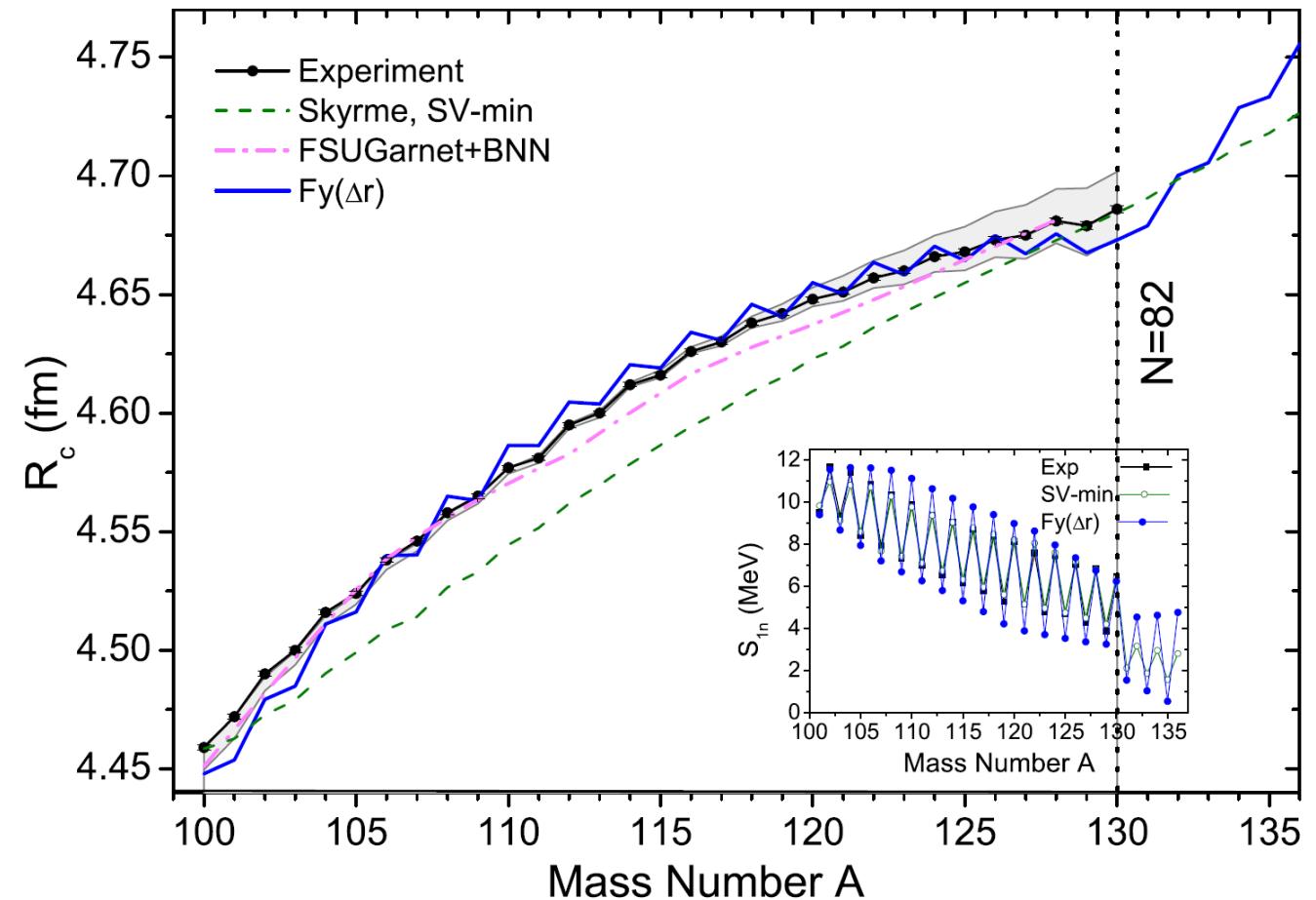
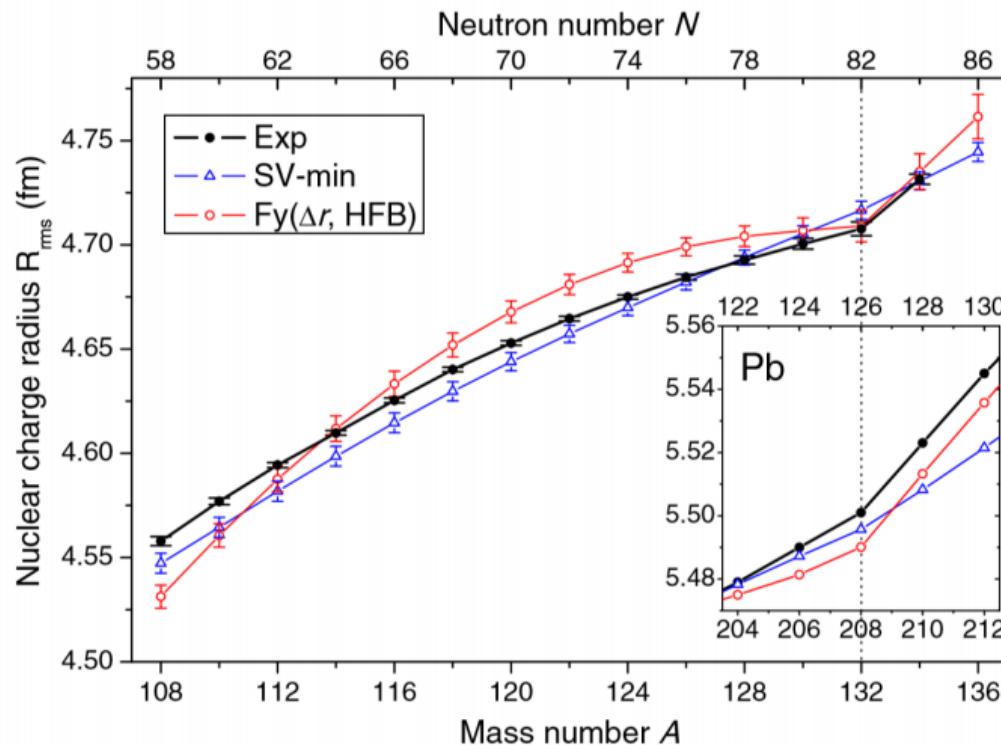
Improved theory explains
Cd radii very well.

