

Charge Densities from Electron Scattering

achievements for stable nuclei, challenges for unstable nuclei

Ingo Sick

Purpose of talk

discuss relation data \leftrightarrow densities

address limitations

particularly with respect to low-L measurements

Start with historical perspective on extracting $\rho(r)$

1. Fifties

first low- q experiments

interpreted in terms of PWBA

$$\sigma(E, \theta) = \sigma_{Mott} F^2(q)$$

momentum transfer $q \sim 2 E \sin(\theta/2)$

σ_{Mott} = cross section for point-nucleus

$$F(q) = \int_0^\infty \rho(r) \frac{\sin(qr)}{qr} r^2 dr$$

Form factor = Fourier-transform of density

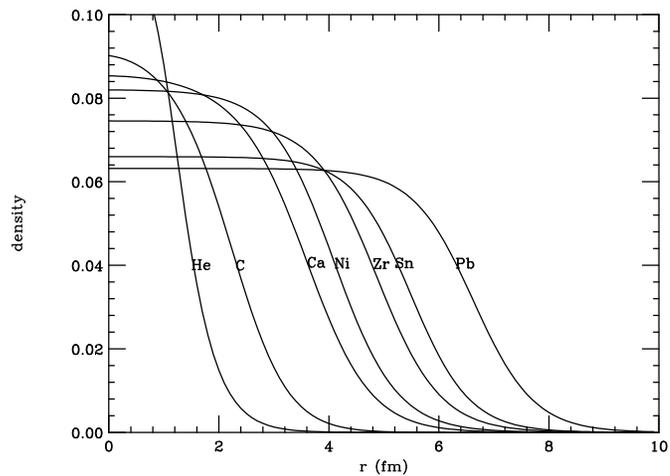
Soon first calculations of σ via solution of Dirac equation
accounts for distortion of e-waves
makes relation $\sigma \leftrightarrow \rho(r)$ less transparent

Determination of model densities

$$\text{3PF } \rho(r) = (1 + w(r/c)^2)/(1 + e^{(r-c)/z})$$

$$\text{3PG } \rho(r) = (1 + w(r/c)^2)/(1 + e^{(r^2-c^2)/z^2})$$

Establish general evolution of $\rho(r)$ with Z, A



Integral properties of interest

- half-density radius c
 - strongly correlated with rms-radius R
- surface thickness z
 - not too correlated with c
- central depression w
 - problematic, discuss below

Knowhow for determination of R

largely lost since
re-activated + improved with recent problems with proton R
new insights (see below)

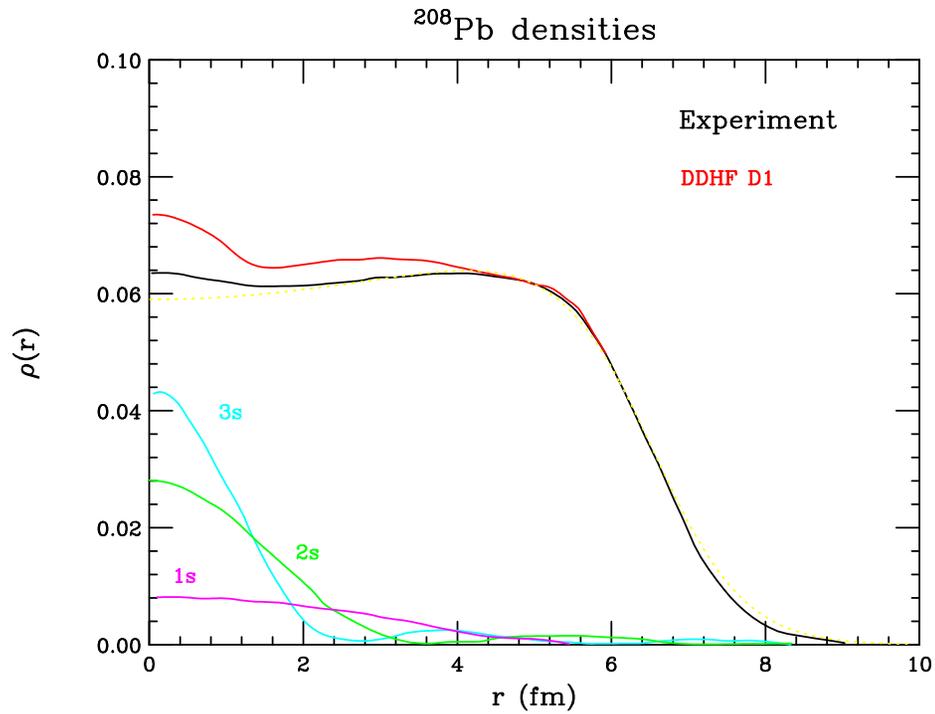
2. Sixties

higher- q data (mainly from HEPL)

more complicated shapes of ρ
Fermi/Gauss + modifications

not always clear if given feature significant
or very model-dependent
example: w

Example: ^{208}Pb : central peak/depression [1]



3PG parameterization yields large $w \sim 0.33$
→ central depression?

