



Many-body observables and effectiveness in nuclear physics

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1. **Effectiveness of observable**
2. Effective field theories (EFTs)
3. Nuclear systems and EFTs
4. Extension to many-body observables



Effectiveness of observable

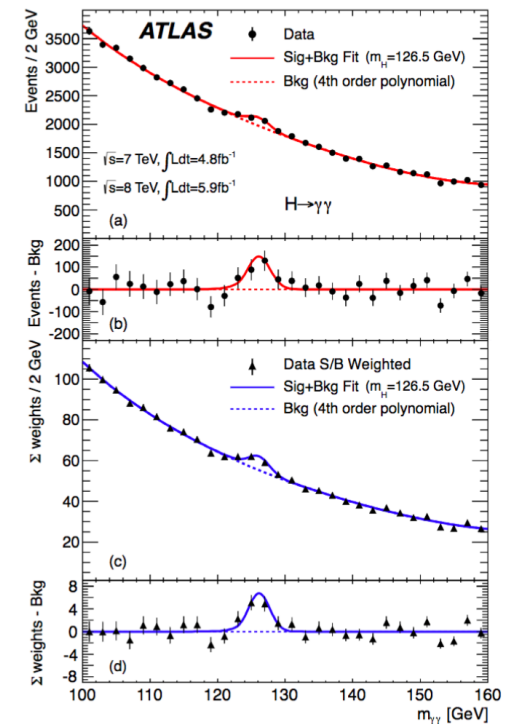
The need for a theory decoupling from UV physics

Observables in experiments



- Experimental observables result of a measurement process
- Extract number from experience
 - Already non-observable/extracted
 - Try to define a number least dependent on measurement process (e.g. length littoral)
- Measurements are limited by the resolution of our tools (accuracy)
- Improve resolution by taking explicitly into account known background processes
 - E.g. subtraction of background by PSD/ToF analysis
- But up to now our measurement capabilities remains finite
 - There is always an uncertainty on the result of an experiment

Experimental observables come in probability distribution of numbers



[Atlas Collaboration, Phys. Lett. B, 2012]

Theoretical prediction of observables



Old-fashion approach

1. Set a model
 - Degrees of freedom
 - Dynamical equations
 - Relation to observables
2. Make predictions
 - Expected value
3. Compare with experimental data
Statistical tests e.g. χ^2

But one never observes an exact value

Effective approach

1. Set a model
 - Idem
 - + Effective model \neq elementary degree of freedoms
 - Scale of unresolved physics M_{hi}
 - Estimation of corrections on observables
2. Make predictions
 - Mean value
 - Standard deviation (+ higher-moments)
3. Compare with experimental data
Statistical tests \leftarrow theoretical uncertainty

But predictions on what ?

Effective approach in Quantum Mechanics



Observables of a system with particular symmetry

- Amplitude of transition associated to unitary evolution operator (S-matrix)
- Eigenvalues of self-adjoint operator $\mathcal{O} \rightarrow$ e.g. Hamiltonian for energy
- Eigenvalues in a **subspace** of the Fock space \mathcal{F}

General system

$$H = \sum_{p,A=0}^{+\infty} H_A^{(p)} \quad \langle H \rangle = \left\langle \sum_{p,A=0}^{+\infty} H_A^{(p)} \right\rangle_{\mathcal{F}}$$

System with A particles

$$H^{\text{eff}} = \sum_{A_1=0}^A \sum_{p=0}^{+\infty} H_{A_1}^{(p)} \quad \langle H \rangle = \left\langle \sum_{A_1=0}^A \sum_{p=0}^{+\infty} H_{A_1}^{(p)} \right\rangle_{\mathcal{F}_{A \subset \mathcal{F}}}$$

	2-body	3-body	4-body
LO			X
NLO		X	?
N ² LO	X		?

How to estimate corrections ?

The decoupling assumption

- To be able to make predictions, assume:

H: elt dof ~ ▪ Physics beyond M_{hi} (UV) can be neglected (strong)

