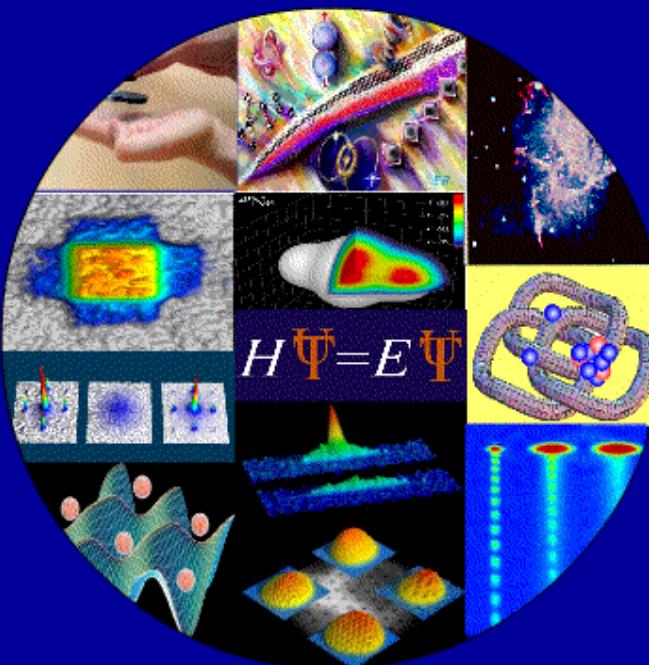


## Stochastic wavefunction approach to Hubbard-like models



Olivier JUILLET

## Hubbard-like models

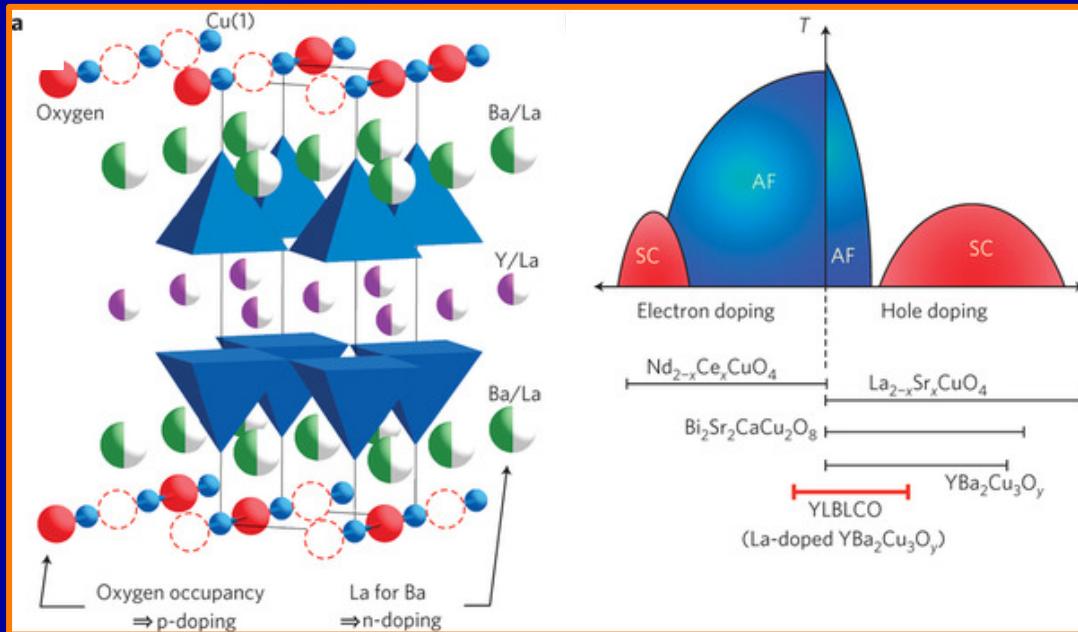
$$\hat{H} = \sum_{\vec{k}} \varepsilon_{\vec{k}} \hat{n}_{\vec{k}} + U \sum_{\vec{r}} \hat{n}_{\vec{r}\uparrow} \hat{n}_{\vec{r}\downarrow} + \sum_{\vec{r}} V_{\vec{r}} \hat{n}_{\vec{r}}$$

Kinetic  
energy

On-site  
interactions

External  
potential

### Cuprates oxides



2D repulsive Hubbard  
Hamiltonian with  
next-neighbour hopping

$$\varepsilon_{\vec{k}} = -2t [\cos(k_x) + \cos(k_y)]$$

$U > 0, V_{\vec{r}} = 0$

Low energy effective  
model of electrons in  
 $\text{CuO}_2$  planes

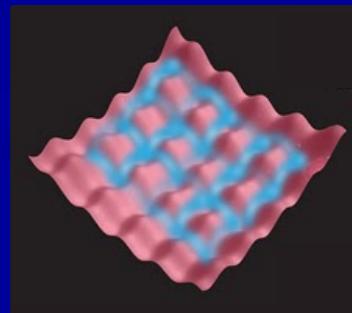
## Hubbard-like models

$$\hat{H} = \sum_{\vec{k}} \epsilon_{\vec{k}} \hat{n}_{\vec{k}} + U \sum_{\vec{r}} \hat{n}_{\vec{r}\uparrow} \hat{n}_{\vec{r}\downarrow} + \sum_{\vec{r}} V_{\vec{r}} \hat{n}_{\vec{r}}$$

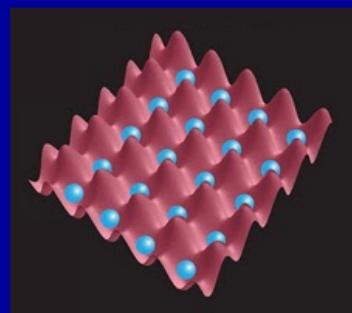
Fermionic atoms loaded in optical lattices

