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# Lambda, the fifth foundational constant considered by Einstein

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## Abstract

The cosmological constant, usually referred to as Lambda, was introduced by Einstein in 1917 and then abandoned by him as his biggest ‘blunder’. It currently seems to have made a spectacular comeback in the framework of the new cosmological standard model. I will explain why, together with Planck’s constant, Boltzmann’s constant, the celerity of light, and Newton’s constant, also considered by Einstein, the cosmological constant may play a foundational role in the conceptual framework, and in the metrological framework in a role comparable to the one attributed to the Avogadro constant.

Keywords: cosmological constant, fundamental constants, redefinition SI, Einstein

(Some figures may appear in colour only in the online journal)

## 1. Introduction

When I was invited to participate in the symposium devoted to the role of some of the fundamental constants in the redefinition of the International System of Units (SI), I was asked to comment on the role of the fundamental constants, as seen by Einstein, and to try and explain what is for Einstein (and for me) the status and the profound meaning of the five constants that will redefine the SI:  $h$ ,  $k$ ,  $c$ ,  $e$ ,  $N_A$ . I realized that Einstein also considered five constants:

- The first three of them,  $h$ ,  $k$ ,  $c$ , coincide with the ones considered for the revision of the SI.
- Instead of  $e$ , he considered  $G$  (the Newton’s gravitational constant).
- And instead of Avogadro’s constant  $N_A$ , he considered the cosmological constant  $\Lambda$ .

My intent in this paper is to show that the purpose of Einstein, one century ago was not far from the purpose of the SI community and that, somehow, the cosmological constant can be related to an ‘Avogadro constant of spacetime.’

In the second section, I will review the interpretation by Einstein of the dual role played by the universal constants: as conversion factors in a metrological framework and as constants reflecting some foundational principles. In the third section, I will focus on the cosmological constant  $\Lambda$ , in relation to

the principle of the relativity of inertia which is at the foundation of the general theory of relativity. Einstein introduced  $\Lambda$  in 1917 and then abandoned it qualifying it as his biggest blunder. Now,  $\Lambda$  currently seems to be making a spectacular comeback. In the fourth section, I will explain why I believe that the five constants considered by Einstein play a foundational role, both in the metrological and conceptual frameworks.

## 2. Einstein and the universal constants

### 2.1. Terminology and classification of the constants

Since the terminology of the constants depends strongly on their interpretation, it may be useful to make precise the terminology I am going to use according to my own interpretation. I tend to distinguish two categories of constants:

1. The ones that I call *universal foundational physical constants*, that are *universal* because they are implied in the whole of physics and *foundational* (I prefer this word to *fundamental*) because they underlie the axiomatic conceptual framework and the basics units of physics.
2. The other constants, such as  $e$ , that are dimensionless once the units have been redefined thanks to the constants of the first category, are *parameters* in the framework either of *models*, such as the standard model of particle physics,

in which case they must be determined experimentally, or of *theoretical conjectures* that would, possibly, allow calculating them.

In principle, all these constants are supposed to be really constant. Any well confirmed variation of any one of them would imply either the existence of some physics *beyond the standard models* (BSM) or the questioning of the general axioms of theoretical physics, a feature that enhances the stakes of a reliable most accurate metrology.

## 2.2. Apogee of classical physics

At the end of the 19th century, the apogee of classical physics, the culmination of the scientific revolution, which had seen the rise and the development of modern science, consisted of several theories that succeeded into three important syntheses or unifications and which allowed modeling in a satisfactory manner all the phenomena that were then observable:

- The *electromagnetic theory of light* of Faraday, Maxwell, and Hertz who unified the electrical, magnetic and optical phenomena.
- The *theory of universal gravitation* of Galileo and Newton who unified terrestrial and celestial mechanics.
- The upgrading of the conceptual status of the *atomistic conception* to the one of a genuine scientific discipline by means of the *molecular theory of matter* and the *statistical thermodynamics* of Maxwell and Boltzmann.

The successes achieved by these theories were such that Lord Kelvin (William Thomson), analyzing in 1900 the field of investigation of physics, announced that it was in the process of completion except for two ‘small clouds’, of which he thought they would only require some adjustments to be resorbed. It was about the failure of the detection of the movement of Earth in the ether (experiment of Michelson and Morley) and the absence of a theoretical explanation to the observed spectrum of the black body, problems to which were added the one of the *photoelectric effect* and the one of the *precession of the perihelion of Mercury*.

## 2.3. Scientific revolution of the 20th century

It is well known that the consideration, in particular by Einstein, of four universal constants, the constant of gravitation  $G$ , the speed of light in vacuum  $c$ , the Planck’s constant  $h$ , and the Boltzmann’s constant  $k$  is at the origin of the relativity and quantum theories, thanks to which the ‘small clouds’ have been dissipated, and form the basis of the scientific revolution of the 20th century. In this respect, these four constants that belong to the first group of constants defined above, have played a foundational role.

The structure in a tripod of the theoretical framework of classical physics is still convenient for the physics of the 20th century. This framework now includes three theories each considering two of these four universal constants and extending the theories of the framework of classical physics:

- The *quantum theory of fields* (constants  $h$  and  $c$ ) extends and encompasses the electromagnetic theory of light and restores it as a quasi-classical approximation; it is the basis of the high energy physics that explores, at sub-atomic or even sub-hadronic scales, the structure of matter (elementary constituents and fundamental interactions). The discovery of the Higgs boson in 2012, using the large hydron collider at European Organization for Nuclear Research (CERN), the keystone of the standard model of this physics, is the mark of a true apogee.
- The theory of *general relativity* (constants  $G$  and  $c$ ) extends, encompasses and restores Newton’s theory of universal gravitation as a non-relativistic approximation. It is at the foundation of modern cosmology, which since the first model of the Big Bang, taking into account the expansion of the Universe at extragalactic scales up to the current apogee of observational astrophysics, is now also equipped with a genuine cosmological standard model (CSM).
- *Quantum statistics* (constants  $h$  and  $k$ ) extends the analytic mechanics, the kinetic theory of matter and the statistical thermodynamics. It serves as the basis for modeling (or for numerically simulating) all phenomena that, from the scale of the atom to that of the galaxy, rely on statistical physics and about which the quantum effects cannot be neglected. This physics is the framework of the phenomenological consolidation of the standard models of the physics of the structure of matter and of the cosmology (CSM), in terms of a *scientific cosmogony* the purpose of which, according to Georges Lemaître, ‘is to search for ideally simple initial conditions from which can result, by the interplay of known physical forces, the current world in its full complexity [1]’. While general relativity and quantum theory of fields form the basis for what could be called the physical laws and fundamental structures of the Universe, quantum statistics forms the basis of the physics of the *emergence* of the observable world at our scales. The role of this third component of the theoretical framework that considers not only the Planck’s  $h$  constant but also the Boltzmann constant  $k$  is, in my opinion, not less important than that of the first two ones. It will be shown in section 5 that, as a consequence of the *informational turn of the interpretation of quantum physics*, this role is rather *foundational* in the sense that it fixes the conditions of the possibility of physics, among them, the existence of a reliable metrology, which, in my opinion, is exactly the aim of the redefinition of the SI thanks to the consideration of universal constants.

