

THE AUSTRALIAN NATIONAL UNIVERSITY

## What we can learn from EO transitions and how?

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## Outline

> Emission of EO transitions
$>$ New tabulation of $\Omega_{C E}(E O)$
> e- $\gamma$ angular correlations of EO + M1+E2 transitions
$>$ EO transitions and the evolution of shape co-existence
W W-Os-Pt-Hg (Z=82, N=126)
$\square \mathrm{Fe}(\mathrm{Z}=\mathrm{N}=28)$
$\square 12 C(Z=N=8)$
$>$ Monopole transitions in atomic nuclei - new review

## E0 transitions - 75 years on

The Absolute Intensities and Internal Conversion Coefficients of the $\gamma$-Rays of Radium B and Radium C.

By C. D. Ellis, F.R.S., Fellow of Trinity College, Cambridge, and G. H. Aston, M.A., B.Sc., Trinity Hall, Cambridge.
Proc. Roy. Soc. (London) 129 (1931) 180-207
Highly converted K, L1, M1 lines corresponding to 1.426 MeV transition

$$
\begin{gathered}
1.4155 \mathrm{MeV} \mathrm{E} 0 \text { in }{ }^{214} \mathrm{Po} \\
\mathrm{~T}_{1 / 2}=99(3) \mathrm{ps} ; \rho^{2}=0.0013(2)
\end{gathered}
$$

### 6.05 MeV EO pairs in ${ }^{16} \mathrm{O}$

Fowler \& Lauritsen, Phys. Rev. 56 (1939) 840
6.05 MeV EO $\left({ }^{16}\right.$ O) E1+M1 double photon Gorodetzky et al., Phys. Rev. Lett. 7 (1961) 170

CONSTITUTION OF


OXFORD
AT THE CLARENDON PRESS 1931

## 1931: <br> $0 \rightarrow 0$ transition: quantum transition forbidden

## Formation of EO transitions


$>$ "Normal conversion" ( $\mathrm{L}>0$ ):

- Small contribution from inside the nucleus $\square$ Inner part of the atom (K-L-M shells)

> EO conversion electrons:
$\square$ Monopole potential localised inside the nucleus
- Point nucleus approx.: $\mathrm{W}_{\mathrm{EO}} \Rightarrow$ vanishes
- Purely penetration effect



## EO transitions

Transition probability

$$
\mathrm{W}_{\mathrm{T}}=\mathbf{W}_{\gamma}+\mathrm{W}_{\mathrm{CE}}+\mathrm{W}_{\pi}
$$

Conversion coefficient

$$
\alpha_{\mathrm{ce}, \pi}=\mathbf{W}_{\mathrm{ce}, \pi} / \mathbf{W}_{\gamma}
$$

CE \& $\pi$

$$
\mathbf{W}_{\mathrm{ce}, \pi}=\mathbf{W}_{\gamma} \quad \times \alpha_{\mathrm{ce}, \pi}
$$

EO

$$
\mathrm{W}_{\mathrm{ce}, \pi}=\rho^{2}\left(\mathbf{0}^{+} \rightarrow \mathbf{0}^{+}\right) \times \Omega_{\mathrm{ce}, \pi}(\mathrm{Z}, \mathrm{\kappa})
$$

Monopole strength parameter
$\rho=\frac{\left\langle\mathbf{0}_{f}^{+}\right| \sum e_{j} r_{r}^{2}\left|\mathbf{0}_{\boldsymbol{i}}^{+}\right\rangle}{e \boldsymbol{R}^{2}}=\frac{\left\langle\mathbf{0}_{\boldsymbol{f}}^{+}\right| \boldsymbol{m}(E \mathbf{E})\left|\mathbf{0}_{\boldsymbol{i}}^{+}\right\rangle}{\boldsymbol{e R ^ { 2 }}}$
Reduced transition probability $B(E O)=e^{2} R^{4} \rho^{2}$

## EO transitions



## Transition probability $\mathbf{W}_{\mathrm{T}}=\mathrm{W}_{\gamma}+\mathrm{W}_{\mathrm{CE}}+\mathrm{W}_{\pi}$

Conversion Coefficient

$$
\alpha_{\mathrm{ce}, \pi}=\mathbf{W}_{\mathrm{ce}, \pi} / \mathbf{W}_{\gamma}
$$

