

Many-body correlations: the relative nature of their definition and the non-observable character of their value



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Outline

⊙ **Part I: Relative nature of many-body correlations**

- Introduction: quasiparticles, correlations & many-body methods
- Examples in nuclei and nuclear matter
- Nuclear Hamiltonians & similarity renormalisation group techniques
- Correlations and resolution scale

⊙ **Part II: Non-observable character of the nuclear shell structure**

- Single-nucleon shells \leftrightarrow correlated nucleon dynamics
- Definition & properties of effective single-particle energies
- Scale dependence & non-observability of effective single-particle energies
- Fermi gaps & spectroscopic factors

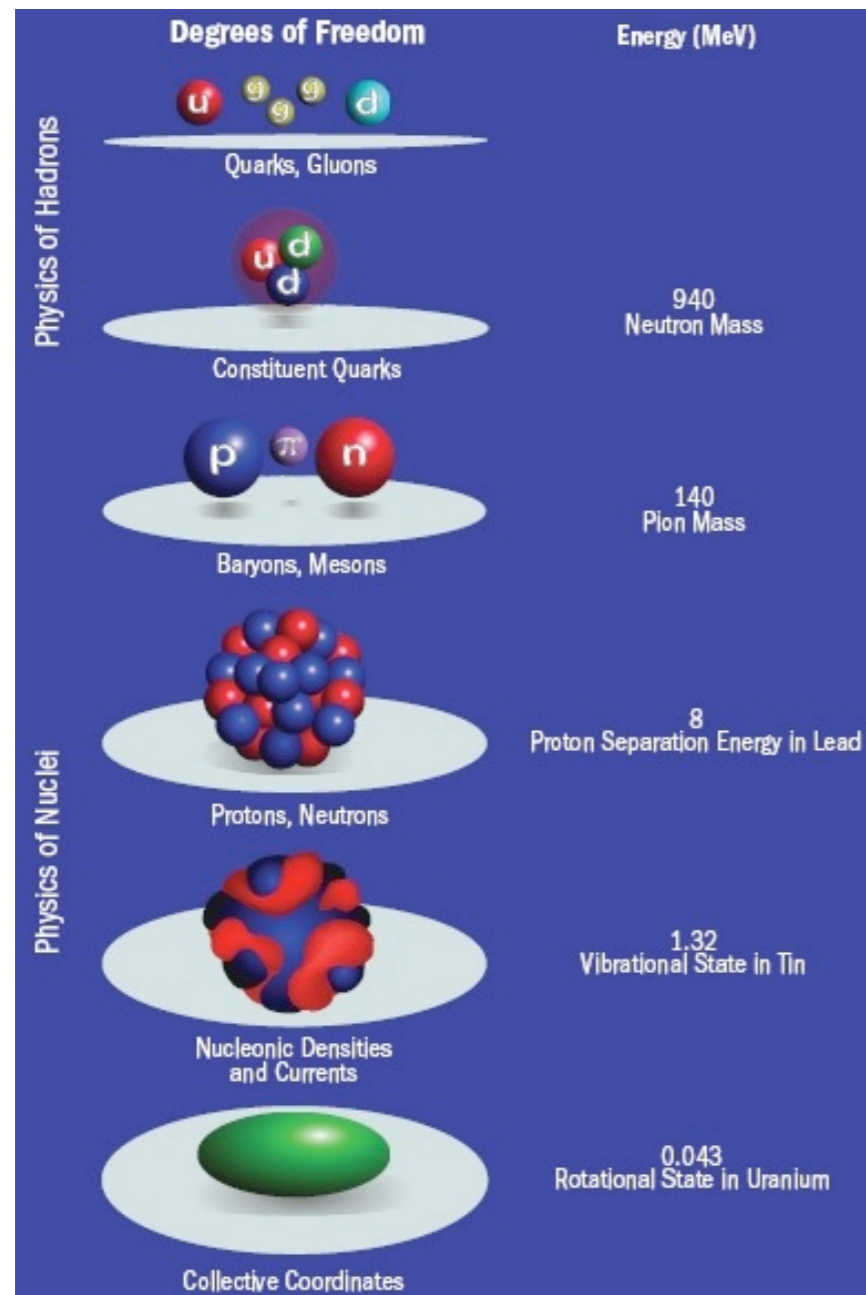
⊙ **Conclusions**

Part I

Relative nature of many-body correlations

Physical systems as a many-body problem

More reductionist/elementary / "fundamental" description



Emergent phenomena amenable to effective descriptions

- ⊙ Quantum/mesoscopic system as many-body problem
- ⊙ Choice of degrees of freedom
 - ↓
 - Physical system in terms of correlations between d.o.f.
- ⊙ Many-body Schrödinger equation
 - Exact solution for $A=2, 3, 4$
 - Approximated solution for $A \gtrsim 5$
 - ↓
 - Accuracy / difficulty depend on correlations
- ⊙ At a given A , how to minimise correlations?
- ⊙ When increasing A , how to monitor the accuracy?

Quasiparticles

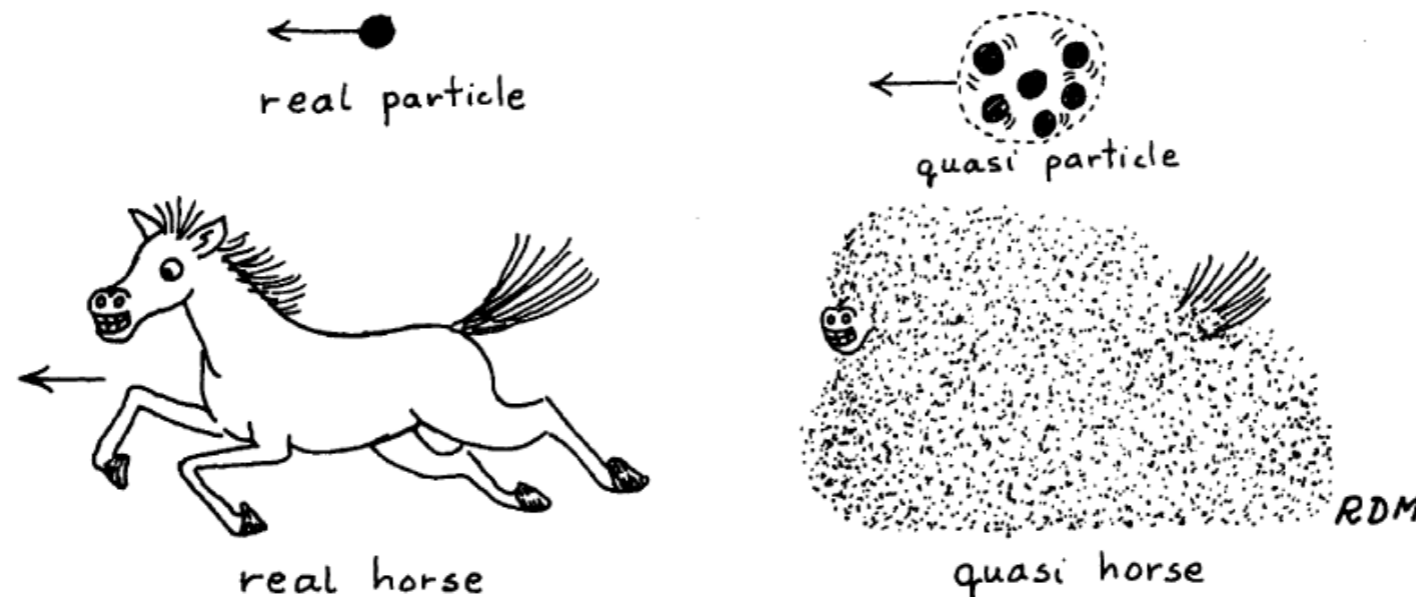
- ⊙ Difficulty as the number of particles increases → how to picture/model many-body correlations?
- ⊙ Easy to deal with independent particles → reformulate in terms of $A \times$ one-body problems
- ⊙ Can we change the (nature of the) chosen degrees of freedom & eliminate many-body correlations?
- ⊙ Concept of (Landau) quasiparticles



Entities with modified (in-medium, renormalised, ...) properties w.r.t. the bare d.o.f.

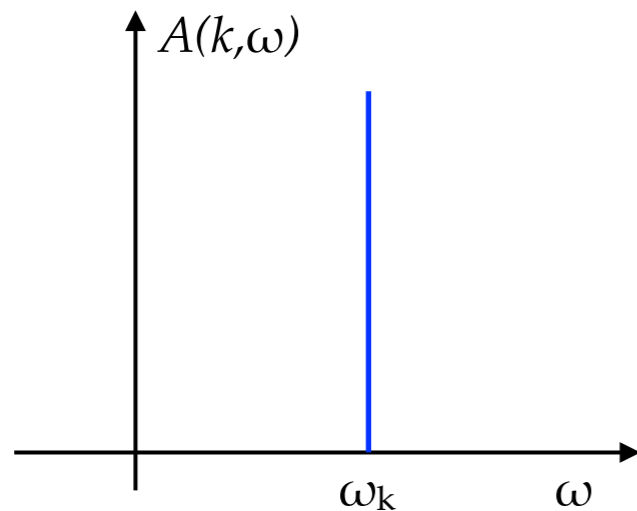


Many-body problem of interacting particles → one-body problem of (independent) quasiparticles

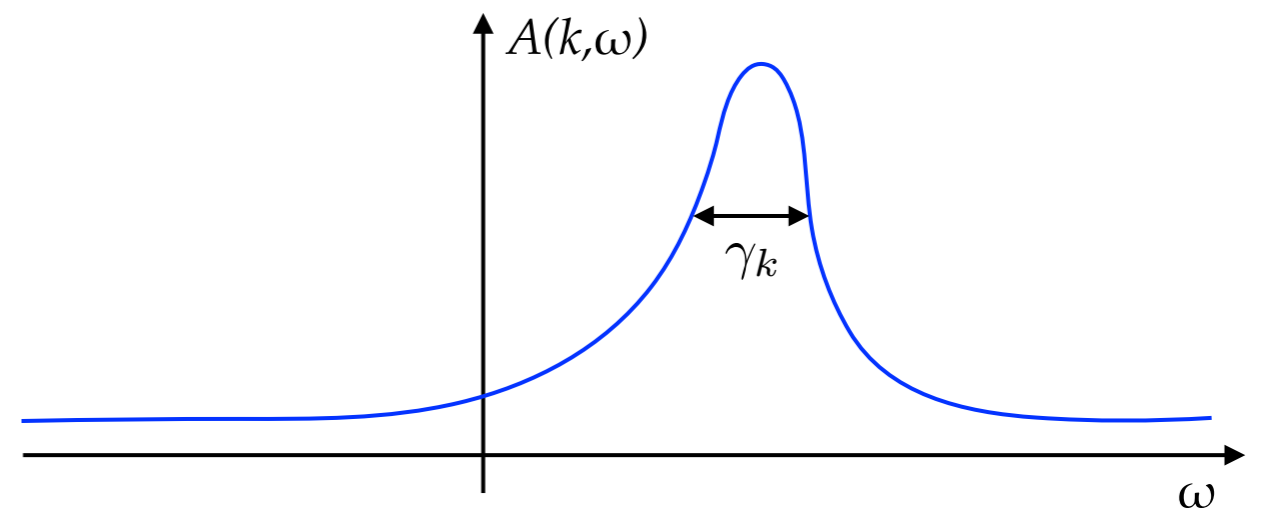


Interacting quasiparticles

- ⊙ In some cases, quasiparticles can be constructed explicitly
- ⊙ In most cases, quasiparticles-like excitations emerge from the many-body dynamics
 - Spectral function $A(k,\omega)$ embodies quasiparticle features & many-body correlations
 - For free particles $A(k,\omega) = \delta(\omega - k^2/2m)$



Infinitely-lived (=independent) quasiparticle



Decaying (=interacting) quasiparticle

- ⊙ Quasiparticles with finite lifetime \rightarrow departure from independent (quasi)particle picture
- ⊙ Many-body correlations as residual interactions between quasiparticles

Particle-hole expansions

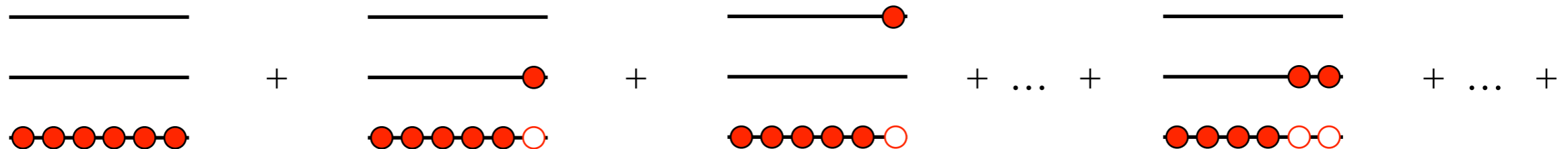
- ⊙ **Independent-particles as 0th-order tenet of numerous many-body methods**

- Perturbation theory
- Density functional theory
- Nuclear shell model

- ⊙ **Hartree-Fock method as an optimised independent-particle description**

- (Many-body) correlations: everything beyond Hartree-Fock

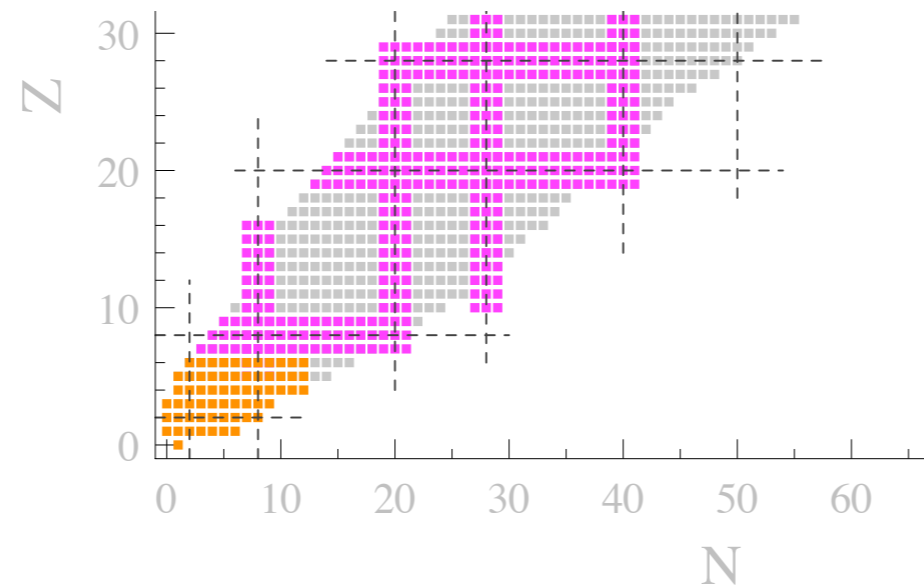
- ⊙ **Beyond-Hartree-Fock methods as expansions in particle-hole excitations**



- Simplest: MBPT
- Exact (= whole expansion): Configuration interaction / No-core shell model
- Freedom to choose the interaction such that HF is the closest to the exact solution?

Different schemes for different correlations

⊙ Methods based on particle-hole expansions face severe scaling



⊙ Why don't include some correlation in the interaction itself? → effective interactions

- One aims at limiting the complications of ph expansions
- Interaction traditionally phenomenological, possible to derive one ab initio?

⊙ Why don't limit ourselves to part of the Hilbert space → valence space methods

- One aims at the exact solution in the limited Hilbert space
- Interaction traditionally phenomenological, recently also ab initio



Correlations in different schemes will be different by construction

Correlations via symmetry breaking & restoration

- ◎ **Correlations can be grasped by exploiting (breaking & restoration of) symmetries**

- For near-degenerate systems essential to expand around a symmetry-breaking reference
- In nuclear physics: $U(1) \leftrightarrow$ pairing correlations; $SU(2) \leftrightarrow$ quadrupole correlations

- ◎ **Can the two types be related?**

- Correlations included via symmetry breaking might be very hard to get via ph expansion
- And viceversa

- ◎ **Can the two types be combined?**

- Gorkov Green's functions [Somà, Duguet, Barbieri 2011]
- Multi-reference IM-SRG [Hergert *et al.* 2013]
- Symmetry broken & restored MBPT and CC [Duguet 2015, Duguet, Signoracci 2016]
- Many-body driven EDF [Duguet *et al.* 2015]
- Symmetry breaking & restoration + truncated CI [Ripoche *et al.* 2017]



see Thomas', Denis', ... talks

Nuclear Hamiltonians

Early Hamiltonians (60's & 70's)

- Soft core
- Could not reproduce nuclear saturation

Phenomenological Hamiltonians (80's & 90's)

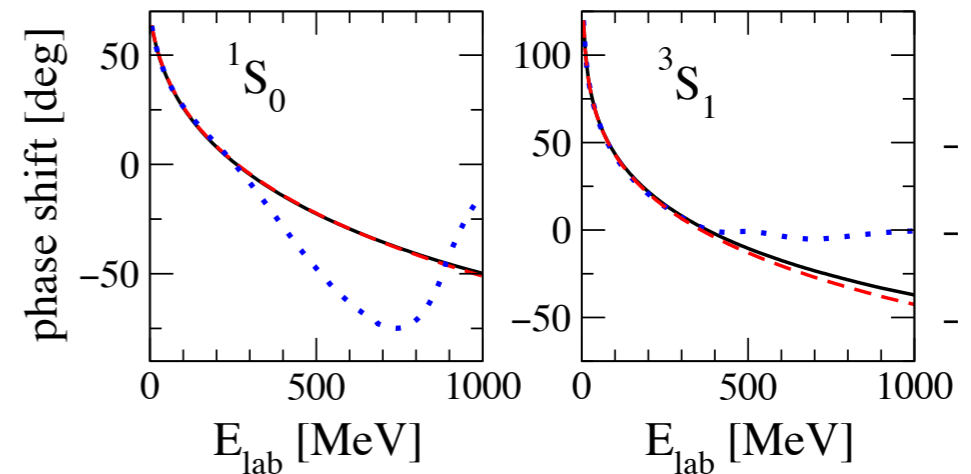
- Hard core
- Three-body forces?