## 3. Brief history of the cosmological constant

### 3.1. At the onset of relativistic cosmology, the Einstein–de Sitter debate over the cosmological constant and the principle of the relativity of inertia

The reason Einstein became interested in cosmology in 1916 was that he wanted to clarify the relation between the theory of general relativity he was establishing and what he called

Mach's principle. This relation is well explained in one of the conferences he gave at Princeton [2], in 1921:

'The assumption of the complete physical equivalence of the systems of coordinates, K and K' we call the "principle of equivalence"; this principle is evidently intimately connected with the theorem of the equality between the inert and the gravitational mass and signifies an extension of the principle of relativity to co-ordinate systems which are in non-uniform motion relatively to each other. In fact, through this conception we arrive at the unity of the nature of inertia and gravitation. For according to our way of looking at it, the same masses may appear to be either under the action of inertia alone (with respect to K) or under the combined action of inertia and gravitation (with respect to K'). The possibility of explaining the numerical equality of inertia and gravitation by the unity of their nature gives to the general theory of relativity, according to my conviction, such a superiority over the conceptions of classical mechanics, that all the difficulties encountered in its development must be considered as small in comparison. (...) Although all of these effects are inaccessible to experiment, because  $\kappa$  is so small, nevertheless they certainly exist according to the general theory of relativity. We must see in them a strong support for *Mach's ideas as to the relativity of all inertial actions*. If we think these ideas consistently through to the end we must expect the *whole* inertia, that is, the whole  $g_{\mu\nu}$ -field, to be determined by the matter of the Universe, and not mainly by the boundary conditions at infinity.'

From this quotation, we can understand why, on the one hand, Einstein became interested in cosmology—because according to 'Mach's ideas', inertia of a test body is but the gravitation exerted on it by the rest of the Universe, and on the other hand, why he considered a *finite Universe*<sup>1</sup>—because he wanted to avoid uncontrollable assumptions about the boundary conditions at infinity.

But, with a finite Universe he faced the problem of a Universe which should eventually collapse under the action of its own gravitation; he thus tried a model involving a static Universe [3], thanks to a *cosmological term* equal to the cosmological constant  $\Lambda$  multiplying the metric field, that in principle can be added in the left-hand side of his equation and that he had previously discarded.

The controversy between Einstein and de Sitter was about Mach's principle that they agreed to name the *postulate of the relativity of inertia*. This controversy is accurately exposed in the three papers that de Sitter published in 1916–1917 [4]. In the first cosmological model that Einstein had proposed in 1917 [3], he had enunciated the principle of the relativity of inertia to which he refers as Mach's principle.

'In any coherent theory of relativity, there cannot be inertia with respect to "space" but only inertia of masses with respect to one another. Consequently, if in space, I

take a mass far enough from all the other masses in the Universe, its inertia must go to zero.'

In the correspondence, they had in March 1917, Einstein and de Sitter agreed on a formulation of this principle which makes of it a genuine foundational principle: in a postscript added by de Sitter at the end of his second paper in [4] he refers to and endorses a statement made (in German) by Einstein:

#### Postscript

'Prof. Einstein, to whom I had communicated the principal contents of this paper, writes "to my opinion, that it would be possible to think of a Universe without matter is unsatisfactory. On the contrary the field  $g_{\mu\nu}$  *must be determined by matter, without which it cannot exist* [underlined by de Sitter]. This is the core of what I mean by the postulate of the relativity of inertia". He therefore postulates what I called above the logical impossibility of supposing matter not to exist. I can call this the "material postulate" of the relativity of inertia. This can only be satisfied by choosing the system A, with its world matter, i.e. by introducing the constant  $\lambda^2$ , and assigning to the time a separate position amongst the four coordinates.

On the other hand, we have the "mathematical postulate" of the relativity of inertia, i.e. the postulate that the  $g_{\mu\nu}$  shall be invariant at infinity. This postulate, which, as has already been pointed out above, has no physical meaning, makes no mention of matter. It can be satisfied by choosing the system B, without a world matter, and with complete relativity of the time. But here also we need the constant  $\Lambda$ . The introduction of this constant can only be avoided by abandoning the postulate of the relativity of inertia altogether.'

In this postscript, de Sitter also summarizes all the issues of the debate he had with Einstein, which are discussed in the core of his three papers. What he calls 'system A' refers to the Einstein's cosmological model of [3], i.e. a spatially finite Universe obeying Einstein's equation to which has been added a cosmological term (the 'constant  $\Lambda$ ') allowing to satisfy the 'material principle of the relativity of inertia', thus playing the role of a hypothetical matter, 'of which the total mass is so enormously great, that compared with it all matter known to us is utterly negligible.'<sup>3</sup> This hypothetical matter I will call the *world matter*'.

The principle of the relativity of inertia is indeed related to Mach's principle:

'To the question: If all matter is supposed not to exist, with the exception of one material point which is to be used as a test-body, has then this test-body inertia or not? The school of Mach requires the answer *No*. Our experience however decidedly gives the answer *Yes*, if by "all matter" is meant all ordinary physical matter:

<sup>2</sup> In Einstein's and de Sitter's articles, the cosmological constant was written as lower case  $\lambda$ , in modern cosmology it is usually written as an upper case  $\Lambda$ , which is the convention we shall adopt in the present paper.

<sup>3</sup> It is interesting to note that, already in 1917, it was realized that the known matter represents only a negligible part of the content of the Universe.

<sup>1</sup> Actually, as noted by Cormac O'Raifeartaigh (whom I acknowledge for his remarks), in the original version of [3], in German, Einstein refers to his model as 'spatially closed' rather than 'finite'.

stars, nebulae, clusters, etc. The followers of Mach are thus compelled to assume the existence of still more matter: the “world matter”. If we place ourselves in this point of view, we must necessarily adopt system A, which is the only one that admits a world matter.’

The last statement of this quotation is reinforced in the postscript in which the world matter of system A is identified by the constant  $\Lambda$ .

‘System A’ is opposed to ‘system B’, which is the well-known de Sitter Universe containing no matter ( $\rho = 0$ ) and that nevertheless is a solution to Einstein’s equation with a cosmological constant. He proposed this solution to Einstein in a letter dated 20 March and received on 24 March. Einstein’s answer, that he commented in the quoted postscript, was communicated at the Royal Netherland Academy of Arts and Sciences (KNAW) meeting on 31 March. Einstein’s criticism of de Sitter’s ‘system B’ solution was based on a three-fold argument: (i) the corresponding Universe is spatially finite, (ii) it is bounded by a singularity<sup>4</sup>, and (iii) this singularity is at finite distance. de Sitter’s critical response to Einstein’s ‘system A’ was that it does not satisfy complete time relativity, but he had to recognize that his ‘system B’ solution satisfies only a ‘mathematical’ principle of the relativity of inertia that he formulates in the following way:

‘Once the system of reference of space- and time-variables has been chosen, Einstein’s equations determine the  $g_{\mu\nu}$  apart from *constants of integration* [underlined by me], or boundary conditions. Only the deviations of the actual  $g_{\mu\nu}$  from these values at infinity are thus due to the effect of matter. (...) If at infinity all  $g_{\mu\nu}$  were zero, we could truly say that the whole of inertia, as well as gravitation, is thus produced. This is the reasoning which has led to the postulate that at infinity all  $g_{\mu\nu}$  shall be zero. I have called this the *mathematical* postulate of relativity of inertia.’

### 3.2. World matter as an ether in the general theory of relativity

The necessity of including a world matter in the theory of general relativity is stressed by Einstein in the address he gave in Leyden [5] in 1920, in which the world matter is rather called an *ether*:

‘Recapitulating, we may say that according to the general theory of relativity space is endowed with physical qualities; in this sense, therefore, there exists an ether. According to the general theory of relativity space without ether is unthinkable; for in such a space there not only would be no propagation of light, but also no possibility of the existence for standards of space and time (measuring-rods and clocks), nor therefore any spacetime intervals in the physical sense. But this ether may not be thought of as endowed with the quality characteristic of ponderable media, as consisting of parts which may be tracked through time. The idea of motion may not be applied to it.’

<sup>4</sup> Actually, thanks to a remark by Cormac O’Raifeartaigh (see footnote 1), one should note that Einstein eventually conceded that the de Sitter model did not contain a singularity.

