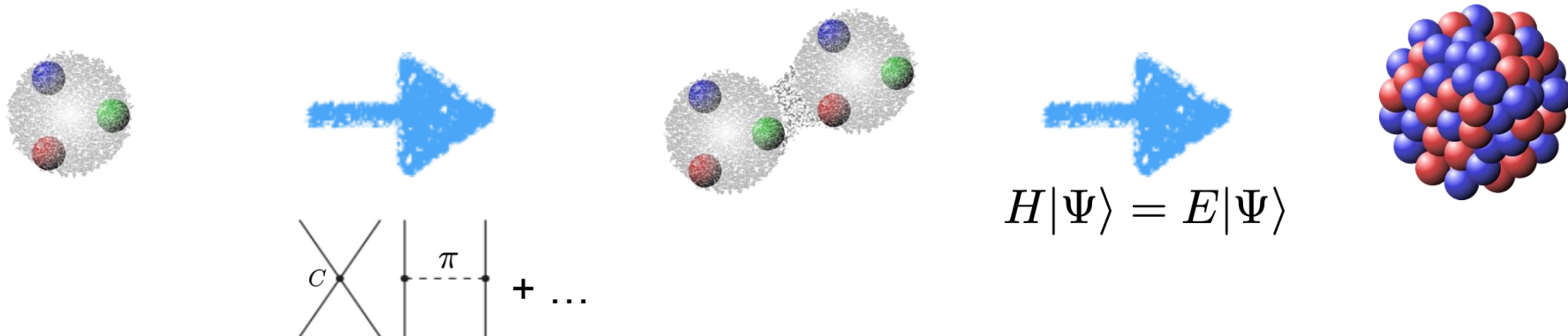


Valence-space in-medium similarity renormalization group description of calcium isotopes



Takayuki Miyagi (University of Tsukuba, Japan)

Ab initio many-body calculations: where has the nuclear pairing gone?

@ CEA, Saclay, France, May 14, 2025

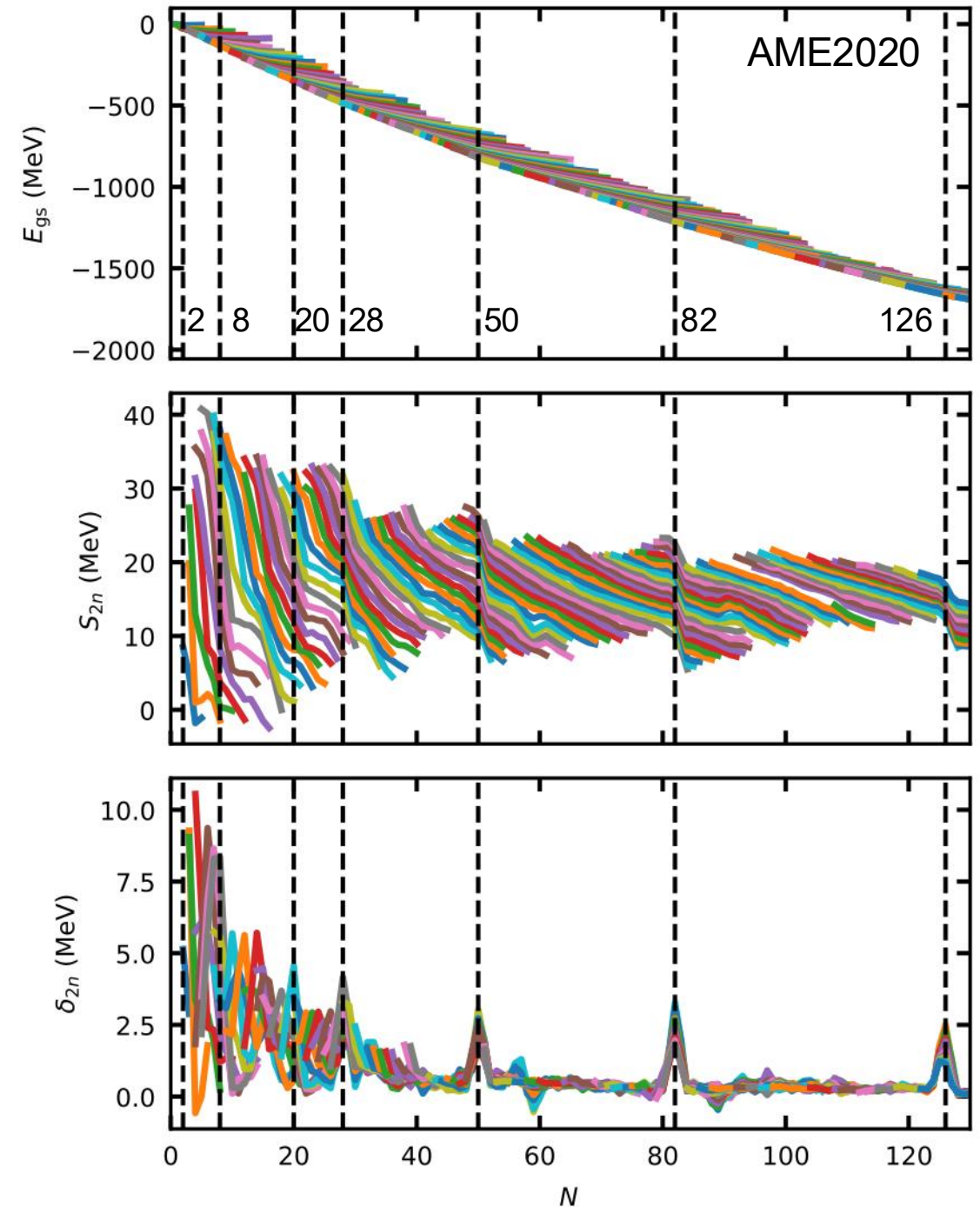
Energies and radii

- Ground-state energies are fundamental properties of nuclei

$$S_{2n}(Z, N) = E_{\text{g.s.}}(Z, N) - E_{\text{g.s.}}(Z, N - 2)$$

$$\delta_{2n}(Z, N) = \frac{1}{2} [S_{2n}(Z, N) - S_{2n}(Z, N + 2)]$$

- The magic signature can be found!



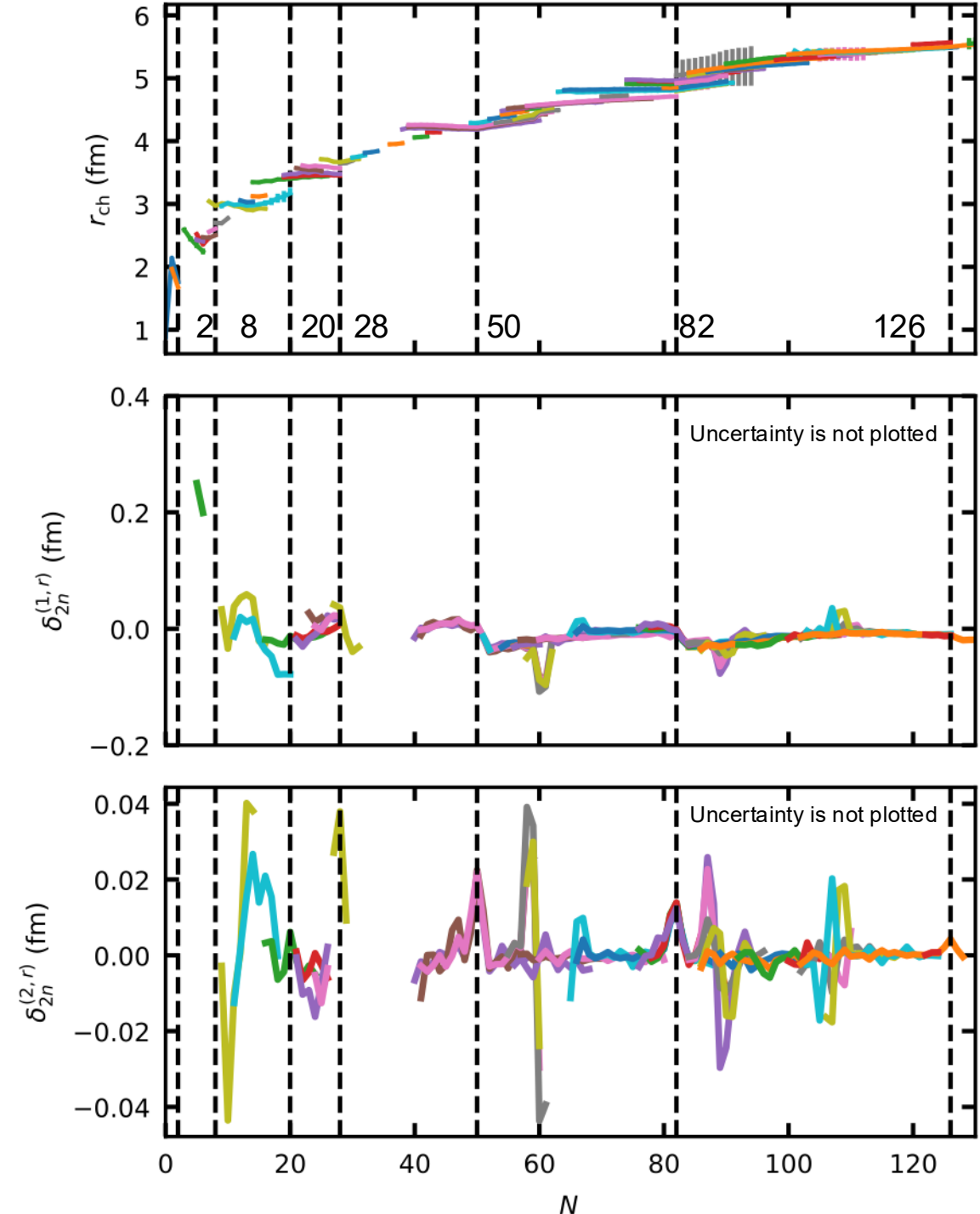
Energies and radii

- Radii are also fundamental properties of nuclei

$$\delta_{2n}^{(1,r)}(Z, N) = r_{\text{ch}}(Z, N) - r_{\text{ch}}(Z, N - 2)$$

$$\delta_{2n}^{(2,r)}(Z, N) = \frac{1}{2} [\delta_{2n}^{(1,r)}(Z, N) - \delta_{2n}^{(1,r)}(Z, N + 2)]$$

- Some shell closures, such as near ^{132}Sn and ^{208}Pb , can be detected.
- Computing Egs' and radii would be a good test of a theoretical model.



Ca radii puzzle

- Charge radii of Ca isotopes
 - 48 to 52
 - 40 to 48

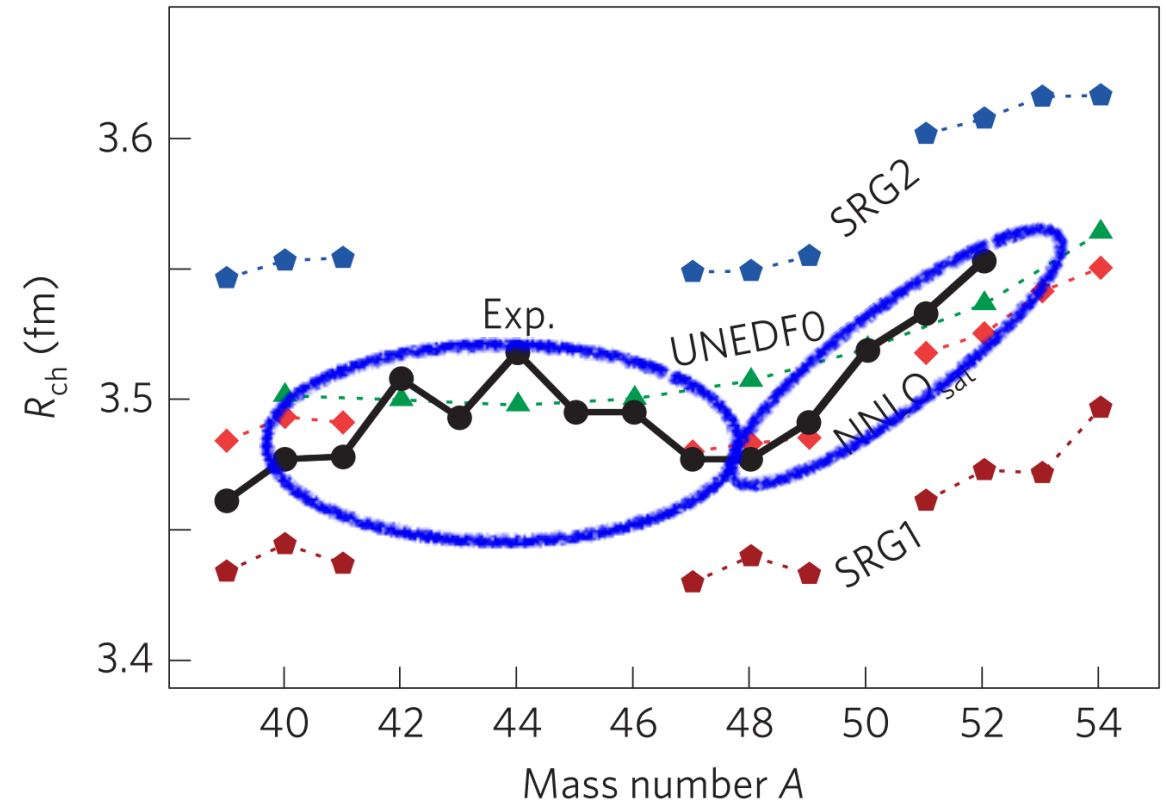
ARTICLES

PUBLISHED ONLINE: 8 FEBRUARY 2016 | DOI: 10.1038/NPHYS3645

nature
physics

Unexpectedly large charge radii of neutron-rich calcium isotopes

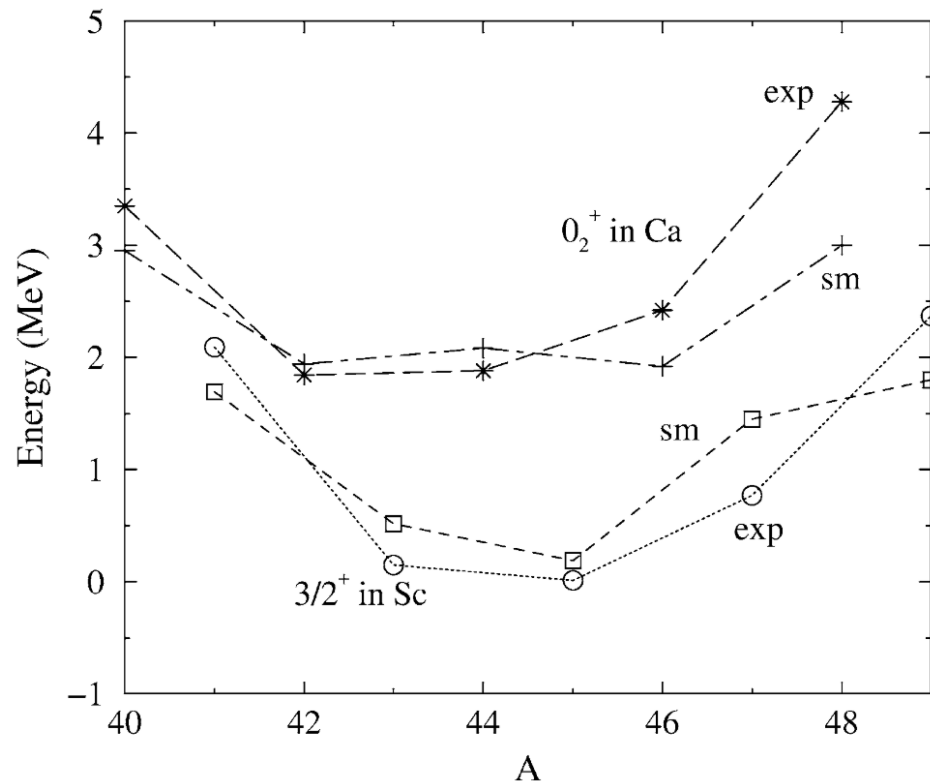
R. F. Garcia Ruiz^{1*}, M. L. Bissell^{1,2}, K. Blaum³, A. Ekström^{4,5}, N. Frömmgen⁶, G. Hagen⁴, M. Hammen⁶, K. Hebeler^{7,8}, J. D. Holt⁹, G. R. Jansen^{4,5}, M. Kowalska¹⁰, K. Kreim³, W. Nazarewicz^{4,11,12}, R. Neugart^{3,6}, G. Neyens¹, W. Nörtershäuser^{6,7}, T. Papenbrock^{4,5}, J. Papuga¹, A. Schwenk^{3,7,8}, J. Simonis^{7,8}, K. A. Wendt^{4,5} and D. T. Yordanov^{3,13}