Chiral EFT interactions (from 00's)

- Softer core
- Three-body forces consistent

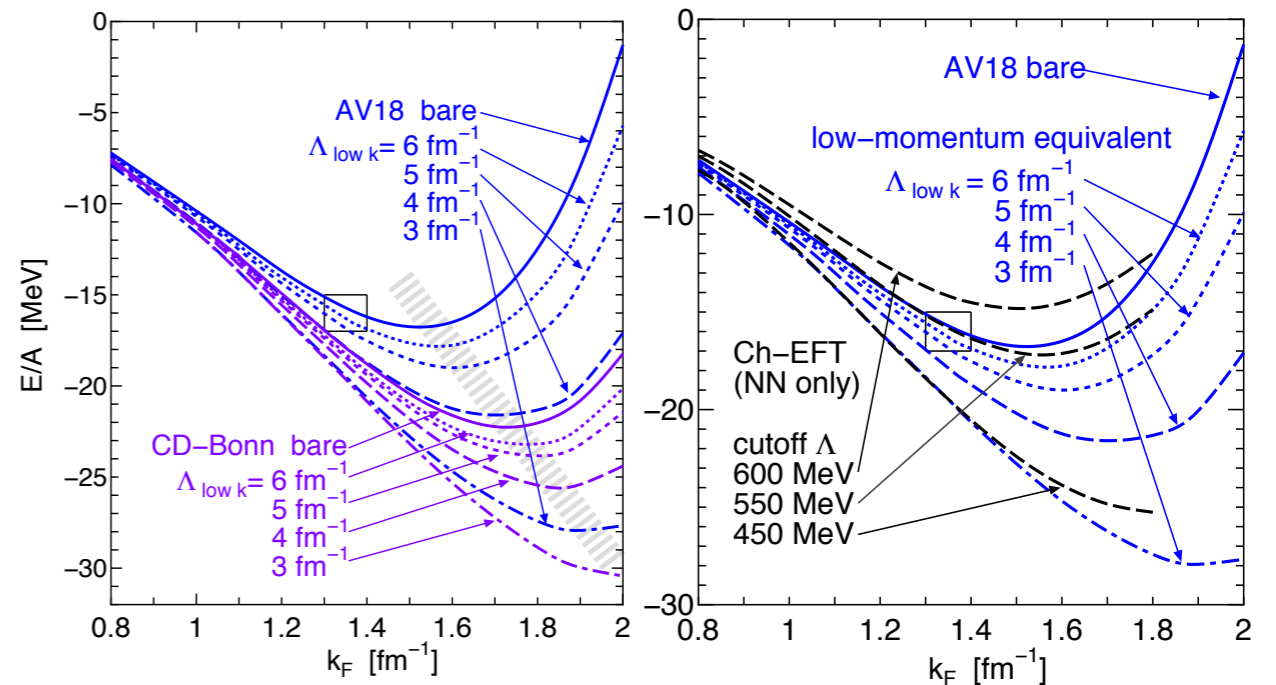
SRG techniques

- Unitary transformation of the Hamiltonian
- Trade hard core for higher-body forces
- Universality at low energy scales



[Bogner et al. 2010]

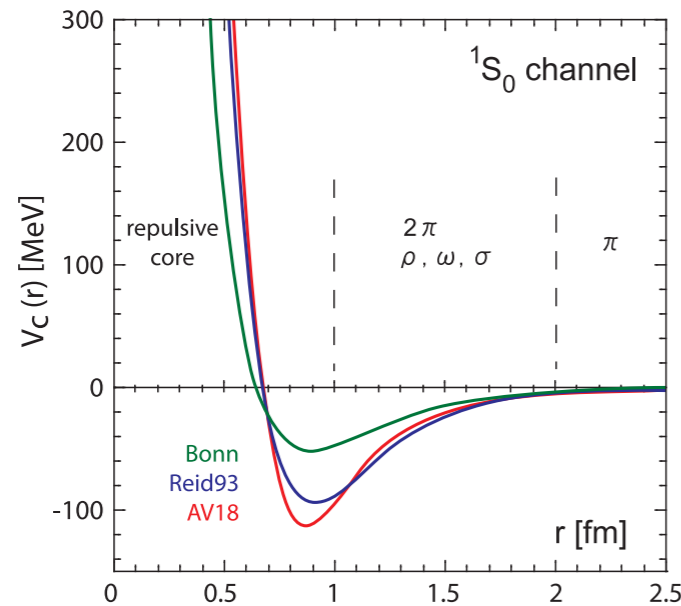
Coester band



[Kohno 2015]

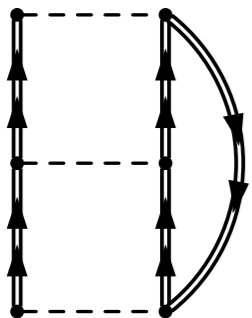
Long- vs short-range correlations

◎ Hard core induces strong short-range correlations

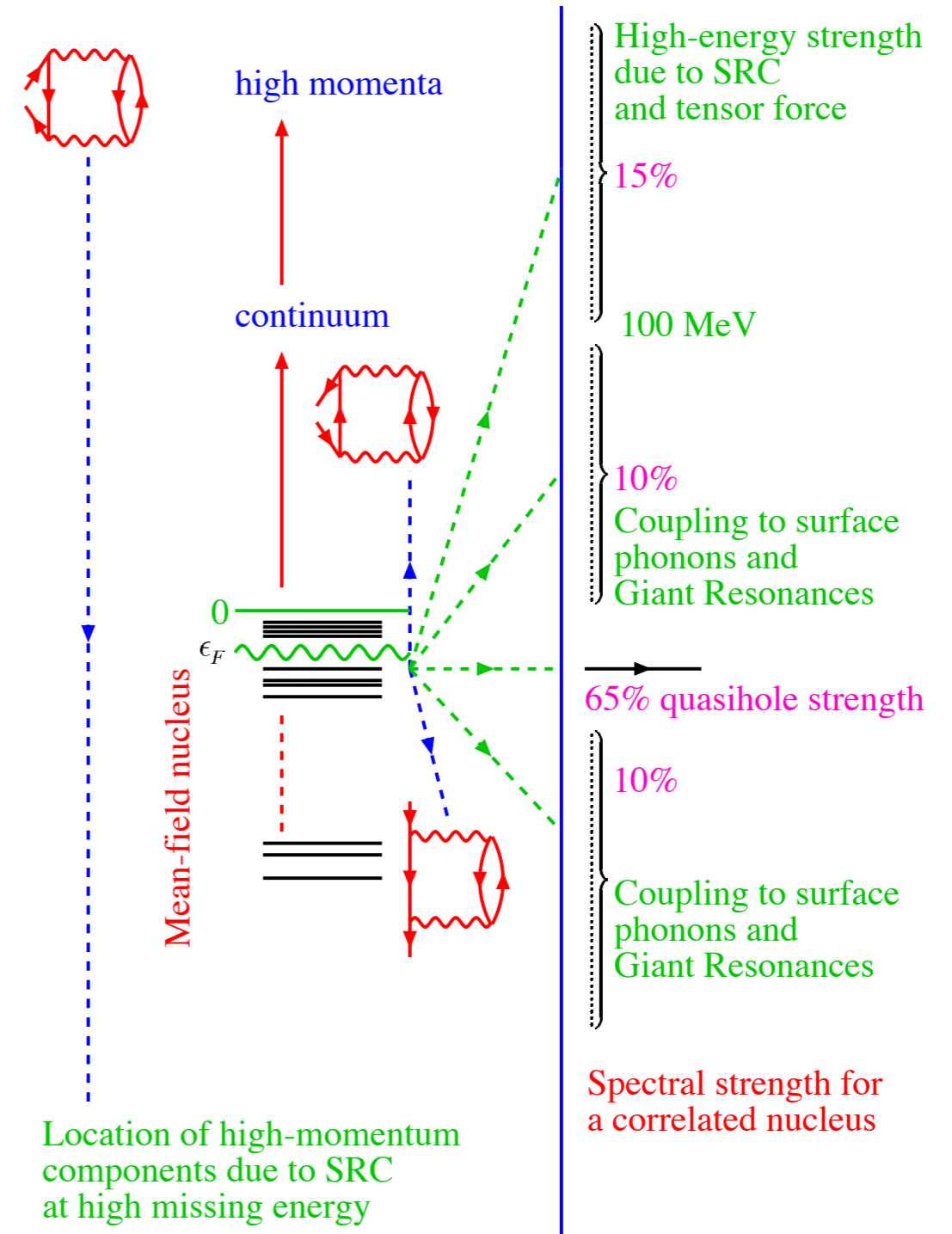
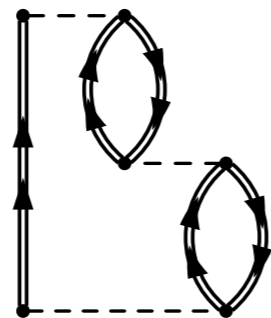


- Sophisticated many-body methods needed
- Strong correlations fragmentation of s.p. strength
- pp/ph excitation \leftrightarrow short-/long-range physics

Short-range

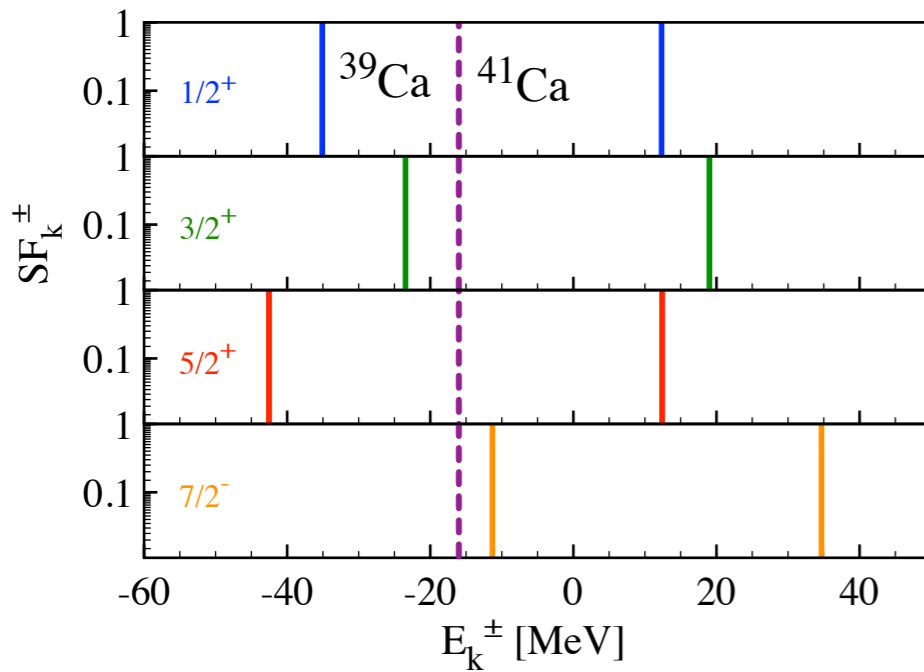


Long-range

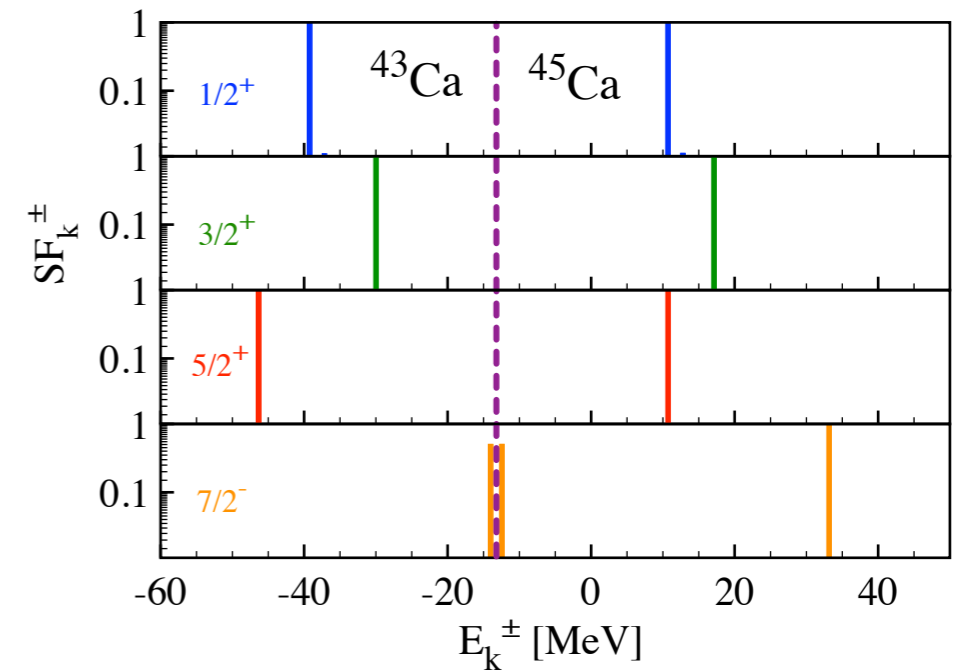


Fragmentation of single-particle strength in nuclei

Dyson 1st order (HF)



Gorkov 1st order (HFB)

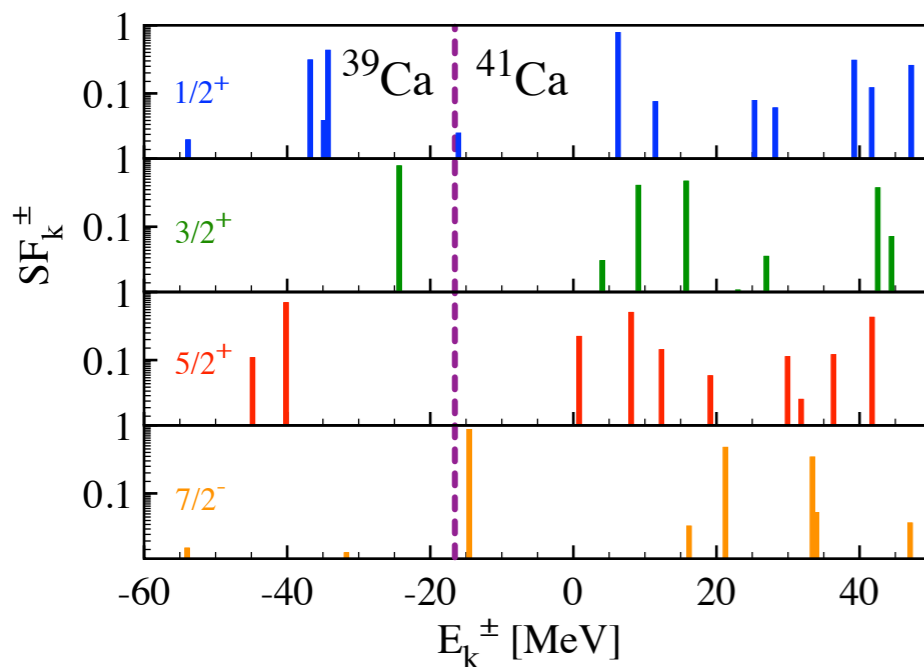


Fragmentation

Static pairing



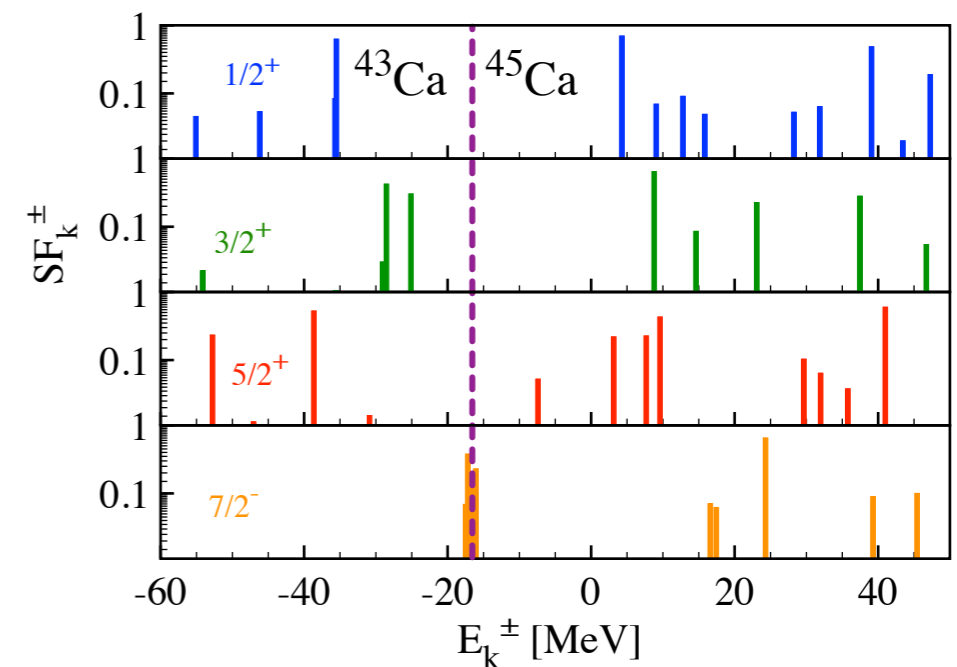
Dyson 2nd order



Dynamical
fluctuations



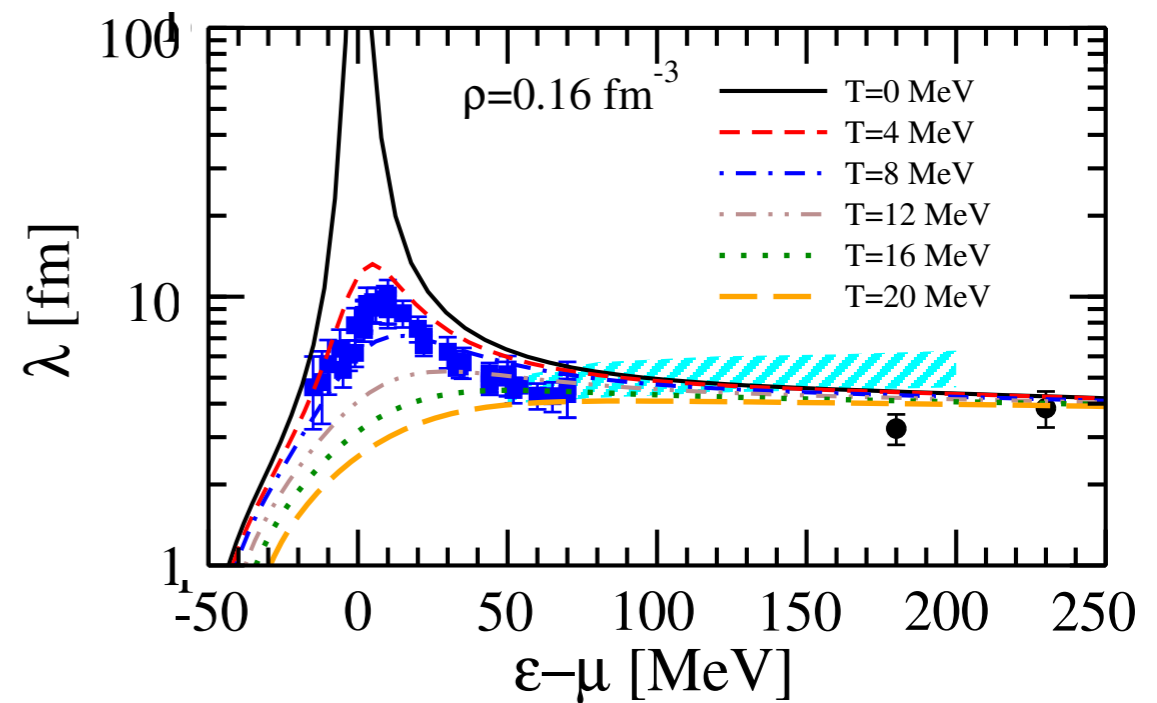
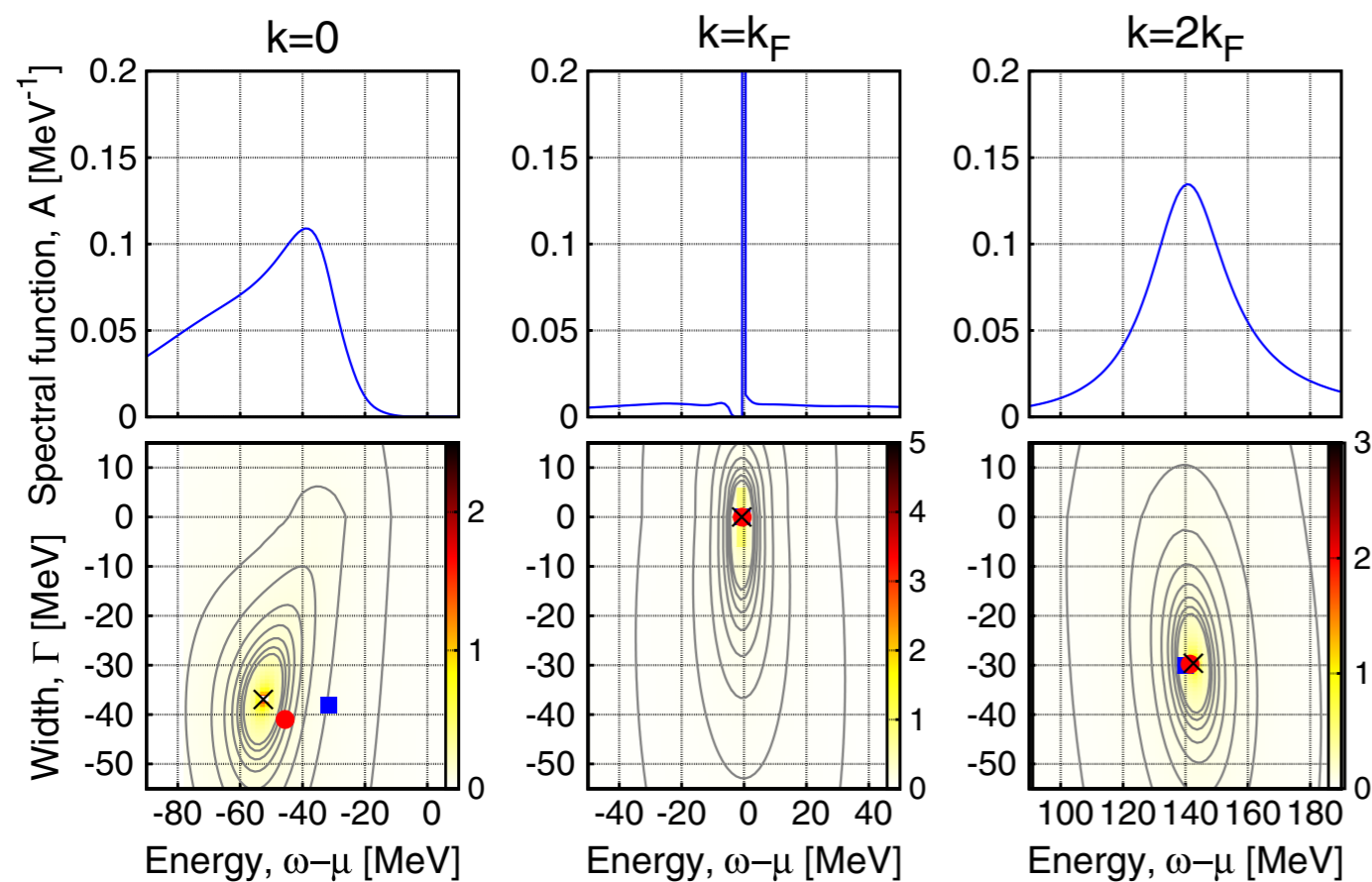
Gorkov 2nd order