CE \& $\pi$

$$
\mathbf{W}_{\mathrm{ce}, \pi}=\mathbf{W}_{\gamma} \quad \times \alpha_{\mathrm{ce}, \pi}
$$

EO

$$
\mathbf{W}_{\mathrm{ce}, \pi}=\rho^{2}\left(\mathbf{O}^{+} \rightarrow \mathbf{0}^{+}\right) \times \Omega_{\mathrm{ce}, \pi}(\mathrm{Z}, \mathrm{~K})
$$

Separation is less complete! nuclear atomic

Piet Van Isacker (GANIL) talk on Monday

$$
\rho=\sum_{k \in \text { proons }}^{Z}\left\langle\mathrm{f}\left(\frac{r_{k}}{R}\right)^{2}-\sigma\left(\frac{r_{k}}{R}\right)^{4}+\cdots\right) \quad\left(R=r_{0} A^{1 / 3}, r_{0}=1.2 \mathrm{fm}\right)
$$



Transition probability

$$
W_{T}=W_{\gamma}+W_{C E}+W_{\pi}
$$

Conversion Coefficient

$$
\alpha_{\mathrm{ce}, \pi}=\mathbf{W}_{\mathrm{ce}, \pi} / \mathbf{W}_{\gamma}
$$

CE \& $\pi$


$$
\mathbf{W}_{\mathrm{ce}, \pi}=\rho^{2}\left(\mathbf{0}^{+} \rightarrow \mathbf{0}^{+}\right) \times \Omega_{\mathrm{ce}, \pi}(\mathrm{Z}, \mathrm{k})
$$

nuclear atomic
$\mathrm{L} \neq 0$ : $\mathrm{W}_{\gamma}$ and $\mathrm{W}_{\text {ce, }, ~}$
$L=0$ i.e. EO: only $W_{\text {ce }, \pi}$ !
Need to measure electrons and/or electron-positron pairs
$>\Omega_{\kappa, \pi}(\mathrm{E} 0) \sim I_{\kappa, \pi}(\mathrm{E} 0)$
> Only ratios of $\Omega_{\mathrm{K}, \pi}(\mathrm{E} 0)$ could be measured

- K/ $\pi$
- K/L, K/LM, K/MN
- L/M

New $\Omega_{C E}$ calculations
Takahe (Jackson Dowie, ANU)

- Modified version of the CATAR code (Puli \& Raff, 1975)
> Relativistic Hartree-Fock-Slater atomic model
> Sliv`s surface current model
- Directional \& polarization particle parameters
> Penetration parameters
> $\Omega_{C E}$ electronic factors
$>Z$ up to 126
$\mathrm{S}_{1 / 2}, \mathrm{P}_{1 / 2}$ and $\mathrm{P}_{3 / 2}$ shells only

| $\Omega_{C E}(\mathrm{EO}), \mathrm{Mo}(\mathrm{Z}=42) 1000 \mathrm{keV}$ |  |  |  |
| :--- | :---: | :---: | :---: |
|  | Takahe <br> (2017) | Bell et al. <br> (1972) |  <br> seltrer <br> (1968) |
| K | $1.454 \mathrm{E}+9$ | $1.459 \mathrm{E}+9$ | $1.438 \mathrm{E}+9$ |
| L1 | $1.629 \mathrm{E}+8$ | $1.574 \mathrm{E}+8$ | $1.611 \mathrm{E}+8$ |
| L2 | $1.358 \mathrm{E}+6$ | $1.152 \mathrm{E}+6$ | $1.346 \mathrm{E}+6$ |
| M1 | $2.913 \mathrm{E}+7$ |  |  |
| M2 | $2.566 \mathrm{E}+5$ |  |  |
| N1 | $4.983 \mathrm{E}+6$ |  |  |
| N2 | $3.788 \mathrm{E}+4$ |  |  |
| O1 | $3.807 \mathrm{E}+5$ |  |  |