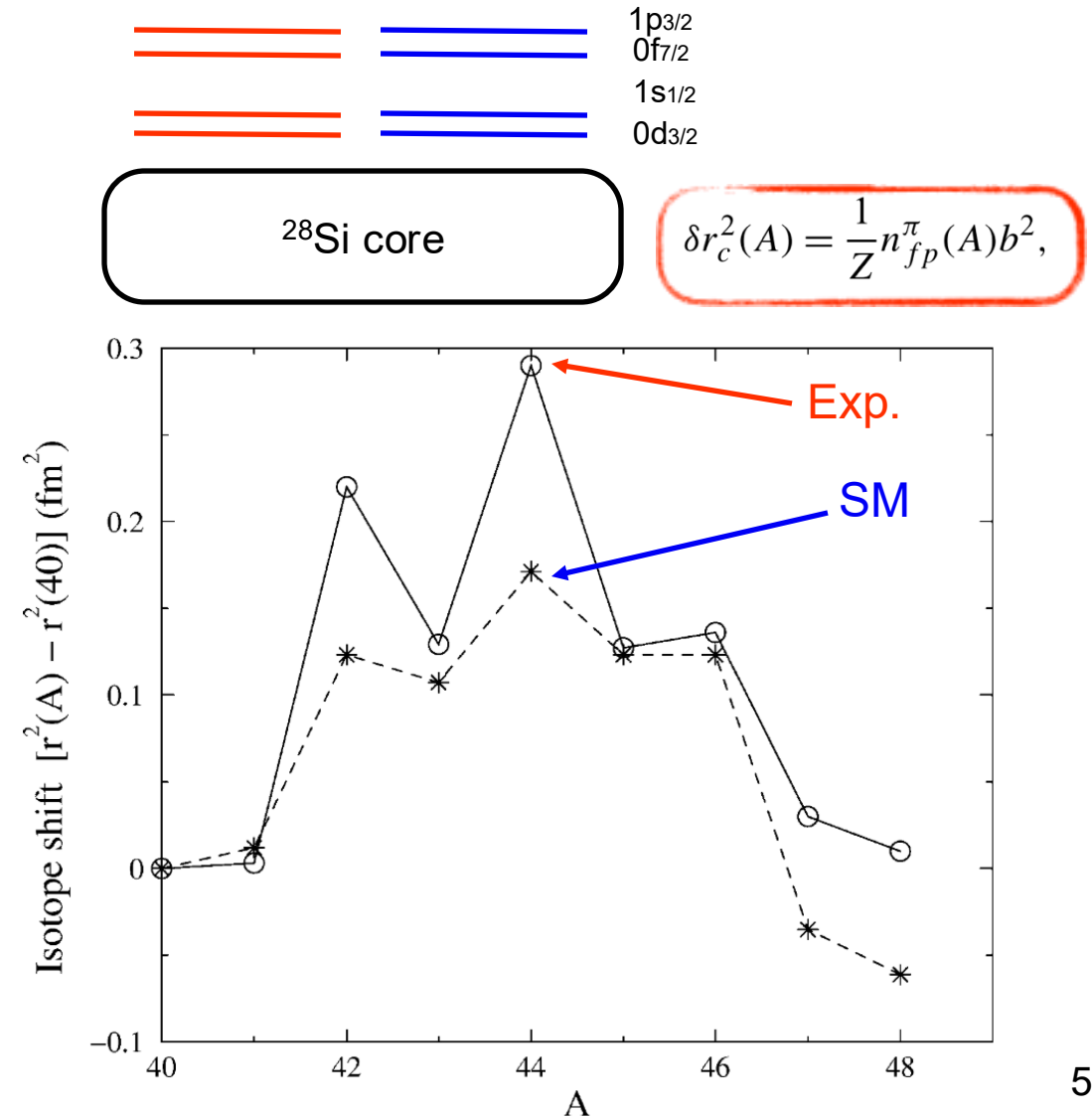
Ca radii puzzle

■ Shell-model calculation

Explicit inclusion of excitations across the N=Z=20 gap seems essential.

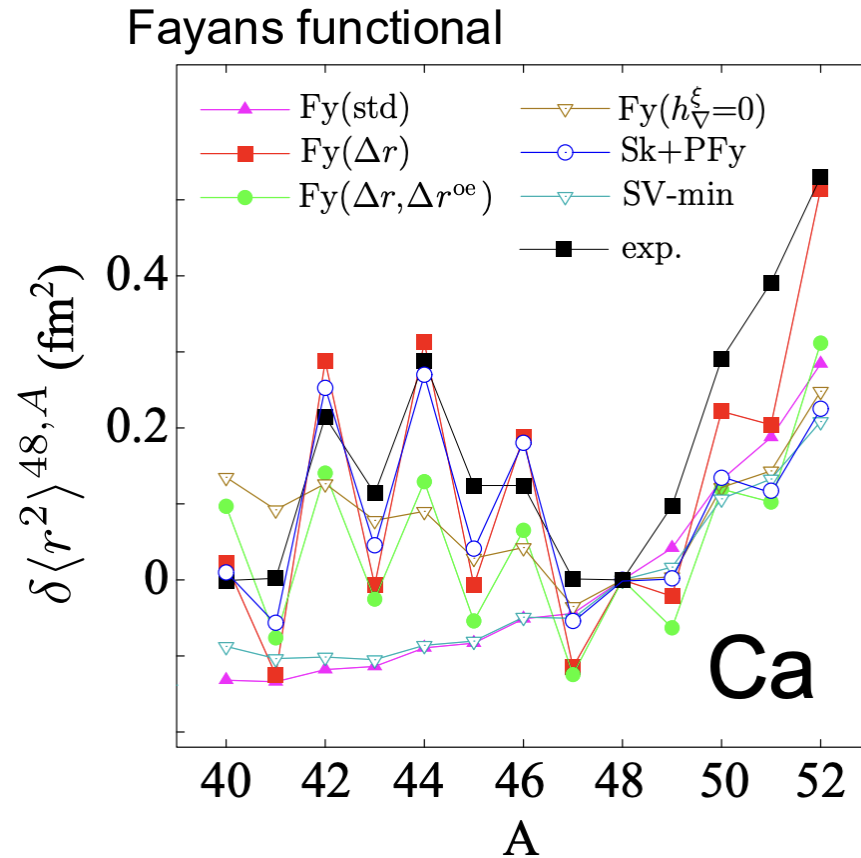


E. Caurier et al., Phys. Lett. B 522, 240 (2001).



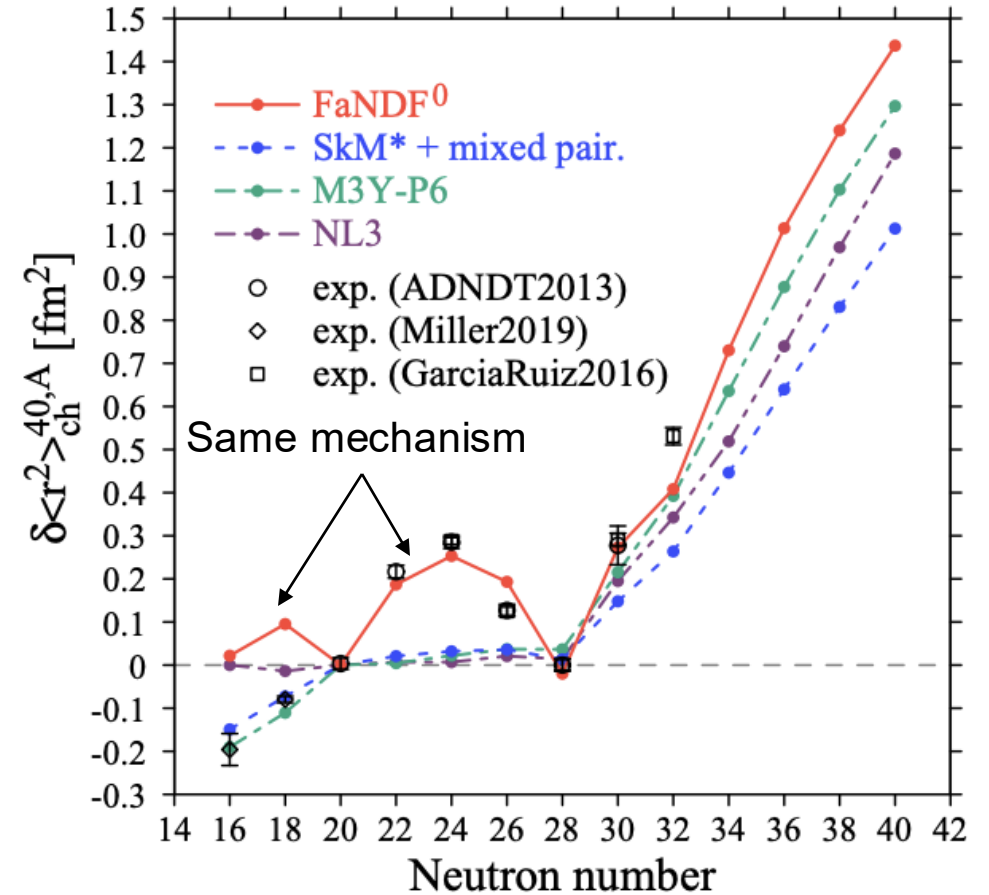
Ca radii puzzle

DFT results



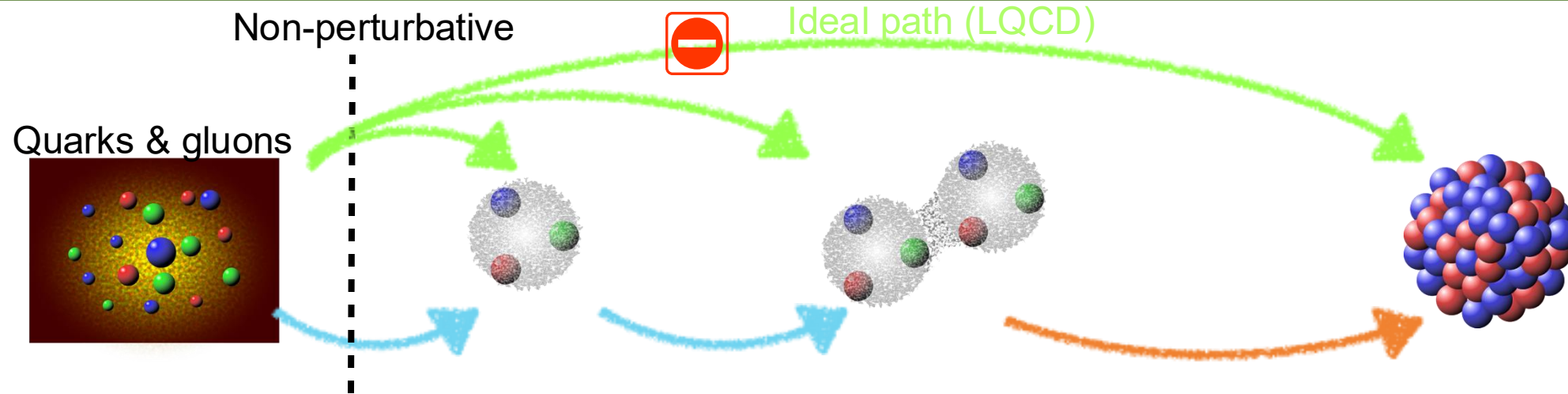
P.-G. Reinhard and W. Nazarewicz, Phys. Rev. C 95, 064328 (2017).

$20 \leq N \leq 48$ and $N < 20$ cannot be explained simultaneously.



T. Inakura, N. Hinohara, and H. Nakada,
Phys. Rev. C 110, 054315 (2024).

Nuclear ab initio calculations



	Two-nucleon force	Three-nucleon force	Four-nucleon force
LO (Q^0)		—	—
NLO (Q^2)		—	—
N ² LO (Q^3)			—
N ³ LO (Q^4)			
N ⁴ LO (Q^5)			

Nuclear many-body problem

- ♦ Green's function Monte Carlo
- ♦ No-core shell model
- ♦ Nuclear lattice effective field theory
- ♦ Self-consistent Green's function
- ♦ Coupled-cluster
- ♦ In-medium similarity renormalization group
- ♦ Many-body perturbation theory
- ♦ ...

Chiral Effective Field Theory

Weinberg, van Kolck, Kaiser, Epelbaum, Glöckle, Meißner, Entem, Machleidt, ...