Fragmentation of single-particle strength in infinite matter

◉ Spectral function depicts correlations

- ◉ Broad peak signals depart from
- ◉ Well-defined (long-lived) quasiparticles at the Fermi surface
- ◉ Long mean free path for $E < E_F$



Renormalisation-group techniques for nuclear forces

- ◎ SRC generated by couplings between low and high momenta
 - Large model spaces needed to converge → applicability limited to light nuclei
- ◎ Are high momenta, i.e. high resolution, necessary to compute **low-energy observables**?



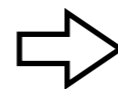
- Interested in long-wavelength information



- Small-distance details irrelevant

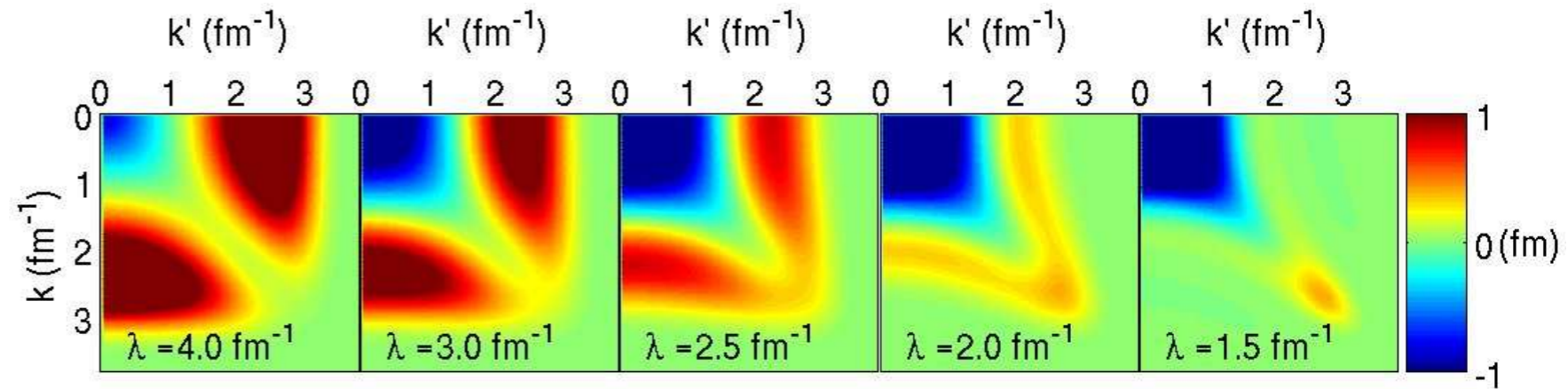
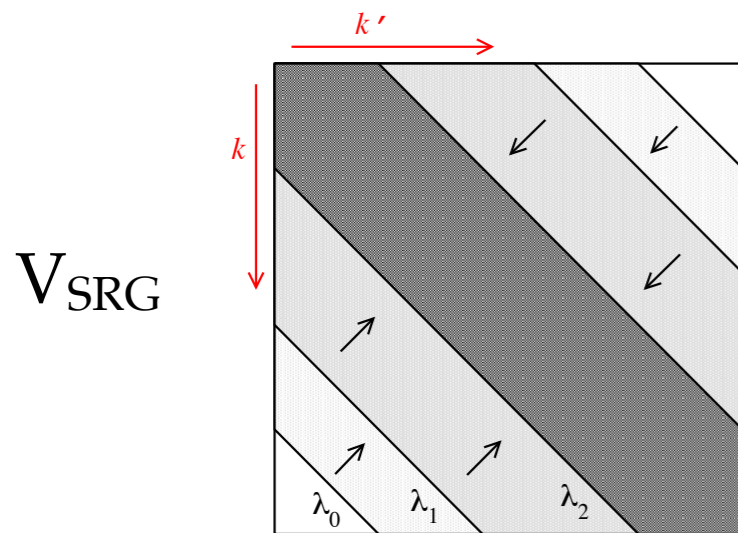
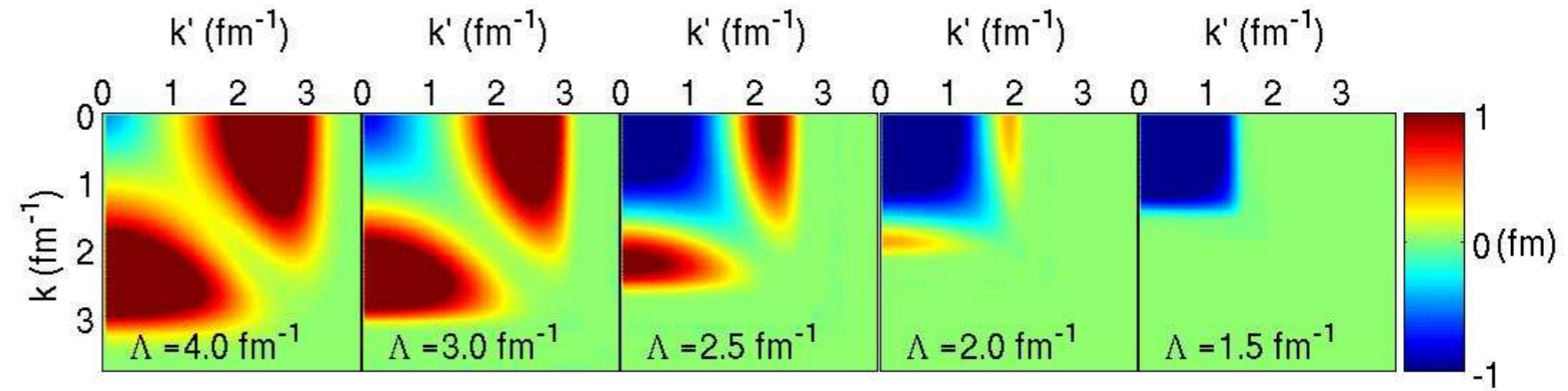
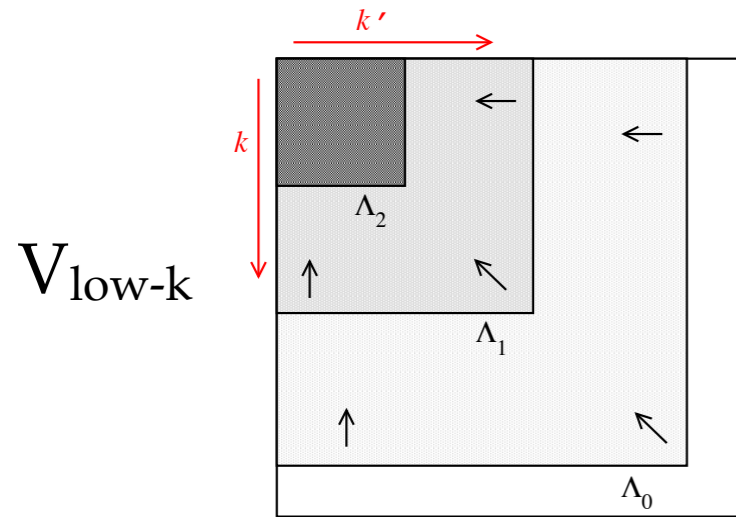


- **Change the resolution** → “integrate out” unnecessary information



Low-momentum evolutions

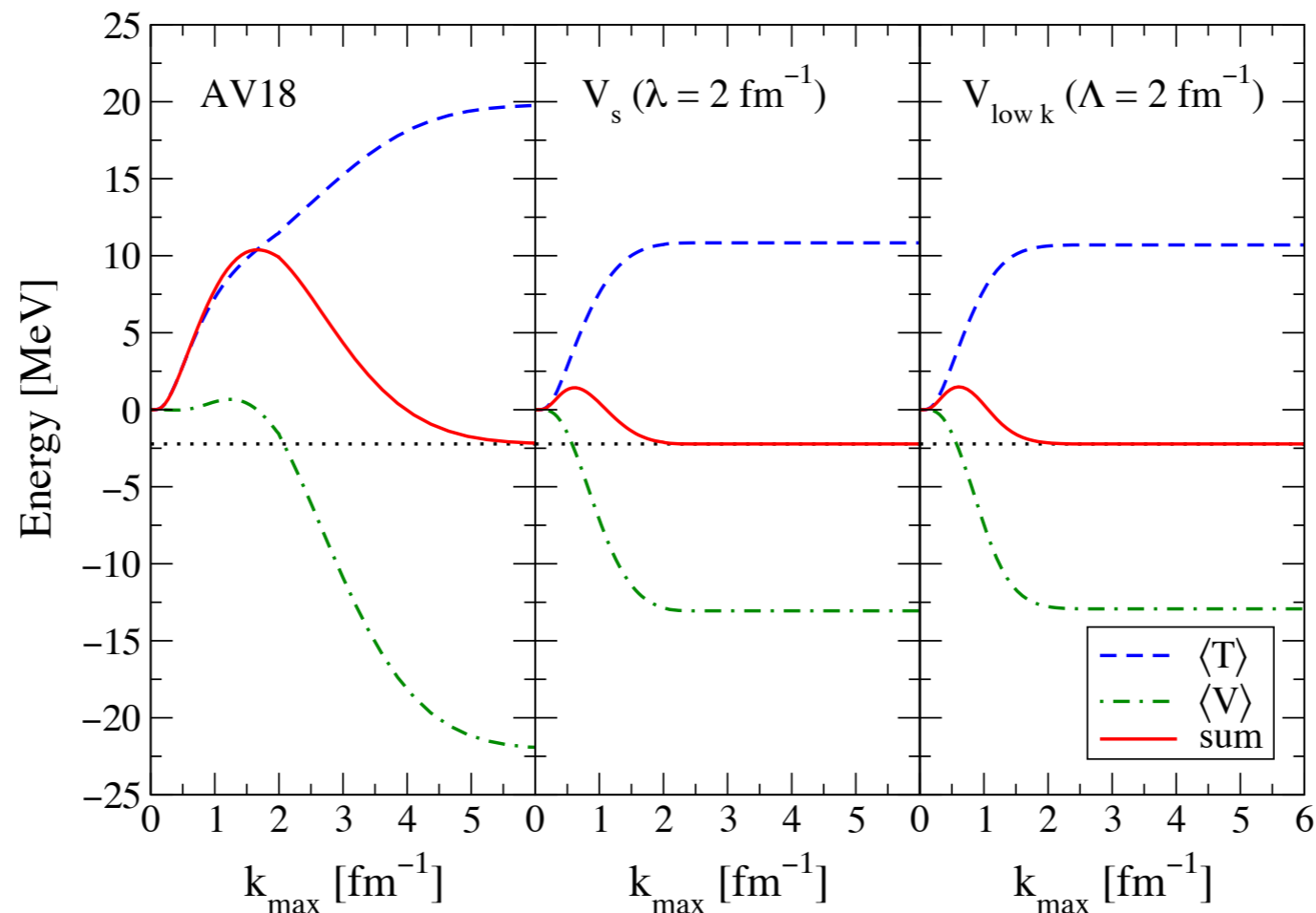
- ⊙ (Unitary) transformation to **change the resolution scale** of the Hamiltonian
- ⊙ Two main types of transformation



Example: deuteron binding energy

- ⊙ Performing RG changes weights of different parts of the Hamiltonian
- ⊙ Observable binding energy remains unchanged
- ⊙ High momenta not needed for softened interactions
- ⊙ Simply cutting off high momenta doesn't work

$$E_d(k < k_{\max}) = \int_0^{k_{\max}} d\mathbf{k} \int_0^{k_{\max}} d\mathbf{k}' \psi_d^\dagger(\mathbf{k}; \lambda) (k^2 \delta^3(\mathbf{k} - \mathbf{k}') + V_s(\mathbf{k}, \mathbf{k}')) \psi_d(\mathbf{k}'; \lambda)$$

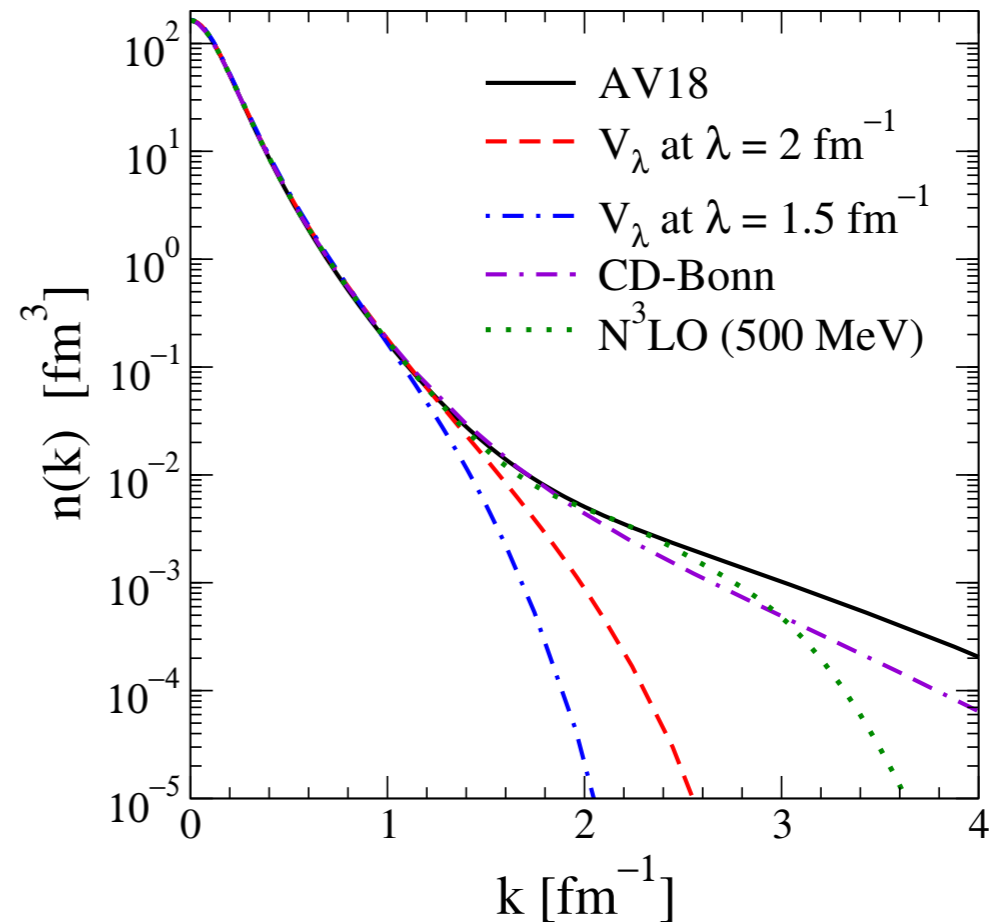
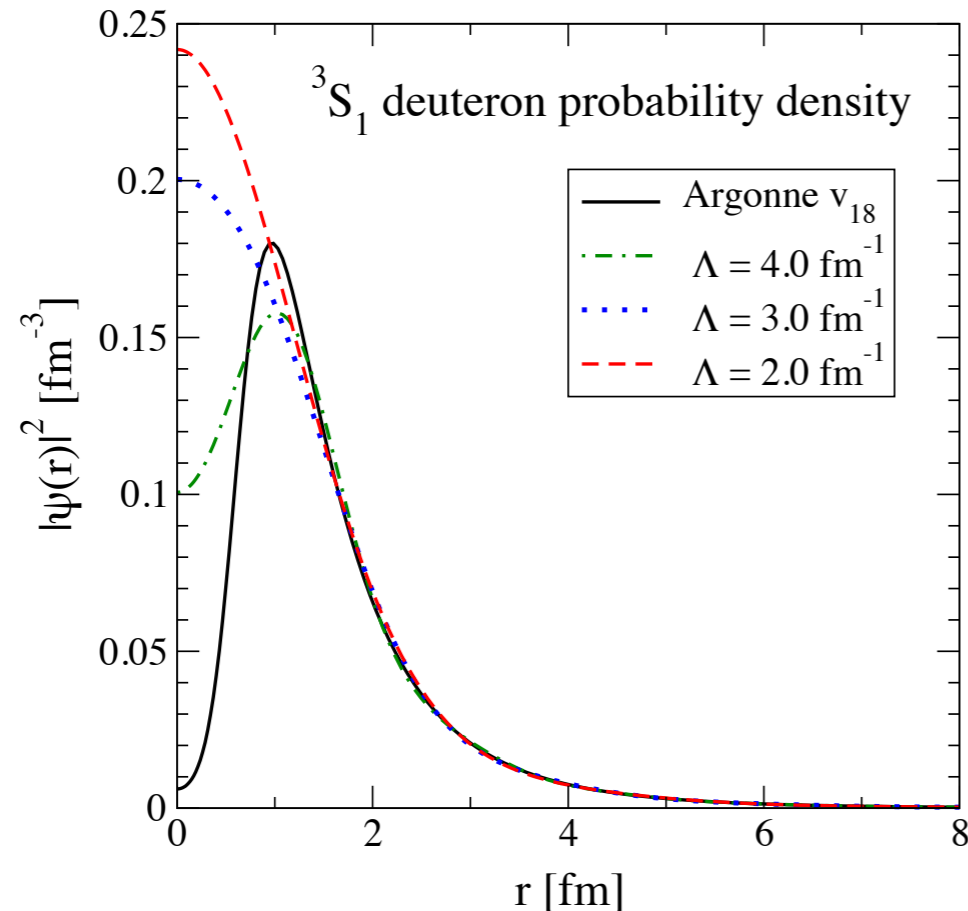