E0 Workshop, CEA 2017

## $\Omega(E 0)$ - how good are they?




## E0 vs E2 transitions



> Z and atomic shell dependent
> Increases with energy
> K/L weak dependence on $Z$ (4 to 9)
> Pair conversion dominant at low $Z$

## E0+M1+E2 transitions

Church, Rose and Weneser (1958)
$>$ EO can proceed in competition of E2 \& M1
> e- $\gamma$ angular correlations: a sensitive test
> First observation: 334.0-356.5 cascade in ${ }^{196} \mathrm{Pt}$ (Gerholm \& Pettersson 1958)


$$
\begin{aligned}
& W(\gamma \gamma, M 1+E 2)=P_{0}+ \\
& \qquad \frac{1}{1+\delta^{2}}\left[A_{2}^{e}+2 \delta A_{2}+\delta^{2} A_{2}^{m}\right] P_{2}+\frac{1}{1+\delta^{2}}\left[A_{4}^{e}\right] P_{4}
\end{aligned}
$$

$$
W(e \gamma, M 1+E 2)=P_{0}+
$$

$$
\frac{1}{1+p^{2}}\left[b_{2}^{e} A_{2}^{e}+2 p b_{2} A_{2}+p^{2} b_{2}^{m} A_{2}^{m}\right] P_{2}+\frac{1}{1+p^{2}}\left[b_{4}^{e} A_{4}^{e}\right] P_{4}
$$

$$
W(e \gamma, E 0+M 1+E 2)=\frac{1+p^{2}}{1+p^{2}+q^{2}} W(e \gamma, M 1+E 2)+
$$

$$
\frac{q^{2}}{1+p^{2}+q^{2}} P_{0}+\frac{q}{1+p^{2}+q^{2}} b_{o} P_{2}
$$

Multi-detector electron- $\gamma$ arrays
Need to evaluate numerically the e-e/e- $\gamma$ correlation!

Mixing ratios:

$$
\begin{aligned}
& \delta^{2}=\frac{N_{\gamma}^{E 2}}{N_{\gamma}^{M 1}} \\
& p^{2}=\frac{\alpha_{K}^{M 1}}{\alpha_{K}^{E 2}} \delta^{2}
\end{aligned}
$$

$$
q^{2}=\frac{N_{K}^{E 0}}{N_{K}^{E 2}}
$$

## E0+M1+E2 transitions

Mixed EO+M1+E2 conversion coefficients

$$
\begin{aligned}
\alpha_{K}^{e x p}(E 0+M 1+E 2)= & \frac{1}{1+\delta^{2}}\left[\delta^{2}\left[1+q^{2}\right] \alpha_{K}(E 2)+\alpha_{K}(M 1)\right] \\
& \text { Mixing ratios: } \quad \delta^{2}=\frac{N_{\gamma}^{E 2}}{N_{\gamma}^{M 1}} \quad \boldsymbol{q}^{2}=\frac{N_{K}^{E 0}}{N_{K}^{E 2}}
\end{aligned}
$$

Reduced EO matrix element

$$
\rho^{2}(E 0)=\frac{1}{\tau(E 0) \times \Omega(E 0)}=\boldsymbol{q}^{2} \frac{\alpha_{K}(E 2) W_{\gamma}(E 2)}{\Omega(E 0)}
$$

$B(E O) / B(E 2)$ ratio (Rasmussen 1960)

$$
X(E 0 / \mathrm{E} 2)=\frac{\rho^{2}(e 0) e^{2} R_{o}^{4}}{B(E 2)}=\frac{2.54 \times 10^{9} \times A^{4 / 3} E_{\gamma}^{5} q^{2} \alpha_{K}(E 2)}{\Omega_{K}(E 0)}
$$

## © ANU <br> Evolution of shape co-existence



## Evolution of shape co-existence W-Os-P†



From radioactive decay: $\gamma, \gamma-9$, CE

174Pt: Dracoulis, et al., PRC 44, R1246 (1991)
176Pt: Dracoulis, et al., J. Phys. G 12, L97 (1986)
${ }^{172}$ Os: Davidson, et al., Nucl. Phys. A568, 90 (1994)
${ }^{174-182}$ Os: Kibedi, et al., Nucl. Phys. A567, 183 (1994)
172-178W: Kibedi, et al., Nucl. Phys. A 688669 (2001)