In this quotation we can see that the role of the principle of the relativity of inertia is foundational not only on a conceptual ground (‘the general theory of relativity space without ether in unthinkable’) but also on a metrological ground (‘no possibility of existence for standards of space and time’).

### 3.3. Abandonment by Einstein of the cosmological term

The severe criticism made by Eddington to Einstein’s cosmological model based on the instability of the equilibrium induced by the cosmological term, and the remark made by Friedman that the general theory of relativity implies a dynamical rather than a static Universe began to lead Einstein to give up his model. Finally, when the expansion of the Universe was discovered by Hubble in 1929, two years after the theoretical prediction made by Lemaitre of the expansion of the Universe, Einstein abandoned the constant  $\Lambda$ , qualifying it as his ‘biggest blunder’.

### 3.4. Principle of the relativity of inertia in modern cosmology

Now, it turns out that in modern cosmology based on the expansion of the Universe and on the quantum physical description of matter, the issues raised both by Einstein and de Sitter can be addressed, and that ‘system A’ and ‘system B’ solutions can be reconciled in an inflationary cosmology such as  $\Lambda$ CDM (Lambda cold dark matter, which is the name of the current standard model of cosmology) provided that the two ‘dark’ components of the content of the Universe (dark energy and dark matter) can be assimilated to two components of the de Sitter’s world matter, related to the vacuum:

- Because of expansion, the part of the Universe that is visible to us, and not the whole Universe, is spatially finite: it is a sphere of radius equal to the inverse of the Hubble constant (multiplied by  $c$ ). This boundary of the visible Universe is not a singularity, it is a *horizon*.
- Although, as noted by de Sitter in the second paper in [4] ‘In fact, there is no essential difference between the nature of ordinary gravitating matter and the world matter. Ordinary matter, the sun, stars, etc, are only condensed world matter, and it is possible, though not necessary, to assume all world matter to be so condensed’, *darkness, namely the absence of non-gravitational interactions*, allows distinguishing observationally world matter from ordinary matter.
- In a description of non-gravitational interactions of matter based on quantum field theory (QFT), the quantum vacuum, namely the ground state of the system of interacting quantum fields with the vanishing of all the occupation numbers, is not the nothingness and can allow modeling the world matter necessary to add to the known visible matter to satisfy the material principle of the relativity of inertia.

The starting point of modern cosmology considering the expansion of the Universe and the possible existence of a cosmological constant is Einstein’s equation, which, following the definitions, conventions (the vacuum velocity of light is put to 1), notations of [6] and the proposal of Gliner [7] and

Zeldovich [8] to take the cosmological term to the right-hand side<sup>5</sup>, reads:

$$\mathcal{R}_{\mu\nu} - \frac{1}{2}g_{\mu\nu}\mathcal{R} = 8\pi GT_{\mu\nu} + \Lambda g_{\mu\nu}, \quad (1)$$

where  $\mathcal{R}_{\mu\nu}$  is the Ricci curvature tensor,  $\mathcal{R}$  is the scalar curvature,  $g_{\mu\nu}$  is the metric tensor,  $\Lambda$  is the cosmological constant,  $G$  is Newton’s gravitational constant and  $T_{\mu\nu}$  is the stress-energy tensor.

The Robertson–Walker metric allows describing a homogeneous and isotropic Universe compatible with Einstein’s equation in terms of two cosmological parameters: the spatial curvature index  $k$ , an integer equal to  $-1, 0$ , or  $+1$  and the overall dimensional expansion (or contraction) radius of the Universe  $R(t)$ , depending only on time; note that due to the homogeneity, the geometry actually does not depend on the radial relative coordinate  $r$ , which is dimensionless:

$$ds^2 = dt^2 - R^2(t) \left[ \frac{dr^2}{1 - kr^2} + r^2 (d\theta^2 + \sin^2\theta d\phi^2) \right]. \quad (2)$$

One often uses a dimensionless scale factor  $a(t) = R(t)/R_0$  where  $R_0 \equiv R(t_0)$  is the radius at present-day.

In order to derive the Friedman–Lemaître equations of motion, one assumes that the matter content of the Universe is a perfect fluid for which the energy momentum tensor is expressed in terms of the isotropic pressure  $P$ , the energy density  $\rho$ , the space time metric  $g_{\mu\nu}$ , and of the velocity vector  $u = (1, 0, 0, 0)$  for the isotropic fluid in co-moving coordinates

$$T_{\mu\nu} = -Pg_{\mu\nu} + (P + \rho) u_\mu u_\nu. \quad (3)$$

Since, in all existent homogeneous and isotropic cosmologies, the metric appearing in (1) and in (3) is conformally flat [9], namely proportional to the flat Minkowski metric  $\eta_{\mu\nu}$  with a coefficient depending on time, the cosmological term, taken to the right-hand side of Einstein’s equation, can be interpreted as an effective energy-momentum tensor for the vacuum possibly playing the role of a world matter, rather than as a cosmological constant appearing in the action of the theory. If, as in equation (3), one associates a perfect fluid to the cosmological term, its pressure  $P_\Lambda(x)$  and energy density  $\rho_\Lambda(x)$  sum to zero and the world-matter energy tensor of the vacuum reduces to the pressure term.

## 4. $\Lambda$ CDM and the comeback of Lambda

### 4.1. From the hot Big Bang model to the new standard model of cosmology

The review by Debono and Smoot [10], in which one can find pedagogical explanations and useful references, shows how theories and observations motivated the adaptation of cosmological models from the first CSM, the hot Big Bang model (HBBM) to the new CSM the *concordance cosmology* or  $\Lambda$ CDM.

The successes of the HBBM include:

- The verification of the Hubble Law on the recession of distant galaxies, established through the measurement of their speeds (Doppler effect) and of their distances (by means of the Cepheids).
- The relative abundance of light elements explained thanks to the theory of primordial nucleosynthesis.
- The observation of the cosmic microwave background radiation (CMB) at about 3 K, predicted by Gamow and detected, by chance, in 1965 by Penzias and Wilson, then measured with more and more precision (cosmic background explorer (COBE), Wilkinson microwave anisotropy probe (WMAP), Planck).
- The pure and simple abandonment of the cosmological constant and of all cosmological models denying the expansion of the Universe (stationary Universe, continuous creation of matter, ‘tired light’, ...).

However, this model suffered from two problems linked to the primordial Universe, the horizon problem and the flatness problem.

The horizon problem stems from the fact that, in the HBBM, the primordial Universe, is, as Gabriele Veneziano says, ‘too large for its age’. As a singularity, the Big Bang implies an expansion of space that is so large that different regions of the Universe that can be observed today could not have been in causal contact in the primordial Universe; now it turns out that all the CMB regions seem to be in thermal equilibrium.

The flatness problem is known as a ‘fine tuning problem’. In the 1960s, the observations suggested that the density of the total content of the present-day Universe is compatible with the so-called critical density, namely the one corresponding to the vanishing of the spatial curvature index  $k$ , a time independent index; but maintaining such a vanishing throughout the history of the Universe implies a huge fine tuning that prevents any reliable cosmological modeling.