[Bogner et al. 2010]

Short-range correlations & momentum distribution

Short-range correlations change drastically with resolution scale

[Bogner *et al.* 2010]



[More *et al.* 2015]

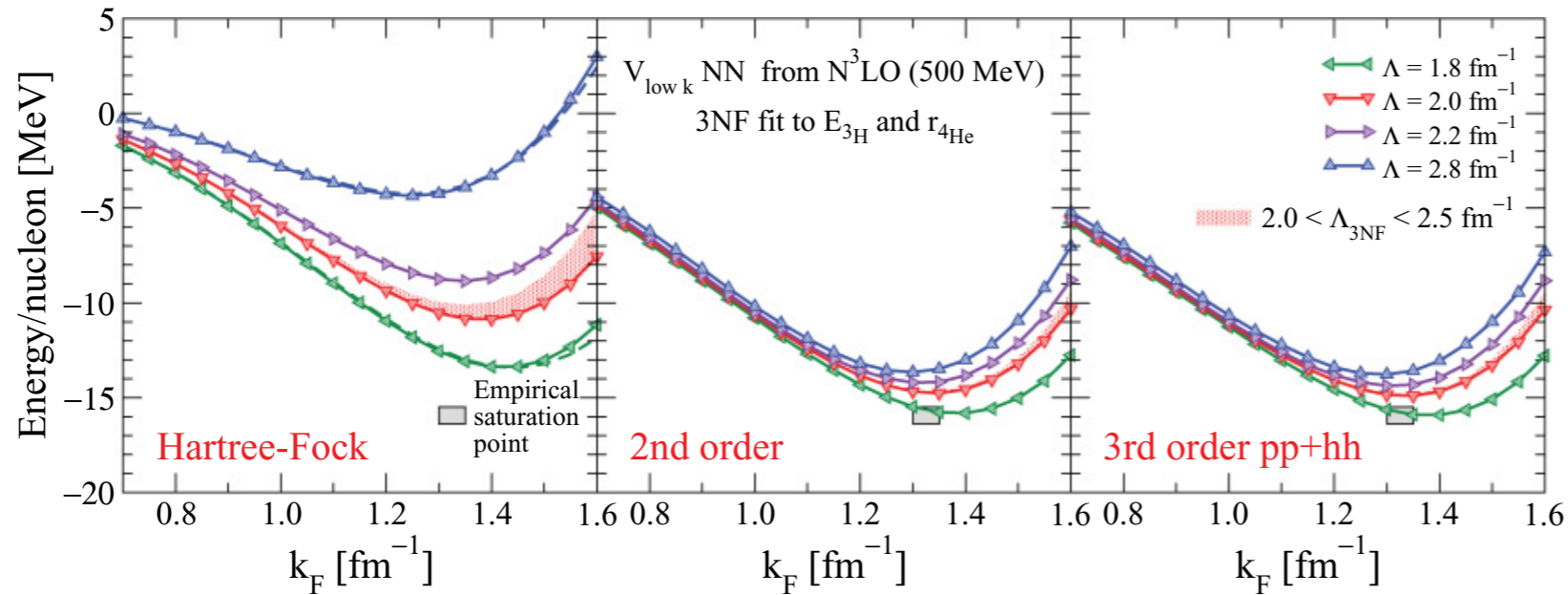
How to explain the momentum distribution “extracted” from experiment?

- Separation between structure and reaction is scale-dependent
- Operators & currents have to evolve consistently with the Hamiltonian
- E.g. what is a one-body current at one scale, gets shifted in two-body currents at another

Benefits in many-body systems

Improved convergence of many-body calculations

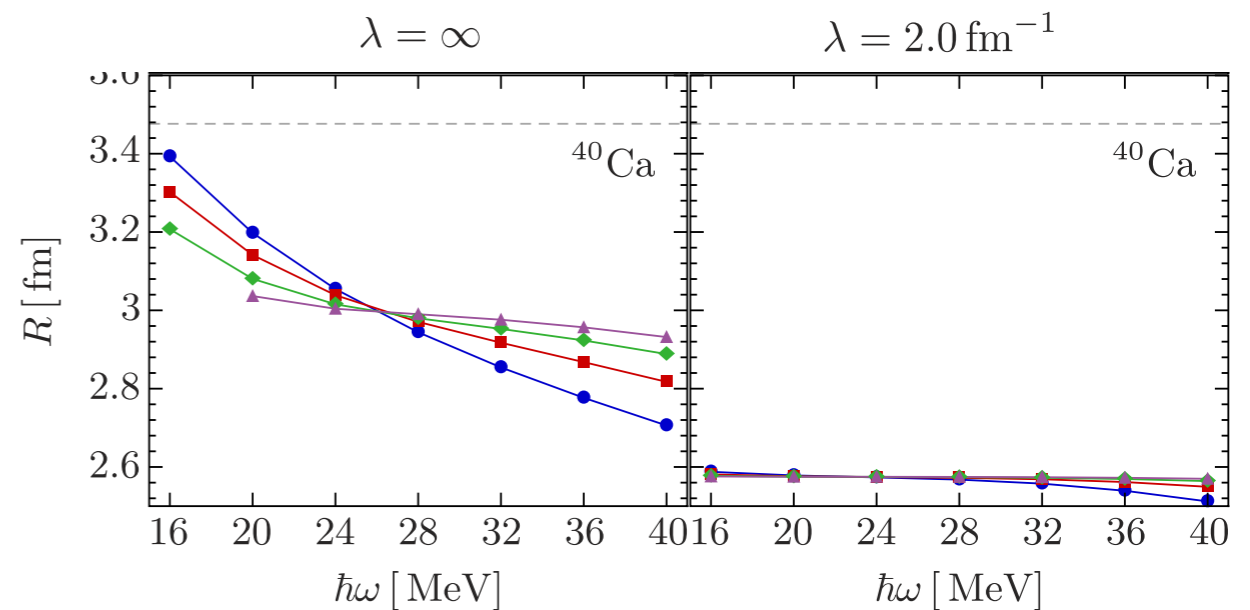
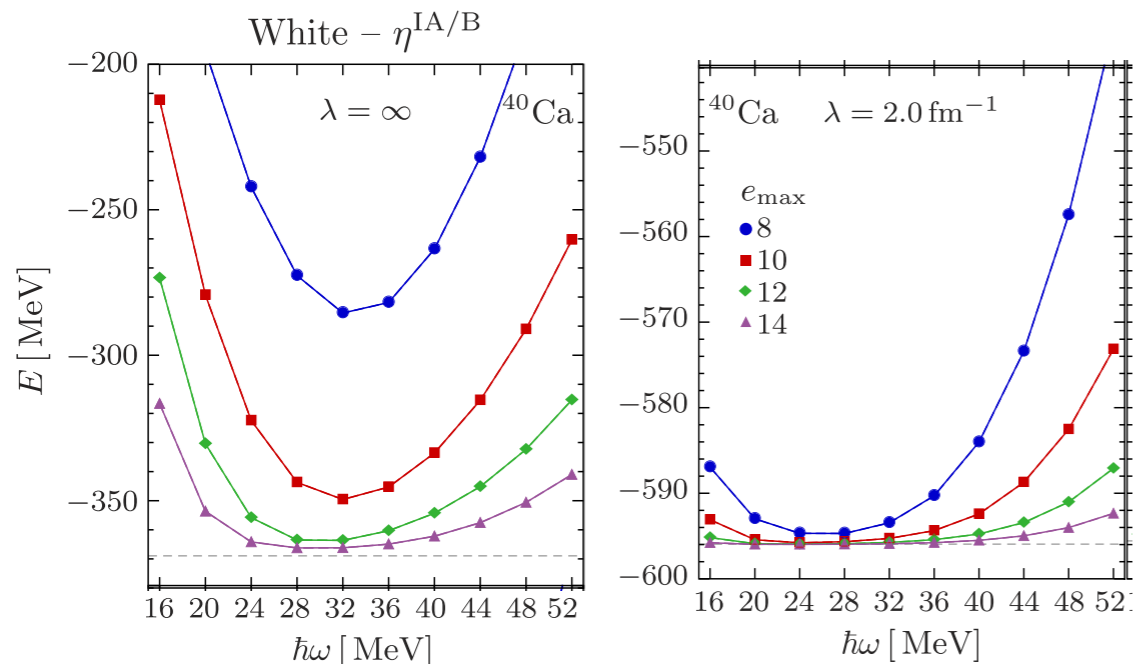
- Smaller model spaces & less refined many-body truncations needed



[Hebelner et al. 2011]

[Bogner et al. 2010]

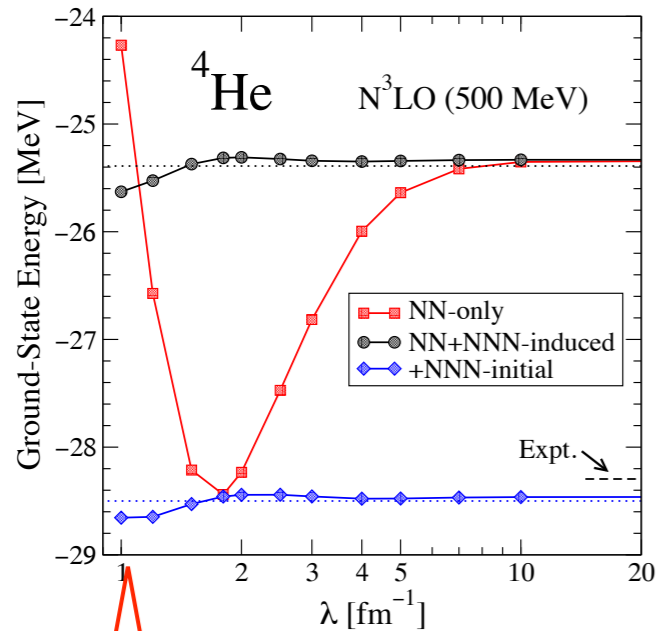
Drawback: additional many-body forces generated through unitary transformation



[Hergert et al. 2016]

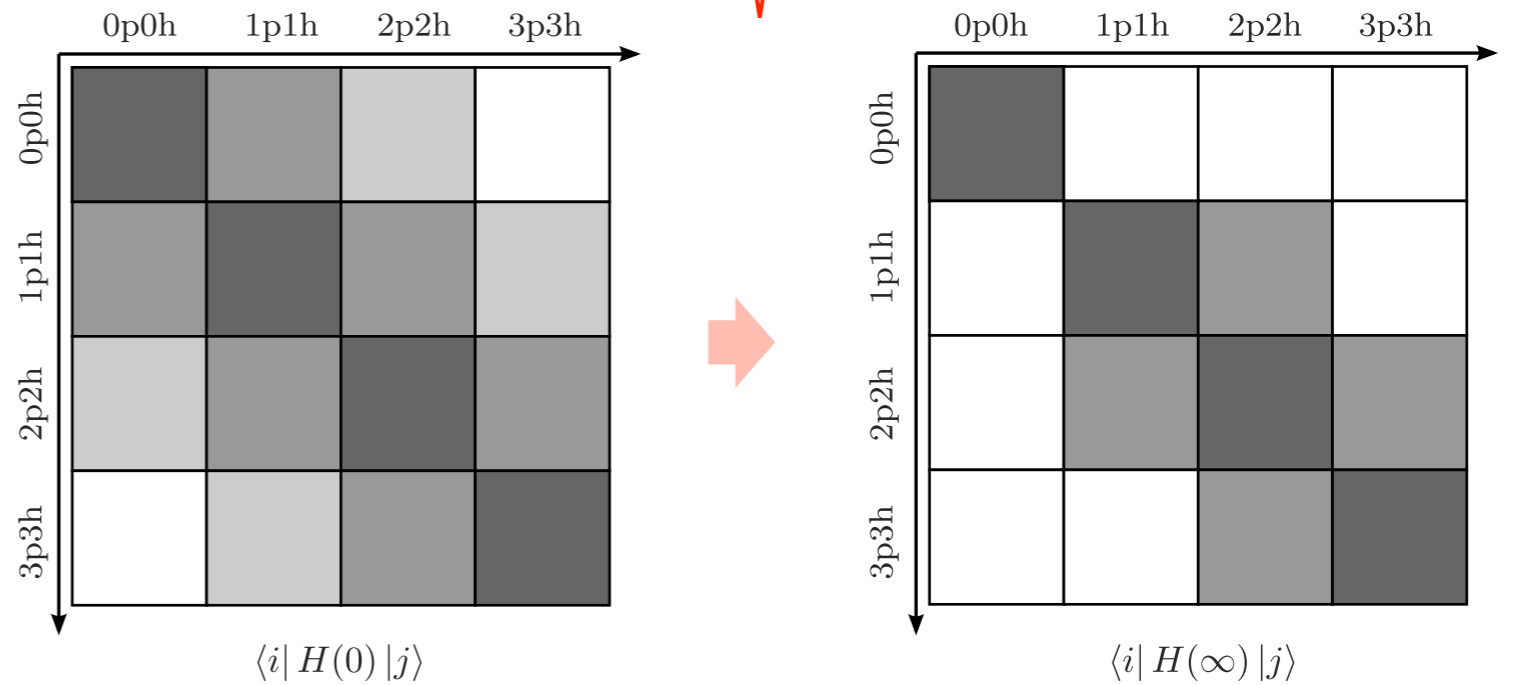
From free space to in medium

- Why don't evolve to the point where correlations have disappeared?



At very low scales, many-body ($A > 3$) forces explode

However, if done step by step keeping normal-ordered parts at each step...



In-medium Similarity Renormalisation Group

- Unpractical to evolve in medium every operator we are interested in
- Combine with another many-body method (e.g. NCSM) to access wide range of observables

Part II

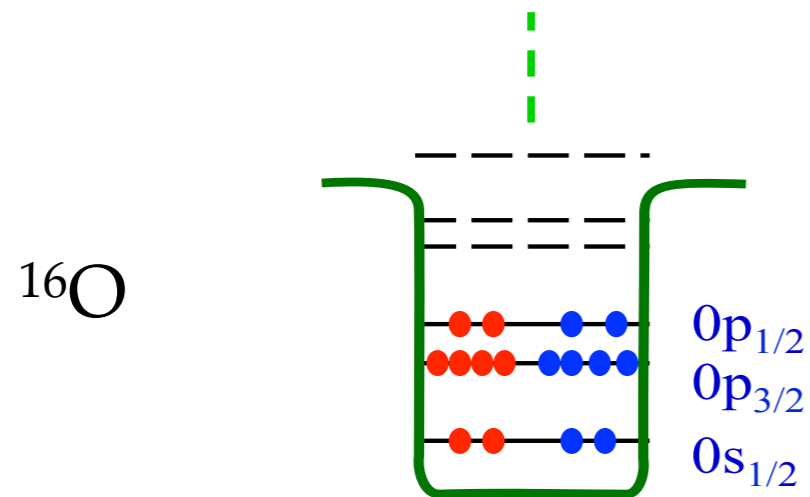
Non-observable character of the nuclear shell structure

T. Duguet, H. Hergert, J.D. Holt, V. Somà, Phys. Rev C 92 034313 (2015)

Single-nucleon shell structure

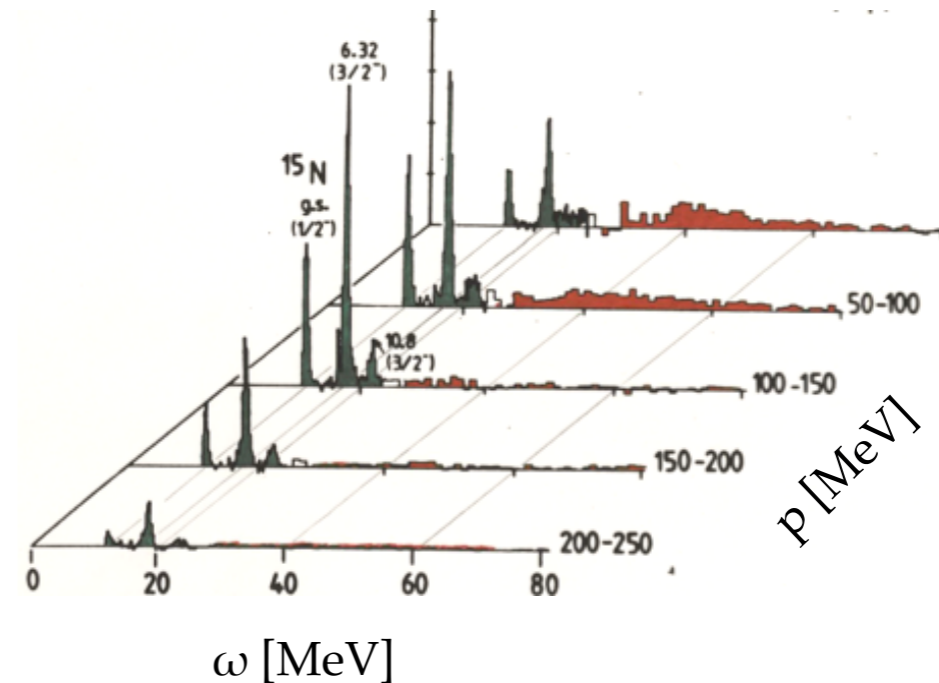
⊙ **Correlated many-body system** ↔ **description in terms of independent particles**

- Can a one-to-one correspondence be established?



?

↔



[Mougey *et al.* 1980]

⊙ **Concept of single-nucleon shells**

- Basic pillar of the shell model
- Provides interpretation of nuclear (low-energy) observables
- Leads to considering a single-particle spectrum (magicity, shell evolution, ...)



Useful interpretation, but which degree of reality?