- Lagrangian construction
 - ◆ Chiral symmetry
 - ◆ Power counting
- Systematic expansion
 - ◆ Unknown LECs
 - ◆ Many-body interactions
 - ◆ Uncertainty estimation

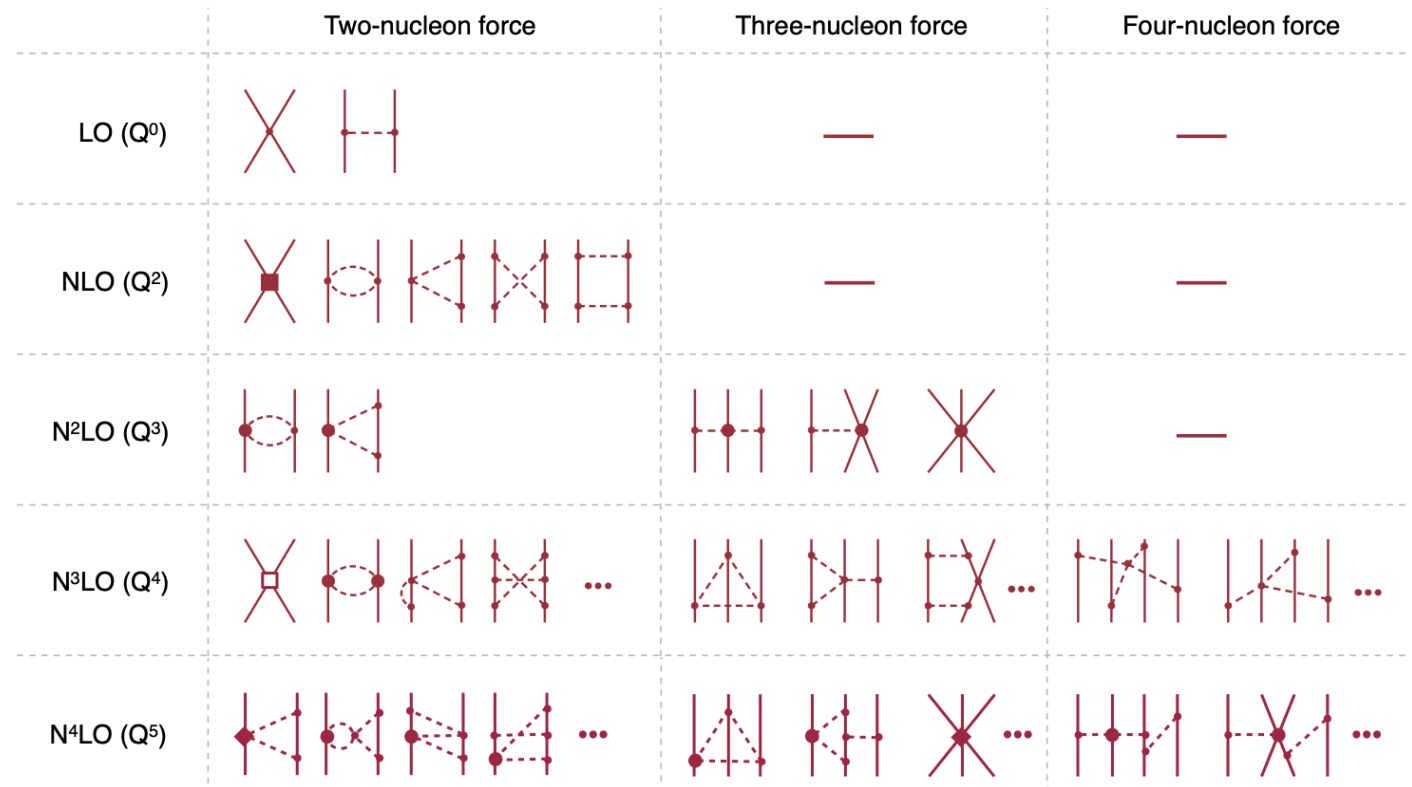
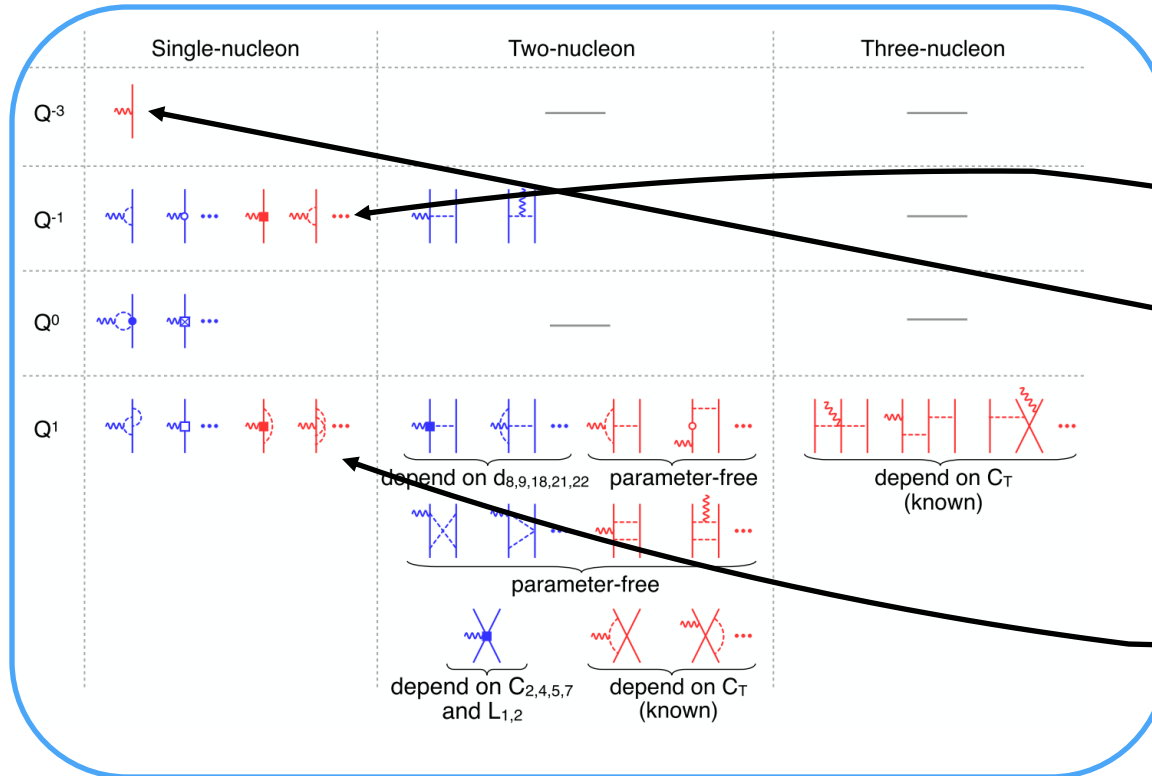


Figure is from E. Epelbaum, arXiv: 1510.07036

The 1.8/2.0 (EM) interaction will be used

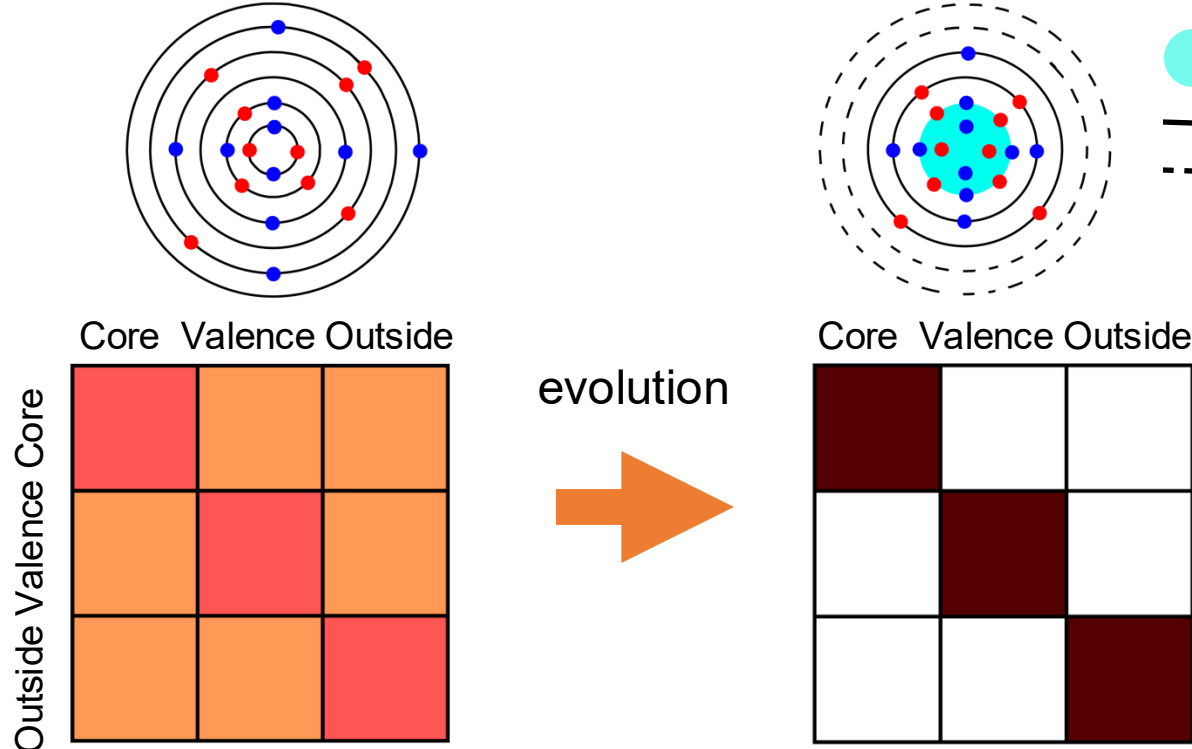
Nuclear currents from chiral EFT

- The charge radius is measure through EM probes.
- Chiral EFT allows us a systematic expansion for **charge** and **current** operators.



$$r_{\text{ch}}^2 = -\frac{6}{Z} \lim_{Q \rightarrow 0} \frac{dF_{\text{ch}}(Q^2)}{dQ^2}, \quad F_{\text{ch}}(Q) = \frac{1}{4\pi} \int d\hat{Q} \tilde{\rho}(\mathbf{Q})$$

$$r_{\text{ch}}^2 = r_{pp}^2 + r_p^2 + \frac{N}{Z} r_n^2 + r_{\text{DF}}^2 + r_{\text{SO}}^2$$



● : frozen core
— : valence
--- : outside

$$\frac{d\Omega}{ds} = \eta(s) - \frac{1}{2}[\Omega(s), \eta(s)] + \dots$$

$$\eta(s) = \sum_{12} \eta_{12}(s) \{a_1^\dagger a_2\} + \sum_{1234} \eta_{1234}(s) \{a_1^\dagger a_2^\dagger a_4 a_3\}$$

$$\eta_{12} = \frac{1}{2} \arctan \left(\frac{2f_{12}}{f_{11} - f_{22} + \Gamma_{1212}} \right)$$

$$\eta_{1234} = \frac{1}{2} \arctan \left(\frac{2\Gamma_{1234}}{f_{11} + f_{22} - f_{33} - f_{44} + A_{1234}} \right)$$

$$A_{1234} = \Gamma_{1212} + \Gamma_{3434} - \Gamma_{1313} - \Gamma_{2424} - \Gamma_{1414} - \Gamma_{2323}$$

Similarity transformation

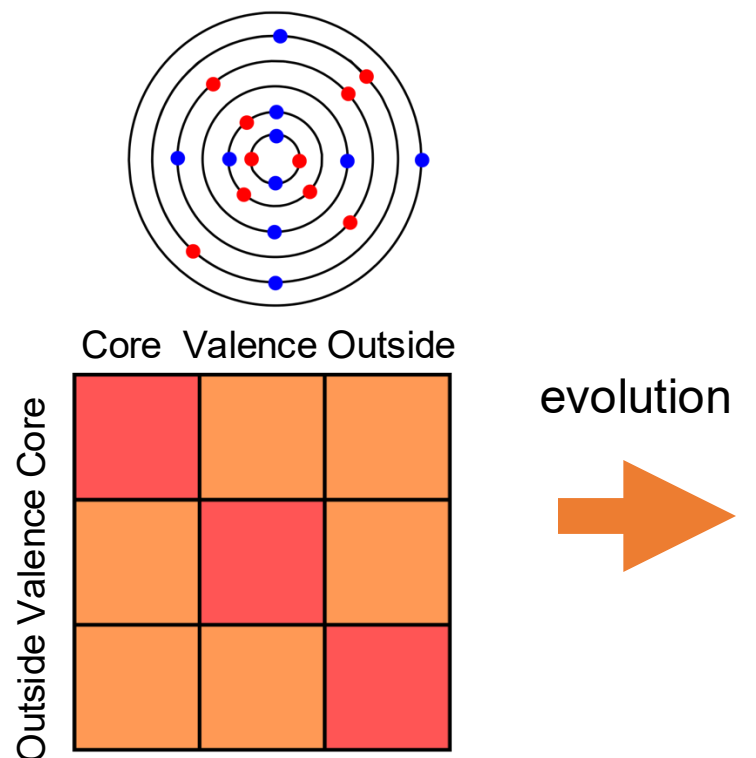
H

$$H(s) = e^{\Omega(s)} H e^{-\Omega(s)}$$

f_{12}, Γ_{1234} : matrix elements to be suppressed

$$H(s) \approx E(s) + \sum_{12} f_{12}(s) \{a_1^\dagger a_2\} + \frac{1}{4} \sum_{1234} \Gamma_{1234}(s) \{a_1^\dagger a_2^\dagger a_4 a_3\} \quad \mathcal{O}(s) = e^{\Omega(s)} \mathcal{O} e^{-\Omega(s)} \approx \mathcal{O}^{[0]}(s) + \sum_{12} \mathcal{O}_{12}^{[1]}(s) \{a_1^\dagger a_2\} + \frac{1}{4} \sum_{1234} \mathcal{O}_{1234}^{[2]}(s) \{a_1^\dagger a_2^\dagger a_4 a_3\}$$