Need better understanding of
odd-even staggering.

Fayans EDF for Cd and Sn isotopes

Fayans EDF for Ca works for Cd and Sn

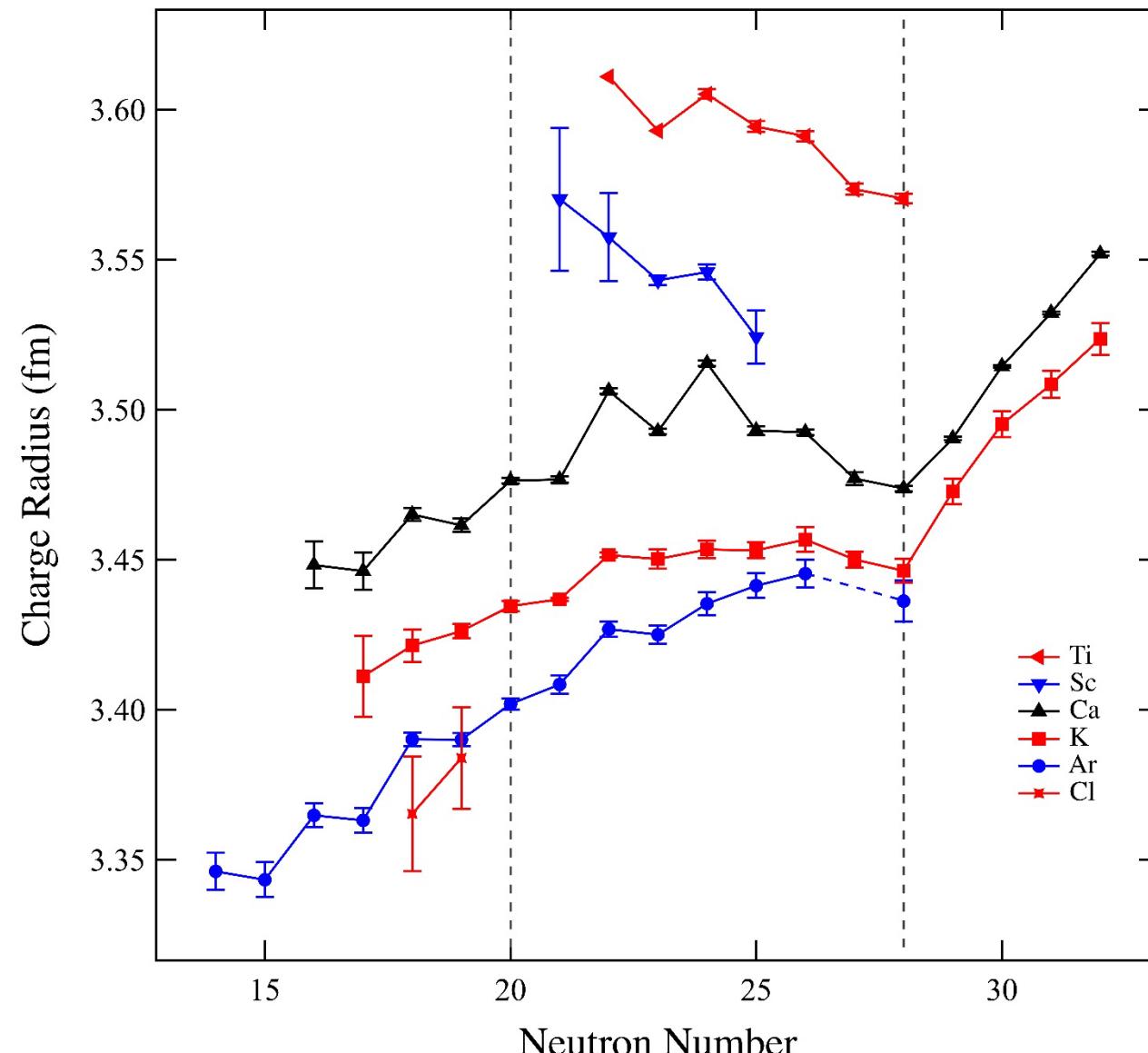
- Fayans EDF captures important mechanism that works at (far) different locations in the nuclear chart.
- Important step towards a global model