Delocalized states

$$U = 0$$

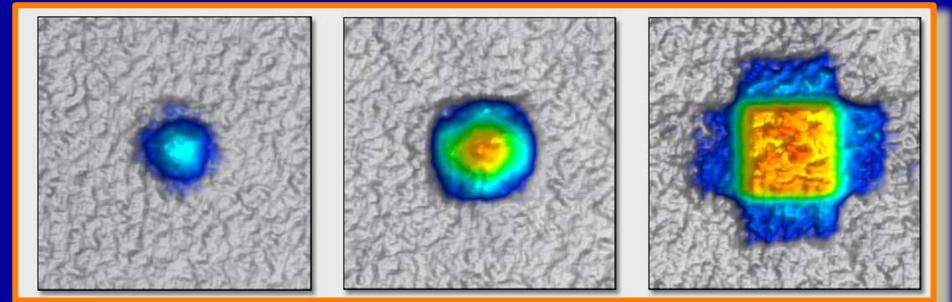


$$U \gg t$$

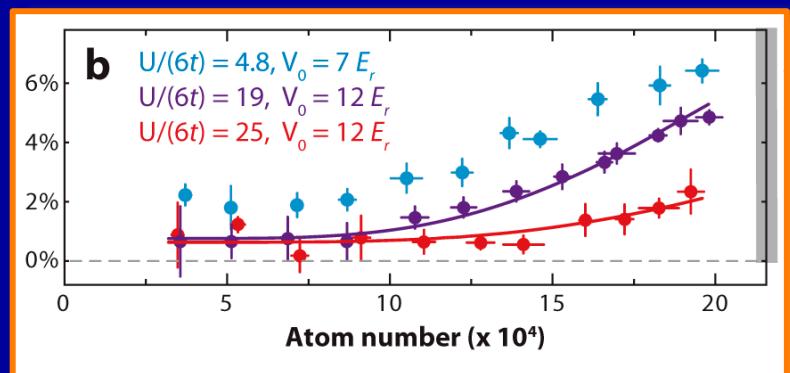


Mott insulator  
One atom per site

M. Köhl & al, Phys. Rev. Lett. (2005)



R. Jördens & al, Nature (2008)



## Hubbard-like models

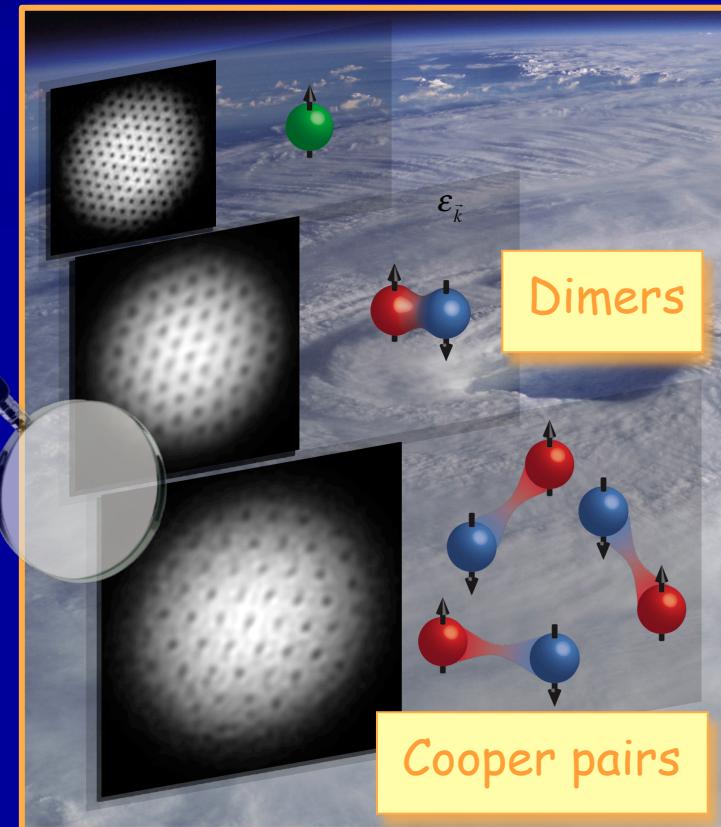
$$\hat{H} = \sum_{\vec{k}} \varepsilon_{\vec{k}} \hat{n}_{\vec{k}} + U \sum_{\vec{r}} \hat{n}_{\vec{r}\uparrow} \hat{n}_{\vec{r}\downarrow} + \sum_{\vec{r}} V_{\vec{r}} \hat{n}_{\vec{r}}$$

BEC-BCS crossover

Unitary limit



$$\sigma = \sigma_{\max} = 4\pi/k^2$$



The on-site coupling constant has to be tuned to reproduce low-energy scattering in continuous space



Dilute limit

## The stochastic wavefunction approach : an alternative to exact diagonalization

→ Hamiltonian eigenproblem : exponential complexity  $\dim(\mathcal{E}) \propto 4^{N_r}$

$$\hat{D} = \sum_{\Phi^{(a)}, \Phi^{(b)}} D_{\Phi^{(a)}, \Phi^{(b)}} |\Phi^{(a)}\rangle\langle\Phi^{(b)}|$$

Orthonormal basis of independent particle states

$$\left. \begin{array}{l} |\Phi^{(a)}\rangle = \hat{c}_{\phi_1^{(a)}}^+ \hat{c}_{\phi_2^{(a)}}^+ \cdots \hat{c}_{\phi_N^{(a)}}^+ | \rangle \\ |\Phi^{(b)}\rangle = \hat{c}_{\phi_1^{(b)}}^+ \hat{c}_{\phi_2^{(b)}}^+ \cdots \hat{c}_{\phi_N^{(b)}}^+ | \rangle \end{array} \right\} \in \left[ \begin{array}{c} \text{up} \\ \text{down} \end{array} \right] , \left[ \begin{array}{c} \text{up} \\ \text{down} \end{array} \right] , \dots , \left[ \begin{array}{c} \text{up} \\ \text{down} \end{array} \right]$$

→ Other solution : use of unrestricted single-particle wavefunctions

Positive reals

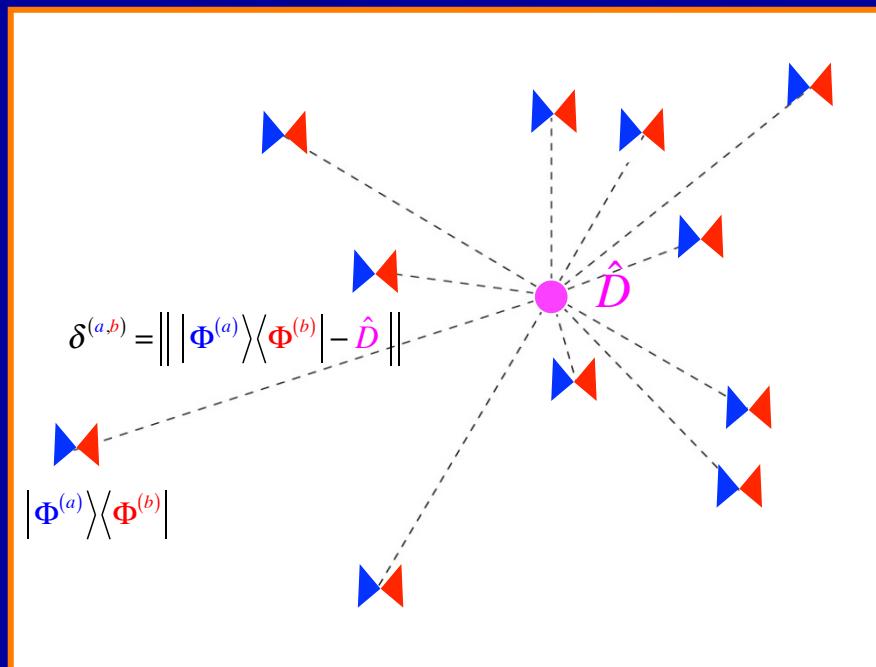
$$\hat{D} = \int_{\Phi^{(a)}, \Phi^{(b)}} \mathcal{D}\Phi^{(a)} \mathcal{D}\Phi^{(b)} P(\Phi^{(a)}, \Phi^{(b)}) |\Omega^{(a,b)}| |\Phi^{(a)}\rangle\langle\Phi^{(b)}|$$

$$= \mathbb{E} \left[ |\Omega^{(a,b)}| |\Phi^{(a)}\rangle\langle\Phi^{(b)}| \right]$$

Weighted average of  
stochastic dyadics

## The stochastic wavefunction approach : validity

The tail of the distribution for the norm of dyadics controls the statistical spread in the simulation



$$\mathbb{E}_{\Omega} \left[ \delta^{(a,b)} \right] \geq \mathbb{E}_{\Omega} \left[ \left\| \Phi^{(a)} \right\| \left\| \Phi^{(b)} \right\| \right] - \sqrt{\text{Tr}(\hat{D}^+ \hat{D})}$$

$$\mathbb{E}_{\Omega} \left[ (\delta^{(a,b)})^2 \right] = \mathbb{E}_{\Omega} \left[ \left\| \Phi^{(a)} \right\|^2 \left\| \Phi^{(b)} \right\|^2 \right] - \text{Tr}(\hat{D}^+ \hat{D})$$

Long tail scaling as  $1/\left(\left\| \Phi^{(a)} \right\| \left\| \Phi^{(b)} \right\| \right)^{1+\mu}$  implies an infinite mean error for  $\mu \leq 1$  and an infinite variance on the error if  $\mu \leq 2$ . Monte-Carlo sampling requires  $\mu > 2$  :

$$\langle \hat{O} \rangle = \text{Tr}(\hat{D} \hat{O}) = \mathbb{E}_{\Omega} \left[ \langle \Phi^{(b)} | \hat{O} | \Phi^{(a)} \rangle \right]$$

$$\sigma_O^2 = \mathbb{E}_{\Omega} \left[ \langle \Phi^{(b)} | \hat{O} | \Phi^{(a)} \rangle^2 \right] - \langle \hat{O} \rangle^2$$

$$\leq \mathbb{E}_{\Omega} \left[ \left\| \Phi^{(a)} \right\|^2 \left\| \Phi^{(b)} \right\|^2 \right] \text{Tr}(\hat{O}^2)$$

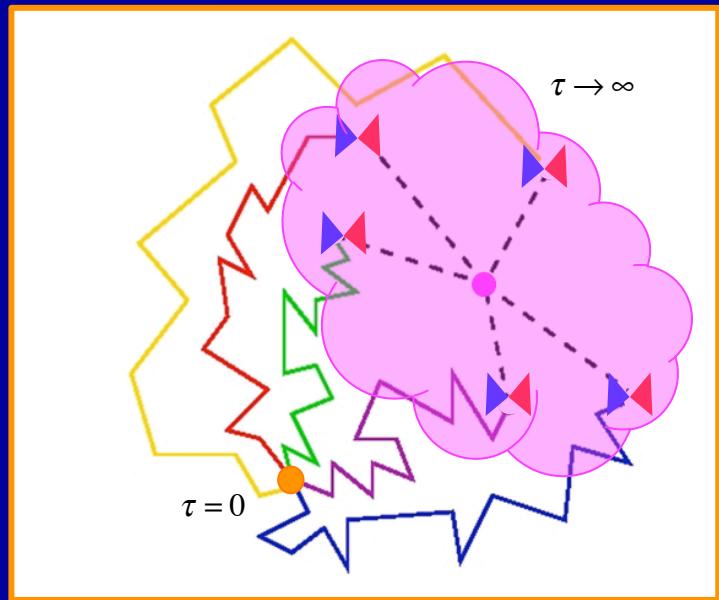
## The stochastic wavefunction approach : Ground-state reconstruction