Single-nucleon shell structure

⊙ Quantum mechanical nuclear many-body problem

- Many-body Schrödinger equation → one-nucleon addition/removal energies

$$H|\Psi_k^A\rangle = E_k^A|\Psi_k^A\rangle \qquad E_k^\pm \equiv \pm(E_k^{A\pm 1} - E_0^A)$$

To what extent the single-particle energy spectrum relates to low-energy observables?

$$\underbrace{E_k^\pm}_{\text{Outcome of Schr. equation}} = \underbrace{e_p}_{\text{Ind. particles}} + \underbrace{\Delta E_{p \rightarrow k}}_{\text{Correlations}}$$

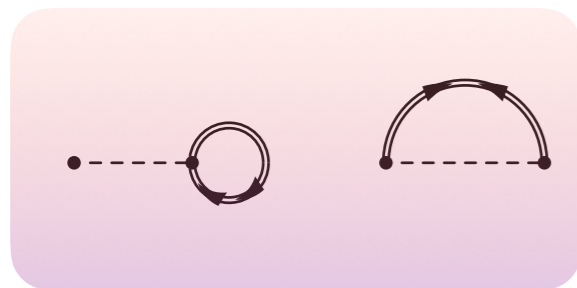
⊙ In the following:

- Reminder of Green's function theory
- Is there a proper/unique definition of single-particle energy? → Baranger ESPEs
- Scale dependence of the above partitioning, i.e. of ESPEs
- Illustration of the scale dependence from ab initio calculations

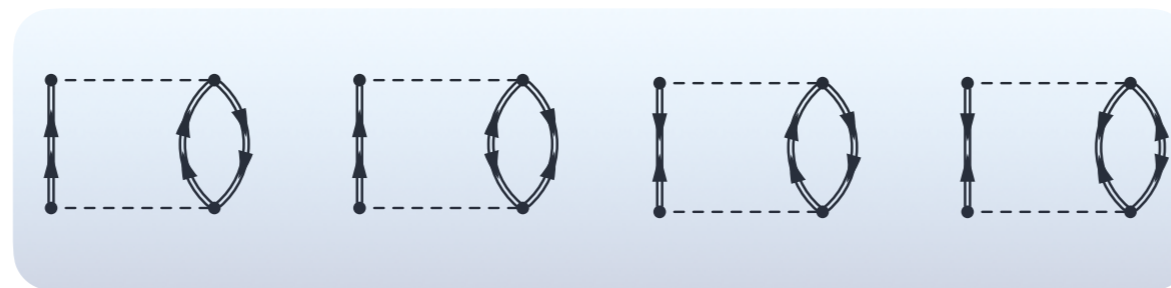
Self-consistent Green's function approach

⊙ **Solution of the A -body Schrödinger equation** $H|\Psi_k^A\rangle = E_k^A|\Psi_k^A\rangle$ **achieved by**

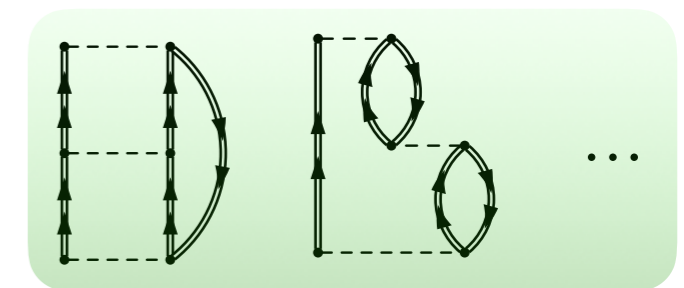
- 1) Rewriting it in terms of 1-, 2-, ... A -body objects $G_1=G, G_2, \dots, G_A$ (**Green's functions**)
 - 2) Expanding these objects in perturbation (in practise only $G \rightsquigarrow$ **one-body observables**)
- \rightsquigarrow **Self-consistent** schemes resum (infinite) subsets of perturbation-theory contributions



ADC(1)=HFB



ADC(2)



ADC(3)

⊙ **Here we employ the Algebraic Diagrammatic Construction (ADC) method**

- Systematic, improvable scheme for the one-body Green's functions, truncated at order n
- ADC(1) = Hartree-Fock(-Bogolyubov); ADC(∞) = exact solution
- At present **ADC(1)**, **ADC(2)** and **ADC(3)** are implemented and used

⊙ **Extension to open-shell nuclei: (symmetry-breaking) Gorkov scheme**

○ Developed at Saclay & Surrey 2010-today

[Somà, Duguet & Barbieri 2011]

Spectral representation

◉ Numerator contains spectroscopic information

$$G_{ab}(z) = \sum_{\mu} \frac{\langle \Psi_0^A | a_a | \Psi_{\mu}^{A+1} \rangle \langle \Psi_{\mu}^{A+1} | a_b^{\dagger} | \Psi_0^A \rangle}{z - E_{\mu}^+ + i\eta} + \sum_{\nu} \frac{\langle \Psi_0^A | a_b^{\dagger} | \Psi_{\nu}^{A-1} \rangle \langle \Psi_{\nu}^{A-1} | a_a | \Psi_0^A \rangle}{z - E_{\nu}^- - i\eta}$$

spectroscopic amplitudes

$$U_{\mu}^b \equiv \langle \Psi_0^A | a_b | \Psi_{\mu}^{A+1} \rangle$$

$$V_{\nu}^b \equiv \langle \Psi_0^A | a_b^{\dagger} | \Psi_{\nu}^{A-1} \rangle$$

spectroscopic probabilities matrices

$$S_{\mu}^{+ab} \equiv \langle \Psi_0^A | a_a | \Psi_{\mu}^{A+1} \rangle \langle \Psi_{\mu}^{A+1} | a_b^{\dagger} | \Psi_0^A \rangle$$

$$S_{\nu}^{-ab} \equiv \langle \Psi_0^A | a_a^{\dagger} | \Psi_{\nu}^{A-1} \rangle \langle \Psi_{\nu}^{A-1} | a_b | \Psi_0^A \rangle$$

spectral function

$$\mathbf{S}(z) \equiv \sum_{\mu \in \mathcal{H}_{A+1}} \mathbf{S}_{\mu}^+ \delta(z - E_{\mu}^+) + \sum_{\nu \in \mathcal{H}_{A-1}} \mathbf{S}_{\nu}^- \delta(z - E_{\nu}^-)$$

spectroscopic factors

$$SF_{\mu}^+ \equiv \text{Tr}_{\mathcal{H}_1} [\mathbf{S}_{\mu}^+] = \sum_{a \in \mathcal{H}_1} |U_{\mu}^a|^2$$

$$SF_{\nu}^- \equiv \text{Tr}_{\mathcal{H}_1} [\mathbf{S}_{\nu}^-] = \sum_{a \in \mathcal{H}_1} |V_{\nu}^a|^2$$

spectral strength distribution

$$\mathcal{S}(z) \equiv \text{Tr}_{\mathcal{H}_1} [\mathbf{S}(z)]$$

$$= \sum_{\mu \in \mathcal{H}_{A+1}} SF_{\mu}^+ \delta(z - E_{\mu}^+) + \sum_{\nu \in \mathcal{H}_{A-1}} SF_{\nu}^- \delta(z - E_{\nu}^-)$$

Spectral representation

◉ Combine numerator and denominator of Lehmann representation

$$G_{ab}(z) = \sum_{\mu} \frac{U_a^{\mu} (U_b^{\mu})^*}{z - E_{\mu}^+ + i\eta} + \sum_{\nu} \frac{(V_a^{\nu})^* V_b^{\nu}}{z - E_{\nu}^- - i\eta}$$

denominator

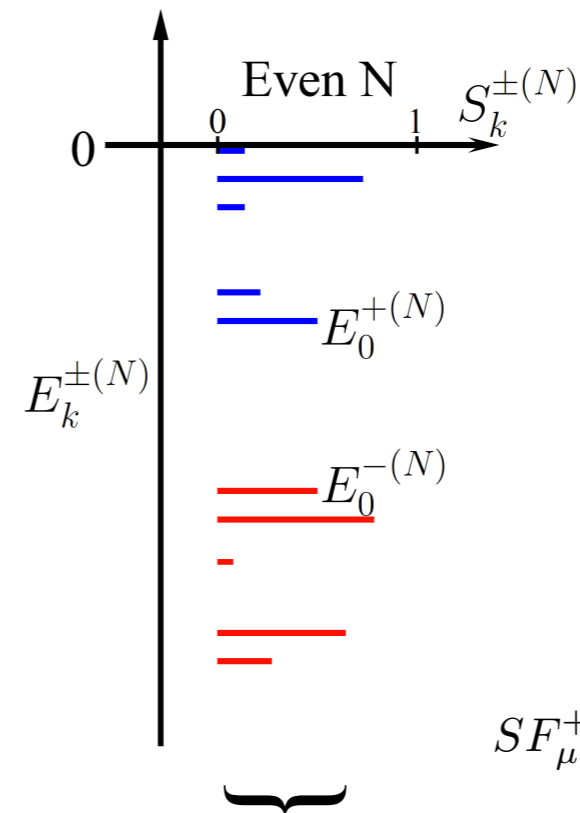
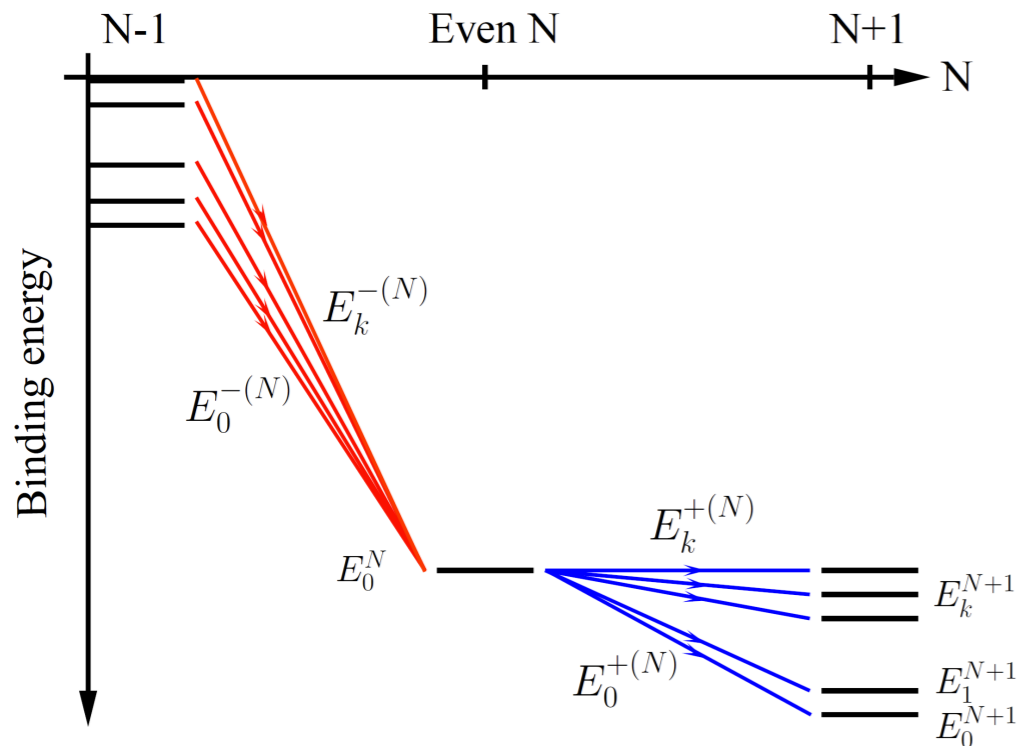
$$E_{\mu}^+ \equiv E_{\mu}^{N+1} - E_0^N$$

$$E_{\nu}^- \equiv E_0^N - E_{\nu}^{N-1}$$

+ numerator

spectral strength distribution

$$\mathcal{S}(z) = \sum_{\mu \in \mathcal{H}_{A+1}} SF_{\mu}^+ \delta(z - E_{\mu}^+) + \sum_{\nu \in \mathcal{H}_{A-1}} SF_{\nu}^- \delta(z - E_{\nu}^-)$$



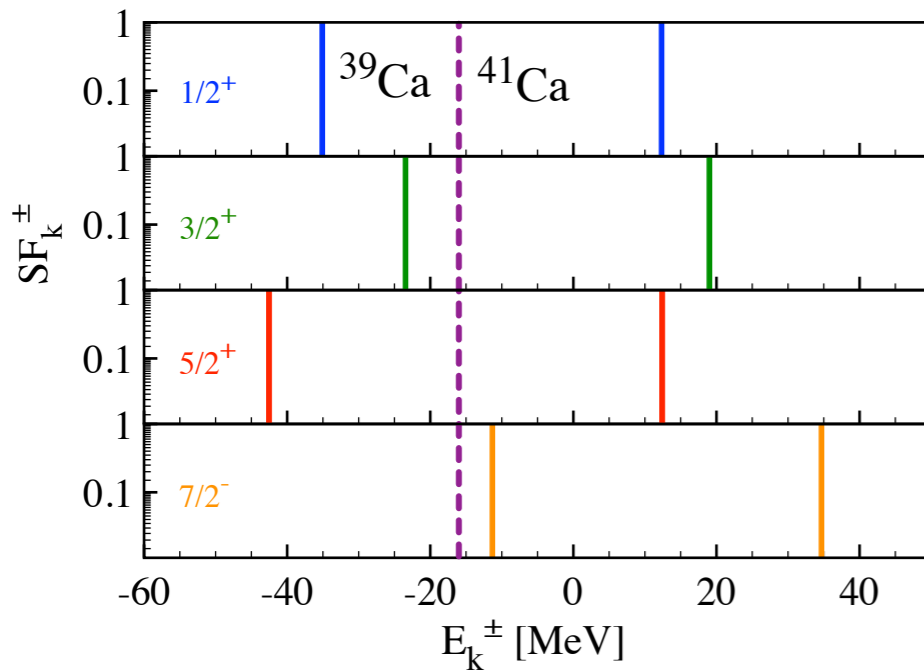
$$SF_{\mu}^+ \equiv \text{Tr}_{\mathcal{H}_1} [\mathbf{S}_{\mu}^+] = \sum_{a \in \mathcal{H}_1} |U_{\mu}^a|^2$$

spectroscopic factors

$$SF_{\nu}^- \equiv \text{Tr}_{\mathcal{H}_1} [\mathbf{S}_{\nu}^-] = \sum_{a \in \mathcal{H}_1} |V_{\nu}^a|^2$$

Spectral strength distribution

Dyson 1st order (HF)

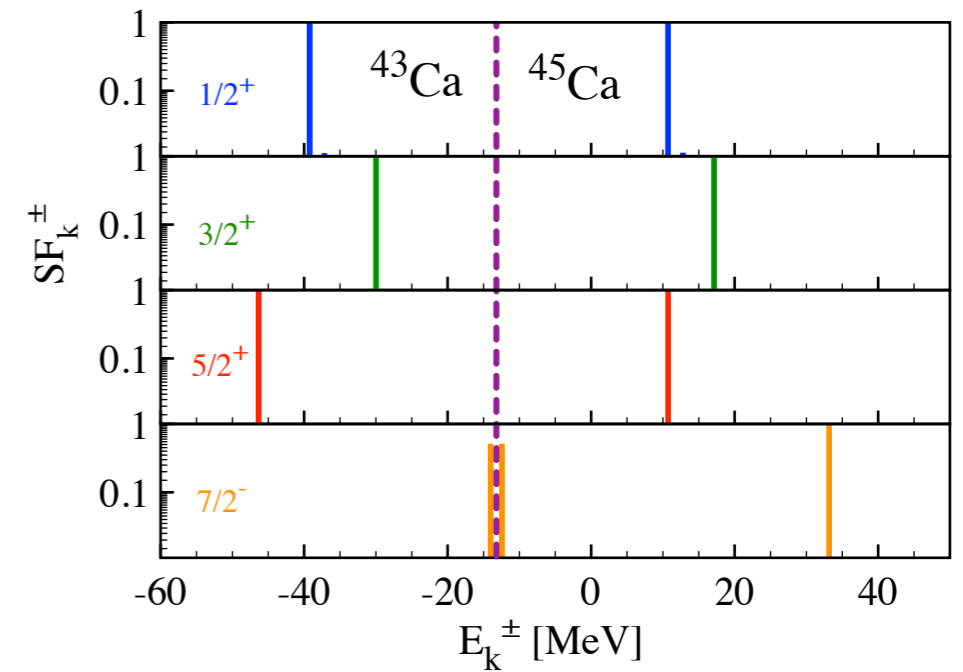


Fragmentation

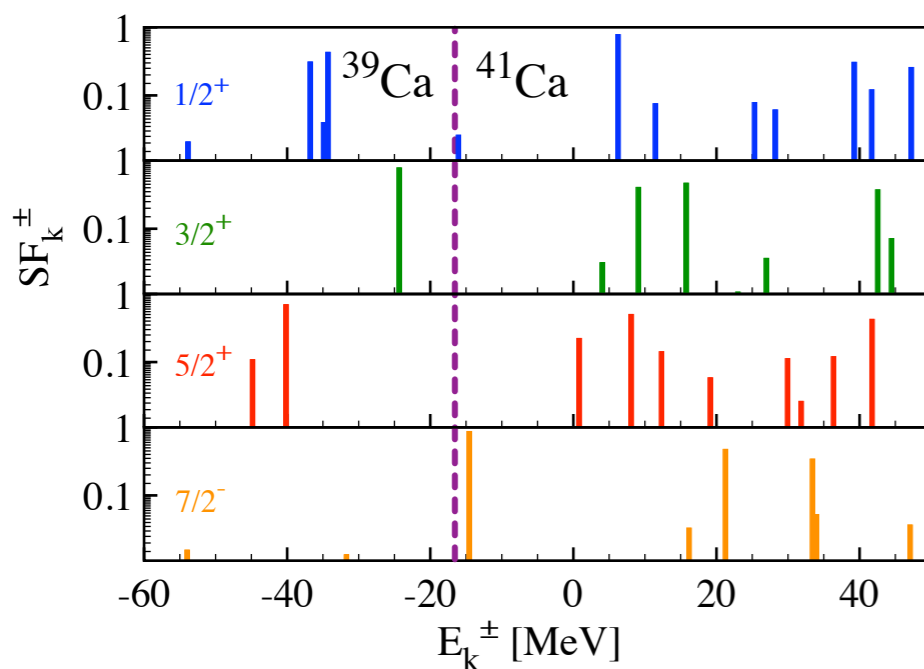
Static pairing



Gorkov 1st order (HFB)