## Evolution of shape co-existence W-Os-Pt

172-178W: Kibedi, et al.,
Nucl. Phys. A 688669 (2001)
$>0+-0+$ EO transitions (4)
$>\mathrm{J}+-\mathrm{J}+(\mathrm{J}=2,4,6) \mathrm{EO}+\mathrm{M} 1+\mathrm{E} 2$ transitions (19)

No T1/2 - only $X=B(E O) / B(E 2)$

Table 6
E0 component of $J_{\mathrm{i}}^{+} \rightarrow J_{\mathrm{f}}^{+}$transitions in $A=170$ to 178 tungsten isotopes (only for $K$-conversion electron lines have been considered)

| $J_{\mathrm{i}}^{+} \rightarrow J_{\mathrm{f}}^{+}$ | A | $\begin{gathered} E_{\gamma} \\ {[\mathrm{keV}]} \end{gathered}$ | $\begin{gathered} \alpha_{K}(\exp ) \\ \times 100 \end{gathered}$ | $\begin{gathered} \alpha_{K}(\mathrm{E} 2) \\ \quad \times 100 \end{gathered}$ | $\begin{gathered} \alpha_{K}(\mathrm{M} 1) \\ \times 100 \end{gathered}$ | $\delta(\mathrm{E} 2 / \mathrm{M} 1)$ | $q^{2}(\mathrm{E} 0 / \mathrm{E} 2)$ | $X(\mathrm{E} 0 / \mathrm{E} 2)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $0_{\beta}^{+} \rightarrow 0_{g}^{+}$ | 172 | 761.6 | E0 |  |  |  | 3.6(11) | 0.060(18) |
|  | 174 | 792.2 | E0 |  |  |  | 2.1(4) | 0.041(2) |
|  | 176 | 844.0 | E0 |  |  |  | 2.5(6) | 0.058(15) |
| $\begin{aligned} 0_{3}^{+} & \rightarrow 0_{g}^{+} \\ 2_{\beta}^{+} & \rightarrow 2_{g}^{+} \end{aligned}$ | 178 | 1294.4 | E0 |  |  |  | 26(2) | 1.73(12) |
|  | 172 | 743.7 | 4.52(20) | 0.677 | 1.78 | $-10.3{ }_{-7.0}^{+3.0}$ | 5.7(3) | 0.153(8) |
|  | 174 | 777.0 | 8.4(9) | 0.618 | 16.0 | $-4.5{ }_{-1.3}^{+0.9}$ | 13.1(15) | 0.39(5) |
|  | 176 | 822.2 | 5.6(4) | 0.551 | 1.38 | $-2.7_{-0.5}^{+0.4}$ | $10.2_{-1.1}^{+1.3}$ | $0.346_{-0.038}^{+0.042}$ |
|  | 178 | 976.5 | 0.76(5) | 0.392 | 0.902 | $-12.3_{-6.4}^{+2.8}$ | $0.94{ }_{-0.13}^{+0.13}$ | $0.045_{-0.006}^{+0.006}$ |
| $2_{\gamma}^{+} \rightarrow 2_{g}^{+}$ | 172 | 807.1 | 1.14(15) | 0.572 | 1.45 | $+7.6_{-7.4}^{+2.5}$ | 7.4(19) | 0.17(4) |