### → Imaginary-time propagation method

$$\exp(-\tau \hat{H}/2) |\Phi_0^{(a)}\rangle \langle \Phi_0^{(b)}| \exp(-\tau \hat{H}/2) = \mathbb{E} \left[ \Omega_\tau^{(a,b)} |\Phi_\tau^{(a)}\rangle \langle \Phi_\tau^{(b)}| \right]$$

$$\propto |\Psi_g\rangle \quad \text{as } \tau \rightarrow \infty$$

$$\Rightarrow \exp(-\tau E_g) \langle \Psi_g | \Phi_0^{(a)} \rangle \langle \Phi_0^{(b)} | \Psi_g \rangle | \Psi_g \rangle \langle \Psi_g | = \mathbb{E} \left[ \Omega_\tau^{(a,b)} |\Phi_\tau^{(a)}\rangle \langle \Phi_\tau^{(b)}| \right]$$



$$\langle \hat{O} \rangle_{\Psi_g} = \frac{\mathbb{E} \left[ \Omega_\tau^{(a,b)} \langle \Phi_\tau^{(b)} | \hat{O} | \Phi_\tau^{(a)} \rangle \right]}{\mathbb{E} \left[ \Omega_\tau^{(a,b)} \langle \Phi_\tau^{(b)} | \Phi_\tau^{(a)} \rangle \right]}$$

Provided initial wavefunctions are  
not orthogonal to the ground state

## The stochastic wavefunction approach : Brownian motion in imaginary-time

A toy model : two-sites Ising Hamiltonian  $\hat{H} = \varepsilon(\hat{\sigma}_{1,z} + \hat{\sigma}_{2,z}) - J \hat{\sigma}_{1,z} \hat{\sigma}_{2,z}$

Imaginary-time dependent Schrödinger equation

$$\begin{aligned} (1 - d\tau \hat{H}/2) |\Phi^{(a)}\rangle &= |\phi_1^{(a)}\rangle \otimes |\phi_2^{(a)}\rangle \\ &\quad - \frac{d\tau}{2} (\varepsilon \hat{\sigma}_z |\phi_1^{(a)}\rangle) \otimes |\phi_2^{(a)}\rangle - \frac{d\tau}{2} |\phi_1^{(a)}\rangle \otimes (\varepsilon \hat{\sigma}_z |\phi_2^{(a)}\rangle) \\ &\quad + \frac{J d\tau}{2} (\hat{\sigma}_z |\phi_1^{(a)}\rangle) \otimes (\hat{\sigma}_z |\phi_2^{(a)}\rangle) \end{aligned}$$

Evolution of a non-interacting state

$$\begin{aligned} |\Phi^{(a)} + d\Phi^{(a)}\rangle &= |\phi_1^{(a)}\rangle \otimes |\phi_2^{(a)}\rangle \\ &\quad + |d\phi_1^{(a)}\rangle \otimes |\phi_2^{(a)}\rangle + |\phi_1^{(a)}\rangle \otimes |d\phi_2^{(a)}\rangle \\ &\quad + |d\phi_1^{(a)}\rangle \otimes |d\phi_2^{(a)}\rangle \end{aligned}$$

The exact dynamics can be recovered in average with single-particle Ito

stochastic differential equations:  $|\dot{\phi}_n^{(a)}\rangle = \left[ \left( -\frac{d\tau}{2} \varepsilon + \sqrt{\frac{J}{2}} dW \right) \hat{\sigma}_z \right] |\phi_n^{(a)}\rangle$

Wiener increments over the step  $d\tau$   
 $\mathbb{E}(dW) = 0, dW^2 = d\tau$

Fermion case - Hubbard model

$$|\dot{\phi}_n^{(a)}\rangle = \left[ -\frac{d\tau}{2} \sum_{\vec{k}\sigma} |\vec{k}\sigma\rangle \varepsilon_{\vec{k}} \langle \vec{k}\sigma| + \sqrt{\frac{|U|}{2}} \sum_{\vec{r}\sigma} |\vec{r}\sigma\rangle \text{sgn}(U) dW_{\vec{r}}^{(a)} \langle \vec{r}\sigma| \right] |\phi_n^{(a)}\rangle$$

Imaginary-time discretization



« Auxiliary-field /Determinantal QMC »

## The stochastic meanfield scheme

The solution that minimises the growth of the error on the exact ground-state

O. Juillet, Ph. Chomaz, Phys. Rev. Lett. (2002)

→ Non-linear Ito-stochastic differential equation with multiplicative noise

$$\left\langle d\phi_n^{(a)} \right\rangle = \left[ -\frac{d\tau}{2} h_{HF}(\rho^{(a)}) + (1 - \rho^{(a)}) \sqrt{\frac{|U|}{2}} \sum_{\vec{r}\sigma} |\vec{r}\sigma\rangle \text{sgn}(U) dW_{\vec{r}}^{(a)} \langle \vec{r}\sigma | \right] \left| \phi_n^{(a)} \right\rangle$$

Hartree-Fock Hamiltonian

one-body density matrix of the Slater determinant

$$\rho^{(a)} = \sum_n \left| \phi_n^{(a)} \right\rangle \left( g^{-1} \right)_{nn'} \left\langle \phi_{n'}^{(a)} \right|$$
$$\left( g_{nn'} = \left\langle \phi_n^{(a)} \middle| \phi_{n'}^{(a)} \right\rangle \right)$$

→ The distance between EACH realization and the exact solution is finite at any imaginary-time

$$\delta_\tau^{(ab)} \leq \left[ 1 + \exp\left( \frac{\tau N|U|}{2} \left( 1 - \frac{1}{N_r} \right) \right) \right] \exp(-\tau E_g)$$

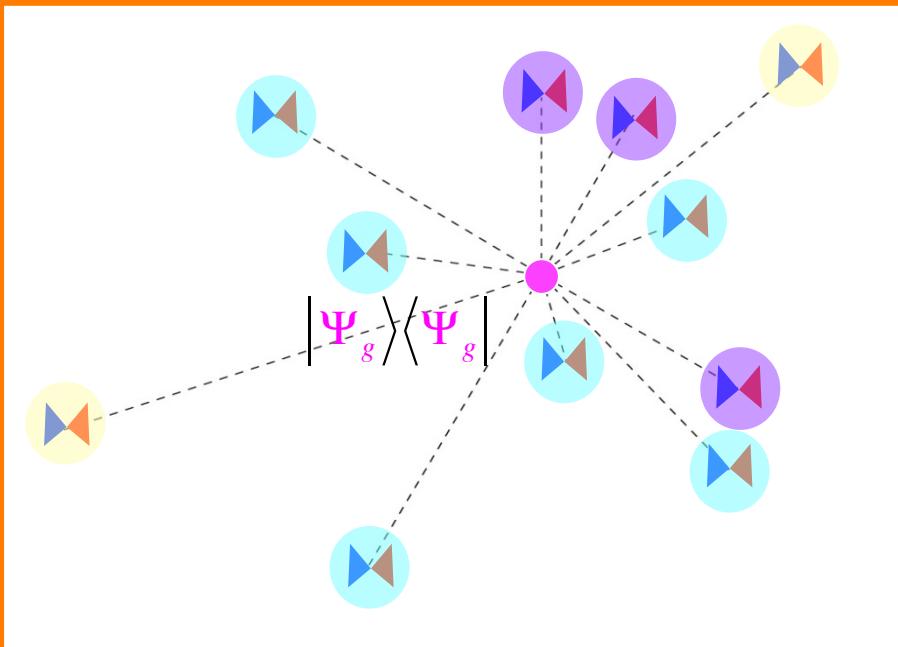
## The stochastic wavefunction approach : Sign problem

Sign fluctuations in the overlap between walkers and the ground state strongly contaminate the sampling

$$\exp(-\Delta\tau E_g) \langle \Psi_g | \Phi_{\tau_o}^{(a)} \rangle \langle \Psi_g | \Phi_{\tau_o}^{(b)} \rangle | \Psi_g \rangle \langle \Psi_g | = \mathbb{E} \left[ \Omega_{\tau_o + \Delta\tau}^{(a,b)} | \Phi_{\tau_o + \Delta\tau}^{(a)} \rangle \langle \Phi_{\tau_o + \Delta\tau}^{(b)} | \right], \quad \Delta\tau \rightarrow \infty$$

$$\Rightarrow \mathbb{E} \left[ \Omega_{\tau_o + \Delta\tau}^{(a,b)} \langle \Phi_{\tau_o + \Delta\tau}^{(b)} | \Phi_{\tau_o + \Delta\tau}^{(a)} \rangle \right] > 0 \text{ as long as } \langle \Psi_g | \Phi_{\tau_o}^{(a)} \rangle \langle \Psi_g | \Phi_{\tau_o}^{(b)} \rangle > 0$$