OR

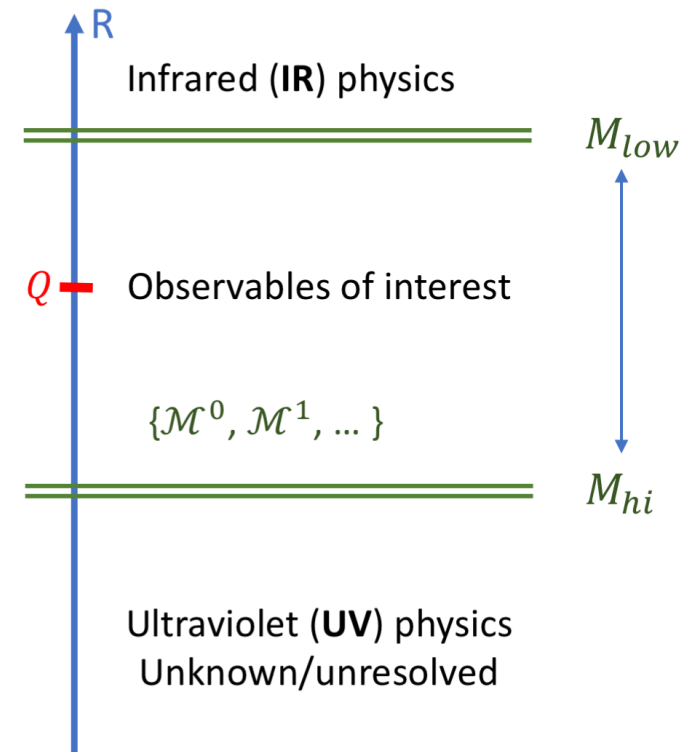
H: eff dof ~ ▪ UV can be taken into account effectively (weak)



- From one model \mathcal{M} to a class of models $\{\mathcal{M}^p, p \geq 0\}$:

- $\mathcal{M}^0 \rightarrow$ Leading contribution
- $\mathcal{M}^p \rightarrow$ Refined models at order p
- Convergence/divergence with $p \rightarrow +\infty \Leftrightarrow$ **Range of validity**
- UV effectively taken into account in parameters of \mathcal{M}^p
- **But** each order needs more & more parameters !

In range
of
validity



Outline



1. Effectiveness of observable
- 2. Effective field theories (EFTs)**
3. Nuclear systems and EFTs
4. Extension to many-body observables



Effective Field Theories

Observables and their decoupling from UV physics

Effective field theory implementation



- Folk theorem : [S. Weinberg 79]

The quantum field theory generated by the most general Lagrangian with some assumed symmetries will produce the most general S matrix incorporating quantum mechanics, Lorentz invariance, unitarity, cluster decomposition and the assumed symmetries.

- Inputs

- Degrees of freedom
 - type of particles (e.g. nucleon, pion etc)
- Symmetries
 - Internal + Space-time
- Coupling constants
 - e.g. fitted to experimental data (or to an underlying theory)
 - Models already contains uncertainty through their fit

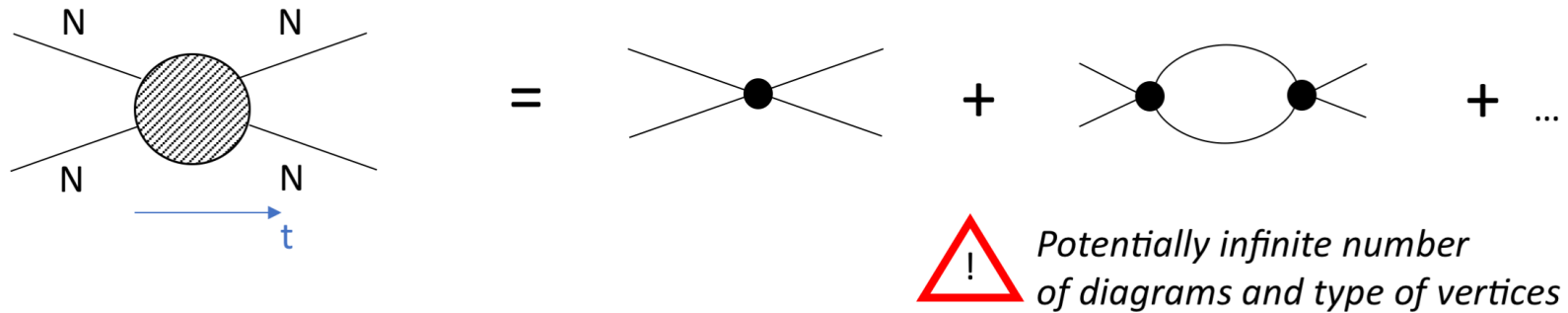


**Gives you unsolvable dynamical equations !
(nice but insufficient for observable predictions)**

Perturbation theory



- **Feynman diagrams.** Physical process decomposed as a sum of perturbations



- **Power-counting rules.** Which diagrams contribute to different order ?