### 4.2. Assets of $\Lambda$ CDM

**4.2.1. Friedman–Lemaître equations.** In terms of the pressure and the density of the perfect fluid describing matter, the Friedman equations, with a non-vanishing cosmological constant (CC) possibly playing the role of a world matter, reads, in terms of the Hubble constant  $H = \dot{R}/R$

$$H^2 \equiv \left( \frac{\dot{R}}{R} \right)^2 = \frac{8\pi G\rho}{3} - \frac{k}{R^2} + \frac{\Lambda}{3}, \quad (4)$$

and

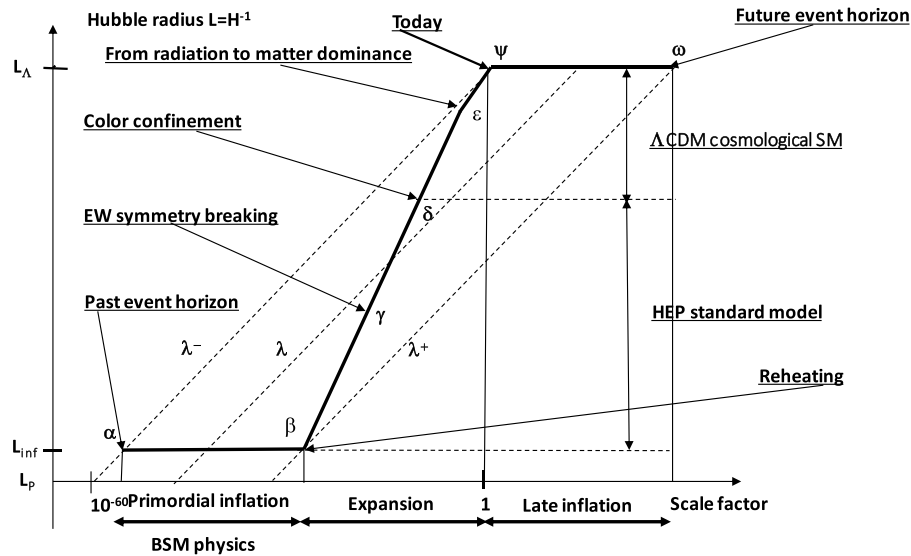
$$\frac{\ddot{R}}{R} = \frac{\Lambda}{3} - \frac{4\pi G}{3} (\rho + 3P). \quad (5)$$

Energy conservation leads to a third equation:

$$\dot{\rho} = -3H(\rho + P). \quad (6)$$

**4.2.2. The three stages of cosmic evolution in  $\Lambda$ CDM.** The phases of the cosmic evolution are well represented by the thick line in figure 1 in which the Hubble radius  $L = H^{-1}$ ,

<sup>5</sup> Actually, thanks to a remark by Cormac O’Raifeartaigh (see footnotes 1 and 4) it may be useful to note that Schrödinger was the first to suggest putting Lambda in the rhs.



**Figure 1.** The three stages of  $\Lambda$ CDM. The  $\Lambda$ CDM cosmology represented in a graphic in which the Hubble radius  $L = H^{-1}$  is plotted versus the scale factor  $a(t)$  (set to 1 today) in logarithmic scale. The cosmic evolution is schematized on the thick line, on which the cosmic time grows linearly in the inflation phases (horizontal parts from point  $\alpha$  to  $\beta$  and from  $\psi$  to  $\omega$ ) and exponentially in the expansion phase (from  $\beta$  to  $\psi$ ). All quantum fluctuations with generic wave-length  $\lambda$  exit from the Hubble horizon in the primordial inflation phase enter it in the expansion phase, and re-exit it in the late inflation phase. No information carrying quantum fluctuation with a wave-length smaller than  $\lambda^-$  or larger than  $\lambda^+$  enters the Hubble horizon. From this limitation, one infers that the total content of information in the observable Universe, which is proportional to inverse of the cosmological constant is **finite**.

the inverse of the time dependent Hubble parameter, is plotted versus the scale parameter  $a(t)$  (set to 1 today) in logarithmic scale, in such a way that the value zero of the scale parameter  $a$  (which would lead to a singularity as in the HBBM) is sent to minus infinity. In this figure, one can distinguish three stages<sup>6</sup>:

- The first stage of the evolution is the *primordial inflation stage*, namely a first de Sitter stage occurring at an energy of about  $10^{16}$  GeV, during which the Hubble radius is constant (about  $10^3$  Planck lengths) whereas the scale factor grows exponentially from  $\alpha$  to  $\beta$  in figure 1 by about 30 orders of magnitude.
- At point  $\beta$ , after the end of the primordial inflation is supposed to have occurred, which is usually called the *reheating* of the Universe. Point  $\beta$ , when the radius of expansion  $R(t)$  reaches about  $10^{30}$  Planck's lengths corresponding to the energy scale of CC or to the Compton wave length of a particle with a mass of about 2.5 meV, is the starting point of the second stage of  $\Lambda$ CDM, an expansion phase such as the one in a HBBM, in which point  $\beta$  plays the role of a physical non-singular Big Bang and can be called the *Big Bang ignition point* [11]. In the expansion phase, during which the content of the Universe obeys the standard Friedman–Lemaître cosmological equations of evolution, with a time dependence of the cosmic scale  $a(t)$  determined by the equation of state parameter  $w = P/\rho$  of the component that dominates the evolution at a given epoch, namely;

- an epoch of dominance by radiation ( $w = 1/3$ ) from point  $\beta$  to point  $\epsilon$  in figure 1 in which  $L \propto a^2$  followed by
- an epoch of dominance by pressure-less matter ( $w = 0$ ) from point  $\epsilon$  to point  $\psi$  (i.e. today<sup>7</sup>) in which  $L \propto a^{1.5}$ .
- In the third stage, extending from point  $\psi$  to point  $\omega$  in figure 1, the Universe will be dominated by the cosmological constant  $\Lambda$ . This stage, like the first one (from point  $\alpha$  to point  $\beta$ ) is an inflation one (a second de Sitter stage with a scale factor growing exponentially with the cosmic time, and a Hubble radius slowly increasing asymptotically to  $\sqrt{3/\Lambda}$ ) called the *late inflation stage* characterized by an equation of state  $w = -1$  compatible with the present day observation, ( $w = -1.019^{+0.075}_{-0.080}$ ) [20].

4.2.3. *Flatness sum rule.* It is useful to define a density, called the *critical density*

$$\rho_c \equiv \frac{3H^2}{8\pi G}, \tag{7}$$

which would be a solution to Friedman's equation (4) if the curvature index  $k$  and CC were zero. With respect to this critical density one defines for each component, including the one of CC, the relative contribution to the critical density, called its cosmological parameter, and rewrite the present day Friedman's equation (4) as

<sup>6</sup> According to a remark by Cormac O’Raifeartaigh (see footnotes 1, 4, 5) one should mention that Lemaître anticipated the three-speed model in the 1930s.

<sup>7</sup> Actually, due to the compressing of time scales in the cosmic evolution in the late epoch, the point  $\psi$  in figure 1 does not correspond to ‘today’, but rather to the transition from expansion to ‘re-inflation’ ( $\rho + 3P = 0$ ); the point corresponding to ‘today’ would be slightly on the right of  $\psi$ .

$$\begin{aligned} \Omega_{\text{tot}} &= \rho/\rho_c \\ k/R^2 &= H^2 (\Omega_{\text{tot}} - 1) \\ k/R_0^2 &= H_0^2 (\Omega_M + \Omega_R + \Omega_\Lambda - 1) \end{aligned} \quad (8)$$

where the subscript  $M$  stands for pressureless matter, (or ‘dust fluid’), the subscript  $R$  stands for radiation (or relativistic particles), and  $\Omega_\Lambda = \Lambda/3H^2$ . Since the curvature index  $k$  does not depend on time, its vanishing at present day implies its vanishing at all epochs, so in terms of time-dependent densities, Friedman’s equation (4) takes the form of the *flatness sum rule* that, in terms of density, reads as

$$\rho_M + \rho_R + \rho_\Lambda = \rho_c. \quad (9)$$

4.3. Emergent perspective of gravity and the first stage of the comeback of  $\Lambda$ : an integration constant in the solution of the Friedman–Lemaître equation

It is tempting to interpret the flatness sum rule as expressing the vanishing of the total (gravitational plus kinetic) energy density, equal to  $\rho_c$ , to which the vacuum energy density  $\rho_\Lambda$  has to be subtracted, thus qualifying  $\Lambda$ CDM as a cosmology of emergence out of the vacuum, a so-called ‘free lunch cosmology’. Actually, such an interpretation is suggested in [6] in the comment made about Friedman’s equation (4): ‘By interpreting  $-k/R^2$  Newtonianly as a “total energy”, we see that the evolution of the Universe is governed by a competition between the potential energy,  $8\pi G\rho/3$ , and the kinetic term  $(\dot{R}/R)^2$ . But, this suggestion is criticized a few lines below in the following way: ‘Note that the quantity  $-k/R_0^2 H_0^2$  is sometimes referred to as  $\Omega_k$ . This usage is unfortunate: it encourages one to think of curvature as a contribution to the energy density of the Universe, which is not correct’.