s: flow parameter

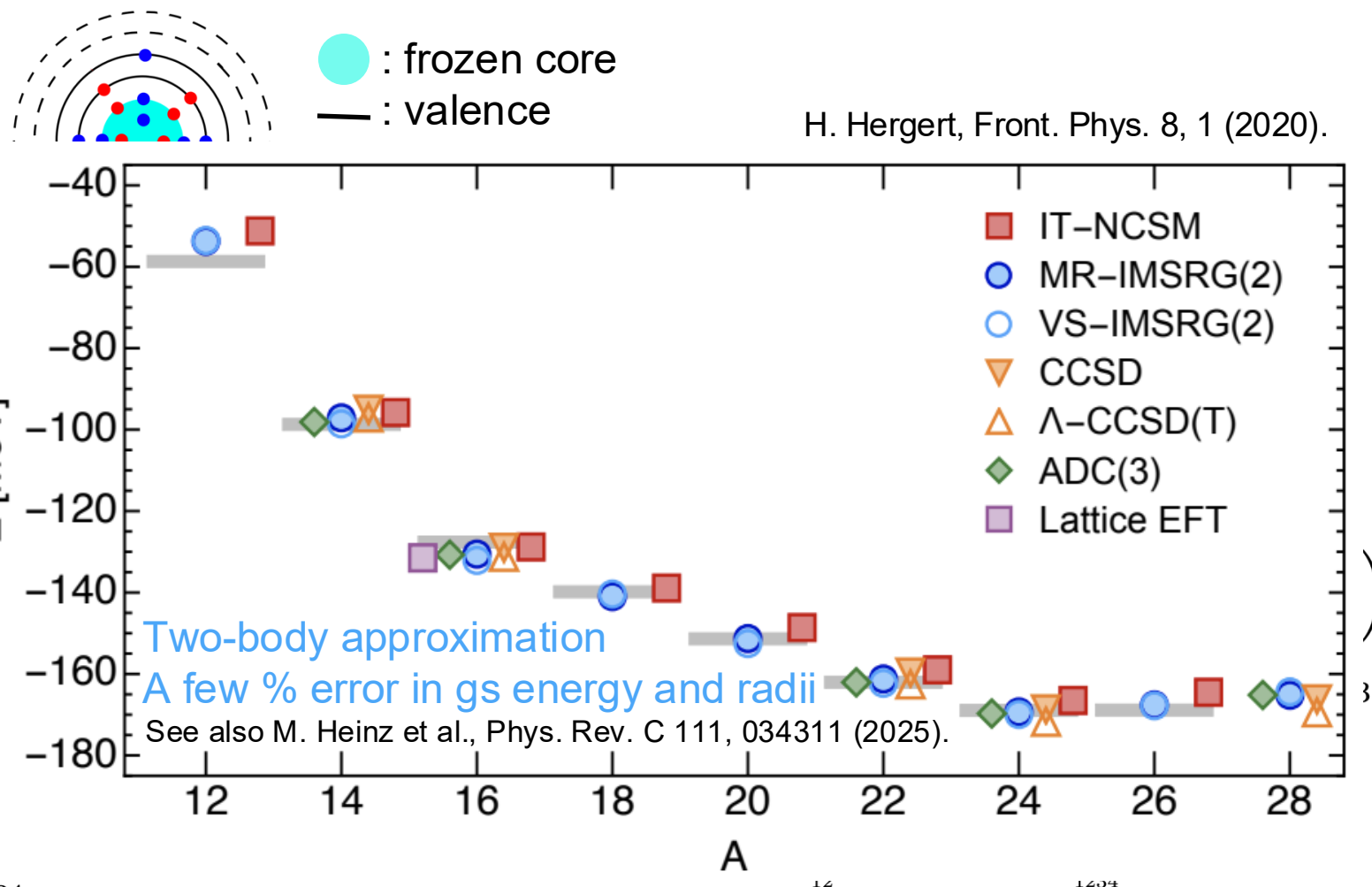


Similarity transform

H

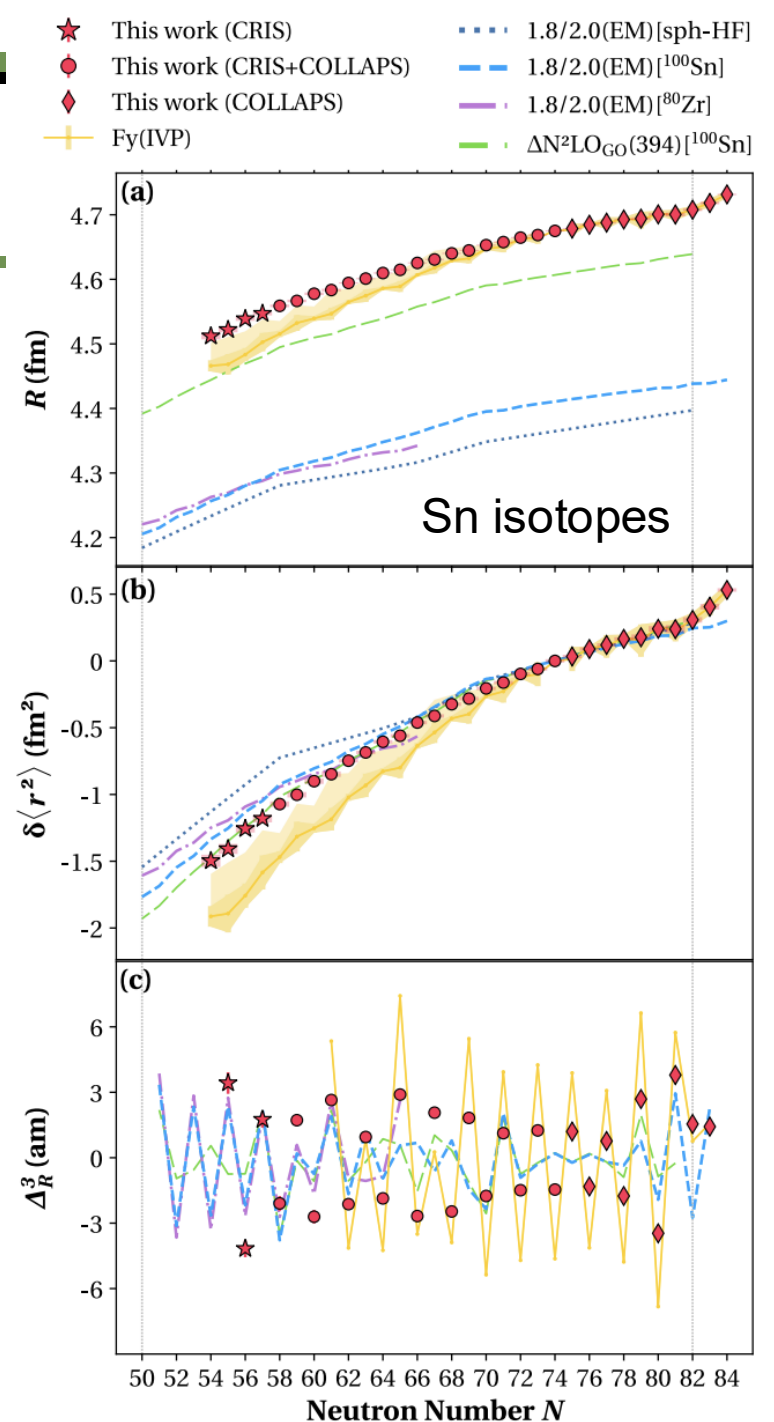
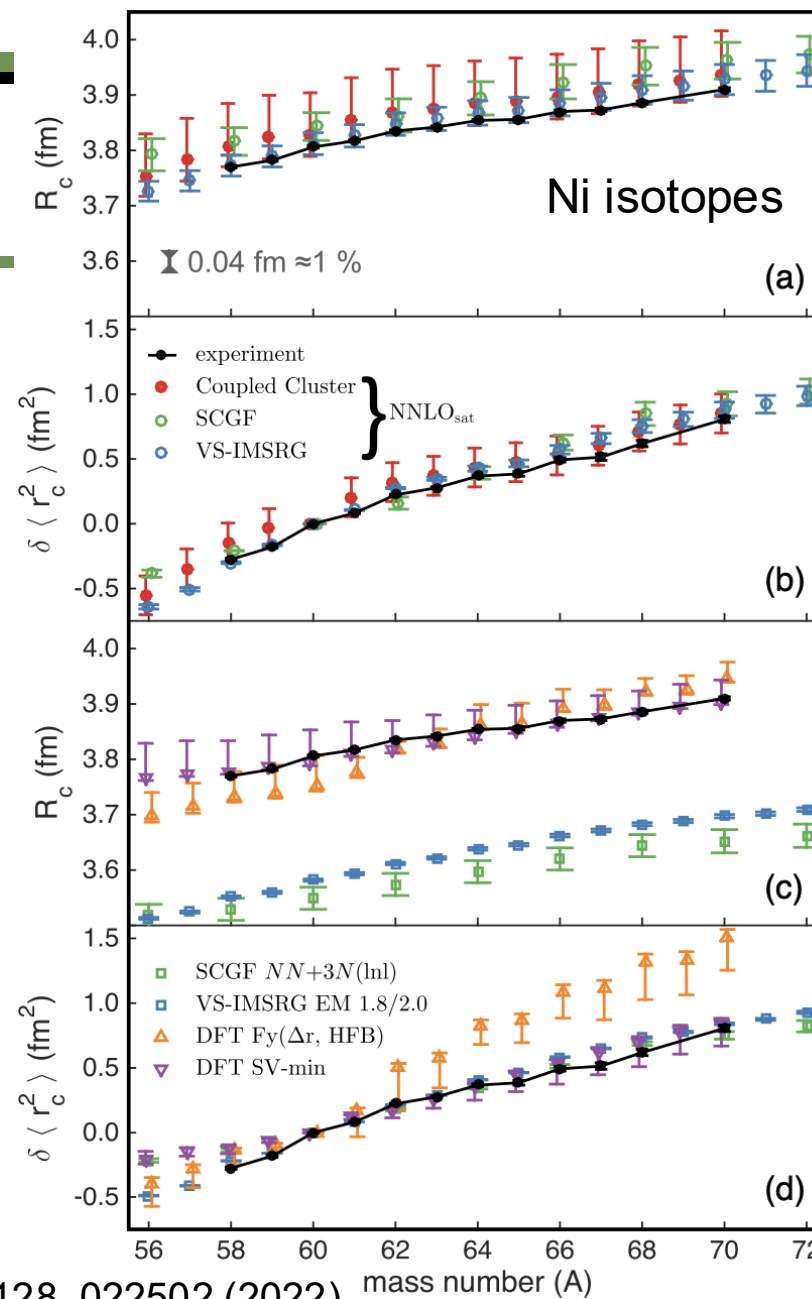
$$H(s) \approx E(s) + \sum_{12} f_{12}(s) \{a_1^\dagger a_2\} + \frac{1}{4} \sum_{1234}$$

s: flow parameter



Some radii results

- Ni and Sn isotopes
- Isotope shifts are well reproduced with the 1.8/2.0 (EM) interaction

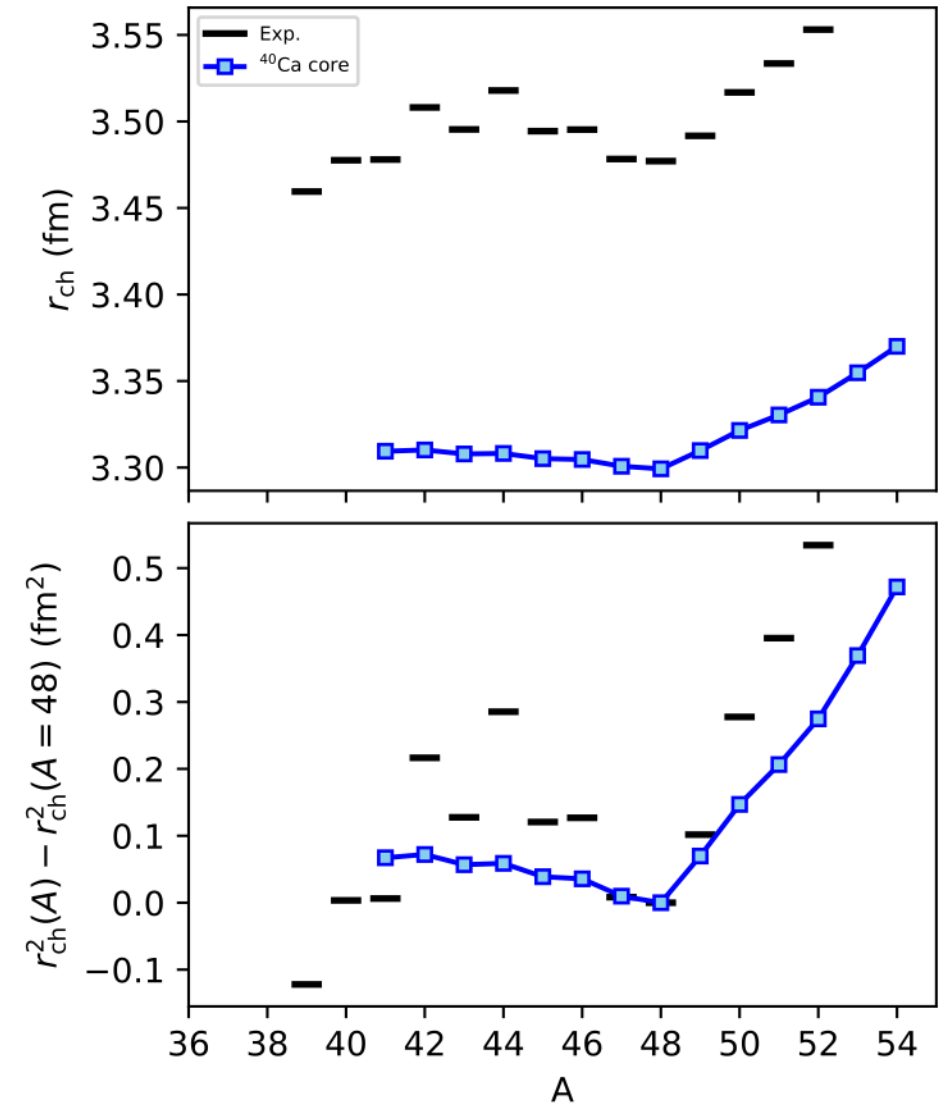


S. Malbrunot-Ettenauer et al., Phys. Rev. Lett. 128, 022502 (2022).

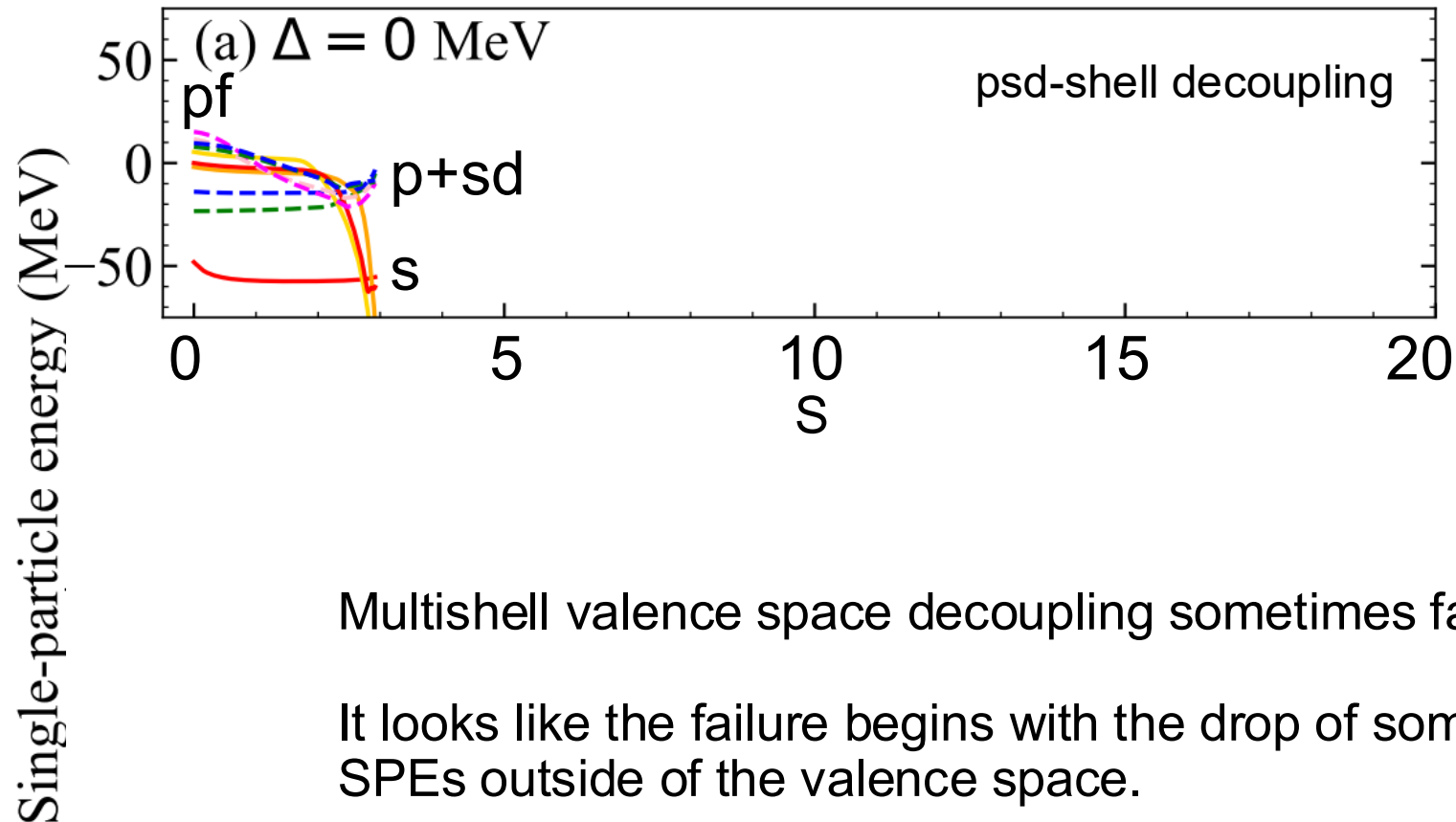
F. P. Gustafsson et al., arXiv: 2504.17060 (2025).