|  | 176 | 932.4 | 0.83(16) | 0.429 | 1.01 | $\begin{array}{r} -1.4 \\ +3.0_{-0.7}^{+1.0} \end{array}$ | $0.89_{-0.47}^{+0.42}$ | $0.039_{-0.018}^{+0.018}$ |
|  | 178 | 1004.6 | 0.90(8) | 0.370 | 0.840 | $>+2$ | $1.5_{-0.3}^{+0.3}$ | $0.076_{-0.013}^{+0.014}$ |
| $\begin{aligned} 2_{4}^{+} & \rightarrow 2_{g}^{+} \\ 4_{\beta}^{+} & \rightarrow 4_{g}^{+} \end{aligned}$ | 178 | 1311.5 | 1.49(13) | 0.223 | 0.436 | $1>\left.2\right\|^{\text {a }}$ | $6.9{ }_{-1.7}^{+0.7}$ | $0.61_{-0.15}^{+0.07}$ |
|  | 170 | 739.8 | 4.6(11) | 0.684 | 1.81 | $-3.3+1.6$ | 6.1(18) | 0.16(5) |
|  | 172 | 715.0 | 7.0(7) | 0.735 | 1.97 | $-4.1_{-1.9}^{+3.6}$ | 8.9(11) | 0.221(26) |
|  | 174 | 739.4 | 5.8(9) | 0.685 | 1.81 | -4.2 ${ }_{-1.1}^{+0.7}$ | 7.8(14) | 0.21(4) |
|  | 176 | 768.7 | 6.6(7) | 0.632 | 1.64 | $-2.2_{-1.2}^{+0.6}$ | $11.1_{-2.2}^{+3.0}$ | $0.33_{-0.06}^{+0.09}$ |
|  | 178 | 932.4 | 1.76(11) | 0.429 | 1.01 | $-6.6_{-3.0}^{+1.5}$ | $3.15{ }_{-0.28}^{+0.30}$ | $0.140_{-0.013}^{+0.013}$ |
| $4_{\gamma}^{+} \rightarrow 4_{g}^{+}$ | 172 | 865.1 | 1.4(4) | 0.497 | 1.22 | $+4.2_{-2.6}^{+1.1}$ | $1.8(9)$ | 0.007(3) |
|  | 178 | 1037.4 | 0.51(5) | 0.497 | 1.22 | $-1.9_{-1.2}^{+0.7}$ | 0.7(5) | 0.014(18) |
| $\begin{aligned} 4_{4}^{+} & \rightarrow 4_{g}^{+} \\ 6_{\beta}^{+} & \rightarrow 6_{g}^{+} \end{aligned}$ | 178 | 1255.1 | 1.01(8) | 0.242 | 0.485 | $1>\left.2\right\|^{\text {a }}$ | $3.7{ }_{-0.9}^{+0.4}$ | $0.30_{-0.07}^{+0.03}$ |
|  | 170 | 702.8 | 6.7(15) | 0.762 | 2.06 | $-1.7_{-2.5}^{+0.8}$ | 10(3) | 0.23(7) |
|  | 172 | 694.1 | 6.4(14) | 0.782 | 2.12 | $-5.0_{-}^{+3.2}$ | 10(3) | 0.23(7) |
|  | 176 | 696.6 | 4.9(9) | 0.776 | 2.10 | $1>\left.2\right\|^{\text {a }}$ | $6.2{ }_{-2.1}^{+1.5}$ | $0.15{ }_{-0.05}^{+0.04}$ |