However, I think that such an interpretation of Friedman’s equation can be made correct, if, as advocated by Padmanabhan [12] and as I am going to explain now, one adopts the emergent perspective of gravity (EPG) per which the quantity that is conserved in the cosmic evolution is not a ‘total energy’ but rather a thermodynamic potential (i.e. defined up to an arbitrary additive constant), namely an *enthalpy* or total *heat* content.

The idea underlying the EPG is that in general relativity, *horizons* are unavoidable, and that, since horizons block information, *entropy* can be associated, through them, to spacetime, and thus that spacetime *has a micro-structure*. The Davies–Unruh [13] effect, the thermodynamics of black holes of Hawking [14] and Bekenstein [15], the Jacobson [16] interpretation of Einstein’s equation as an equation of state, or the interpretation of gravity as an entropic force by Verlinde ([17] and more recently [18]) say, rely on this idea which provides a possible thermodynamic route toward quantum gravity.

In the conventional approach, gravity is treated as a field which couples to the energy density of matter. The addition of a cosmological constant—or equivalently, shifting of the zero level of the energy—is not a symmetry of the theory since the gravitation field equations (and their solutions) change under such a shift. But in the EPG, rather than the *energy density* it is the *entropy density* which plays the crucial role and shifting the zero level of the entropy is now a symmetry of the theory.

In  $\Lambda$ CDM, the time-dependent null surface, with radius  $H^{-1}$  blocks information and can thus be endowed with an entropy [15] proportional to its area

$$S = (A/4L_P^2) = (\pi/H^2 L_P^2), \quad (10)$$

where  $L_P = (\hbar G/c^3)^{1/2}$  is the Planck’s length, and a temperature [14]

$$T = \hbar H/2\pi. \quad (11)$$

During a time-interval  $dt$ , the change of the gravitational entropy (i.e. the entropy associated with spacetime) is

$$(dS/dt) = (1/4L_P^2) (dA/dt), \quad (12)$$

and the corresponding *heat flux* is

$$T (dS/dt) = (H/8\pi G) (dA/dt). \quad (13)$$

For the matter contained in the Hubble volume, the classical (Gibbs–Duhem) thermodynamic relation tells us that the entropy density is  $s_m = (1/T) (\rho + P)$ , corresponding to a heat flux through the boundary equal to

$$TS_m A = (\rho + P) A. \quad (14)$$

Balancing gravitational (12) and matter (14) heat flux equations leads to  $\frac{H}{8\pi G} \frac{dA}{dt} = (\rho + P) A$ , which, with  $A = 4\pi/H^2$  gives

$$\dot{H} = -4\pi G (\rho + P). \quad (15)$$

Now, energy conservation for matter leads to

$$\begin{aligned} \frac{d(\rho a^3)}{dt} &= -P \frac{da^3}{dt} \\ \dot{\rho} &= -3H (\rho + P) \end{aligned} \quad (16)$$

which is nothing, but equation (6), combined with equation (15) and integrating over time leads to

$$\frac{3H^2}{8\pi G_N} \equiv \rho_c = \rho + \text{arbitrary constant}. \quad (17)$$

Comparing this last equation with equation (9), one sees that since the entropy density vanishes for the cosmological constant, the arbitrary constant can be put to  $\rho_\Lambda$  that acts as an integration constant because  $\rho_c \xrightarrow[t \rightarrow \infty]{} \rho_\Lambda$ . This feature is the main interest of the EPG: instead of being a parameter in the action of the theory, the cosmological constant appears as an integration constant in a particular solution of the equation. Now, as Padmanabhan says in the conclusion of his textbook on gravitation [19]

‘The integration constants which appear in a particular solution have a completely different conceptual status compared with the parameters that appear in the action describing the theory. It is much less troublesome to choose a fine tuned value for a particular integration constant in the theory if observations require us to do so. From this point of view, the cosmological constant problem is considerably less severe when we view from gravity from the emergent perspective.’



Equation (9) thus becomes

$$\rho_M + \rho_R = \rho_c - \rho_\Lambda = \Lambda_{\text{eff}}/8\pi G_N, \quad (18)$$

where the left-hand side is the sum of energy densities of all the components of the Universe (baryonic, relativistic and dark matters) contributing to gravitation and the right-hand side equated to an *effective cosmological constant*  $\Lambda_{\text{eff}}$ , which can be interpreted as the energy density of the world matter corresponding to a negative pressure and contributing to inertia, just as the ‘constant  $\Lambda$ ’ term in system A or system B of Einstein and de Sitter expresses the principle of equivalence of gravitation and inertia, so equation (18) can be re-written as

$$\rho(\text{matter}) + P(\text{world-matter}) = 0. \quad (19)$$

#### 4.4. Primordial inflation and the second stage of the comeback of $\Lambda$ : as a fifth foundational constant.

The primordial inflation phase cures the defects of the simple Big Bang model implied by the existence of a singularity. During the primordial inflation phase, the ‘young Universe’ is no more ‘too large for its age’, the whole content of the Universe is in causal contact, which solves the horizon problem. Inflation also solves the flatness problem: it flattens space in such a way that we can assume that at its end the spatial curvature is already compatible with zero, in agreement with the present-day observations ( $\Omega_k \equiv -k/R_0^2 H_0^2 = 0.0008^{+0.0040}_{-0.0039}$ ) [20].

Although this primordial inflation process remains conjectural (is it induced by a new ad hoc quantum field, the ‘inflaton’ or strictly related to the metric field?), the main asset of  $\Lambda$ CDM is to have provided inflation with the credibility it was lacking beforehand.

The fact that the scale, at which this primordial inflation is compatible with the one where grand unified theories (GUTs) are expected to be at work, suggests that it could also allow solving a third problem of the HBBM, the one of the absence of magnetic monopoles that would be predicted to be produced in large numbers in the GUT framework: inflation would expel the GUT magnetic monopoles beyond the horizon. More generally, there is a wide consensus to consider the primordial inflation as belonging to the realm of all the issues that are BSMs of elementary particles and of cosmology (BSM), including the ones related to pre-geometric or quantum gravity, namely gravity at the Planck’s scales of length  $L_P = \sqrt{\hbar G/c^3}$ , time  $T_P = \sqrt{\hbar G/c^5}$ , and energy  $E_P = \sqrt{\hbar c^5/G}$ .