Ca radii with one major-shell valence space

- The computed radii are significantly smaller than the experiments due to the 1.8/2.0 (EM) interaction.
- $40 < A \leq 48$ behavior is totally off from the data.
- The enhancements in $A > 48$ can be observed, but not enough.
- Explicit excitations across $Z=N=20$ gap?
E. Caurier et al., Phys. Lett. B 522, 240 (2001).



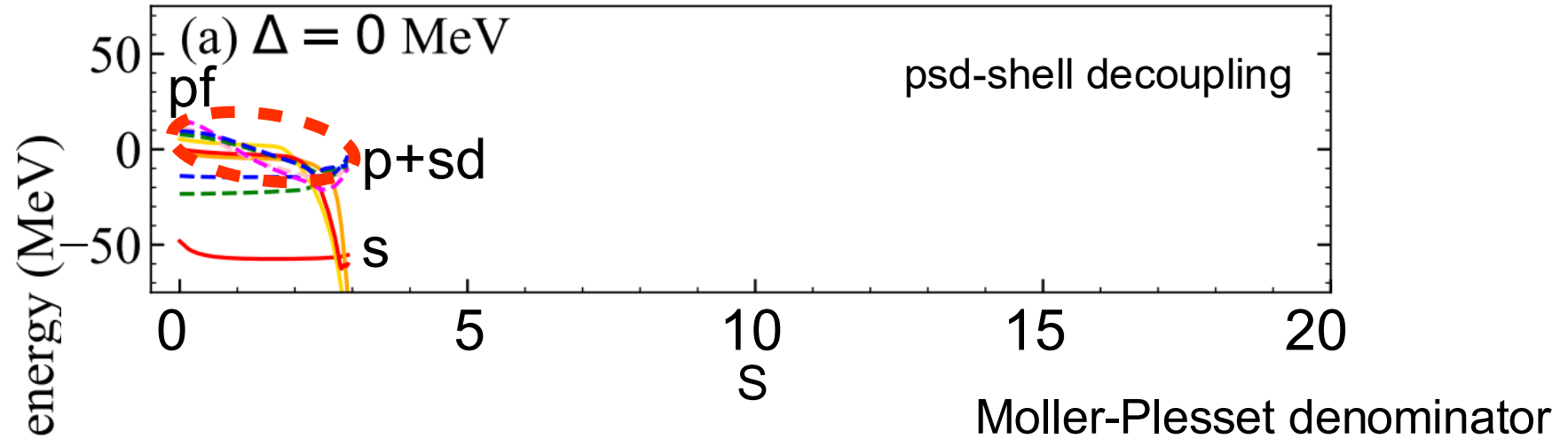
Multishell valence-space decoupling



Multishell valence space decoupling sometimes fails...

It looks like the failure begins with the drop of some SPEs outside of the valence space.

Multishell valence-space decoupling



$$\frac{df_o}{ds} = \sum_p \sum_{hh'} \frac{|\Gamma_{ophh'}|^2}{f_o + f_p - f_h - f_{h'}} \quad \text{positive}$$

$$+ \sum_{pp'} \sum_h \frac{|\Gamma_{ohpp'}|^2}{f_o + f_h - f_p - f_{p'}} \quad \text{can be negative}$$

Multishell valence-space decoupling

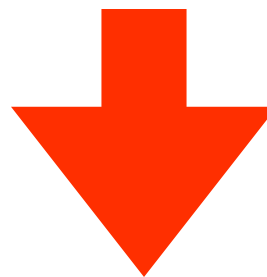
Moller-Plesset denominator

$$\eta_{12} = \frac{1}{2} \arctan \left(\frac{2f_{12}}{f_{11} - f_{22} + \Gamma_{1212}} \right)$$

$$\eta_{1234} = \frac{1}{2} \arctan \left(\frac{2\Gamma_{1234}}{f_{11} + f_{22} - f_{33} - f_{44} + A_{1234}} \right)$$

$$A_{1234} = \Gamma_{1212} + \Gamma_{3434} - \Gamma_{1313} - \Gamma_{2424} - \Gamma_{1414} - \Gamma_{2323}$$

$$\frac{df_o}{ds} = \sum_p \sum_{hh'} \frac{|\Gamma_{ophh'}|^2}{f_o + f_p - f_h - f_{h'}} + \sum_{pp'} \sum_h \frac{|\Gamma_{ohpp'}|^2}{f_o + f_h - f_p - f_{p'}}$$



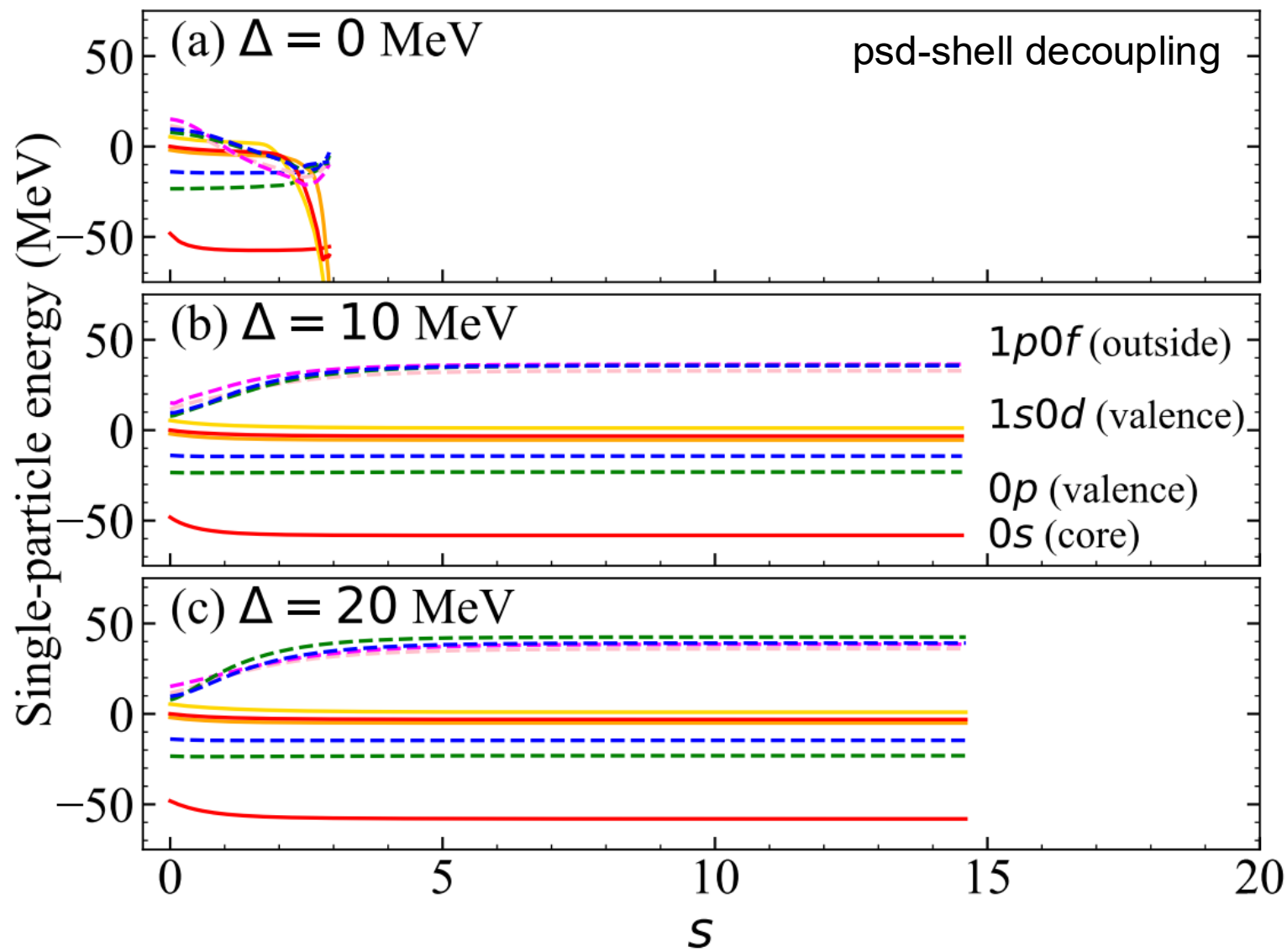
$$\eta_{12} = \frac{1}{2} \arctan \left(\frac{2f_{12}}{f_{11} - f_{22} + \Gamma_{1212} + \Delta} \right)$$

$$\eta_{1234} = \frac{1}{2} \arctan \left(\frac{2\Gamma_{1234}}{f_{11} + f_{22} - f_{33} - f_{44} + A_{1234} + \Delta} \right)$$

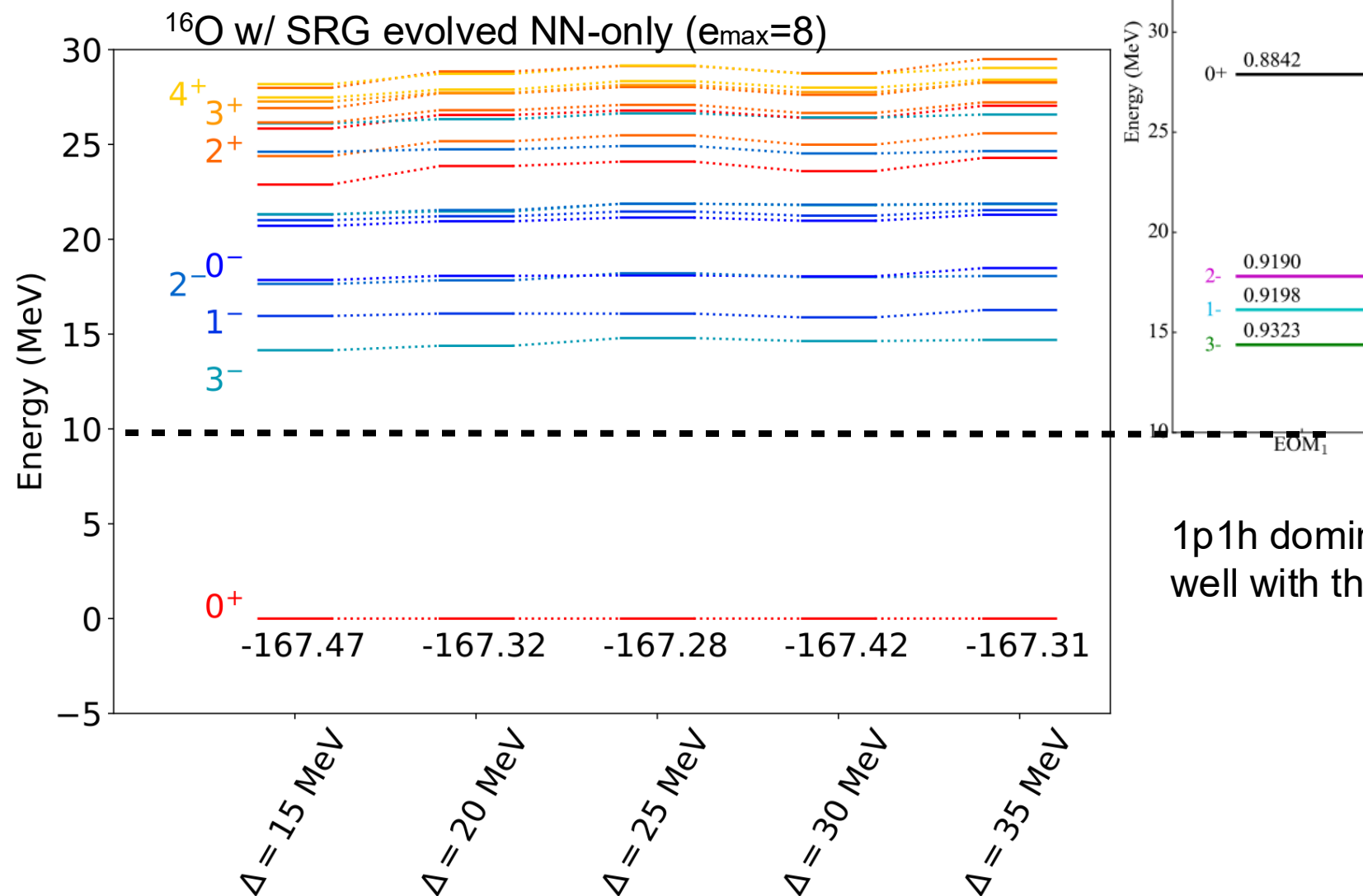
$$A_{1234} = \Gamma_{1212} + \Gamma_{3434} - \Gamma_{1313} - \Gamma_{2424} - \Gamma_{1414} - \Gamma_{2323}$$

$$\frac{df_o}{ds} = \sum_p \sum_{hh'} \frac{|\Gamma_{ophh'}|^2}{f_o + f_p - f_h - f_{h'} + \Delta} + \sum_{pp'} \sum_h \frac{|\Gamma_{ohpp'}|^2}{f_o + f_h - f_p - f_{p'} + \Delta}$$

Multishell valence-space decoupling



Multishell valence-space decoupling



1p1h dominant lowest 3^- , 1^- , 2^- agree well with the EOM-IMSIRG results.