## Evolution of shape co-existence W-Os-Pt

## Level systematics



## Evolution of shape co-existence W-Os-Pt



Kibedi, et al., Nucl. Phys. A 688669 (2001)

4-band mixing calculations
$>\mathrm{K}=0$ g.s. unperturbed ground-state rotation
$>\mathrm{K}=0$ "deformed" band
$>\mathrm{K}=2 \mathrm{~g}$-band
> K=0 "s"-band, unperturbed rotation-aligned band, back banding observed in (HI,xn)

P Parameters to fit excitation energies:
] Moment of inertia (VMI)
U Unperturbed band-head energies
Spin-independent interactions
$\square$ aligned angular momentum of the s-band

## Evolution of shape co-existence W-Os-P†



Kibedi, et al., Nucl. Phys. A 688669 (2001)
> Interactions:

- 150 keV ( $\mathrm{g}-\mathrm{d}$ )
- 30 keV ( $\mathrm{g}-\gamma$ )
- $5 \mathrm{keV}(\mathrm{d}-\gamma)$

4-band mixing calculations rotation
> K=0 "deformed" band
$>\mathrm{K}=2 \mathrm{~g}$-band
> K=0 "s"-band, unperturbed rotation-aligned band, back

$$
5 \operatorname{kev}(d-\gamma)
$$

$>\mathrm{K}=0$ g.s. unperturbed ground-state banding observed in (HI,xn)

## Evolution of shape co-existence W-Os-P†



> Smooth evolution across $\mathrm{Z}=80$ to 74 and $\mathrm{N}=98$ to 106
> Differences in deformation
> Unperturbed $\gamma$ and $d$ bands shifted down in energy as $Z$ >> 82

## Evolution of shape co-existence W-Os-Pt



## Evolution of shape co-existence $x$ (EO/E2) systematics


$\rho^{2}(E O)$
$0_{2}-0_{1}$
$B(E 0) / B(E 2)$
$\mathrm{O}_{2}-\mathrm{O}_{1}$

B(EO)/B(E2)
$2_{2}-2_{1}$
B(E0)/B(E2)
$2_{\gamma}-2_{1}$

15 data

42 data

39 data

24 data

Shape co-existence around $\mathrm{N}=\mathrm{Z}=28$


Heyde \& Wood, Rev. Mod. Phys. 83 (2011) 1467
$\mathrm{N}=\mathrm{Z}=28$ double magic
Shape co-existence "could emerge"
$\square$ Excited $0^{+}$around $\mathrm{N}=\mathrm{Z}=28$ : from mp -mh excitations from the $1 f_{7 / 2}$ to the $1 f_{5 / 2}, 2 p_{1 / 2}$ and $2 p_{3 / 2}$ orbits

E0 transitions: not very well known

A Aim: Characterise E0 transitions in $\mathrm{Z}=26, \mathrm{~N}=28,30,32$

- E0`s in Ni isotopes talk by Adam Garnsworthy

$\square\left(p, p^{\prime}\right)$ reaction, $E_{p}=6.7-7 \mathrm{MeV}$ DC beam
- $1-2 \mathrm{mg} / \mathrm{cm}^{2} 54,56,58 \mathrm{Fe}$ targets
$\square$ Singles gamma, conversion electron and pair conversion (Super-e)
$\varepsilon_{\mathrm{CE}}=0.1-0.5 \%$
$\varepsilon_{\mathrm{IPF}}=0.01 \%$

| E <br> $[\mathrm{keV]}$ | CEK/IPF <br> $(\mathbf{Z}=26)$ |
| :---: | :---: |
| 1500 | 1.6 |
| 2500 | 0.066 |
| 4000 | 0.009 |

## Extended level schemes



## Extended level schemes



## Angular correlation of $0-2-0$ cascades




$$
\begin{gathered}
\mathrm{W}_{\mathrm{yY}}(\theta)=\mathrm{N}\left(1+\mathrm{a}_{2} \cos ^{2}(\theta)+\mathrm{a}_{4} \cos ^{4}(\theta)\right) \\
0^{+} \rightarrow 2^{+} \rightarrow 0^{+} \\
\mathrm{a}_{2}=-3, \mathrm{a}_{4}=4
\end{gathered}
$$

|  | $\mathrm{a}_{2}$ | $\mathrm{a}_{4}$ |
| :---: | :---: | :---: |
| ${ }^{54} \mathrm{Fe}$ | $-2.8(6)$ | $3.9(7)$ |
| ${ }^{56} \mathrm{Fe}$ | $-2.6(9)$ | $3.7(10)$ |
| ${ }^{58} \mathrm{Fe}$ | $-2.8(8)$ | $3.8(9)$ | EO transitions $-{ }^{54} \mathrm{Fe}$

 EO transitions - 56 Fe


## EO transitions in $54,56,58 \mathrm{Fe}$

Experimental monopole strength

$$
\rho^{2}(E 0)=\frac{1}{\Omega(E 0) \times \tau(E 0)}
$$

$$
\varrho^{2}(E 0)=\frac{Z^{2}}{R_{0}^{4}} a^{2}\left(1-a^{2}\right)\left[\Delta\left\langle r^{2}\right\rangle\right]^{2}
$$

|  |  | $10^{3} \mathrm{r}^{2}(\mathrm{E} 0)$ |
| :---: | :---: | :---: |
| 54 Fe | $2561\left(0_{2}-\mathrm{O}_{1}\right)$ | $<80$ |
|  | $4291\left(0_{3}-\mathrm{O}_{1}\right)$ | $92(+22-23)$ |
| 56 Fe | $2942\left(0_{2}-\mathrm{O}_{1}\right)$ | $3.2(11)$ |
| 58 Fe | $2258\left(0_{2}-\mathrm{O}_{1}\right)$ | $<12$ |