→ Sign-problem in observables



$$\langle \hat{O} \rangle_{\Psi_g} = \frac{\mathbb{E} \left[ \Omega_{\tau}^{(a,b)} \langle \Phi_{\tau}^{(b)} | \hat{O} | \Phi_{\tau}^{(a)} \rangle \right]}{\mathbb{E} \left[ \Omega_{\tau}^{(a,b)} \langle \Phi_{\tau}^{(b)} | \Phi_{\tau}^{(a)} \rangle \right]}$$

Balance between 3 contributions  
from previous populations :

$$\langle \Psi_g | \Phi_{\tau_o}^{(a)} \rangle \langle \Psi_g | \Phi_{\tau_o}^{(b)} \rangle - 0 +$$

## The sign-free stochastic mean-field scheme

O. Juillet, New J. Phys (2007)

$$\exp(-\Delta\tau E_g) \langle \Psi_g | \Phi_{\tau_o}^{(a)} \rangle \langle \Psi_g | \Phi_{\tau_o}^{(b)} \rangle = \mathbb{E} \left[ \Omega_{\tau_o + \Delta\tau}^{(a,b)} \langle \Phi_{\tau_o + \Delta\tau}^{(b)} | \Phi_{\tau_o + \Delta\tau}^{(a)} \rangle \right], \quad \Delta\tau \rightarrow \infty$$

$$\langle \hat{O} \rangle_{\Psi_g} = \frac{\mathbb{E} \left[ \Omega_{\tau}^{(a,b)} \langle \Phi_{\tau}^{(b)} | \hat{O} | \Phi_{\tau}^{(a)} \rangle \right]}{\mathbb{E} \left[ \Omega_{\tau}^{(a,b)} \langle \Phi_{\tau}^{(b)} | \Phi_{\tau}^{(a)} \rangle \right]}$$

= 1 with biorthogonal  
wavefunctions  $\langle \phi_n^{(b)} | \phi_{n'}^{(a)} \rangle = \delta_{nn'}$



No sign problem

→ Non-linear Ito-stochastic differential equation with multiplicative noise

$$d\phi_n^{(a)} = \left(1 - \mathcal{R}^{(ab)}\right) \left[ -\frac{d\tau}{2} h_{HF}(\mathcal{R}^{(ab)}) + \sqrt{\frac{|U|}{2}} \sum_{\vec{r}\sigma} |\vec{r}\sigma\rangle \text{sgn}(U) dW_{\vec{r}}^{(a)} \langle \vec{r}\sigma | \right] \phi_n^{(a)}, \quad d\Omega_{\tau}^{(a,b)} = -d\tau \Omega_{\tau}^{(a,b)} \langle \Phi_{\tau}^{(b)} | \hat{H} | \Phi_{\tau}^{(a)} \rangle$$

Interstate one-body  
density matrix

Hartree-Fock Hamiltonian

$$\mathcal{R}^{(ab)} = \sum_n |\phi_n^{(a)}\rangle \langle \phi_n^{(b)}|$$

(Semi-implicit Euler algorithm - Adaptive step with Brownian trees)

→ Zero-temperature and canonical version of the « Gaussian QMC » method

J. F. Corney, P.D. Drummond, Phys. Rev. Lett. (2004)

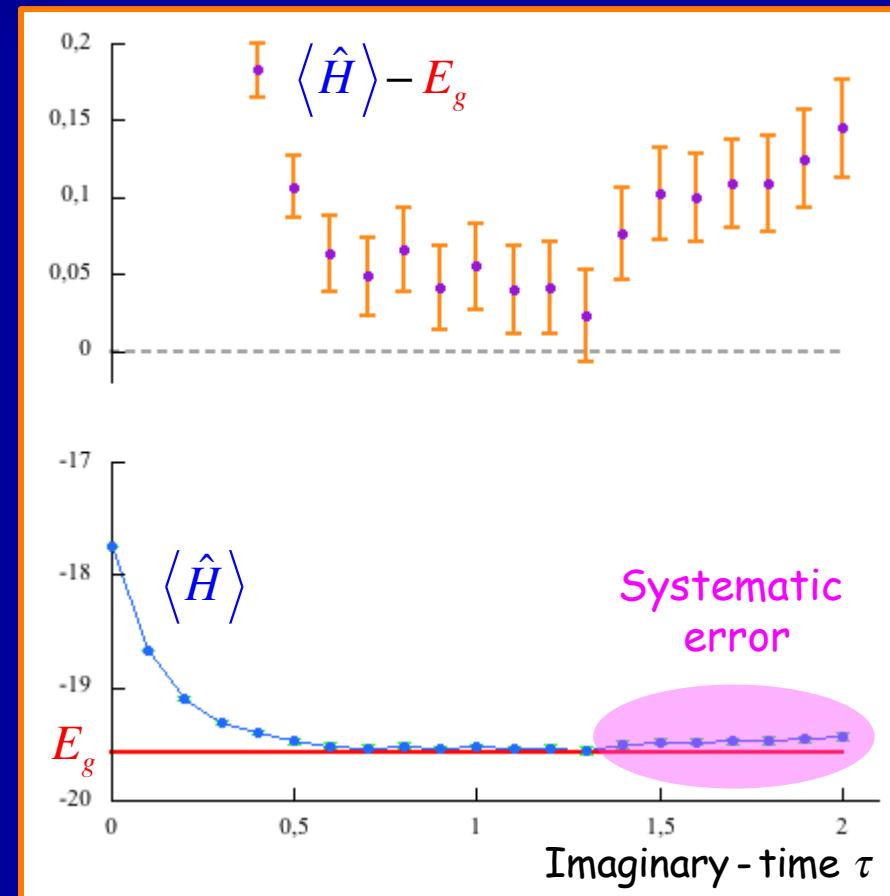
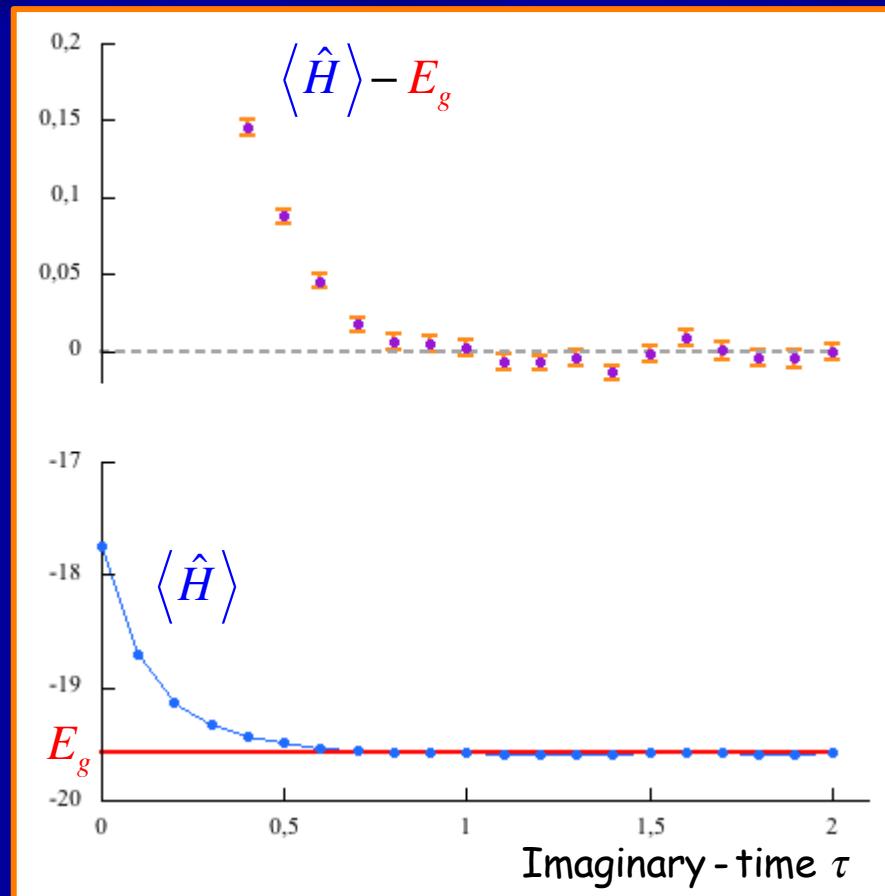
## The sign-free stochastic mean-field scheme

Fixed

$$|\Psi_g\rangle\langle\Phi_0^{(b)}| \underset{\tau \rightarrow \infty}{\propto} \mathbb{E}\left[\Omega_{\tau}^{(a,b)}|\Phi_{\tau}^{(a)}\rangle\langle\Phi_0^{(b)}|\right]$$

$$|\Psi_g\rangle\langle\Psi_g| \underset{\tau \rightarrow \infty}{\propto} \mathbb{E}\left[\Omega_{\tau}^{(a,b)}|\Phi_{\tau}^{(a)}\rangle\langle\Phi_{\tau}^{(b)}|\right]$$