It is remarkable that the primordial inflation ends when the radius of expansion  $R(t)$  reaches about  $10^{30}$  Planck’s lengths corresponding to an energy scale that is quite compatible with the one of  $\Lambda$  or to the Compton wave length of a particle with a mass of about 2.5 meV, which strongly suggests that the domain in which one can understand the true meaning of  $\Lambda$  is the one of BSM physics, including the realm of quantum gravity. Now it turns out that such a connection of  $\Lambda$  with quantum gravity can be guessed in the EPG discussed above. According to the EPG, which assumes that spacetime has a

microstructure in terms of ‘spacetime atoms’, the entropy of the Hubble horizon determined by CC is proportional to its area  $1/\Lambda$  expressed in the Planck’s unit of area  $A_P = \hbar G/c^3$ . This means that the area of the Hubble horizon divided by the Planck’s area is proportional to the ‘Avogadro constant of spacetime’,  $N_A^{\text{ST}}$  which, when multiplied by the Boltzmann’s constant  $k$ , represents to the total (very large, but *finite*) content of relevant information in the observable Universe  $\text{cosmIn}$ .

$$\frac{c^3 k}{\hbar G \Lambda} = \text{cosmIn} = N_A^{\text{ST}} k. \quad (20)$$

In [21], Padmanabhan showed that (i) the numerical value of the cosmological constant, as well as (ii) the amplitude of the primordial, scale invariant, quantum fluctuation spectrum can be determined in terms of a single free parameter,  $a_{\text{QG}}$ , which specifies the energy scale at which the Universe makes a transition from a pre-geometric or quantum phase to the classical phase. For a specific value of the parameter, the correct results for both (i) and (ii) were obtained.

## 5. Foundational role played by the five constants considered by Einstein

### 5.1. ‘Clocks and measuring rods’, the metrological concern of Einstein

In several circumstances, Einstein expressed his concerns about clocks and measuring rods in which, based on experience, he had to make some unavoidable assumptions. For instance, in one of his lectures given at Princeton in 1920, he says [22]:

‘In this the physical assumption is essential that the relative lengths of two measuring rods and the relative rates of two clocks are independent, in principle, of their previous history. But this assumption is certainly warranted by experience; if it did not hold there could be no sharp spectral lines; for the single atoms of the same element certainly do not have the same history, and it would be absurd to suppose any relative difference in the structure of the single atoms due to their previous history if the mass and frequencies of the single atoms of the same element were always the same.’

In retrospect, it seems that Einstein’s point of view about the universal constants is not far from the one adopted in the current methodology of the redefinition of the SI. This can be seen in the following quotation of the scientific autobiography of Einstein [23]:

‘The speed of light  $c$  is a quantity which intervenes as a “universal constant” in the equations of physics. But if one takes as a unit of time, not the second, but the time that light takes to go 1 cm,  $c$  no longer appears in the equations. In this sense, the constant  $c$  is only an apparent universal constant. It is manifest, and universally admitted, that it would also be possible to eliminate universal constants by introducing instead of the gram and the centimeter, adequately selected “natural” units

(e.g. mass and the radius of the electron). (...) Imagine that this has been realized [elimination of two universal constants to the benefit, for example of the mass and the radius of the electron]; then there appear in the fundamental equations of physics only dimensionless constants. About them, I would like to enunciate a principle which, provisionally, cannot be based on nothing else than on my confidence in the simplicity, or rather in *the intelligibility of Nature*: there is no arbitrary constant of this type. In other words: Nature is such that it is logically possible to establish laws that are so strongly defined that only constants susceptible of a complete rational determination appear in them (there are therefore no constants whose numerical values can be modified without the theory being destroyed).'

The same idea is also put forward in the following sentences from a 1935 article 'Physics and reality [24]':

'The very fact that the totality of our sense experiences is such that by means of thinking (operations with concepts, and the creation and use of definite functional relations between them, and the coordination of sense experiences to these concepts) it can be put in order, this fact is one which leaves us in awe, but which we shall never understand. One may say "the eternal mystery of the world is its comprehensibility". It is one of the great realizations of Immanuel Kant that the setting up of a real external world would be senseless without this comprehensibility.'

### 5.2. Incompleteness reproach made to 'quantum statistics' by Einstein

It is in the correspondence he exchanged with Schrödinger, a few days after the publication of the famous EPR paper [25] that Einstein expressed in the clearest way his criticism on quantum physics [26]:

'From the point of view of the principles, I absolutely do not believe in the existence of a statistical basis of physics, in the sense that quantum mechanics understands—and this, despite the successes that this formalism has gained case by case (...). I find that giving up a spatio-temporal apprehension of the reality is an idealistic and spiritualistic position.'

And he formulated the requirement that any description of reality should satisfy to be qualified as 'complete':

'In quantum theory, a real state of a system is described by a normed function of the coordinates  $\psi$  (of the configuration space). The evolution over time is unequivocally given by the Schrödinger equation. We would like to be able to say:  $\psi$  is in a one-to-one coordination with the actual state of the real system. The statistical character of the measurement results is exclusively to be due to the measuring devices or the measuring procedures. When it works, I speak of a complete description of reality by theory. But, if such an interpretation proves

impracticable, I say that the theoretical description is "incomplete".'

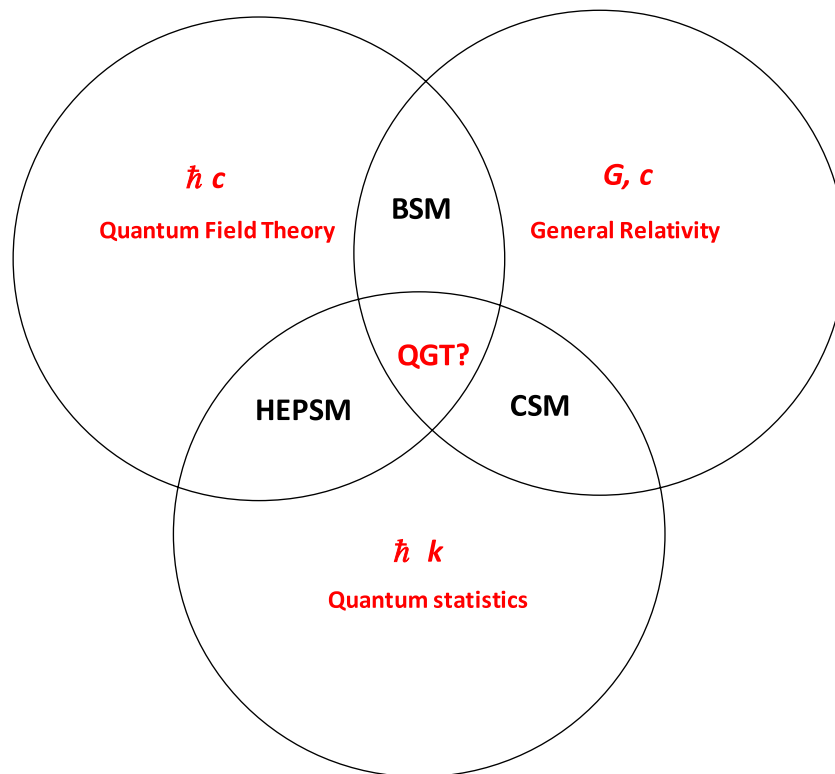
### 5.3. Informational turn: the recognition of the foundational role of information

Until recently, the consensus about the meaning of the universal constants, was that three constants, and only three, the gravitational constant, the speed of light, and the Planck constant, determine both the fundamental laws and the fundamental units of the general framework of physics. Considering these constants, one at a time, two at a time, or all three together, leads, according to this consensus, to structuring physics into subdisciplines that we can symbolically represent as the summits of a cube, 'the cube of theories'. What I call the *informational turn of physics* is the recognition of the foundational role of information and, thus, the inclusion of the Boltzmann's constant  $k$  that has the dimensional content of an information (or an entropy) in the panoply of the foundational constants.