${ }^{54} \mathrm{Fe} 1758\left(2_{2}-2_{1}\right)$ $\alpha_{k}=4.9 E-5(3) M 1+E 2$ Small/negligible EO


## Summary

| Z | 20 | 22 | 24 | 26 | 28 | 30 | 32 | 34 | 36 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Ge | 32 |  |  |  |  |  |  |  |  |
| Zn | 30 |  |  |  |  |  |  | $3.8_{4}$ | $0.19-54$ |
| Ni | 28 |  |  |  |  | $0.0063_{10}$ | $1-27$ |  |  |
| Fe | 26 |  |  |  | $<80$ | $3.2_{11}$ | $<12$ |  |  |
| Cr | 24 |  |  |  |  |  |  |  |  |
| Ti | 22 |  |  |  |  |  |  |  |  |
| Ca | 20 | $25.6_{7}$ | $140_{12}$ | $140_{50}$ | $14.5_{9}$ |  |  |  |  |

$\square$ New results in $54,56,58 \mathrm{Fe}$ : extended level schemes, new $\mathrm{T} 1 / 2, \delta(E 2 / \mathrm{M} 1)$, EO transitions, $\rho^{2}$ (EO)
$\square$ Future: look for EOs between $\mathrm{J}>0$ states to characterise bands built on excited $0^{+}$states
$\square$ Interpretation within the bandmixing approach
$\square E O s$ in $\mathrm{Cr}(Z=24)$ and $\mathrm{Ti}(Z=22)$ to explore $\mathrm{N}=28$ isotones

## The Hoyle state



Heyde \& Wood, Rev. Mod. Phys. 83 (2011) 1467
Hoyle state: not a typical excited $0^{+}$at south-west from the double magic ${ }^{16} \mathrm{O}$

## The Hoyle state


D.J. Marín-Lámbarri, et al., PRL 113 (2014) 012502

- $\Gamma\left(0_{2}\right)=9.3(9) \mathrm{eV} ; \mathrm{T}_{1 / 2}\left(0_{2}\right)=3.5(3) \times 10^{-17} \mathrm{~s}$
[ "Extended object" (Brink 1966) RMS=2.89(4) fm = 1.2* RMS(g.s.) PRC 80 (2009) 054603
- $\rho^{2}(E 0)=500(81)$ ADNDT 89 (2005) 77
- $2^{+}$at 9.8 MeV

Nucl. Phys. A738, (2004) 268; Phys. Rev. C 84 (2011) 054308; 80 (2009) 041303(R); 84 (2011) 027304; 86 (2012) 034320; PRL 113 (2014) 012502

- $4^{+}$at 13.3 MeV

Phys. Rev. C83 (2011) 034314
$\square 0_{3}{ }^{+}$at $10.3 \mathrm{MeV} ; \Gamma\left(0_{3}\right)=2.7 \mathrm{MeV}$ Nucl. Phys. A738, (2004) 268
$\square$ Microscopic $\alpha$-cluster model /exp
$\mathrm{E}\left(\mathrm{O}_{2}\right)-\mathrm{E}_{3 \alpha}=0.23 / 0.38 \mathrm{MeV}$
$\Gamma\left(0_{2}\right)=7.6 / 9.3(9) \mu \mathrm{eV}$
$M(E 0)=6.3 / 5.4(2) \mathrm{fm}^{2}$
Yasuro Funaki, Phys. Rev. C 94 (2016) 024344