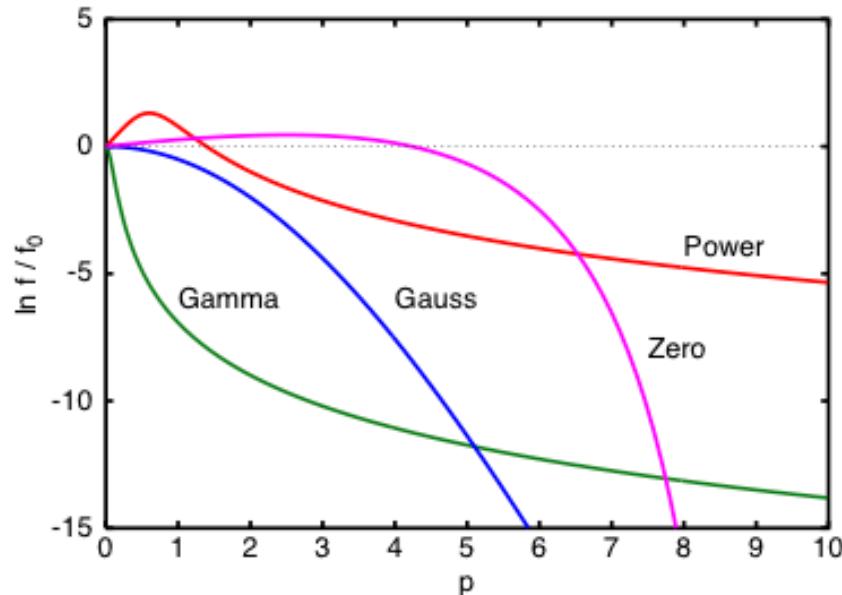
4\*4 lattice - U=4 - t=1 - N=10 electrons



## The sign-free stochastic mean-field scheme



Multiplicative noises generate power-law tail distribution



### Langevin equation

$$dp = (p_o - \gamma p)dt + dW_1 - pdW_2$$

Drift  
term

Additive  
noise

Multiplicative  
noise

### Stationary distributions

#### Gauss

$$dW_1^2 = 2Ddt, dW_2^2 = 0$$

#### Gamma distribution in 1/p

$$dW_1^2 = 0, dW_2^2 = 2Cdt$$

#### Power-law distribution

$$\gamma = 0, dW_1^2 = 2Ddt, dW_2^2 = 2Cdt$$

T.S. Biro, A. Jacovák,  
Phys. Rev. Letters (2004)

## The sign-free stochastic mean-field scheme

Infinite moments of the error on the exact many-body state can be detected through extreme value analysis of dyadics' norm  $\mathcal{N}^{(a,b)} = \left\| |\Phi^{(a)}\rangle\langle\Phi^{(b)}| \right\|$

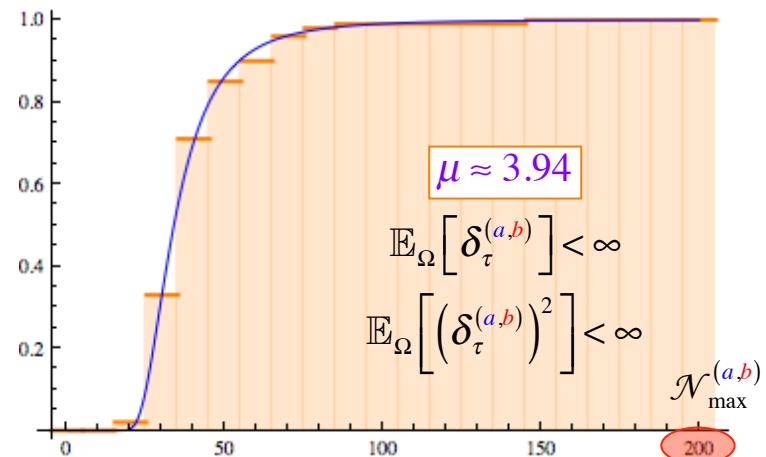


If  $P_\Omega(\mathcal{N}^{(a,b)}) \propto 1/(\mathcal{N}^{(a,b)})^{1+\mu}$  for  $\mathcal{N}^{(a,b)} \rightarrow \infty$ , the distribution of the maximum  $\mathcal{N}_{\max}^{(a,b)}$  over a sufficiently large sample is given by the

$$\text{Frechet distribution } \frac{\mu}{(\mathcal{N}_{\max}^{(a,b)}/\mathcal{N}_o)^{1+\mu}} \exp\left[\left(\mathcal{N}_{\max}^{(a,b)}/\mathcal{N}_o\right)^{-\mu}\right]$$

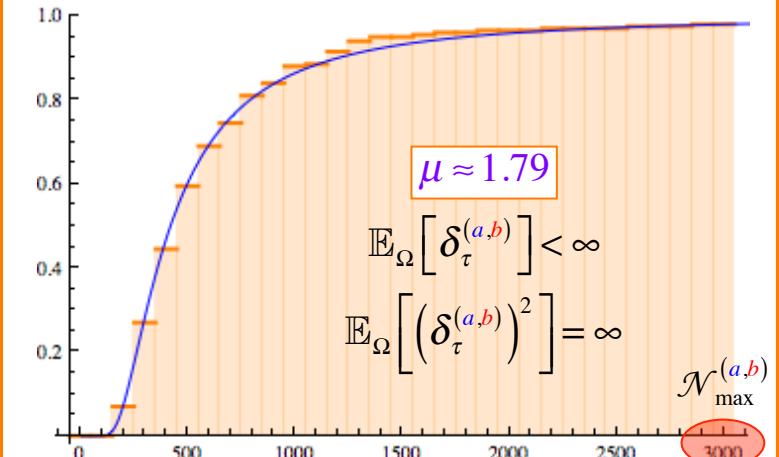
C.D.F of  $P_\Omega(\mathcal{N}_{\max}^{(a,b)})$

$|\Psi_g\rangle\langle\Phi_0^{(b)}| - \tau=2$



C.D.F of  $P_\Omega(\mathcal{N}_{\max}^{(a,b)})$

$|\Psi_g\rangle\langle\Psi_g| - \tau=2$



## The sign-free stochastic mean-field scheme

Distributions for the norm of projected dyadics on ground-state quantum-numbers  $\Gamma = (\vec{K}, S, S_z, \dots)$  decay much faster

Same idea used in Gaussian QMC : T. Aima, M. Imada , J. Phys. Soc. Jpn . (2007)

$$\begin{aligned}
 \exp(-\Delta\tau E_g) \langle \Psi_g | \Phi_0^{(a)} \rangle \langle \Psi_g | \Phi_0^{(b)} \rangle \langle \hat{P}^{(\Gamma)} | \Psi_g \rangle \langle \Psi_g | \hat{P}^{(\Gamma)} &= \mathbb{E}_{\tau \rightarrow \infty} \left[ \Omega_{\tau}^{(a,b)} \hat{P}^{(\Gamma)} | \Phi_{\tau}^{(a)} \rangle \langle \Phi_{\tau}^{(b)} | \hat{P}^{(\Gamma)} \right] \\
 \text{Projection operator} \quad \text{⚡} & \\
 \mathbb{E}_{\tau \rightarrow \infty} \left[ \Omega_{\tau}^{(a,b)} \left| \langle \Phi_{\tau}^{(b)} | \hat{P}^{(\Gamma)} | \Phi_{\tau}^{(a)} \rangle \right| \frac{\hat{P}^{(\Gamma)} | \Phi_{\tau}^{(a)} \rangle \langle \Phi_{\tau}^{(b)} | \hat{P}^{(\Gamma)}}{\left| \langle \Phi_{\tau}^{(b)} | \hat{P}^{(\Gamma)} | \Phi_{\tau}^{(a)} \rangle \right|} \right] \\
 \text{Importance sampling} \quad \text{⚡} & \\
 \mathbb{E}_{\tilde{\Omega}} \left[ \frac{\hat{P}^{(\Gamma)} | \Phi_{\tau}^{(a)} \rangle \langle \Phi_{\tau}^{(b)} | \hat{P}^{(\Gamma)}}{\left| \langle \Phi_{\tau}^{(b)} | \hat{P}^{(\Gamma)} | \Phi_{\tau}^{(a)} \rangle \right|} \right] &
 \end{aligned}$$



Sign problem in observables, but less severe :