➔ Truncation of perturbation series to (hopefully) a computable, at each order, subpart

From UV divergences to UV independence



- **Problem**

- Most diagrams are UV divergent
- Ill-defined theory

- **But**

- Feynman diagrams are not observable ! Only their sum gives observable estimation
- EFT calculations should be independent of UV
 - We can introduce an arbitrary regularization procedure → Modify Hamiltonian in the UV sector
 - Observables should be independent of this regularization

- **Solution**

1. Modify the UV arbitrarily → well-defined theory
2. Modifying Hamiltonian $H = H_b + H_{ct}$ → recover independence of UV
3. Fit the **total** Hamiltonian to experimental data

} Renormalization procedure



- Arbitrariness of regularization compensated by regulator dependent Hamiltonian.
- In practice H_{ct} will be of the same form as H_b → Regulator dependence in coupling constants.
- In the end we should verify

$$\partial_\Lambda \langle \mathcal{O}(\Lambda, C(\Lambda)) \rangle \approx 0 \rightarrow \text{Renormalization group (RG) equations}$$

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Nuclear systems

Fundamental problem of current phenomenological descriptions

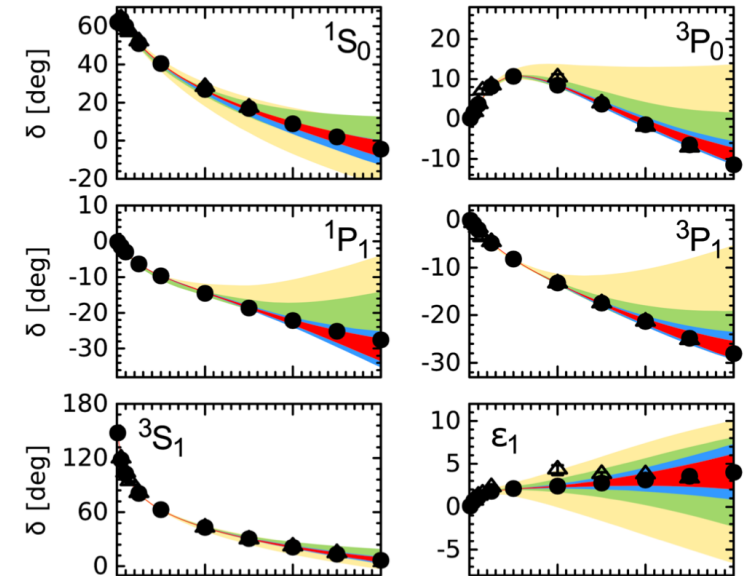
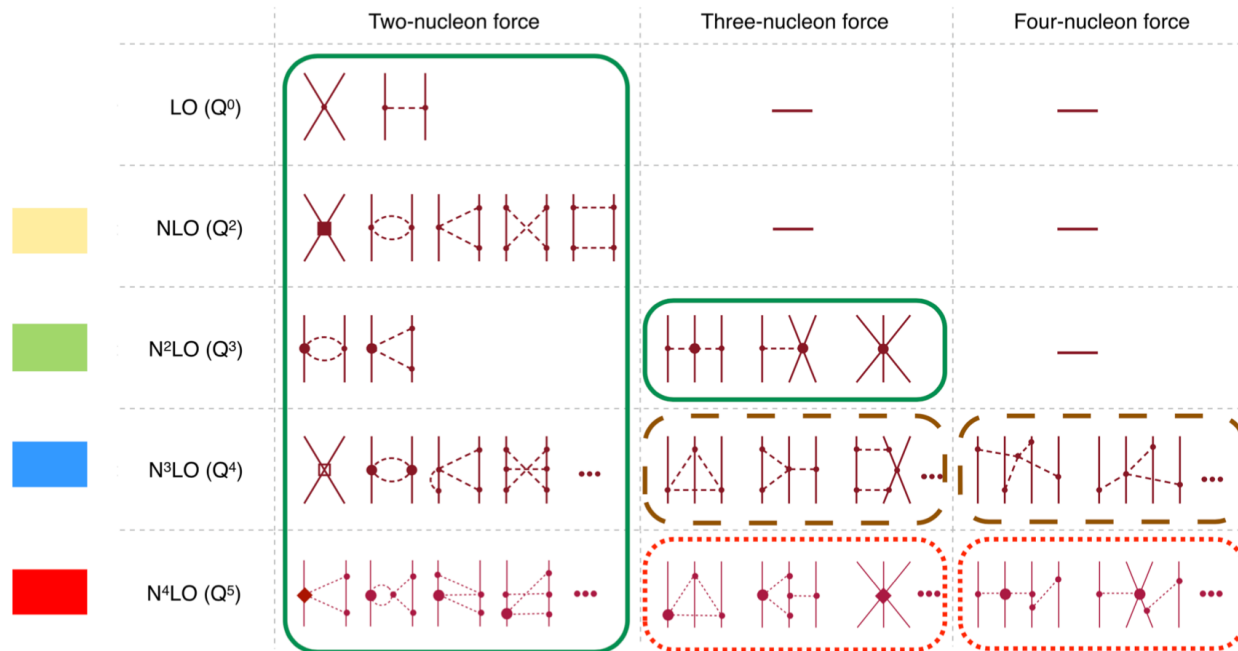
Historical perspective on NN interactions