The informational turn is well explained by Grinbaum in his paper [27] 'On epistemological modesty', in which he quotes another excerpt of Einstein's scientific autobiography:

"One is struck [by the fact] that the theory [of special relativity] introduces two kinds of physical things, i.e. (1) measuring rods and clocks, (2) all other things, e.g. the electromagnetic field, the material point, etc. This, in a certain sense, is inconsistent; strictly speaking measuring rods and clocks would have to be represented as solutions of the basic equations (objects consisting of moving atomic configurations), not, as it were, as theoretically self-sufficient entities. However, the procedure justifies itself because it was clear from the very beginning that the postulates of the theory are not strong enough to deduce from them sufficiently complete equations in order to base upon such a foundation a theory of measuring rods and clocks". (Einstein 1969, p. 59)

Epistemologically, it is unreasonable to expect, as Einstein did, that the theory of measuring rods and clocks could be based on a set of yet stronger postulates that would, at the same time, provide also an account of all physical phenomena measured by means of these rods and clocks. To see why Einstein found himself at an impasse, albeit an unnecessary one, consider the following schematic representation of physical theories. Assume that phenomena are best described by theories that are interconnected in the form of loop. Any particular theory is represented by cutting the loop at some point and thus separating the target object of the theory from the theory's presuppositions. (...) Consider the loop between physical theory and information. Physics and information mutually constrain each other, and every theory will give an account of but a part of the loop, leaving the other part for metatheoretic assumptions. In a first loop cut, information lies in the meta-theory of the physical theory, and



**Figure 2.** The theoretical/phenomenological landscape at the end on the 20th century. BSM = beyond the standard models; QGT? = quantum gravity theory (the question mark means that this theory is yet to be discovered); HEP SM = high energy physics standard model; CSM = cosmological standard model.

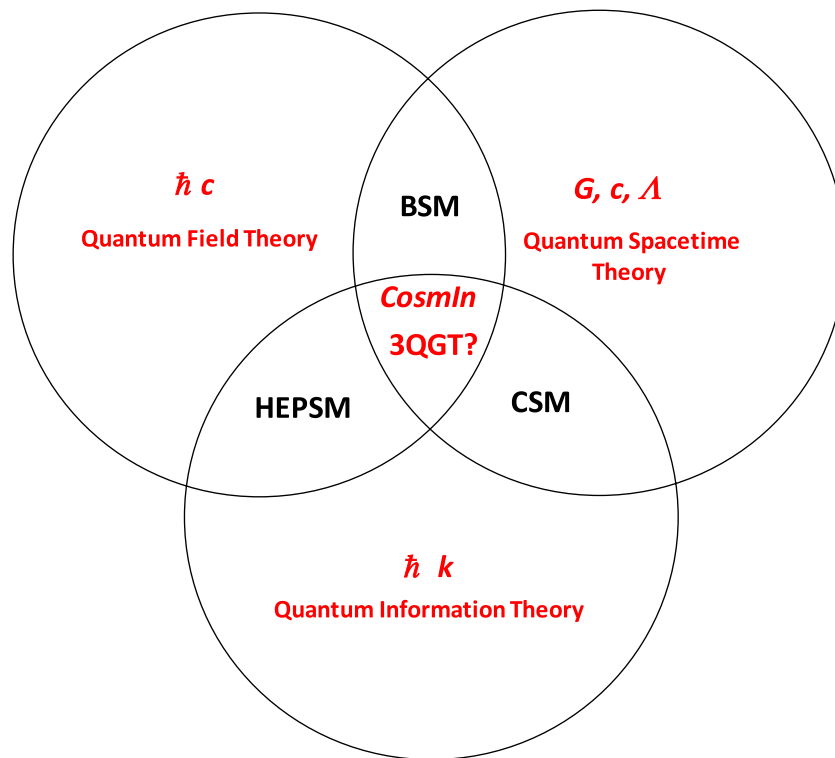
physics is therefore based on information. In a different loop cut, informational agents are physical beings, and one can describe their storage of, and operation with, information, by means of effective theories that are reduced, or reducible in principle, to physical theory.'

The way out of the impasse in which Einstein was stuck is not to adopt the cutting of the epistemic loop that leads to the statement that 'information is physical' but rather the one leading to the statement that 'physics is informational'. Now this is precisely what Bub proposes to do in his paper [28] titled 'Why the quantum?':

- (i) A quantum theory is best understood as a theory about the possibilities and impossibilities of information transfer, as opposed to a theory about the mechanics of non-classical waves or particles.
- (ii) Given the information-theoretic constraints, any mechanical theory of quantum phenomena that includes an account of the measuring instruments that reveal these phenomena must be empirically equivalent to a quantum theory.
- (iii) Assuming the information-theoretic constraints are in fact satisfied in our world, no mechanical theory of quantum phenomena that includes an account of measurement interactions can be acceptable, and the appropriate aim of physics at the fundamental level then becomes the representation and manipulation of information.

The informational turn is also what, according to Anton Zeilinger, provides quantum physics with the 'generally accepted conceptual foundation' it was missing in contrast with the general theory of relativity;

'In contrast to the theories of relativity, quantum mechanics is not yet based on a generally accepted conceptual foundation. It is proposed here that the missing principle may be identified through the observation that all knowledge in physics has to be expressed in propositions and that therefore the most elementary system represents the truth value of one proposition, i.e., it carries just one bit of information. Therefore, an elementary system can only give a definite result in one specific measurement. The irreducible randomness in other measurements is then a necessary consequence. (...) The Universe is participatory at least in the sense that the experimentalist by choosing the measurement apparatus, defines out a set of mutually complementary observables which possible property of a system can manifest itself as reality and the randomness of individual events stems from the *finiteness of information* [underlined by me]. (...) In conclusion, it may be very well said that information is the irreducible kernel from which everything else flows. Then the question why nature appears quantized is simply a consequence of the fact that information itself is quantized by necessity.'



**Figure 3.** Towards a new theoretical/phenomenological landscape. With respect to figure 2, QIT has replaced ‘quantum statistics’; 3QGT (thrice quantized gravity theory) has replaced QGT; mention is made to ‘CosmIn’, the total content of information in the observable Universe—see equation (20) and the caption of figure 1.

#### 5.4. Informational turn and the answer to the incompleteness reproach made by Einstein to ‘quantum statistics’

Schematically the overall outcome of the informational turn can be shown by the replacement of the theoretical tripod of 20th century physics (see figure 2) by the new theoretical tripod of figure 3, in which ‘quantum statistics’ is replaced by ‘quantum information theory’ (QIT), and ‘general relativity’ is replaced by the ‘quantum spacetime theory’ (QSTT). These replacements mean that the three foundational theories of the new tripod that each involve an elementary quantum (the quantum of action in quantum field theory (QFT), a quantum of spacetime area  $A_P = \hbar G/c^4$  in QSTT and a quantum of information in QIT), do not imply any recourse to statistics. This recourse is confined in the intersection of the QFT and QIT domains, which is the realm of quantum mechanics and of the high energy standard model (HEPSM) and in the intersection of the QIT and QSTT domains, which is the realm of astrophysics and CSM. As belonging to the phenomenological realm, these two standard models use statistical methods in the comparison with experiment or observation.

About the interpretation of quantum physics, the result of the informational turn is that the complementarity should be considered as that between the means of observation pertaining to the quantum theory of information and the observed system pertaining to quantum mechanics or to QFT, rather than that between two classically contradictory limits (waves or corpuscles, for example) of quantum physics. According to the proposed new interpretation, *matter/information*

*complementarity* is to be considered as the result of a *double quantization* involving the quantum of action and the quantum of information simultaneously, expressed by the key concept of *probability amplitude*, of which the modulus refers to the probability, i.e. the information and the phase to the coherence, i.e. the action.