## The radiative width of the Hoyle state

Unc. \#
[\%] exp

$$
\begin{aligned}
& r_{3 \alpha} \propto\left[\Gamma_{r a d}\right] \exp \left(-\left[Q_{3 \alpha}\right] / k T\right) \\
& \Gamma_{r a d}=\left[\frac{\Gamma_{r a d}}{\Gamma}\right] *\left[\frac{\Gamma}{\Gamma_{-}(E 0)}\right]<\left[\Gamma_{\pi}(E 0)\right]
\end{aligned}
$$

|  |  | Unc. <br> [\%] | \# <br> exp |
| :--- | :--- | :---: | :---: |
| $\Gamma_{\mathrm{rad}} / \Gamma\left(\times 10^{-4}\right)$ | $4.19(11)$ | 8 |  |
| $\Gamma_{\pi}(\mathrm{E} 0) / \Gamma\left(\times 10^{-6}\right)$ | $6.7(6)$ | 4 |  |
| $\Gamma_{\pi}(\mathrm{E} 0)(\mu \mathrm{eV})$ | $62.3(20)$ | 8 |  |

"traditional approach"
${ }^{12} \mathrm{C}$
$\Gamma_{\text {rad }}=3.9(4) \times 10^{-3} \mathrm{eV}$ EM
branching: $0.042 \%$

## The radiative width from $p-\gamma-\gamma$



$$
\begin{aligned}
& r_{3 \alpha} \propto\left[\Gamma_{r a d}\right] \exp \left(-\left[Q_{3 \alpha}\right] / k T\right) \\
& \Gamma_{r a d}=\left[\frac{\Gamma_{r a d}}{\Gamma}\right] \times\left[\frac{\Gamma}{\Gamma_{\pi}(E 0)}\right] \times\left[\Gamma_{\pi}(E 0)\right]
\end{aligned}
$$

$\frac{\text { Events in } 12 \text { days: }}{\square \text { Total: } 6.0 \mathrm{E}+9}$
p-singles $(7.65): 2.72 \mathrm{E}+8$
$\mathrm{p}-\gamma-\gamma(7.65): 260(16)$
$\left[\frac{\Gamma_{\text {rad }}}{\Gamma}\right]=2.76(21) \mathrm{E}-4$

## The radiative width from pair conversion measurements



$$
\alpha_{\pi}(\mathrm{E} 2)=8.765 \mathrm{E}-4
$$

## The radiative width from pair conversion measurements



# Properties of Monopole transitions in atomic nuclei 



# Properties of Monopole transitions in atomic nuclei 

Motivation:
> New data since the last evaluations First pass of ENSDF: $174 \mathrm{j}-\mathrm{j}$ ( $\mathrm{j}>0$ ) transitions
Eg=[41.9:1877] keV
$\mathrm{J}_{\mathrm{i}}=1[8], 2[109], 3[4], 4[32], 5[3], 6[6]$, 8[2], 9[1], 10[1], 16[1]


Heyde \& Wood., Rev. Mod. Phys. 83 (2011) 1467
EO strength model independent probe of co-existing structures -spin dependence

With J.L. Wood and A. Garnsworthy


Properties of Monopole transitions in atomic nuclei

With J.L. Wood and A. Garnsworthy
Approach
$>$ Consistent treatment of the data known
$\square$ Combine data on $\mathrm{O}^{+} \rightarrow \mathrm{O}^{+}$and $\mathrm{J}^{+} \rightarrow \mathrm{J}^{+}$ ( $J>0$ ) transitions
Collect and adopt T1/2, EM branching ratios, multipolarities, mixing ratios, experimental conversion coefficients from original references
$\square$ Accept data if: T1/2, ICC and mixing ratio are known
$>$ New conversion coefficients, $\Omega(E 0)$ electronic factors

238U $966.1 \mathrm{keV} 2^{+}$
$T_{1 / 2}=2.4(+17-7) \mathrm{ps}$
$921.2 \mathrm{keV} \mathrm{EO}+\mathrm{M} 1+\mathrm{E} 2$
$\alpha_{k}=0.191(30)$
$\alpha_{k}(M 1)=0.0390$
$\alpha_{k}(E 2)=0.00966$


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ANU Major Equipment Grans 2011 ARC Discovery (2014-2016) DP140102986

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