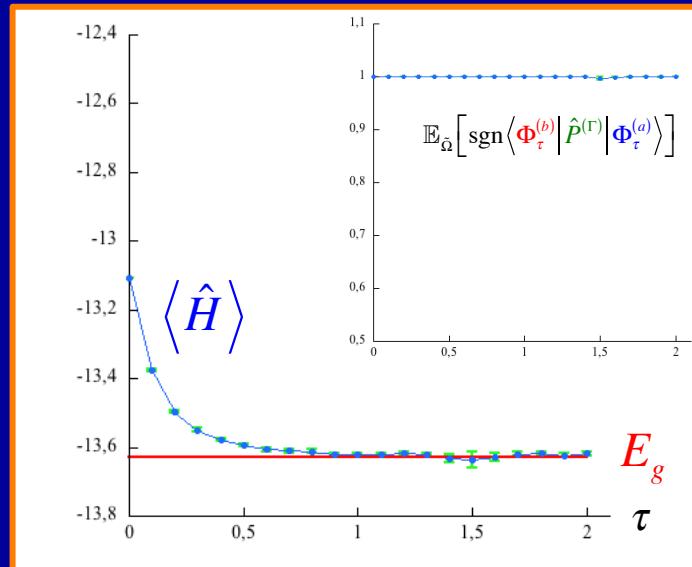
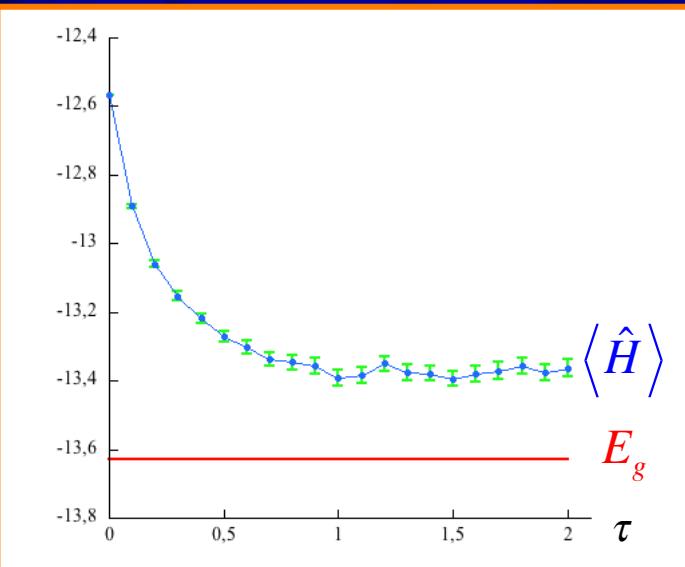
Any walker leads to a positive-sign population in the  $\tau \rightarrow \infty$  limit

$$\begin{aligned}
 \exp(-\Delta\tau E_g) \langle \Psi_g | \Phi_{\tau_0}^{(a)} \rangle \langle \Psi_g | \Phi_{\tau_0}^{(b)} \rangle &= \mathbb{E}_{\tilde{\Omega}} \left[ \frac{\langle \Phi_{\tau_0+\Delta\tau}^{(b)} | \hat{P}^{(\Gamma)} | \Phi_{\tau_0+\Delta\tau}^{(a)} \rangle}{\left| \langle \Phi_{\tau_0+\Delta\tau}^{(b)} | \hat{P}^{(\Gamma)} | \Phi_{\tau_0+\Delta\tau}^{(a)} \rangle \right|} \right] \\
 > 0 &
 \end{aligned}$$

# The sign-free stochastic Hartree-Fock scheme

$$|\Psi_g\rangle\langle\Psi_g| \underset{\tau \rightarrow \infty}{\propto} \mathbb{E}_{\Omega}\left[ |\Phi_{\tau}^{(a)}\rangle\langle\Phi_{\tau}^{(b)}| \right]$$

$$|\Psi_g\rangle\langle\Psi_g| \underset{\tau \rightarrow \infty}{\propto} \mathbb{E}_{\tilde{\Omega}}\left[ \frac{\hat{P}^{(\Gamma)}|\Phi_{\tau}^{(a)}\rangle\langle\Phi_{\tau}^{(b)}|\hat{P}^{(\Gamma)}}{\left\langle\Phi_{\tau}^{(b)}|\hat{P}^{(\Gamma)}|\Phi_{\tau}^{(a)}\right\rangle} \right]$$



4\*4 lattice  
U=4 - t=1  
N=16 electrons

	$\langle \hat{T} \rangle$	$S_m = \frac{1}{3} \sum_{\vec{r}} (-1)^{x+y} \langle \hat{\sigma}_{\vec{0}} \cdot \hat{\sigma}_{\vec{r}} \rangle$	$S_c = \sum_{\vec{r}} (-1)^{x+y} \langle \hat{n}_{\vec{0}} \hat{n}_{\vec{r}} \rangle$
<i>Ex. Diag.</i>	-20.989	3.64	0.385
<i>QMC – No projection</i>	-20.85(3)	3.35(1)	0.385(1)
<i>QMC – projection</i>	-21.00(2)	3.66(1)	0.386(1)

## The sign-free stochastic mean-field scheme

$$|\Psi_g\rangle\langle\Psi_g| \underset{\tau \rightarrow \infty}{\propto} \mathbb{E}_{\Omega}\left[ |\Phi_{\tau}^{(a)}\rangle\langle\Phi_{\tau}^{(b)}| \right]$$

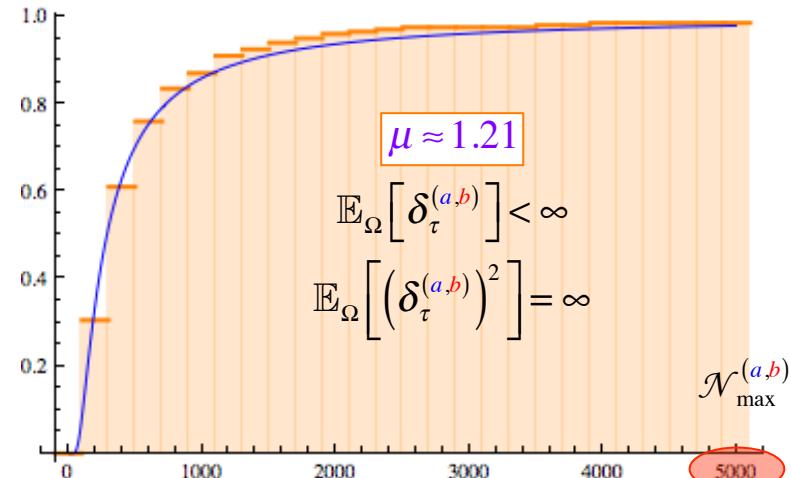
Norm  $\mathcal{N}^{(a,b)}$

$$|\Psi_g\rangle\langle\Psi_g| \underset{\tau \rightarrow \infty}{\propto} \mathbb{E}_{\tilde{\Omega}}\left[ \frac{\hat{P}^{(\Gamma)}|\Phi_{\tau}^{(a)}\rangle\langle\Phi_{\tau}^{(b)}|\hat{P}^{(\Gamma)}}{\langle\Phi_{\tau}^{(b)}|\hat{P}^{(\Gamma)}|\Phi_{\tau}^{(a)}\rangle} \right]$$

Norm  $\tilde{\mathcal{N}}^{(a,b)}$

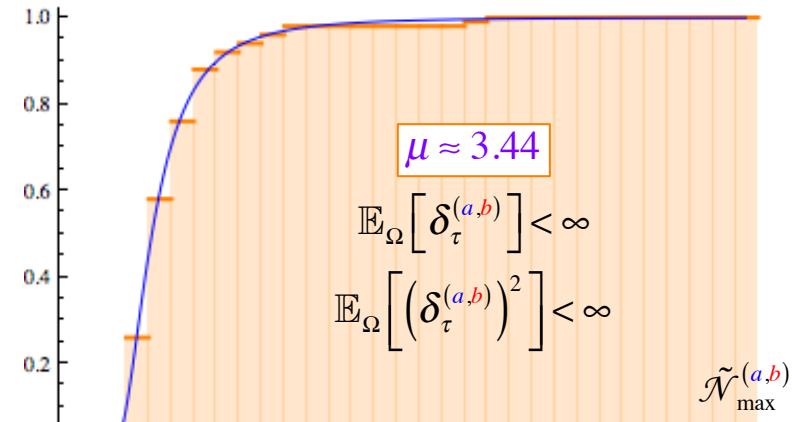
C.D.F of  $P_{\Omega}(\mathcal{N}_{\max}^{(a,b)})$

$\tau=2$



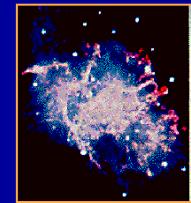
C.D.F of  $P_{\tilde{\Omega}}(\tilde{\mathcal{N}}_{\max}^{(a,b)})$