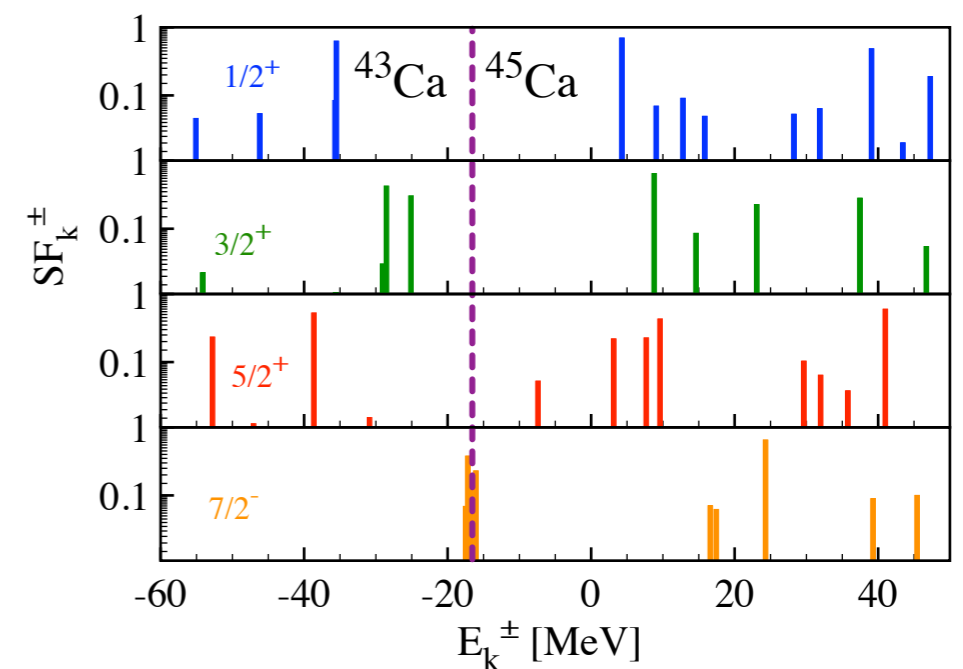
Dyson 2nd order



Dynamical
fluctuations

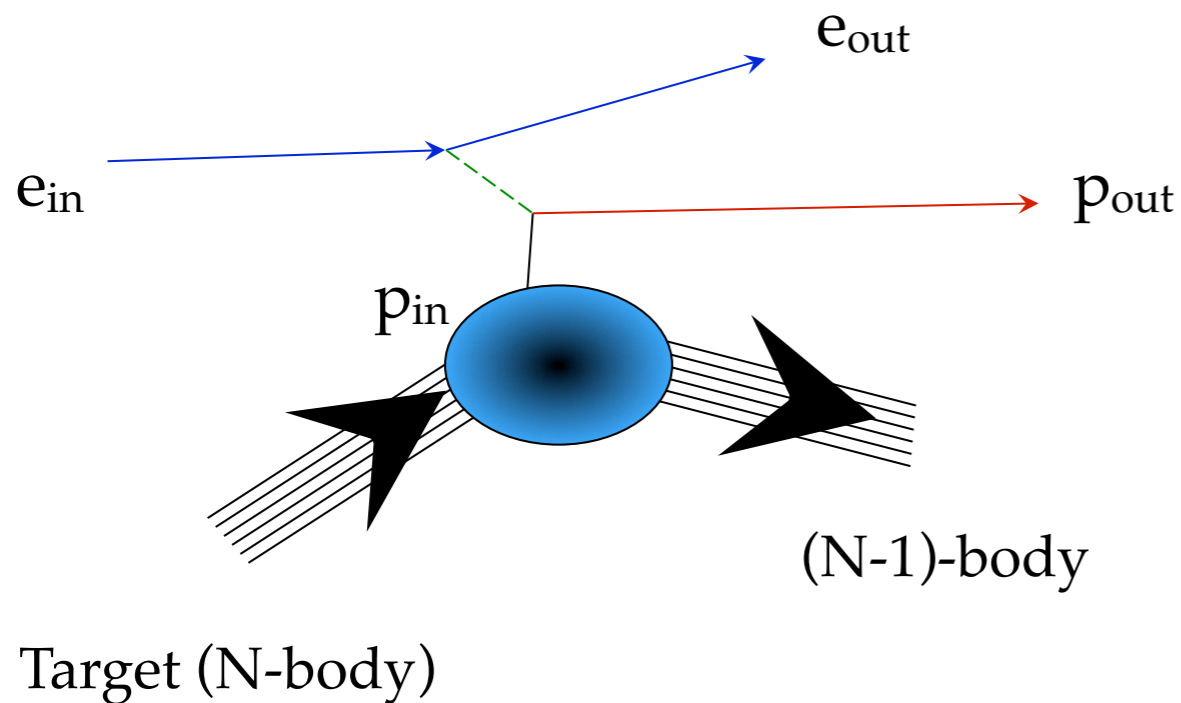


Gorkov 2nd order



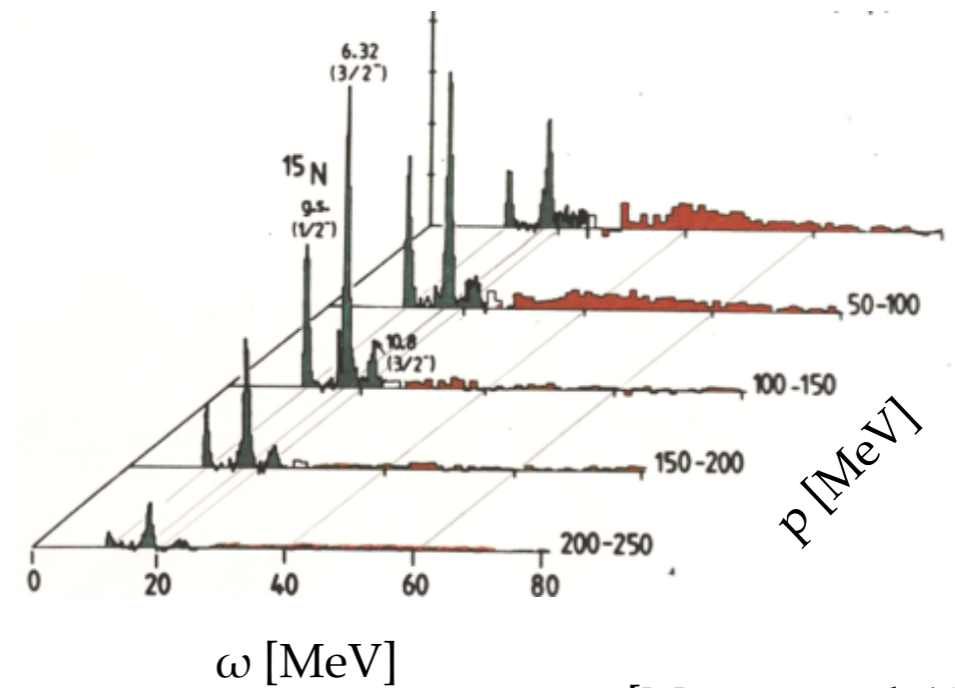
Spectral strength in experiments

⊙ Spectroscopy via knock-out reactions



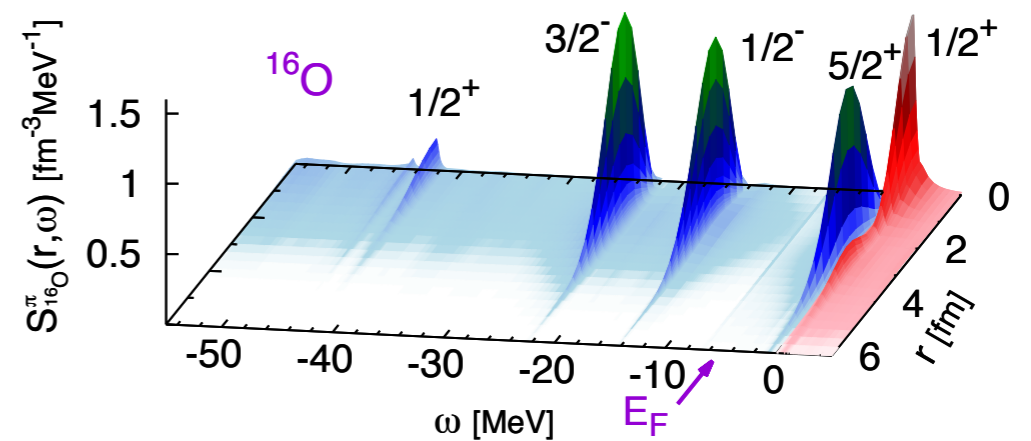
By measuring e_{in} , e_{out} and p_{out}
get information on p_{in}

Results from $(e,e'p)$ on ^{16}O (ALS in Saclay)



[Mougey *et al.* 1980]

SCGF calculations



[Cipollone *et al.* 2015]

Effective Single-Particle Energies (ESPEs)

Spectroscopic probability matrices

$$S_{\mu}^{+pq} \equiv \langle \Psi_0^A | a_p | \Psi_{\mu}^{A+1} \rangle \langle \Psi_{\mu}^{A+1} | a_q^{\dagger} | \Psi_0^A \rangle$$

$$S_{\nu}^{-pq} \equiv \langle \Psi_0^A | a_q^{\dagger} | \Psi_{\nu}^{A-1} \rangle \langle \Psi_{\nu}^{A-1} | a_p | \Psi_0^A \rangle$$

Spectroscopic factors

$$SF_{\mu}^{+} \equiv \text{Tr}[\mathbf{S}_{\mu}^{+}]$$

$$SF_{\nu}^{-} \equiv \text{Tr}[\mathbf{S}_{\nu}^{-}]$$

[Baranger 1970] ↓

Centroid one-body Hamiltonian

$$\mathbf{h}^{\text{cent}} \equiv \sum_{\mu} \mathbf{S}_{\mu}^{+} E_{\mu}^{+} + \sum_{\nu} \mathbf{S}_{\nu}^{-} E_{\nu}^{-} = \mathbf{T} + \mathbf{\Sigma}(\infty)$$

Self-energy

$$\Sigma(\omega) = \Sigma(\infty) + \Sigma^{\text{dyn}}(\omega)$$

Energy-independent part of the self-energy

Effective single-particle energies

$$\mathbf{h}^{\text{cent}} \psi_p^{\text{cent}} = e_p^{\text{cent}} \psi_p^{\text{cent}}$$

Baranger ESPEs

- Defined solely from Schrödinger eq.
- Computable in any many-body scheme
- Relate to the average dynamics of nucleons
- Reduce to HF SPEs in HF approximation

$$e_p^{\text{cent}} \equiv \sum_{\mu \in \mathcal{H}_{A+1}} S_{\mu}^{+pp} E_{\mu}^{+} + \sum_{\nu \in \mathcal{H}_{A-1}} S_{\nu}^{-pp} E_{\nu}^{-}$$

Inverting ESPEs

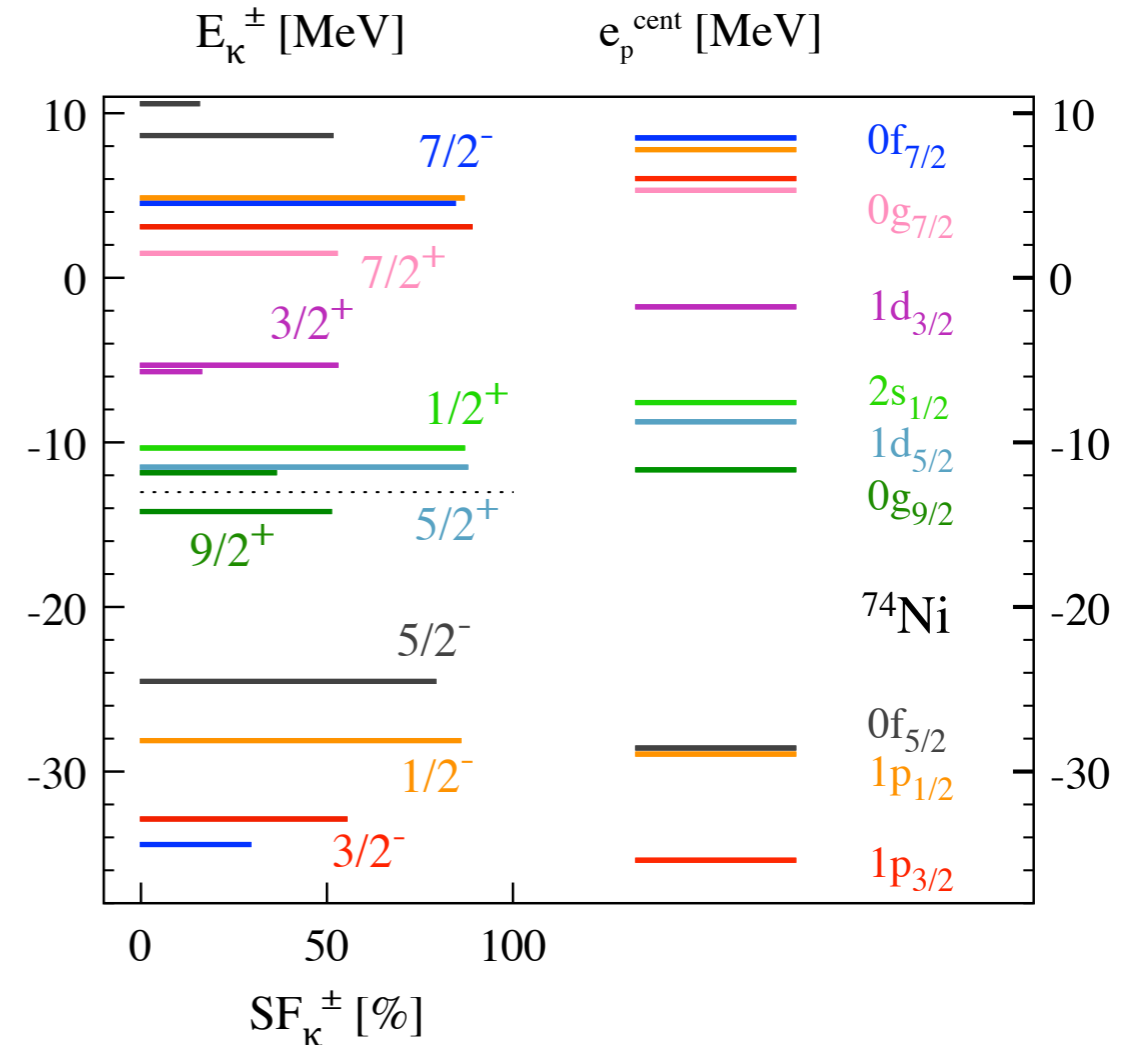
⊙ Baranger ESPEs in the basis associated to h^{cent}

$$e_p^{\text{cent}} \equiv \sum_{\mu \in \mathcal{H}_{A+1}} S_{\mu}^{+pp} E_{\mu}^{+} + \sum_{\nu \in \mathcal{H}_{A-1}} S_{\nu}^{-pp} E_{\nu}^{-}$$

invert \downarrow (same for E_{ν}^{-})

$$E_{\mu}^{+} = \sum_p s_{\mu}^{+pp} e_p^{\text{cent}} + \sum_{pq} s_{\mu}^{+pq} \Sigma_{qp}^{\text{dyn}}(E_{\mu}^{+})$$

with $s_{\mu}^{+} \equiv \mathbf{S}_{\mu}^{+} / SF_{\mu}^{+}$ & $\Sigma^{\text{dyn}}(\omega) \equiv \Sigma(\omega) - \Sigma(\infty)$



⊙ Rigorous partitioning into independent-particle + correlation contributions

- Exact result, no approximations so far
- A given one-nucleon addition energy does not relate to a single ESPE
- Connection between the two spectra is of matrix character

Partitioning & scale dependence

- ⊙ Nuclear Hamiltonian carries an intrinsic scale resolution Λ_{init}
- ⊙ One can further apply a unitary transformation $U(\lambda)$ over Fock space

$$\begin{aligned} H(\lambda) &\equiv U(\lambda) H U^\dagger(\lambda) \\ |\Psi_\mu^A(\lambda)\rangle &\equiv U(\lambda) |\Psi_\mu^A\rangle \end{aligned} \quad \rightarrow \quad H(\lambda) |\Psi_\mu^A(\lambda)\rangle = E_k^A |\Psi_\mu^A(\lambda)\rangle$$

- ⊙ Any other operator transforms accordingly

$$O(\lambda) \equiv U(\lambda) O U^\dagger(\lambda) \equiv O^{1N}(\lambda) + O^{2N}(\lambda) + O^{3N}(\lambda) + \dots$$

- ⊙ Spectroscopic amplitudes defined at any value of λ

$$U_\mu^p(\lambda) \equiv \langle \Psi_0^A(\lambda) | a_p | \Psi_\mu^{A+1}(\lambda) \rangle$$

$$V_\nu^p(\lambda) \equiv \langle \Psi_0^A(\lambda) | a_p^\dagger | \Psi_\nu^{A-1}(\lambda) \rangle$$

- ⊙ Generator of the transformation

$$\eta(\lambda) \equiv \frac{dU(\lambda)}{d\lambda} U^\dagger(\lambda)$$

Partitioning & scale dependence

- ⊙ Spectroscopic probabilities / factors are **scale-dependent**

$$\frac{d}{d\lambda} V_\nu^p(\lambda) = -\langle \Psi_\nu^{A-1}(\lambda) | [\eta(\lambda), a_p] | \Psi_0^A(\lambda) \rangle^*$$

$$\frac{d}{d\lambda} U_\mu^p(\lambda) = -\langle \Psi_\mu^{A+1}(\lambda) | [\eta(\lambda), a_p^\dagger] | \Psi_0^A(\lambda) \rangle^*$$

- ⊙ ESPEs acquire scale dependence via spectroscopic probabilities

$$e_p^{\text{cent}} \equiv \sum_{\mu \in \mathcal{H}_{A+1}} S_\mu^{+pp} E_\mu^+ + \sum_{\nu \in \mathcal{H}_{A-1}} S_\nu^{-pp} E_\nu^-$$



$$\underbrace{\text{many-body observable}}_{\underbrace{E_\mu^+}_{\text{invariant under } U(\lambda)}} \equiv \underbrace{\sum_p \text{single-particle components } s_\mu^{+pp}(\lambda) e_p^{\text{cent}}(\lambda)}_{\text{varies under } U(\lambda)} + \underbrace{\sum_{pq} \text{correlations } s_\mu^{+pq}(\lambda) \Sigma_{qp}^{\text{dyn}}(E_\mu^+; \lambda)}_{\text{varies under } U(\lambda)}$$

- ⊙ A convenient choice of λ maximises the ESPE component
- ⊙ However, **correlations with observables are not absolute**
- ⊙ Scale must be fixed / specified prior to theoretical / experimental comparisons