Cd: M. Hammen et al., Phys. Rev. Lett. 121, 102501 (2018).
Sn: C. Gorges et al. Phys. Rev. Lett. 122, 192502 (2019).

Charge radii around Ca

- No “kink” at $N = 20$
- Change the sign of slope at $N = 20$ between Ca and Sc
- Approved experiment on Sc
(NSCL e18024: A. Klose spokesperson)



Summary/future prospects

**BECOLA is a laser spectroscopy facility at NSCL/FRIB/MSU,
complementing capabilities at ISOL type facilities.**

Experiments on neutron-deficient Sc and Ni before NSCL is shut down.

FRIB operation anticipated to start in 2022.

“Day 1” experiment for BECOLA

- ^{52}Ni for mirror charge radii
- Proton emitters e.g. ^{147}Tm , $^{156}\text{Ta}^g, m$
- Extending radii of key nuclei e.g. Ca
- Light mass elements e.g. O and F
-

and last but not least...

BECOLA collaboration

K. Minamisono^{1,2}, R. Beerwerth^{3,4}, B. A. Brown^{1,2}, N. Everett¹, S. Fritzsche^{3,4}, D. Garand¹, R. P. de Groot⁵, J. D. Holt⁶, P. Imgram⁷, C. Kalman¹, A. Klose⁸, K. König¹, J. D. Lantis^{1,9}, Y. Liu¹⁰, B. Maaß⁷, P. F. Mantica^{9,11}, A. J. Miller^{1,2}, P. Müller¹², W. Nazarewicz^{2,11,13}, W. Nörtershäuser⁷, E. Olsen², M. R. Pearson⁶, S. Pineda^{1,9}, R. C. Powell^{1,2}, P. -G. Reinhard¹⁴, R. Romeo-Romero^{10,15}, D. M. Rossi⁷, E. E. Saperstein^{16,17}, A. Schwenk⁷, F. Sommer⁷, C. Sumithrarachchi¹, A. Teigelhöfer⁶ and S. V. Tolokonnikov^{16,18}, J. Watkins^{1,2}

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²Department of Physics and Astronomy, Michigan State University

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⁴Theoretisch-Physikalisches Institut, Friedrich-Schiller-Universität Jena

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¹⁰Physics Division, Oak Ridge National Laboratory

¹¹Facility for Rare Isotope Beams, Michigan State University

¹²Physics Division, Argonne National Laboratory

¹³Institute of Theoretical Physics, Faculty of Physics, University of Warsaw

¹⁴Institut für Theoretische Physik, Universität Erlangen

¹⁵Department of physics, University of Tennessee

¹⁶National Research Centre “Kurchatov Institute”

¹⁷National Research Nuclear University MEPhI

¹⁸Moscow Institute of Physics and Technology

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Computer Center of Kurchatov Institute.