$\tau=2$



## Hubbard model at unitarity

→ Ideal regime of strong interaction where the cross-section saturates the limit imposed by unitarity :  $\sigma = \sigma_{\max} = 4\pi/k^2$



→ Scale invariance  $E_{\text{unitarity}}(N) = \xi E_{\text{Free}}(N)$

Bertsch many-body  
X challenge,  
Seattle, 1999

$$\xi_{\text{exp}} = \begin{cases} 0.376(5) \\ 0.41(1) \end{cases}$$

M.J.H. Ku & al (2011)

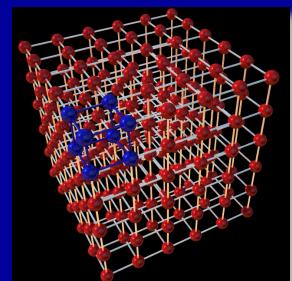
N. Navon & al, Science (2011)

$$\xi_{QMC} = \begin{cases} \leq 0.383(1) \\ = 0.372(5) \end{cases}$$

M. Forbes & al, Phys. Rev. Lett. (2011)  
J. Carlson & al (2011)

→ OK with Hubbard-like models in the dilute limit

$$k \sim k_F = (3\pi^2 \rho)^{1/3} \xrightarrow[\rho \rightarrow 0]{} 0$$



3D cubic lattice  
with period  $b$

$$\sigma \underset{k \rightarrow 0}{\sim} \frac{4\pi}{|a^{-1} + ik - r_e k^2/2 + \dots|^2}$$

Scattering  
length  $a = a(\varepsilon_{\vec{k}}, U)$

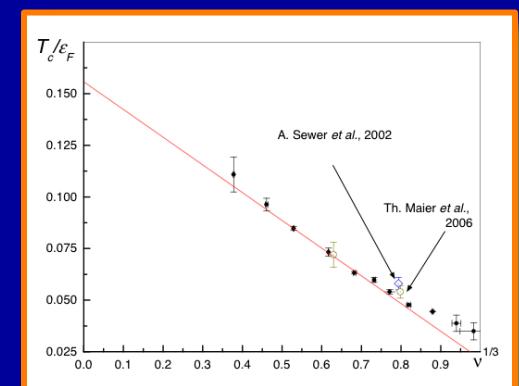
Effective range  
 $r_e = r_e(\varepsilon_{\vec{k}})$

Divergence for finite  $U < 0$



Pairing correlations

E. Burovski & al,  
Phys. Rev. Lett. (2006)



# Lattice modeling of the unitary gas : Sign-free stochastic mean-field with BCS states

O. Juillet, to be published (2011)

**Fixed**

$$\left| \Psi_g \right\rangle \langle \Phi_0^{(b)} \right| \underset{\tau \rightarrow \infty}{\propto} \mathbb{E} \left[ \Omega_\tau^{(a,b)} \left| \Phi_\tau^{(a)} \right\rangle \langle \Phi_0^{(b)} \right| \right]$$

$$\langle \Phi_0^{(b)} | \Phi_\tau^{(a)} \rangle = 1$$

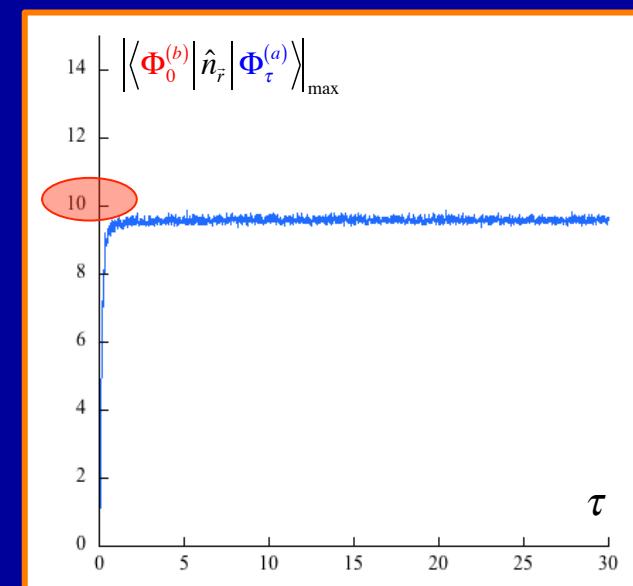
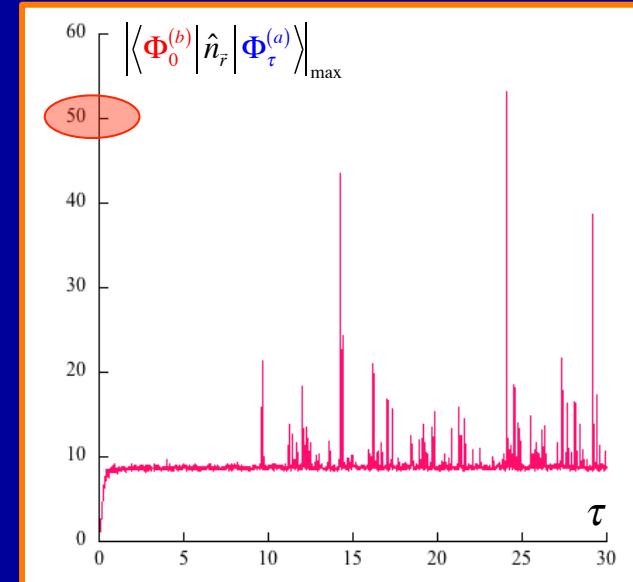
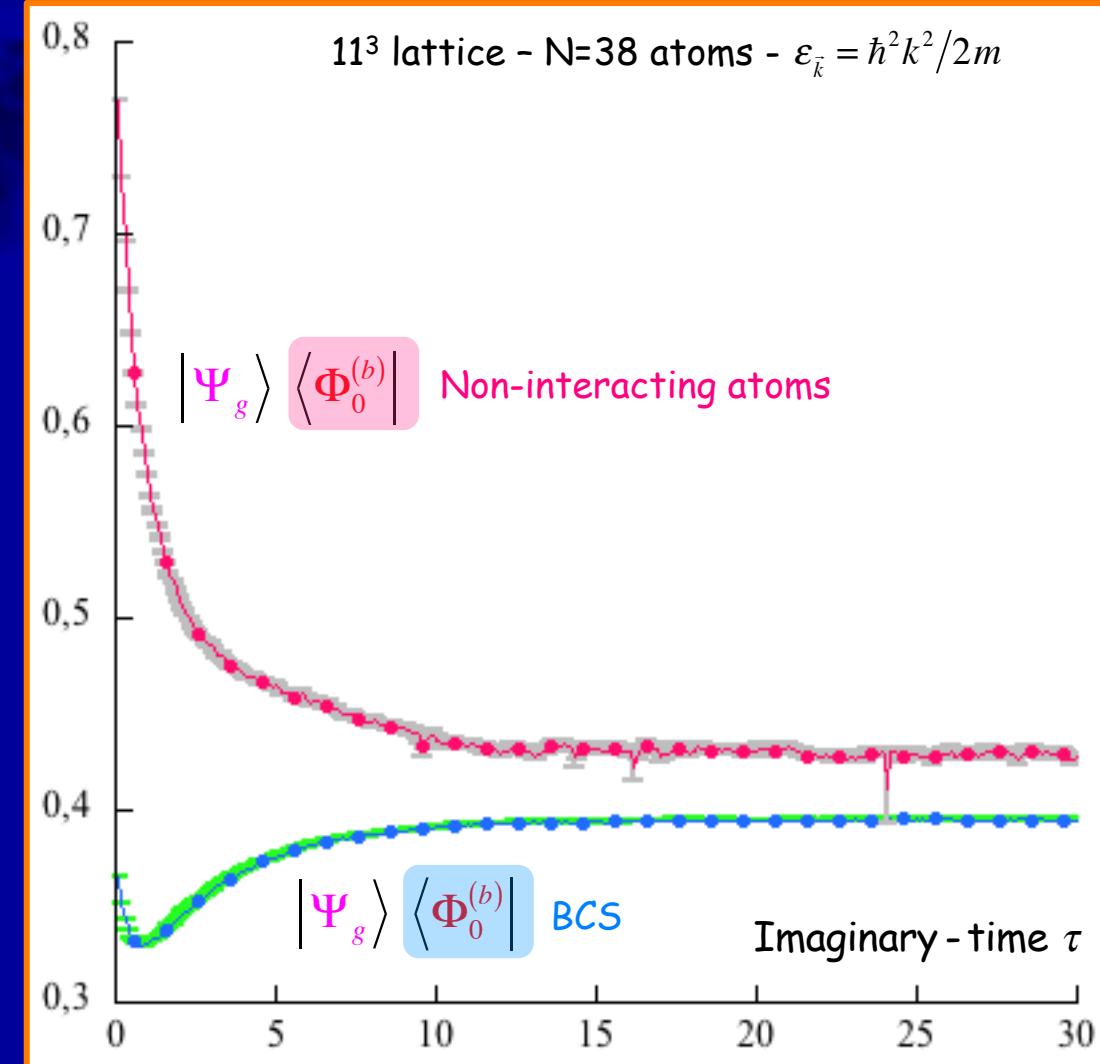
$$\begin{cases} \left| \Phi^{(a)} \right\rangle = \hat{c}_{\phi_1^{(a)}, \uparrow}^+ \cdots \hat{c}_{\phi_{N/2}^{(a)}, \uparrow}^+ \hat{c}_{\phi_1^{(a)}, \downarrow}^+ \cdots \hat{c}_{\phi_{N/2}^{(a)}, \downarrow}^+ | \ \rangle \\ \left| \Phi^{(b)} \right\rangle \propto \exp \left[ \sum_{\vec{k}} f_{\vec{k}}^{(b)} c_{\vec{k}\uparrow}^+ c_{\vec{k}\downarrow}^+ \right] | \ \rangle \end{cases}$$