Multishell valence-space decoupling

Center-of-mass issue

A prescription in usual shell-model diagonalization method:

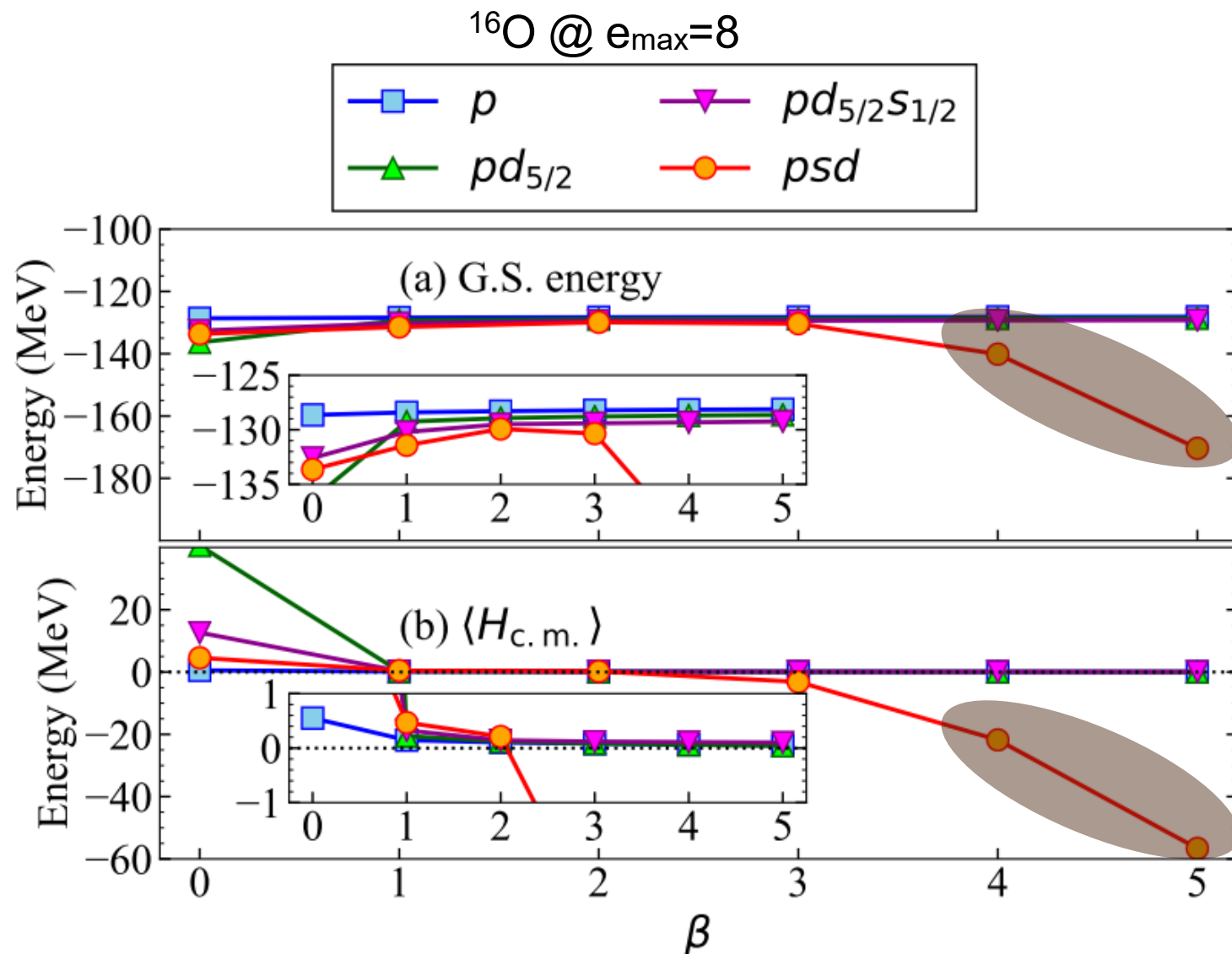
$$H_{\text{valence}} \rightarrow H_{\text{valence}} + \beta H_{\text{c.m.}}^{\text{HO}}$$

Our basis is no longer HO

$$H_{\text{intr}} \rightarrow H_{\text{intr}} + \beta H_{\text{c.m.}}^{\text{HO}} \xrightarrow{\text{IMSRG}} H_{\text{valence}}$$

IMSRG(2) breaks down

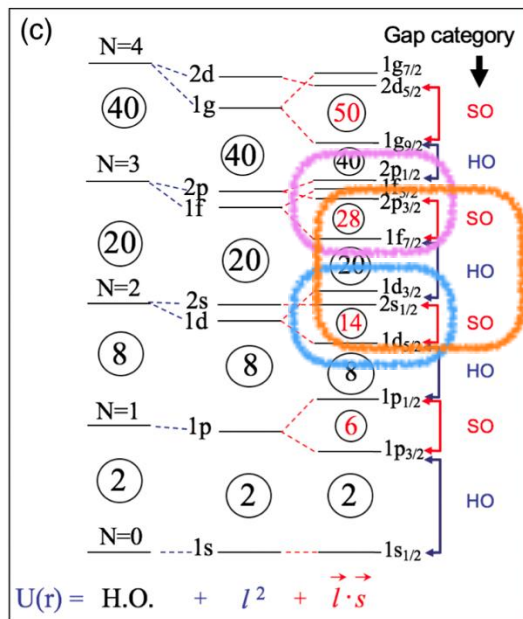
Valence-space choice is crucial



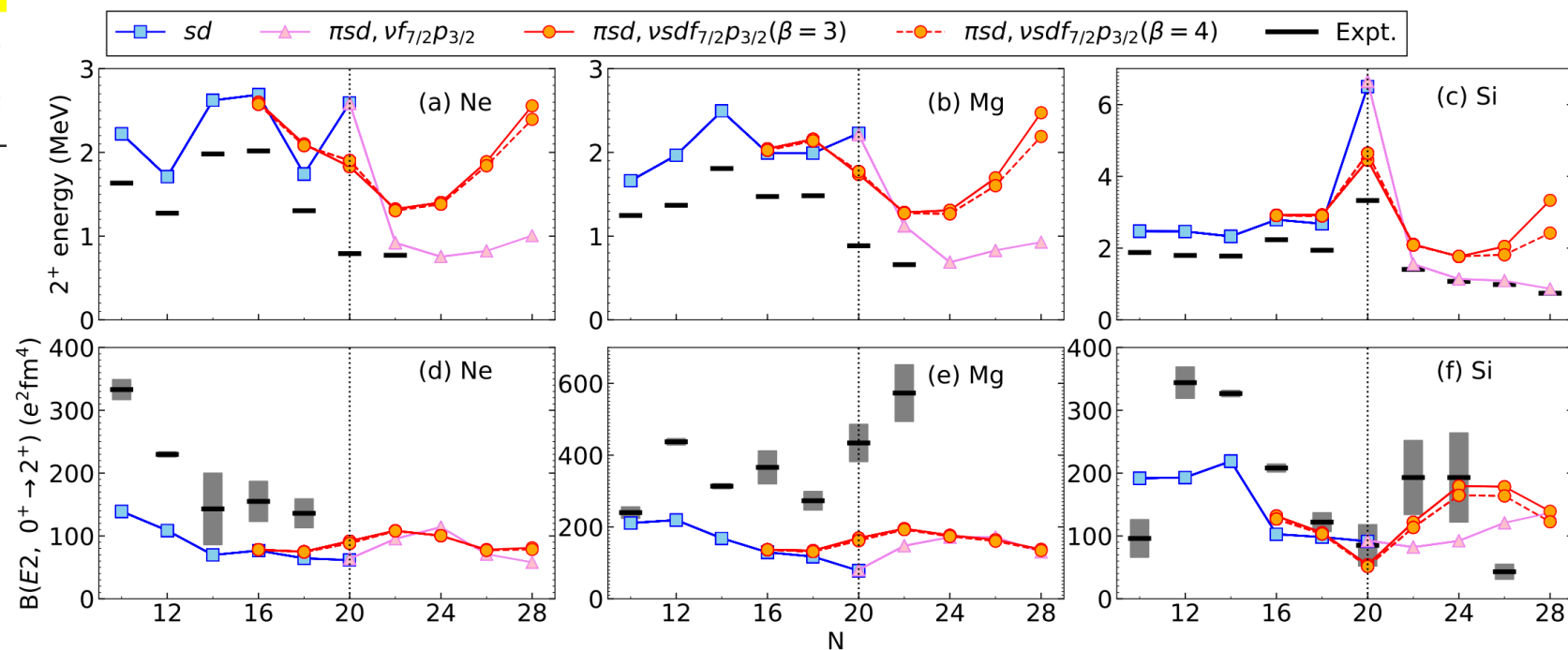
Island of inversion @ N=20



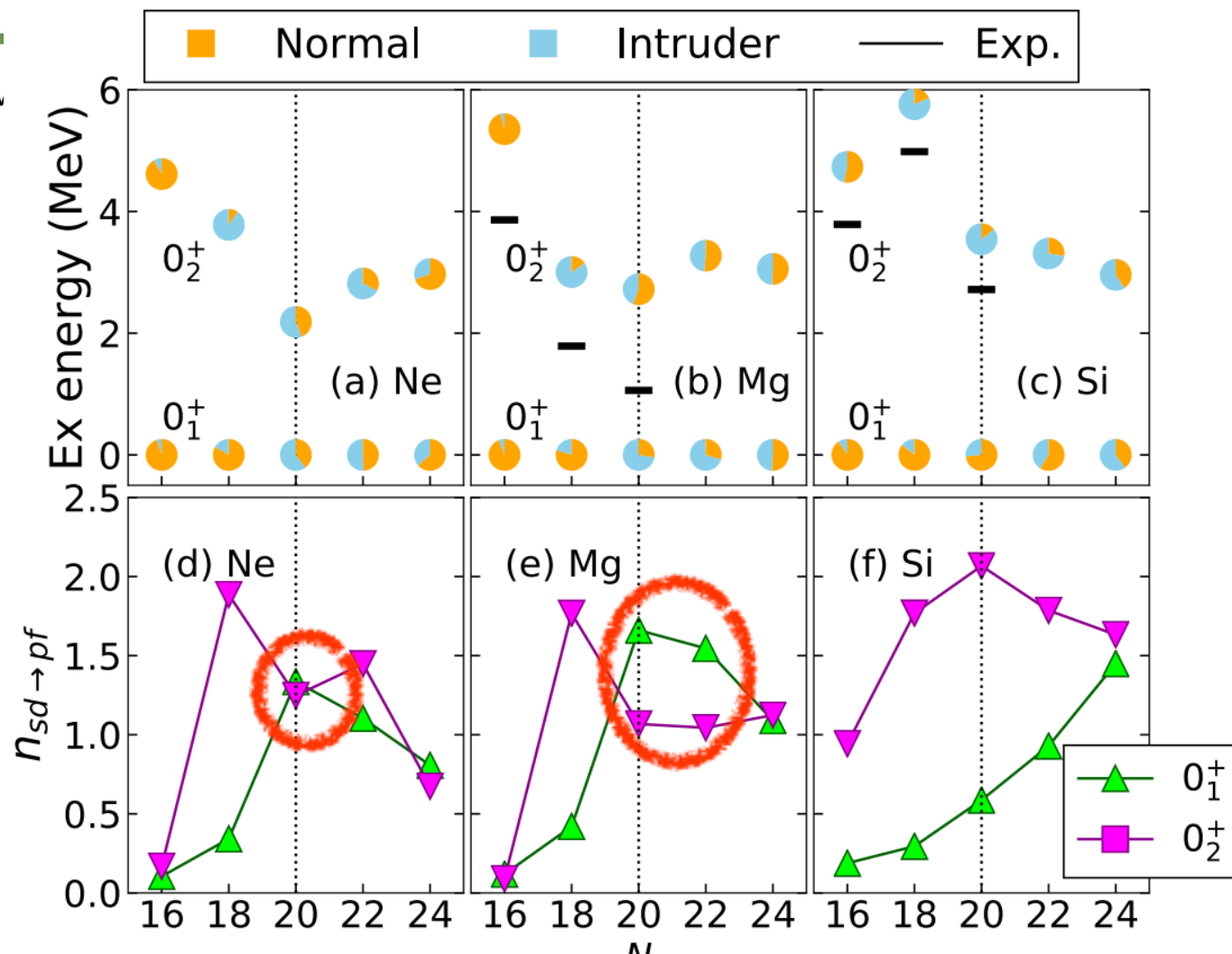
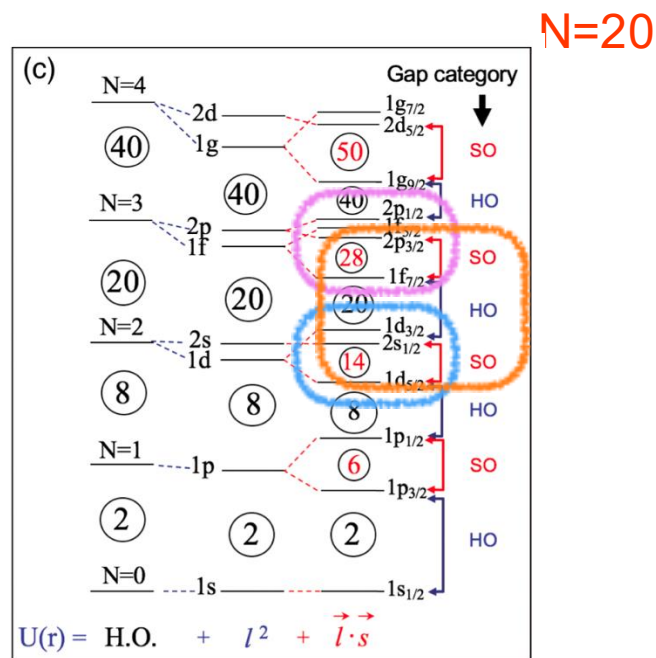
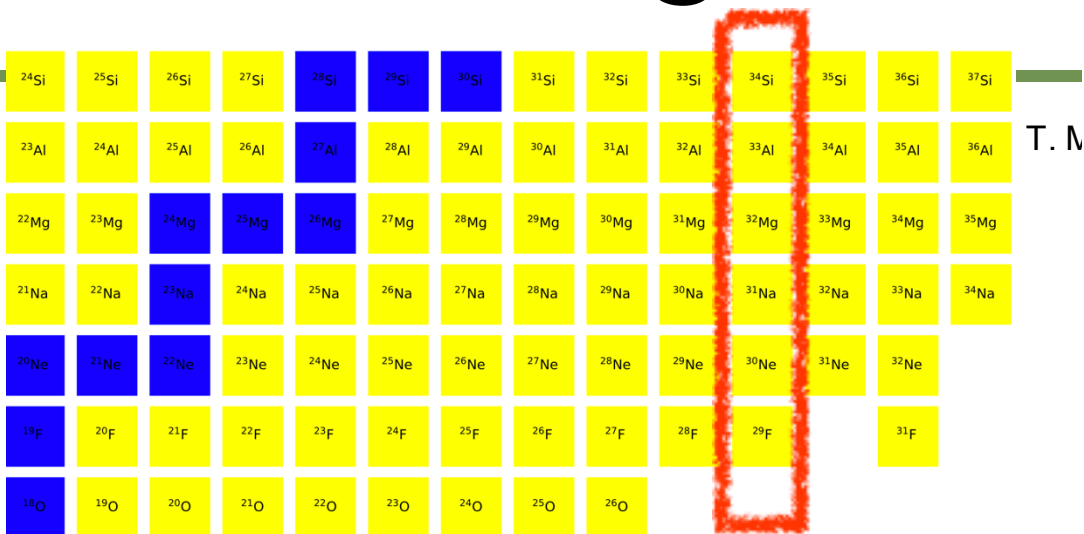
T. Miyagi, S. R. Stroberg, J. D. Holt, and N. Shimizu, Phys. Rev. C 102, 034320 (2020).