- 1935 : Yukawa potential
 - Birth of meson theory
- 40s-50s : Pion theories
 - Troubles with multi-pion, anti-nucleons diagrams
- 60s-00s : One-Boson-Exchange Model ($\rho, \sigma, \omega, \dots$)
 - Good data fitting
 - But no systematic
- 90s – today : EFT based interactions [S.Weinberg 90 91]
 - Chiral potentials → Try to estimate errors from breaking of RG
[Entem, Machleidt 03] [Epelbaum, Glöckle, Meissner 05]
 - EFTs → Take RG invariance as guiding constraint
[Kaplan, Savage, Wise 96] [Nogga, Timmermans, van Kolck 05]

No RG invariance

Chiral potentials



Generally UV dependent observables → problems needs to be addressed

EFTs program in nuclear physics



Start simple

- Pionless theory
- few-body $A \leq 4$ (here only 1S_0 partial wave)

} Gives your Hamiltonian

Power-counting rules

- Naïve Dimensional Analysis ?
- Resum any diagram with LO vertices ?



Back to unsolvable problem in many-body sector



Needs to investigate further truncation scheme many-body sector

	2-body	3-body	4-body
<i>LO</i>			X
<i>NLO</i>		X	?
<i>N²LO</i>	X		?

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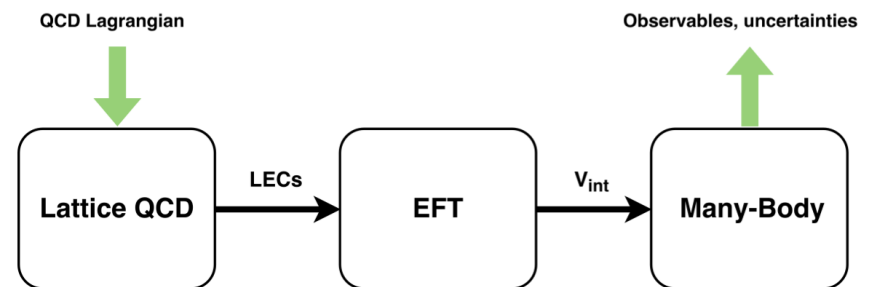
Problems with many-body observables in nuclear systems

Extending EFT requirements to many-body schemes

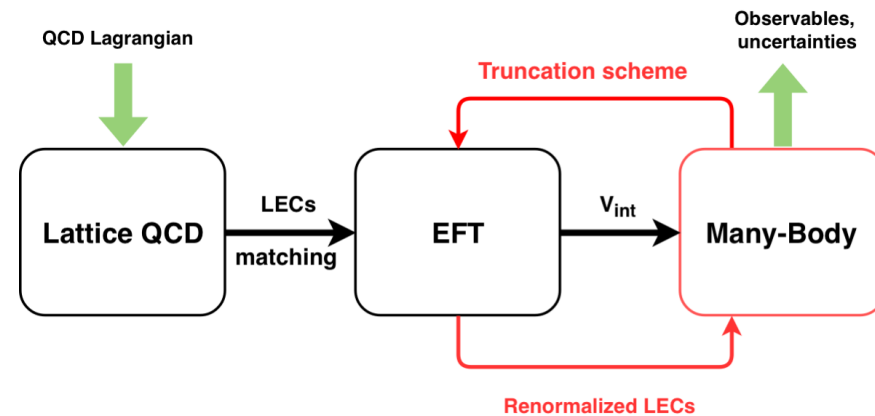
EFTs and many-body observables



- Traditional view : V_{int} as a black box



- Interplay using EFT interactions



Many many-body schemes



- Focus on Many-Body Perturbation Theories

Fock space decomposition

$$\mathcal{F} \equiv \bigoplus_{N=0}^{+\infty} (\mathcal{H}'_1)^{\otimes N}$$



$$\begin{aligned} \forall \mu, \beta_\mu^\dagger |\Phi\rangle &= |\Phi^\mu\rangle \\ \beta_\mu |\Phi\rangle &= 0. \end{aligned}$$



$$\begin{aligned} \{\beta_\mu^\dagger, \beta_\nu\} &= \delta_{\mu\nu} \\ \{\beta_\mu^\dagger, \beta_\nu^\dagger\} &= 0 \\ \{\beta_\mu, \beta_\nu\} &= 0. \end{aligned}$$