$$\begin{aligned} \left| d\phi_n^{(a)} \right\rangle &= \left( 1 - \mathcal{R}^{(ab)} \right) \left[ -d\tau h_{HF} \left( \mathcal{R}^{(ab)} \right) + \sqrt{|U|} \sum_{\vec{r}} | \vec{r} \rangle dW_{\vec{r}}^{(a)} \langle \vec{r} | \right] \left| \phi_n^{(a)} \right\rangle \\ &+ \left[ \frac{U d\tau}{2N} \sum_{\vec{r}} \left\langle \Phi_0^{(b)} \right| \hat{n}_{\vec{r}\uparrow} \hat{n}_{\vec{r}\downarrow} \left| \Phi_\tau^{(a)} \right\rangle - \left\langle \Phi_0^{(b)} \right| \hat{n}_{\vec{r}\uparrow} \left| \Phi_\tau^{(a)} \right\rangle \left\langle \Phi_0^{(b)} \right| \hat{n}_{\vec{r}\downarrow} \left| \Phi_\tau^{(a)} \right\rangle \right] \mathcal{R}^{(ab)} \left| \phi_n^{(a)} \right\rangle \end{aligned}$$

$$d\Omega_\tau^{(a,b)} = -d\tau \Omega_\tau^{(a,b)} \left\langle \Phi_0^{(b)} \right| \hat{H} \left| \Phi_\tau^{(a)} \right\rangle$$

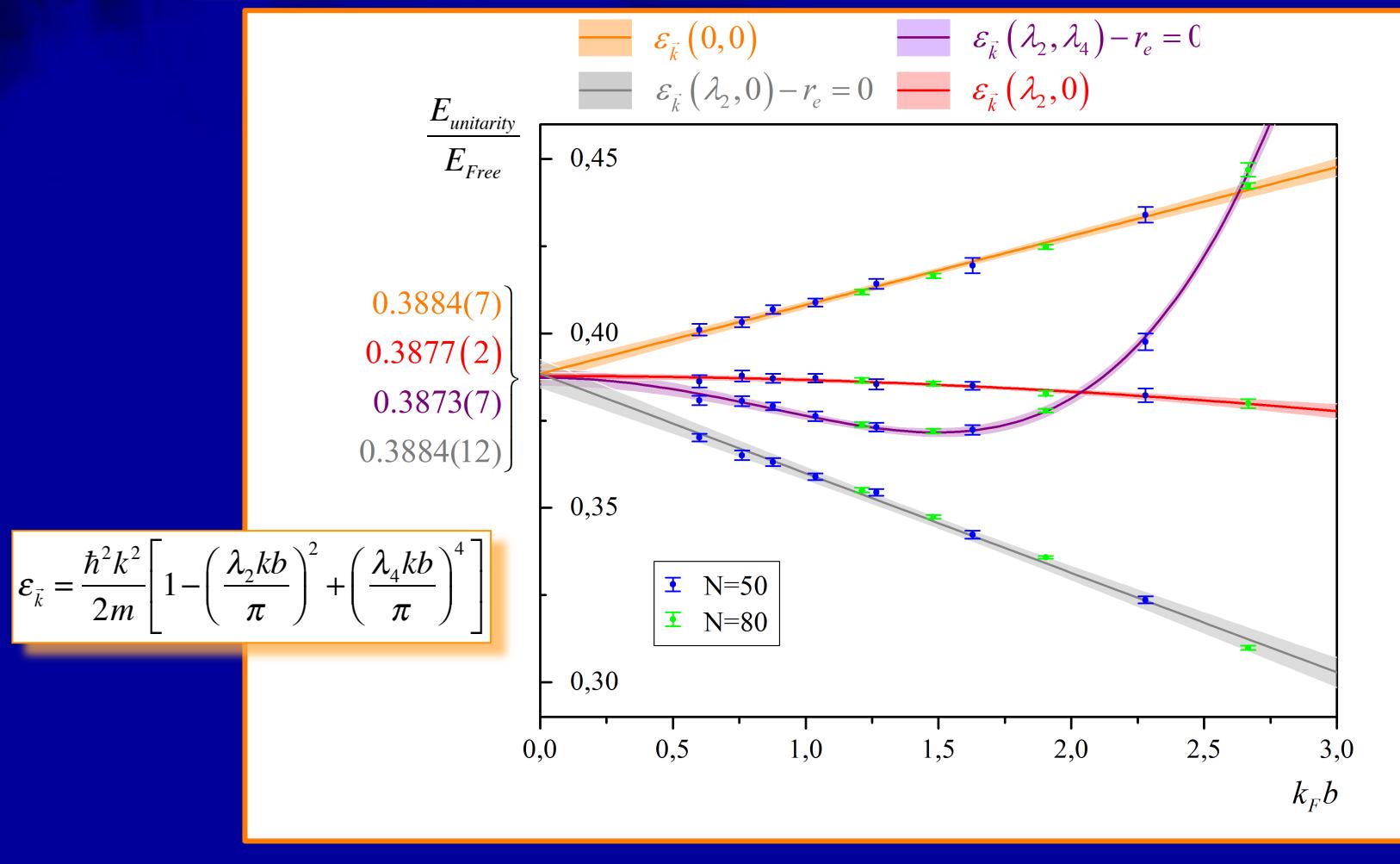
$$\mathcal{R}^{(ab)} = \sum_{n,n'} \left| \phi_n^{(a)} \right\rangle \left( g^{-1} \right)_{nn'} \left\langle \phi_{n'}^{(a)} \right| \quad \text{with} \quad g_{nn'} = \left\langle \phi_n^{(a)} \right| f^{(b)} \left| \phi_{n'}^{(a)} \right\rangle$$

## Hubbard-like model at unitarity



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Different dispersion relations must give the same universal parameter  $\xi$  in the vanishing lattice filling limit



## Stochastic wavefunction approach to Hubbard-like models

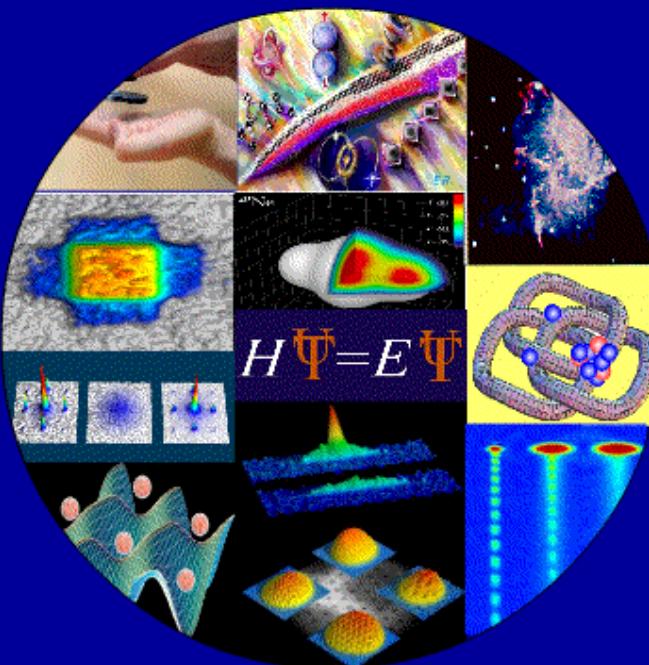


The Hubbard-model can be investigated with stochastic wavefunctions undergoing a Brownian motion in imaginary-time that ensures positive weights in any regime



Positive weights in Quantum Monte-Carlo methods are NOT sufficient to obtain exact results

# Stochastic wavefunction approach to Hubbard-like models



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