In cosmology, it has been realized, thanks to the informational turn, that the entropy that is involved in the thermodynamics of spacetime underlying the EPG is not a statistical entropy but rather an entanglement entropy in such a way that it does not imply the existence of any hidden degree of freedom [29]. On an interpretive ground, the CSM relies on *spacetime/information complementarity* involving the spacetime quantum and information quantum.

In the third intersection, the one of the QFT and QSTT domains, the investigations about physics lying beyond the BSM can ignore statistics as far as they remain as purely theoretical conjectures about an ideal physics at zero temperature.

The intersection of the three domains of QFT, QIT, and QSTT is the domain of the purely theoretical program of building a supposedly complete *thrice quantized gravity theory* (3QGT), considering the five universal constants assembled in *cosmIn*—see equation (20).

## 6. Conclusion

‘Nature is earlier than man, but man is earlier than natural science’, this aphorism of Von Weiszäcker, quoted by Zeilinger to strengthen his argument in favor of the informational turn,

could also have been used by Bitbol in favor of the *reflective turn* [30] in the philosophy (metaphysics) of quantum physics that he proposes:

‘Instead of either formulating new metaphysical images (as realists would do) or rejecting any metaphysical attempt (as empiricists would do), the case of quantum mechanics might well require from us a complete redefinition of the nature and task of metaphysics. The sought redefinition can be performed in the spirit of Kant, according to whom metaphysics is the discipline of the boundaries of human knowledge. This can be called a “reflective” conception of metaphysics. In this paper, each one of the most popular “interpretations” of quantum mechanics is shown to be naturally associated with a variety of Kant-like reflective metaphysics. Then, the two major “paradoxes” of quantum mechanics (the measurement problem and the EPR correlations) are reformulated by way of this reflective attitude, and they are thereby “dissolved”. Along with this perspective, quantum mechanics becomes one of the most elegant and understandable theories of the history of physics in addition of being one of the most efficient. The only point that must be clarified is why it looks culturally so difficult to accept a reflective and non-ontological standpoint on physical theories.’

In conclusion, it must be stressed that this reflective turn is also, as said above, what underlies the redefinition of the SI because the five universal constants  $c$ ,  $h$ ,  $k$ ,  $G$ , and  $\Lambda$  on which it is based are all related to the boundaries of objective knowledge provided by physics.

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## References

- [1] Lemaître G 1946 *L’hypothèse de l’atome: Primitive-Essai de Cosmogonie* (Neuchâtel: Editions du Griffon) p 176
- [2] Einstein A 1923 *The Meaning of Relativity: Four Lectures Delivered at Princeton University* (Princeton, NJ: Princeton University Press) p 64
- [3] Einstein A 1917 *Cosmological (Considerations About the General Theory of Relativity)* (Sitzungsberichte: Preussische Akademie des Wissenschaften) pp 142–52  
Einstein A 1993 *Œuvres choisies, Tome 3, Relativités* (Paris: Le Seuil) p 91 (In French)
- [4] de Sitter W 1917 On the relativity of rotation in Einstein’s theory *KNAW, Proc.* vol 19 (*Amsterdam*) pp 527–32 (first paper)  
de Sitter W 1917 On the relativity of inertia. Remarks concerning Einstein’s latest hypothesis *KNAW, Proc.* vol 19 (*Amsterdam*) pp 1217–25 (second paper)  
de Sitter W 1917 On Einstein’ theory of gravitation and its astronomical consequences *KNAW, Mon. Notices* **78** 3–28 (third paper)
- [5] Einstein A 1922 Ether and the theory of relativity: an address delivered on May 5th, 1920, in the University of Leyden *Sidelights on Relativity* (London: Methuen & Co.) pp 23–4
- [6] Beringer J *et al* 2012 Particle data group *Phys. Rev. D* **86** 010001
- [7] Gliner E B 1966 Algebraic properties of the energy-momentum tensor and vacuum-like states of matter *Sov. Phys.-JETP* **22** 378
- [8] Zeldovich Y B 1967 The cosmological constant and the theory of elementary particles *Sov. Phys. Usp.* **11** 381
- [9] Gürsey F 1963 Reformulation of general relativity in accordance with Mach’s principle *Ann. Phys., NY* **24** 211  
Infeld L and Schild A 1945 A new approach to kinematic cosmology *Phys. Rev.* **68** 250  
Tauber G 1967 Expanding universe in conformally flat coordinates *J. Math. Phys.* **8** 118
- [10] Debono I and Smoot G F 2016 General relativity and cosmology: unsolved questions and future directions *Universe* **2** 23
- [11] Bjorken J D 2004 *The Classification of Universes* (arXiv:[astro-ph/0404233v1](https://arxiv.org/abs/astro-ph/0404233v1))
- [12] Padmanabhan T 2016 *The Atoms of Space, Gravity and the Cosmological Constant* *Int. Jour. Mod. Phys. D* **25** 1630020
- [13] Davies P 1975 Scalar production in Schwarzschild and Rindler metrics *J. Phys. A* **8** 609  
Unruh W 1976 Notes on black-hole evaporation *Phys. Rev. D* **14** 870
- [14] Hawking S 1975 Particle creation by black holes *Commun. Math. Phys.* **43** 199  
Hawking S 1975 Black holes and thermodynamics *Phys. Rev. D* **13** 191
- [15] Bekenstein J D 1973 Core-electron splittings and hyperfine fields in transition-metal compounds *Phys. Rev. D* **7** 2333  
Bekenstein J D 1974 Generalized second law of thermodynamics in black-hole physics *Phys. Rev. D* **9** 3292
- [16] Jacobson T 1995 Thermodynamics of spacetime: the Einstein equation of state *Phys. Rev. Lett.* **75** 1260–3
- [17] Verlinde E 2011 On the origin of gravity and the laws of Newton *J. High Energy Phys.* JHEP04(2011)029
- [18] Verlinde E 2017 *Emergent Gravity and Dark Universe* *SciPost Phys.* **2** 016
- [19] Padmanabhan T 2010 *Gravitation—Foundations and Frontiers* (Cambridge: Cambridge University Press) p 688
- [20] Planck 2015 results. XIII. Cosmological parameters (arXiv.org June 2016:[1502.01589v2](https://arxiv.org/abs/1502.01589v2))
- [21] Padmanabhan T and Padmanabhan H 2017 Cosmic information, the cosmological constant and the amplitude of primordial perturbations *Phys. Letts. B* **773** 81–5
- [22] Einstein A *The Meaning of Relativity op. cit.* pp 70–1
- [23] Einstein A 1970 *Autobiographical Notes* ed S A Paul (New York: MJF Books) p 61
- [24] Einstein A 1936 Physics and reality *J. Franklin Inst.* **221** 349–82
- [25] Einstein A, Podolsky B and Rosen N 1935 *Phys. Rev.* **XLVII** 777–80
- [26] Einstein A 2002 *Letter to Schrödinger* (Paris: Le Seuil) pp 465–70 (Einstein-Physique, philosophie, Politique-textes choisis et commentés par Françoise Balibar)
- [27] Grinbaum A 2010 On epistemological modesty *Philosophica* **83** 139–50

- [28] Bub J 2004 Why the quantum? *Studies in History and Philosophy of Science Part B: Studies in History and Philosophy of Modern Physics* **35** 241–66
- [29] Chirco G, Haggard H M, Riello A and Rovell C 2014 Spacetime thermodynamics without hidden degrees of freedom *Phys. Rev. D* **90** 044044
- [30] Bitbol M 2008 Reflective metaphysics: understanding quantum mechanics from a Kantian standpoint *Philosophica* **83** 53–83
- [31] O’Raifeartaigh C, O’Keeffe M, Nahm W and Mitton S 2018 One hundred years of the cosmological constant: from ‘superfluous stunt’ to dark energy *Eur. Phys. J. H* **43** 73–117