SRG transformation & ab initio calculations

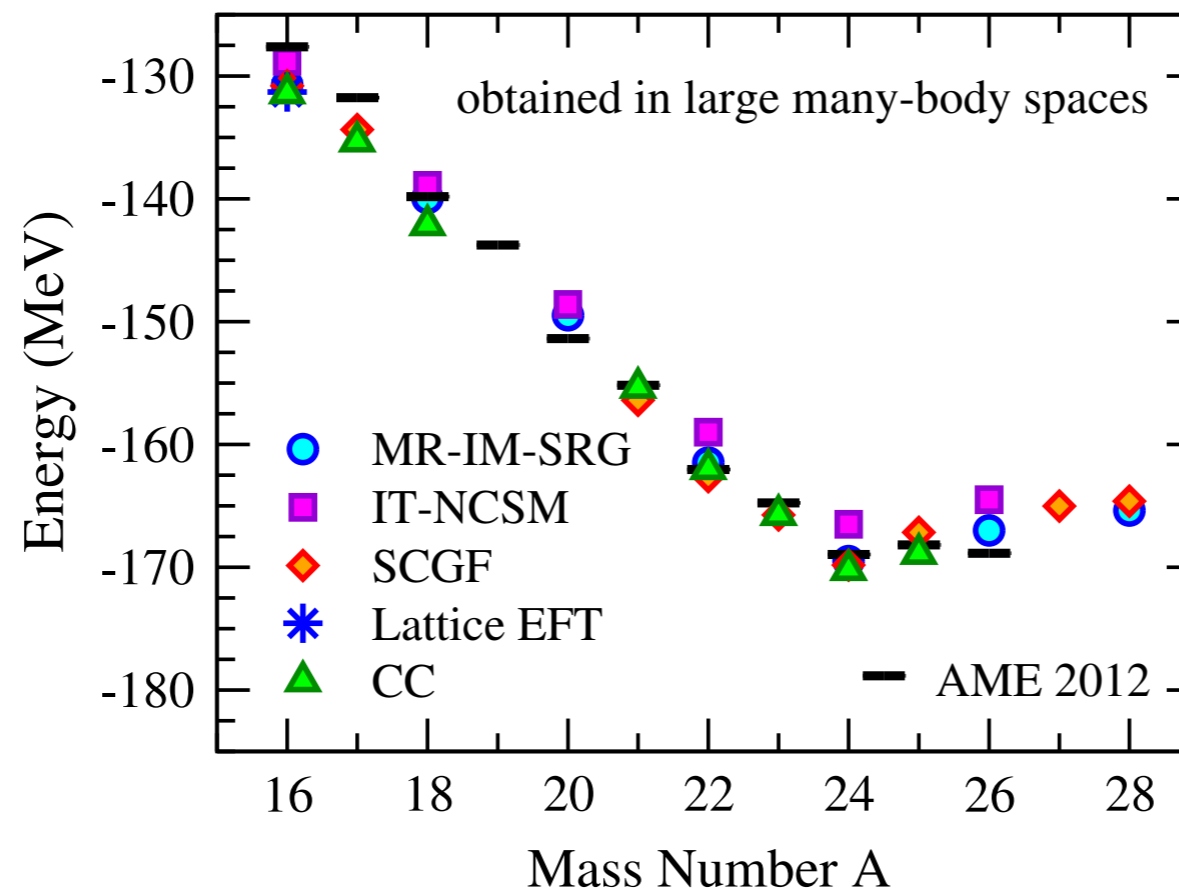
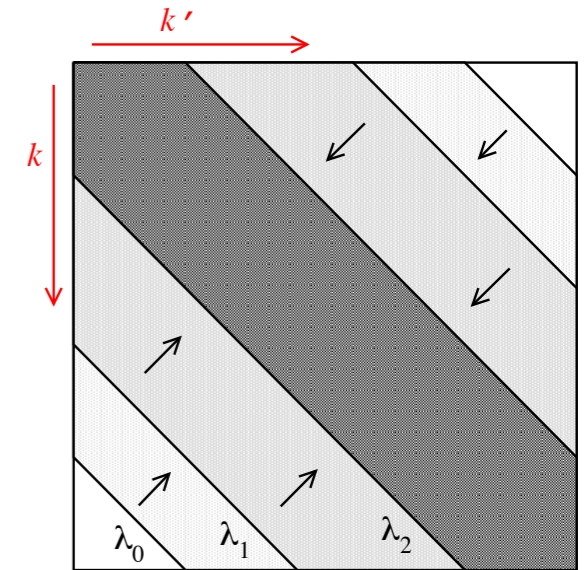
⊙ SRG transformations $U(\lambda)$ applied to the starting Hamiltonian $H(\Lambda_{\text{init}})$

○ Limited range of variation: $\lambda \in \{1.88, 2.0, 2.24\} \text{ fm}^{-1}$

⊙ Two different ab initio methods

○ **Gorkov-Green's functions** [Somà, Barbieri, Duguet 2011, ...]

○ **In-medium SRG** [Tsukiyama, Bogner, Schwenk 2010, Hergert *et al.* 2013, ...]



[Hergert *et al.* 2013]

[Cipollone *et al.* 2013]

[Jansen *et al.* 2014]

[Lähde *et al.* 2014]

Breaking of unitarity

- ⊙ **Unitarity artificially broken**

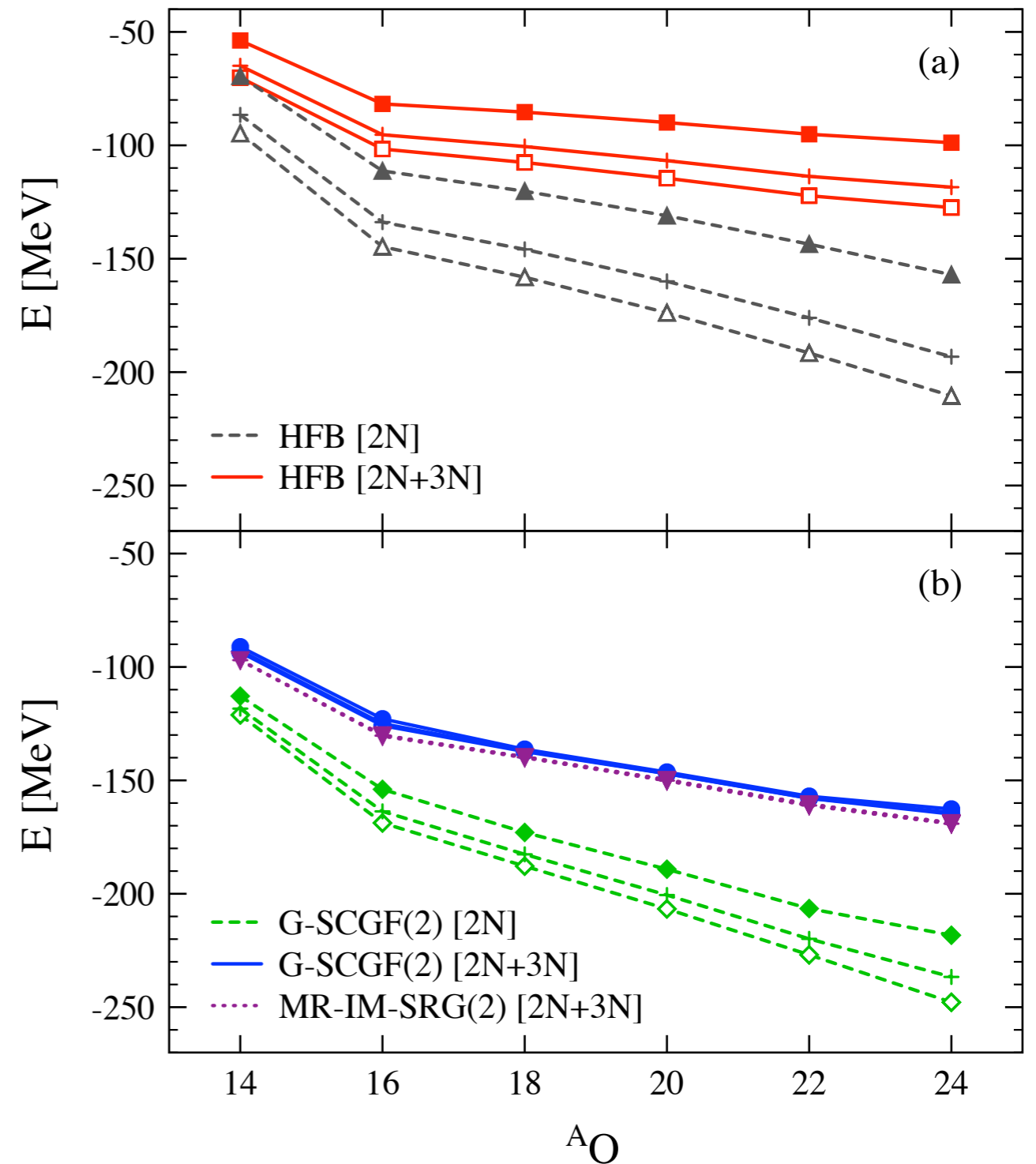
- Omission of A -body operators with $A > 3$
- Many-body truncations

- ⊙ **Breaking can be estimated**

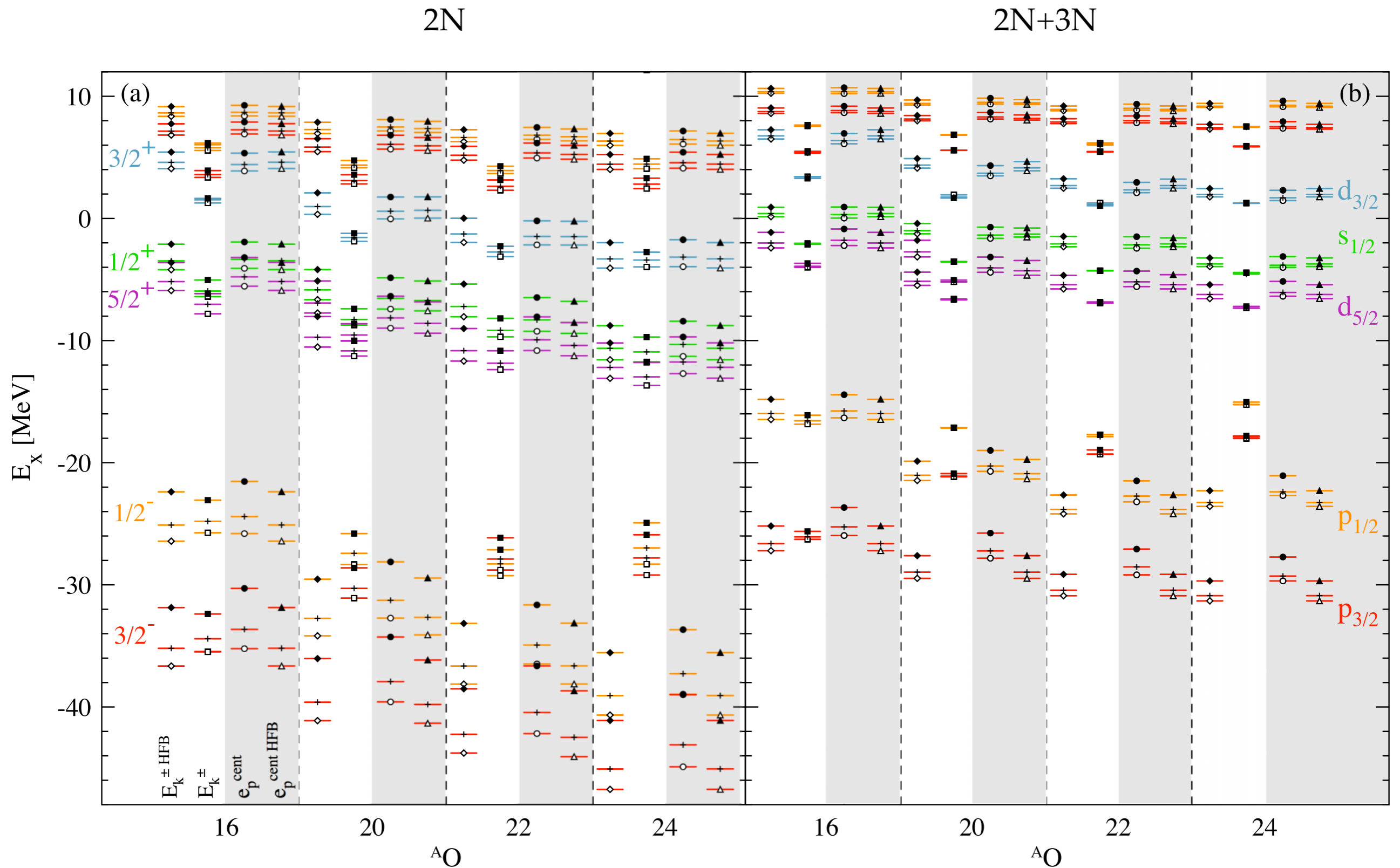
- Omitting 3-body operators
- Degrading the many-body truncation

- ⊙ **Breaking for total energy**

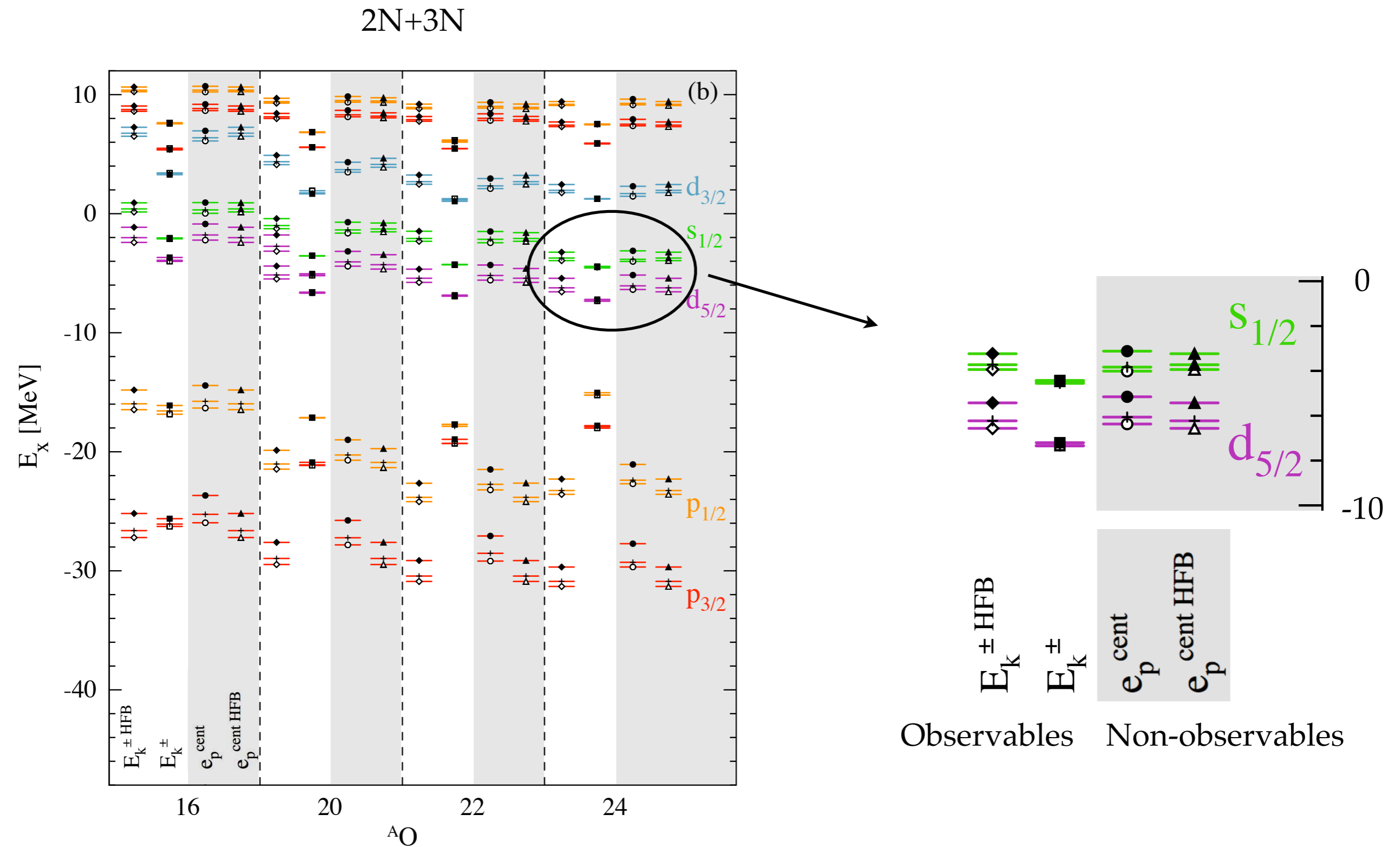
- Around 1 MeV for GGF
- Around 100 keV for IM-SRG



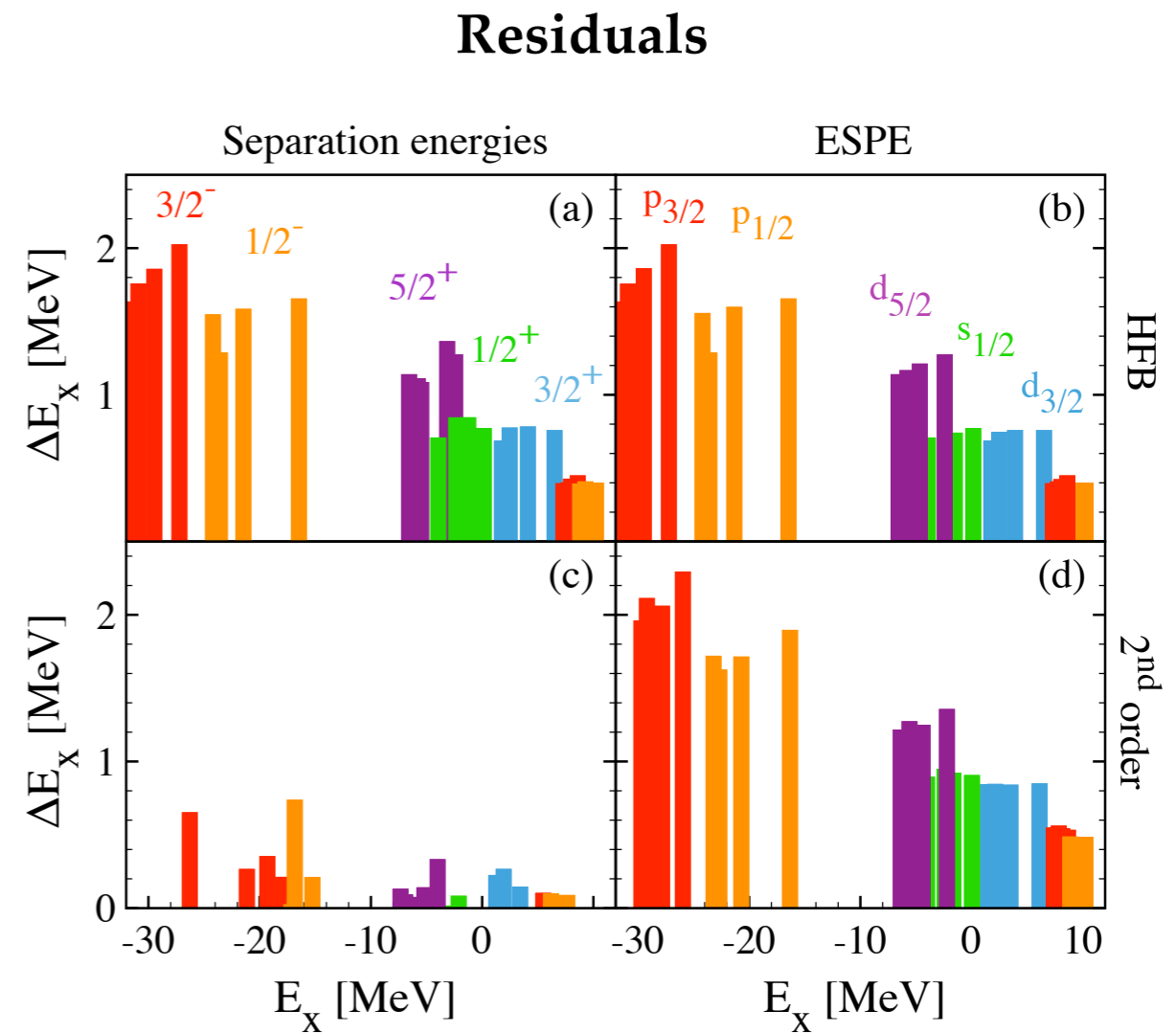
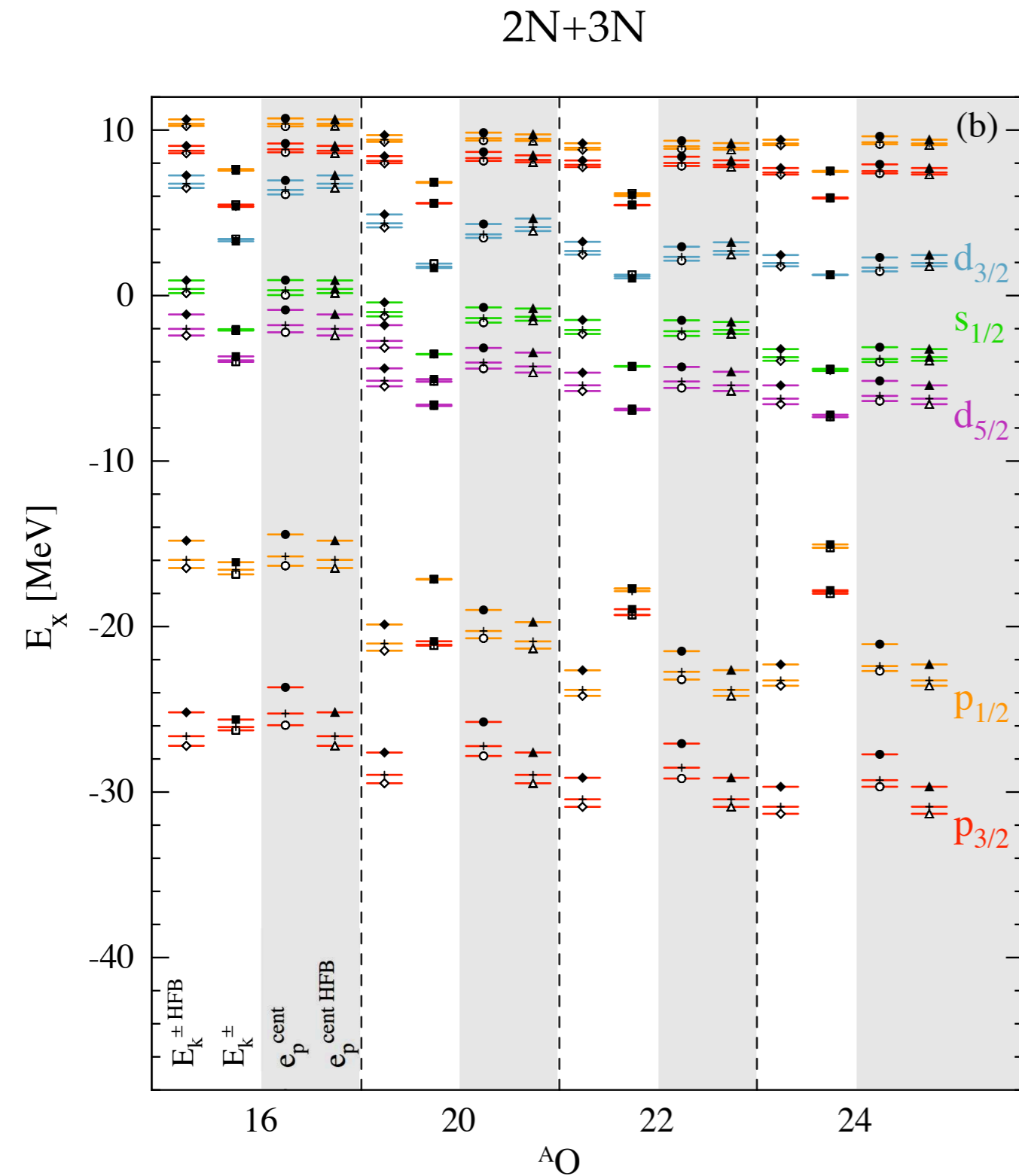
Scale (in)dependence of separation energies & ESPs