N=20



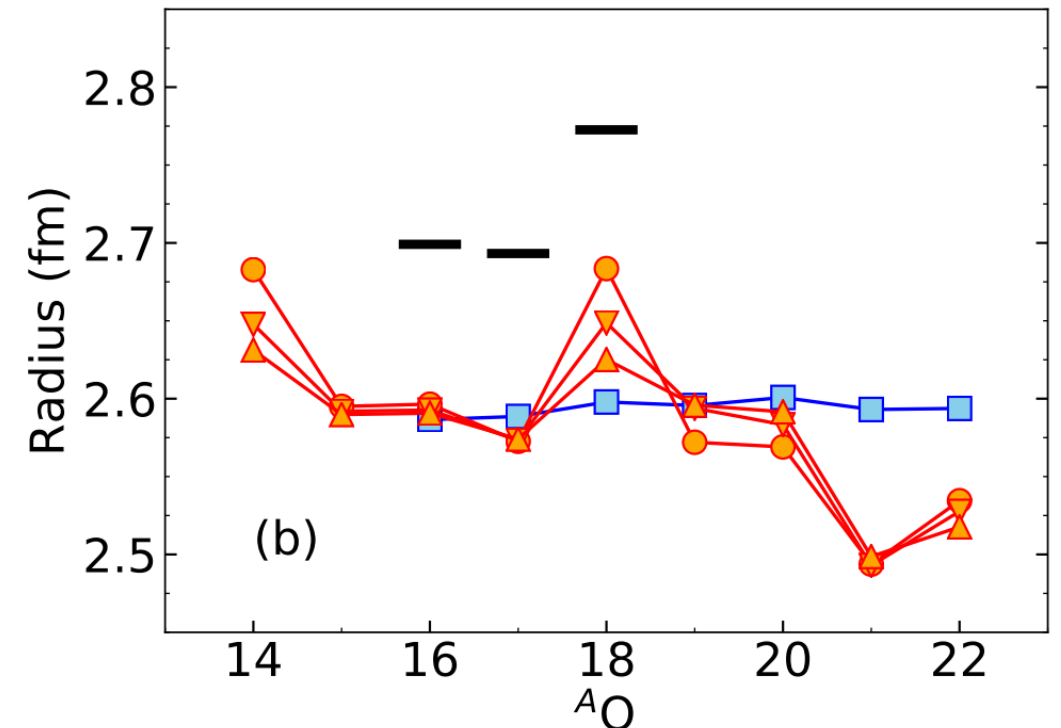
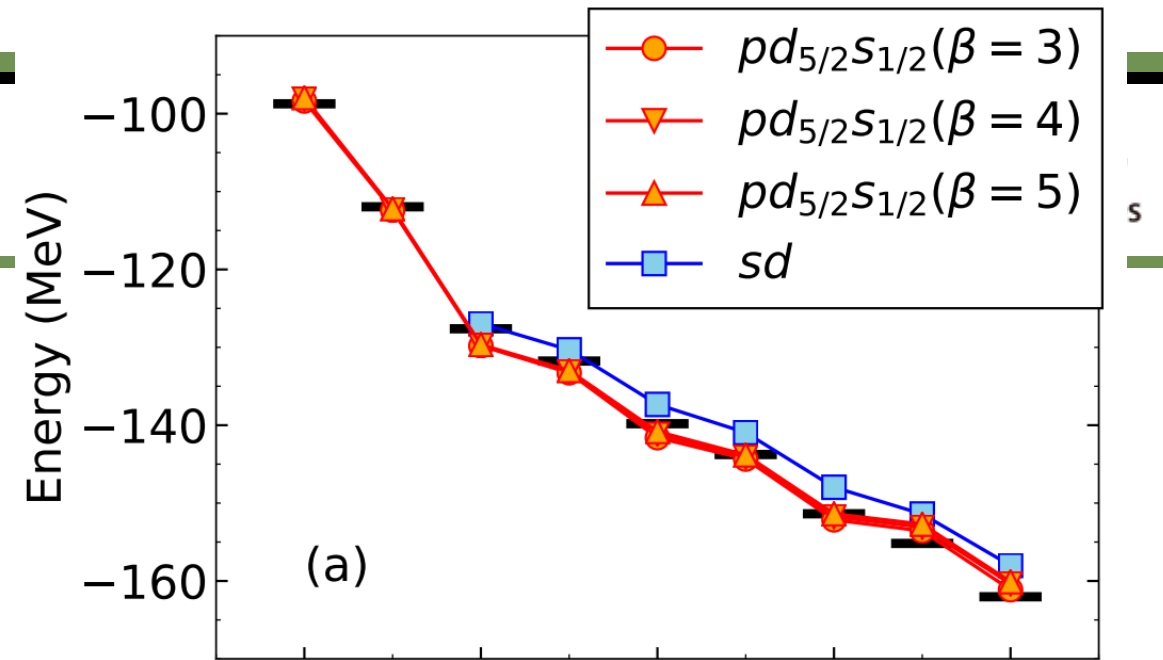
Island of inversion @ N=20



Island of inversion from underlying nuclear forces.

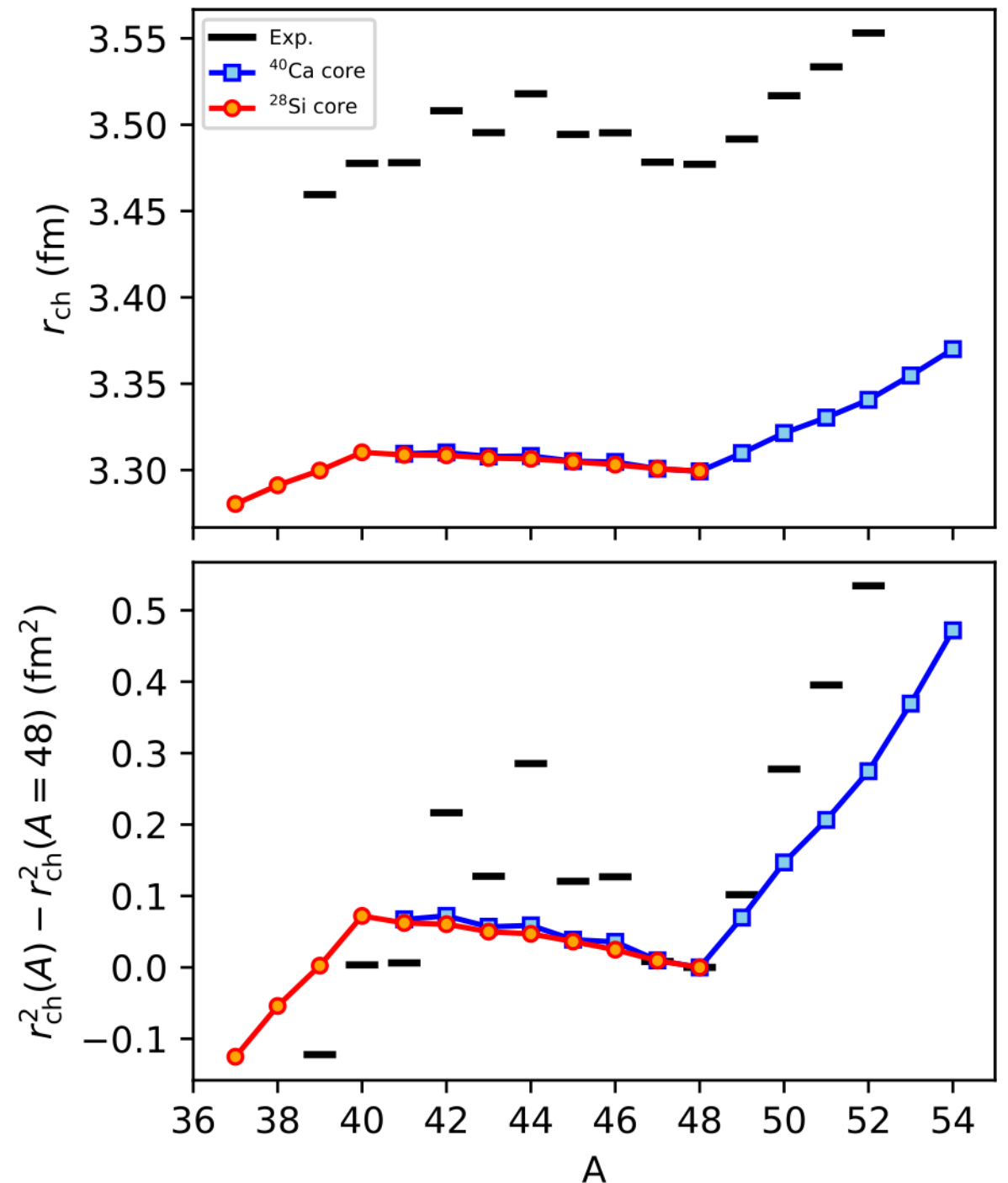
Oxygen radii

- The oxygen radii trend is also non-trivial.
- The computed radii are not fully β -independent
- $R_{\text{ch}}(A=17) < R_{\text{ch}}(A=16) < R_{\text{ch}}(A=18)$ trend can be seen in the multishell valence-space results.
- 0hw components (beta=3):
 - ♦ ^{16}O : 0.71
 - ♦ ^{17}O : 0.71
 - ♦ ^{18}O : 0.42

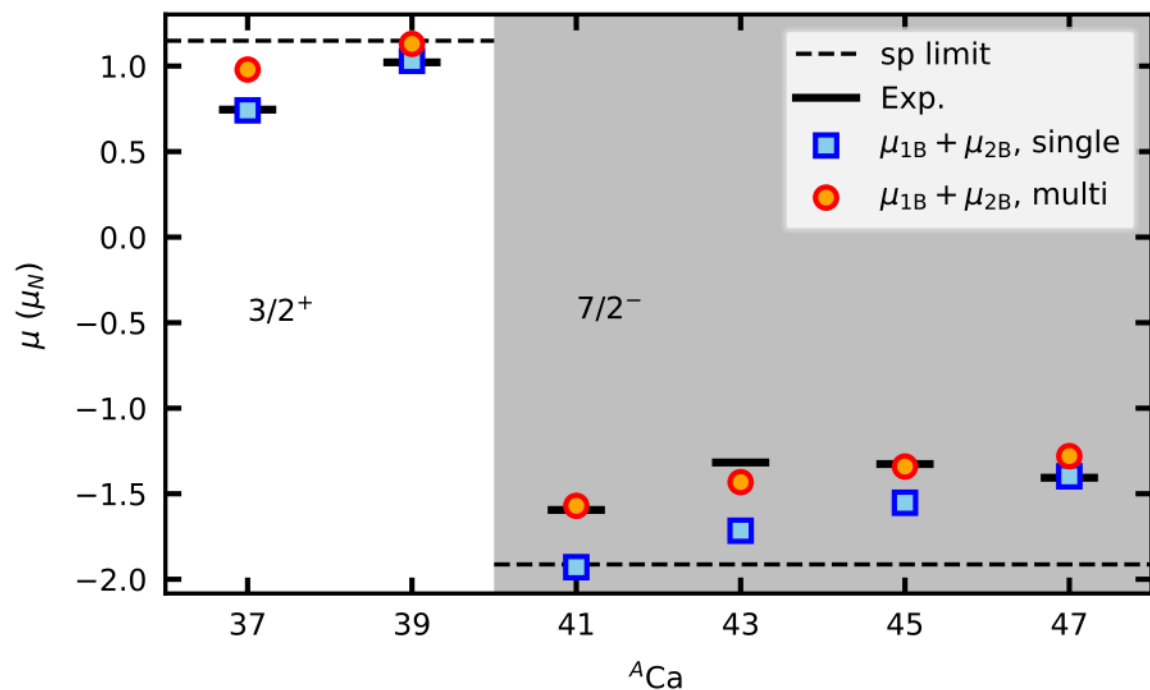


Ca radii

- The ^{40}Ca and ^{28}Si core results are almost the same.
- The beta dependence is weak.
- 0hw components:
 - ◆ ^{40}Ca : 0.49
 - ◆ ^{42}Ca : 0.50
 - ◆ ^{44}Ca : 0.54
 - ◆ ^{46}Ca : 0.65
 - ◆ ^{48}Ca : 0.81