Hamiltonian decomposition

$$H = H_0 + H_1$$



$$\begin{aligned} H_0 &\equiv H^{00} + \bar{H}^{11} \\ H_1 &\equiv H^{20} + \check{H}^{11} + H^{02} \end{aligned}$$



$$\bar{H}^{11} = \sum_{\mu} E_{\mu}^0 \beta_{\mu}^{\dagger} \beta_{\mu}$$

$G_{\mu\nu}^0$

- What EFT in many-body sector ? → Need to test many different truncation schemes

How renormalization procedure translates into MBPTs ?

Fermi-sea MBPT



- Fermi-sea decomposition

Fermi-sea vacuum

$$|\Phi\rangle = \prod_{p < k_f} a_p^\dagger |0\rangle$$

Kinetic basis

$$|p\rangle$$

Kinetic spectrum

$$\left\{ E_p = \frac{p^2}{2m} \right\}$$

- Bare propagator

$$iG_{\alpha\beta}^0(p, \omega) = \delta_{\alpha\beta} \left[\frac{\theta(p - k_f)}{\omega - E_p + i\epsilon} + \frac{\theta(k_f - p)}{\omega - E_p - i\epsilon} \right]$$

- Rewrite it in a medium-insertion approach

$$iG_{\alpha\beta}^0(p, \omega) = \delta_{\alpha\beta} \left[\frac{1}{\omega - E_p + i\epsilon} + 2i\pi\delta(\omega - E_p)\theta(k_f - p) \right]$$

Importance of renormalization consistency



Hartree-Fock : (expanded in $\frac{2\sqrt{2}k_F}{\Lambda}$ for convenience)

$$\frac{E^{HF}}{A}(\Lambda; k_F) = \frac{k_F^2}{2m} \left\{ \frac{3}{5} - \frac{mC_0(\Lambda)}{(2\pi)^2} 8k_F \left[\frac{1}{12} + O\left(\left(\frac{2\sqrt{2}k_F}{\Lambda}\right)^2\right) \right] \right\}$$

Ladder pp/hh resummation :

$$\frac{E^{Ld}}{A}(\Lambda; k_F) = \frac{k_F^2}{2m} \left\{ \frac{3}{5} - \frac{48}{\pi} \int_0^1 ds s^2 \int_0^{\sqrt{1-s^2}} d\kappa \kappa \arctan \left(\frac{k_F I_{k_F}^{\frac{k_F}{\Lambda}}(s, \kappa)}{\frac{4\pi}{mC_0(\Lambda)} + \frac{k_F}{\pi} \tilde{R}_{k_F}^{\frac{k_F}{\Lambda}}(s, \kappa)} \right) \right\}$$

+

$C_0(\Lambda)$ matched at **first order** to a_0

$$C_0(\Lambda) = \frac{4\pi}{M} a_0$$

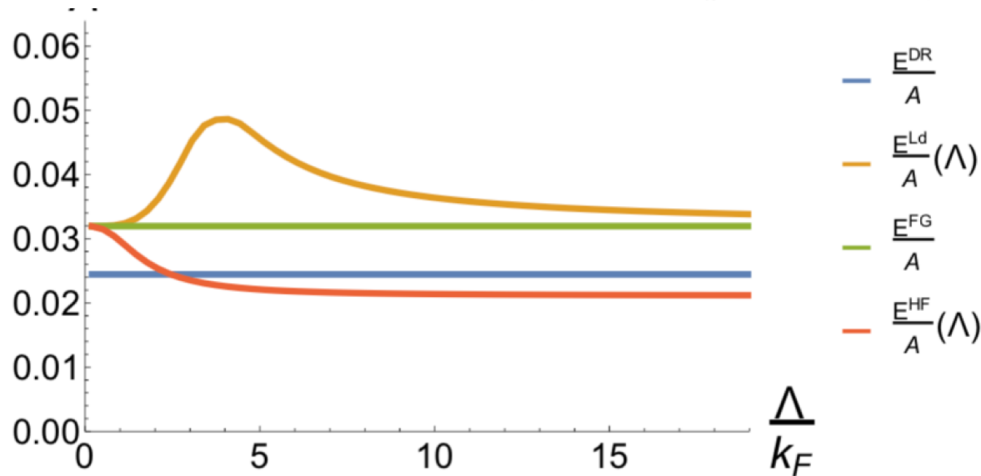
$C_0(\Lambda)$ **exactly** matched to a_0