Scale (in)dependence of separation energies & ESPs



Scale (in)dependence of separation energies & ESPEs



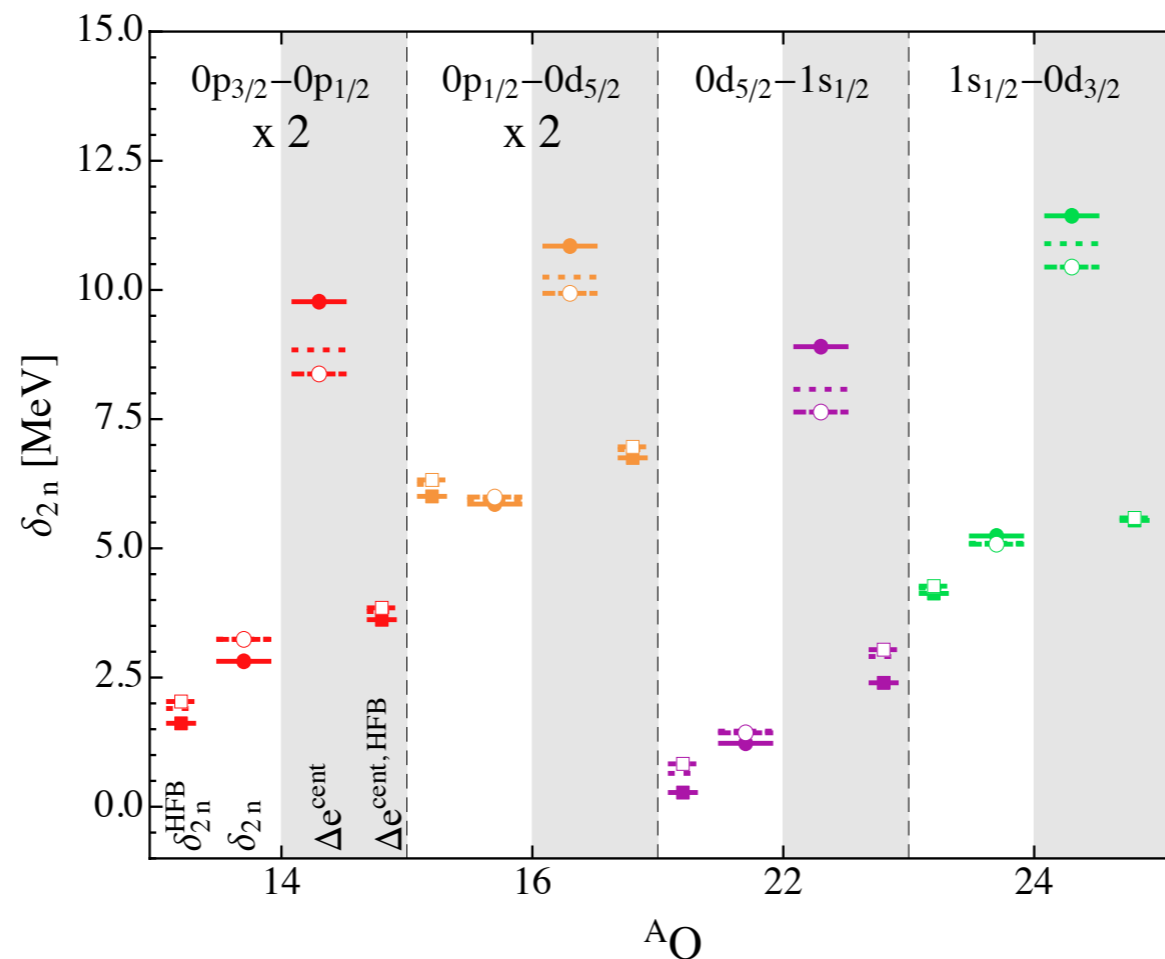
- Spread of sep. en. reduced significantly
- Spread of ESPEs unchanged
- ESPEs less sensitive to correlations

Shell gaps

⊙ **Gaps across the Fermi energy (equal in the HF limit)**

○ (Observable) two-neutron shell gap $\delta_{2n}(N, Z) \equiv \frac{1}{2} [E(N+2, Z) - 2E(N, Z) + E(N-2, Z)]$

○ (Non-observable) ESPE Fermi gap $\Delta e_F^{\text{cent}}(N, Z) \equiv e_p^{\text{cent}}(N, Z) - e_h^{\text{cent}}(N, Z)$



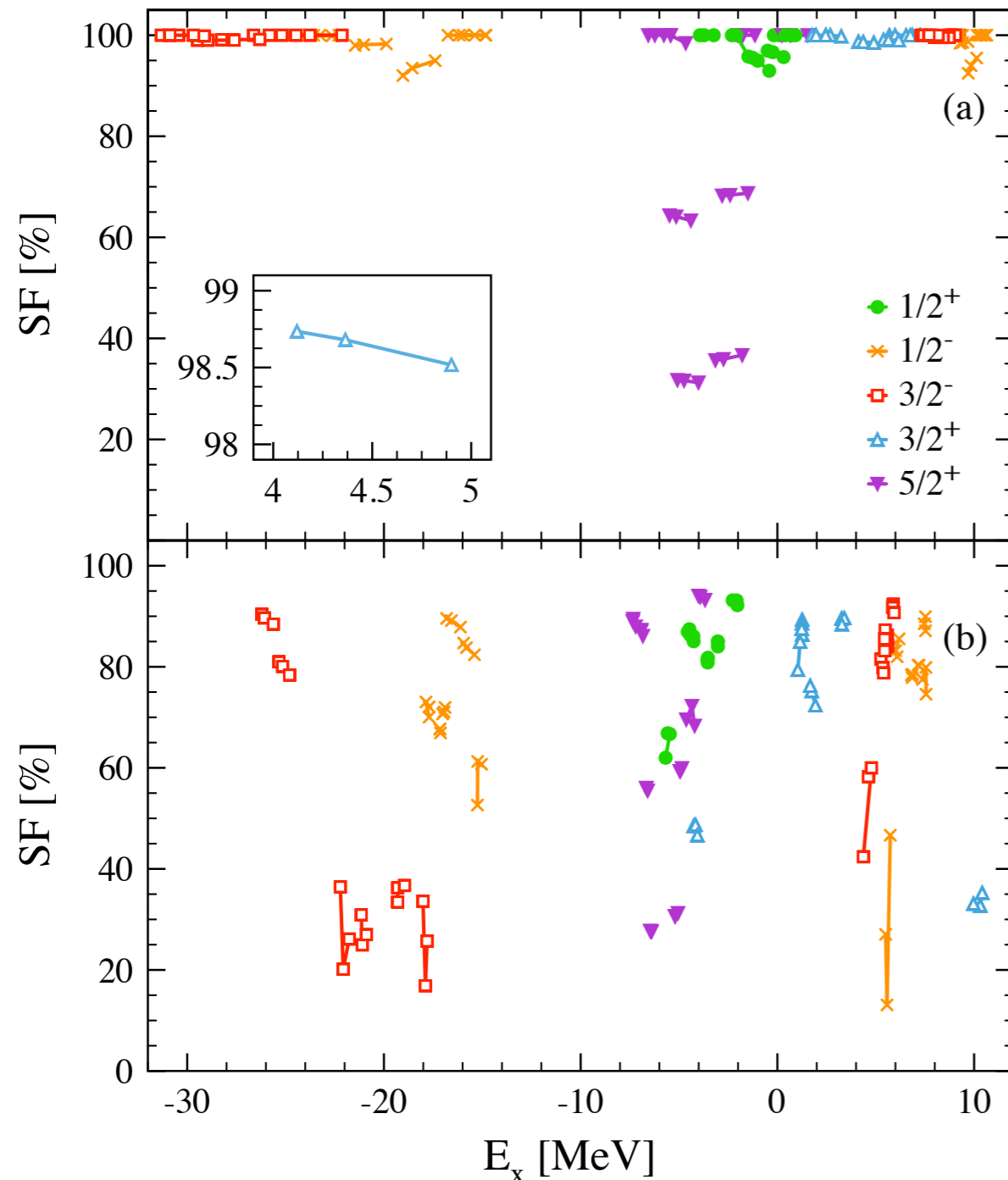
○ **HFB level:** ESPE & 2N gaps similar, the former well captures the latter

○ **Correlated calculation:** scale dependence of ESPE gaps is systematically large

Spectroscopic factors

◎ Compilation of SF for one-neutron addition/removal on $^{14-24}\text{O}$

- Limited running with $\lambda \in \{1.88, 2.0, 2.24\} \text{ fm}^{-1}$!



HFB

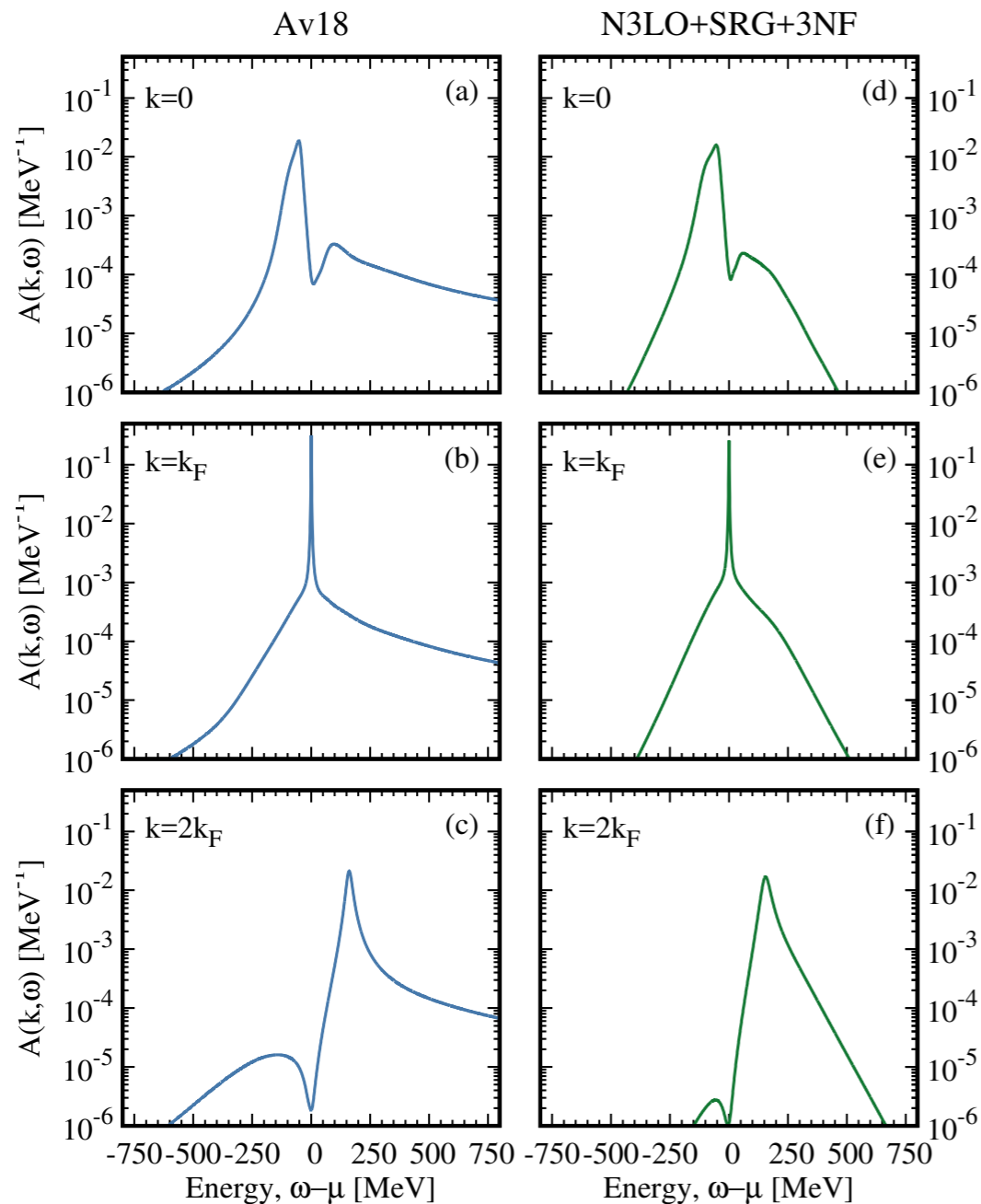
- Independent (quasi-)particle picture
- SFs ~ 1
- Spread is horizontal

Correlated

- Horizontal spread minimised
- Spectroscopic strength now fragmented
- Some scale dependence of SFs appears

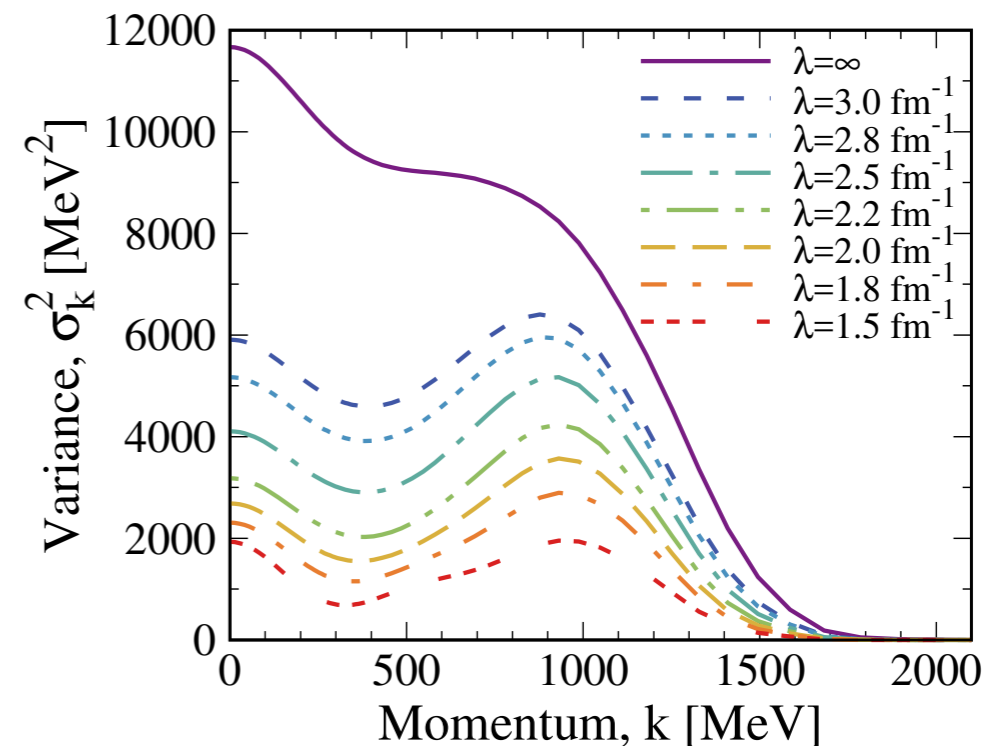
SRG & correlations in infinite matter

- ◎ Larger range of scales can be explored in infinite nuclear matter
 - Momentum tails in spectral function depend on the interaction
 - Variance depicts amount of correlations



$$\sigma_k^2 = \int_{-\infty}^{+\infty} \frac{d\omega}{2\pi} [\omega - m_k^{(1)}]^2 \mathcal{A}_k(\omega) = m_k^{(2)} - (m_k^{(1)})^2$$

$$= - \int_{-\infty}^{+\infty} \frac{d\omega}{\pi} \text{Im} \Sigma_k(\omega)$$



[Rios, Carbone, Polls 2017]

Summary

◎ Part I

- **Correlations are scheme and scale dependent**
- What balance between different ways of accounting for correlations?
- Similarity renormalisation group as a knob for (short-range) correlations

◎ Part II

- **Non-observability of shell structure** formally revisited
- Ab initio calculations corroborate formal analysis
- Correlations between observables & shell structure **depend on the resolution scale**
- Scale/scheme dependence should be explicit & consistent

◎ Perspectives

- **Quantification of scale dependence** interesting from a pragmatic point of view
- Focus should be on **consistency** to combine structure & reactions