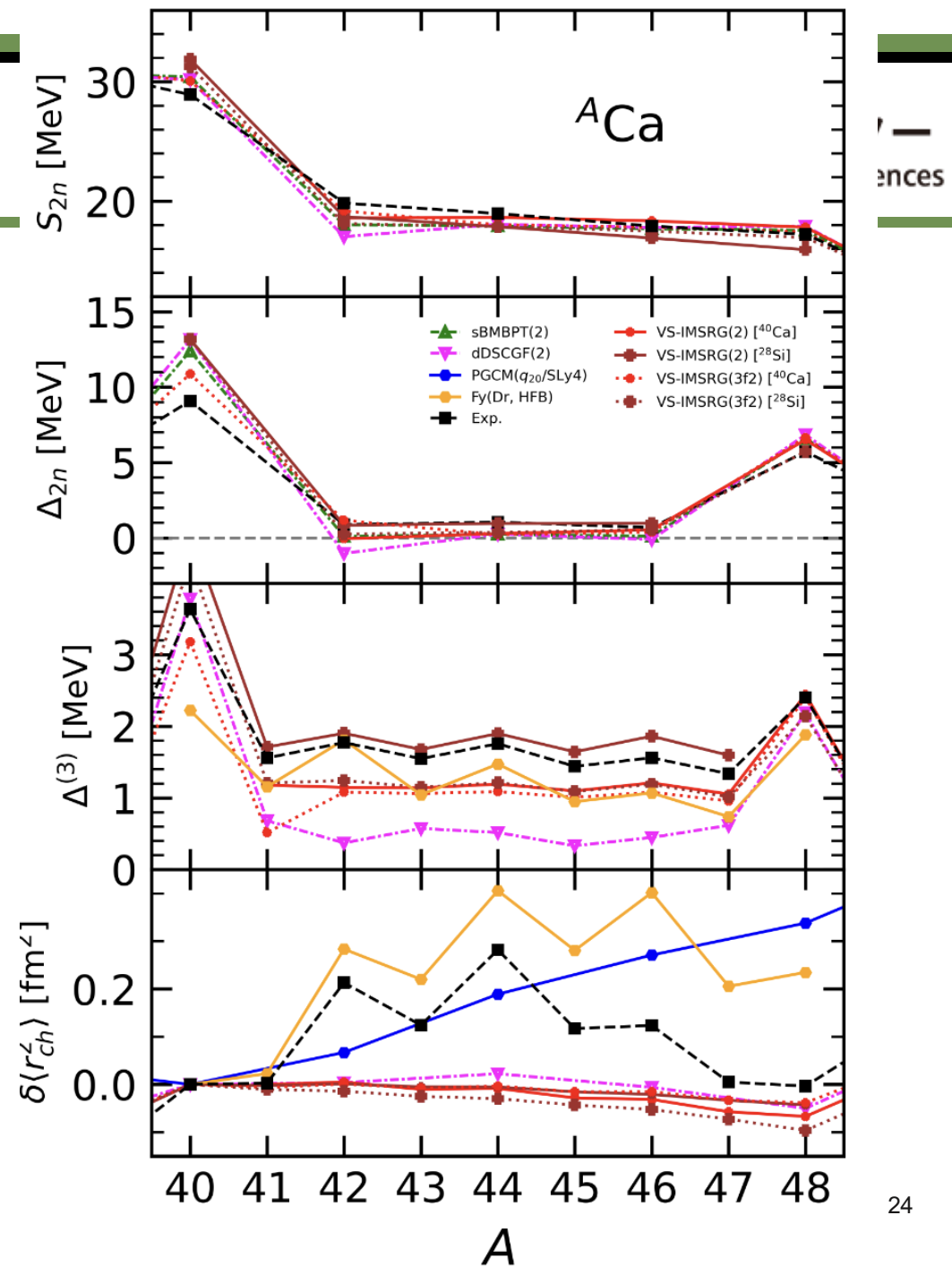


Other observables



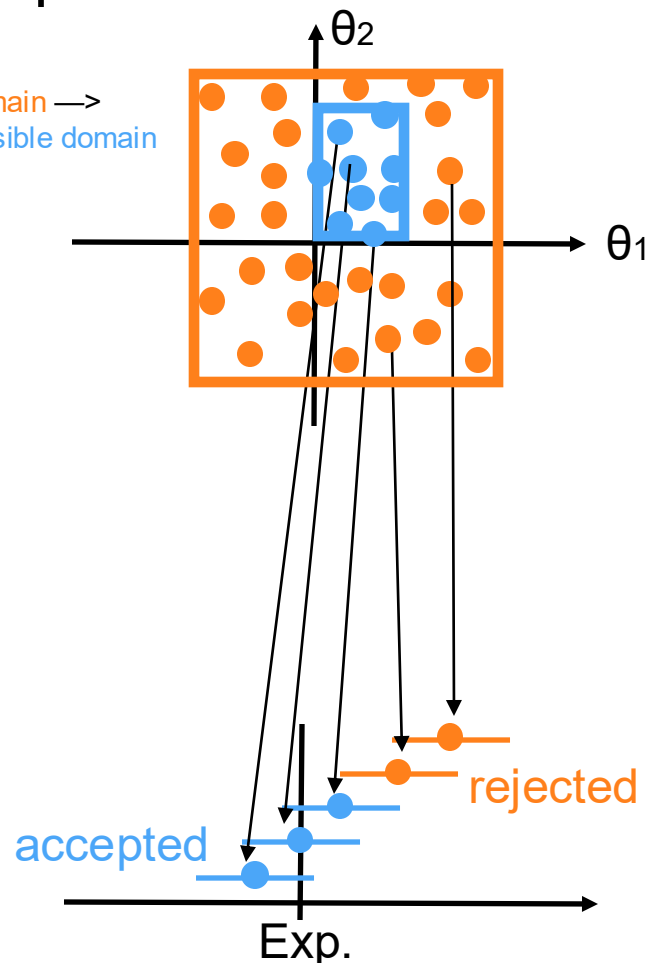
VS-IMSRG(2), 1.8/2.0 (EM)

TM et al., Phys. Rev. Lett. 132, 232503 (2024).

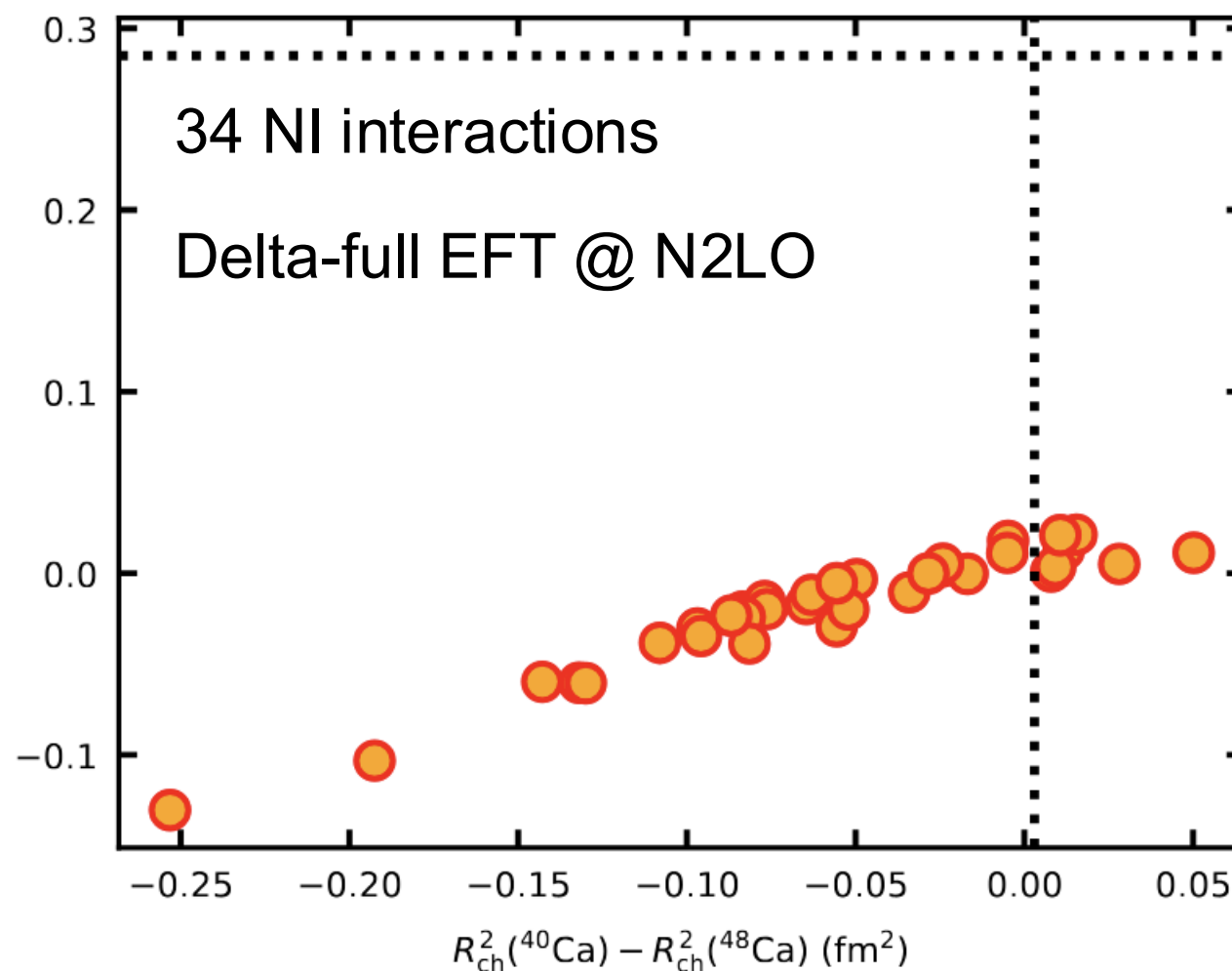


Hamiltonian issue?

Non-implausible interactions

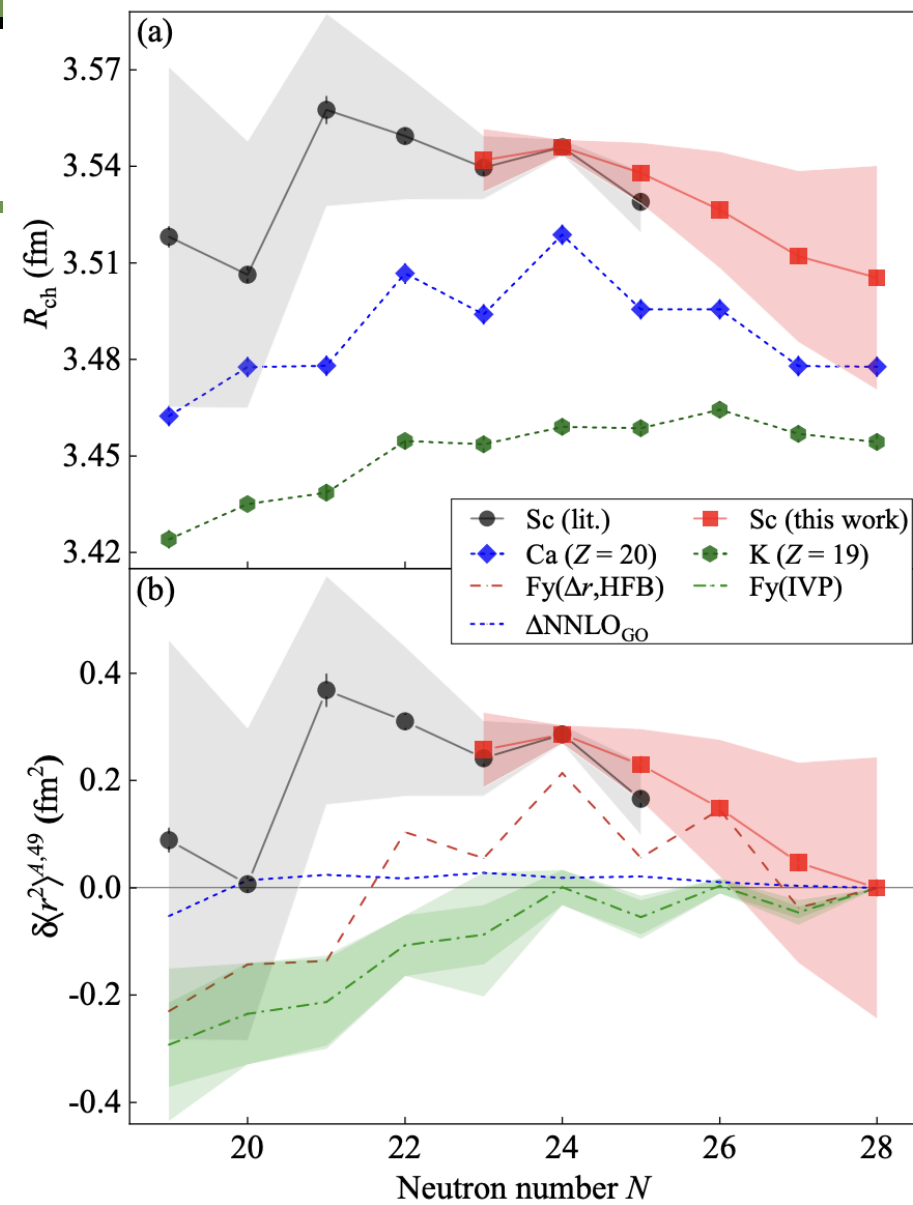
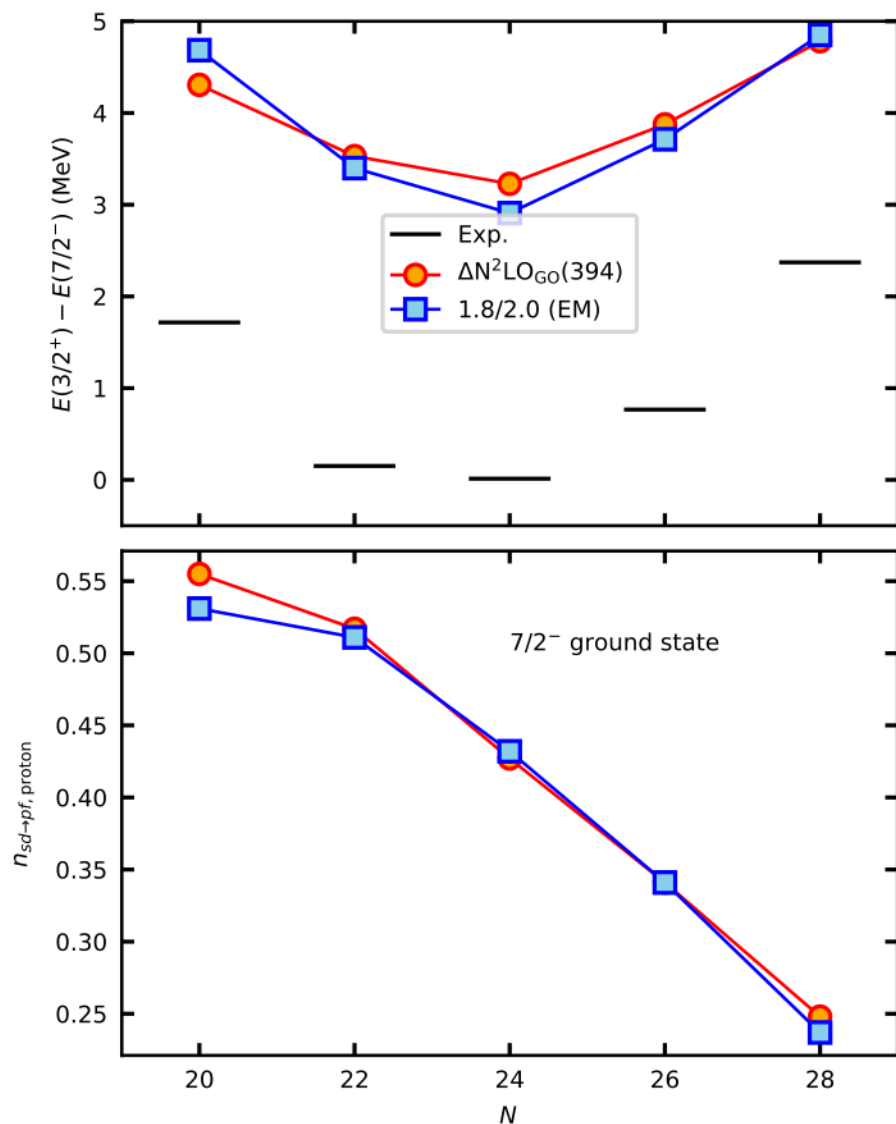


Target observable (few-body systems & ^{16}O)



VS-IMSRG(2), 34 Non-implausible interactions

Potentially related issues



S. W. Bai et al., Phys. Rev. Lett. 134, 182501 (2025).

Maybe Isomer shifts as well

- Nuclear charge radius is a fundamental observable and provides us with an insight into the nuclear structure.
- The Ca radii behavior in $40 \leq A \leq 48$ is a long standing issue.
- With VS-IMSRG, we have included explicit excitations across $Z=N=20$ gap.
 - ◆ The results are almost the same as the pf-shell calculation results
- Other many-body calculations with deformation also fail to reproduce the trend.
- Hamiltonian dependence not seems to explain it.
- What is the missing physics?