$$C_0(\Lambda) = \frac{4\pi}{M} \frac{1}{\frac{1}{a_0} - \frac{\Lambda}{\sqrt{2\pi}}}$$

Importance of renormalization consistency

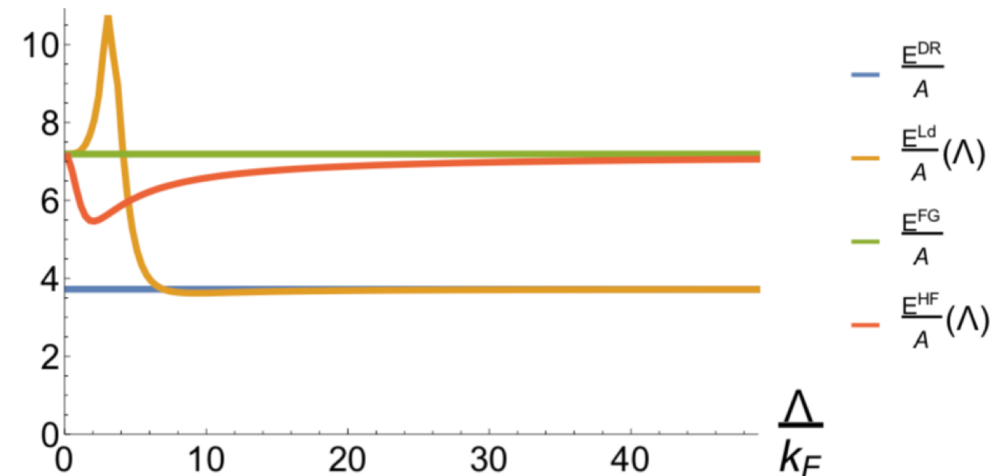


Hartree-Fock : (expanded in $\frac{2\sqrt{2}k_F}{\Lambda}$ for convenience)



$k_F = 60$ MeV

Ladder pp/hh resummation :



$k_F = 150$ MeV

$\frac{E^{DR}}{A}$ from [Kaiser 11]

Toward general rules for renormalizing MBPTs

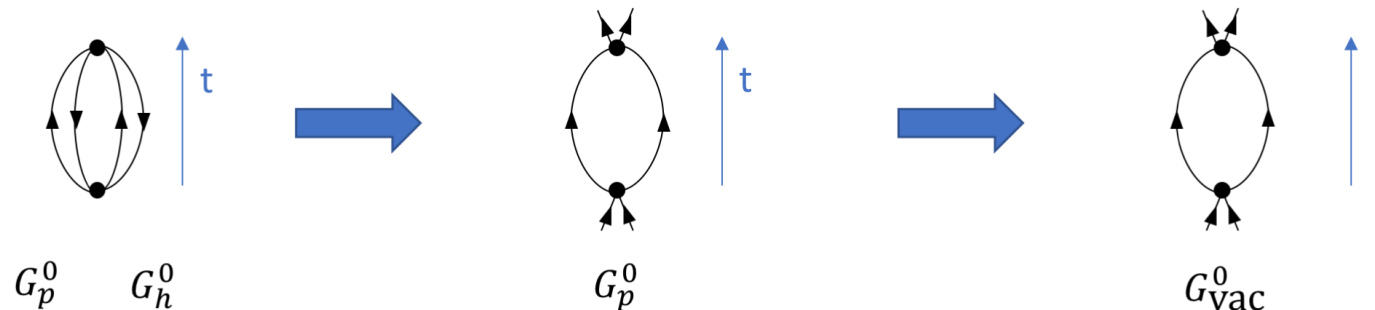


- Weinberg's asymptotic theorem

UV divergences of Feynman diagrams depends only on asymptotic coefficient of the used bare propagator.

- Fermi-sea MBPT revisited

- Hole propagator \rightarrow No UV contributions \rightarrow factorize out of integral
- Particle propagator \rightarrow Same UV behaviour as propagator in particle vacuum



- On going generalization to more general MBPTs



Merci !