SOG (model-independent density) does not really show it

3. Seventies

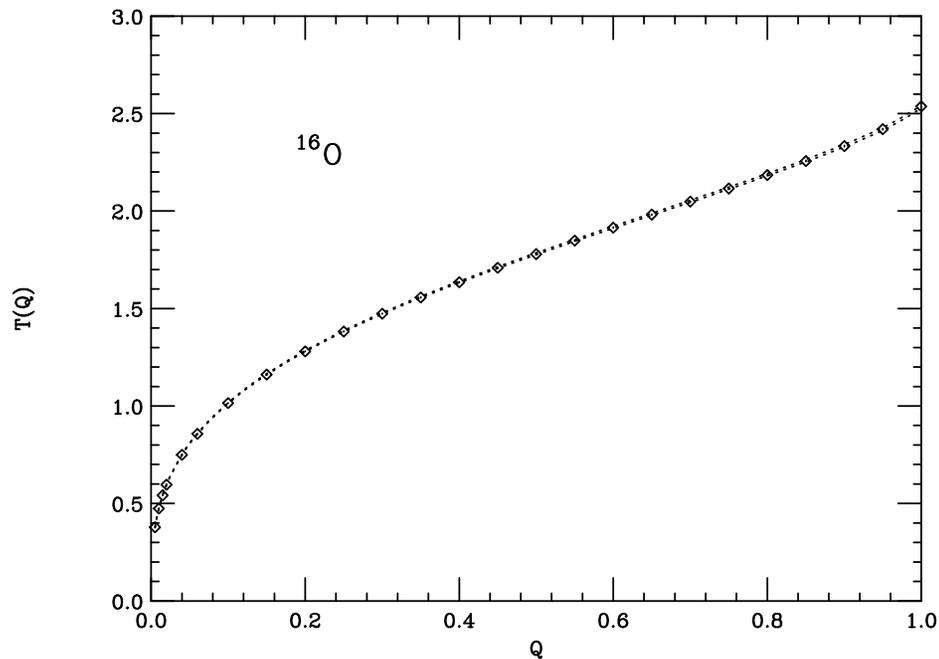
Extraction of model-independent information

possible only for *integral* quantities due to finite q_{max}

Partial moment function (F. Lenz [2])

$T(Q) = \int_0^Q r(Q')dQ'$ with Q =fraction of charge out to radius r

$T(Q)/Q$ = linear moment averaged over charge Q



Strictly model-independent, welldefined δT , not intuitive, never caught on

Less model-dependent densities

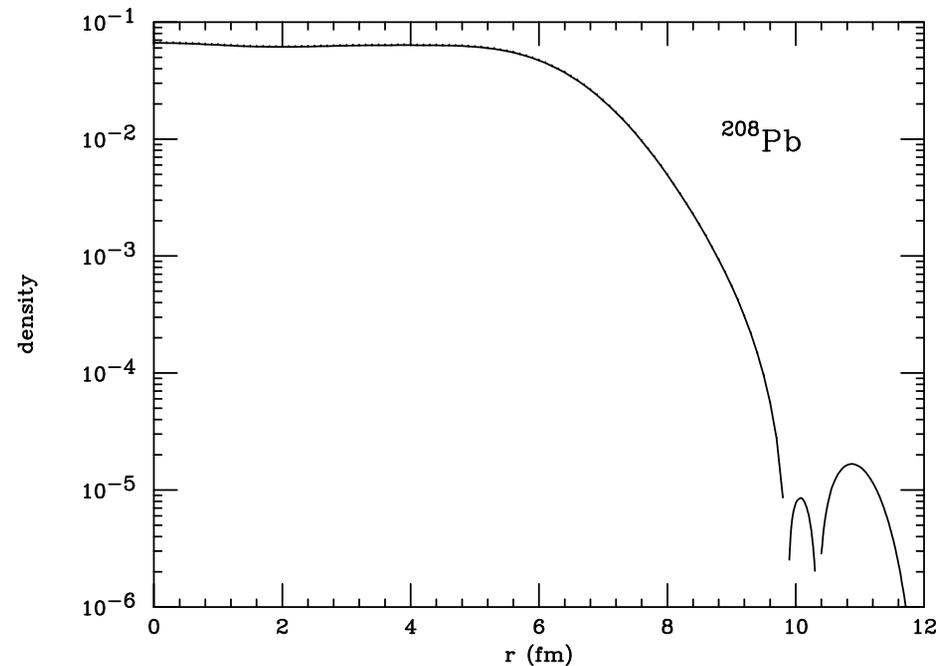
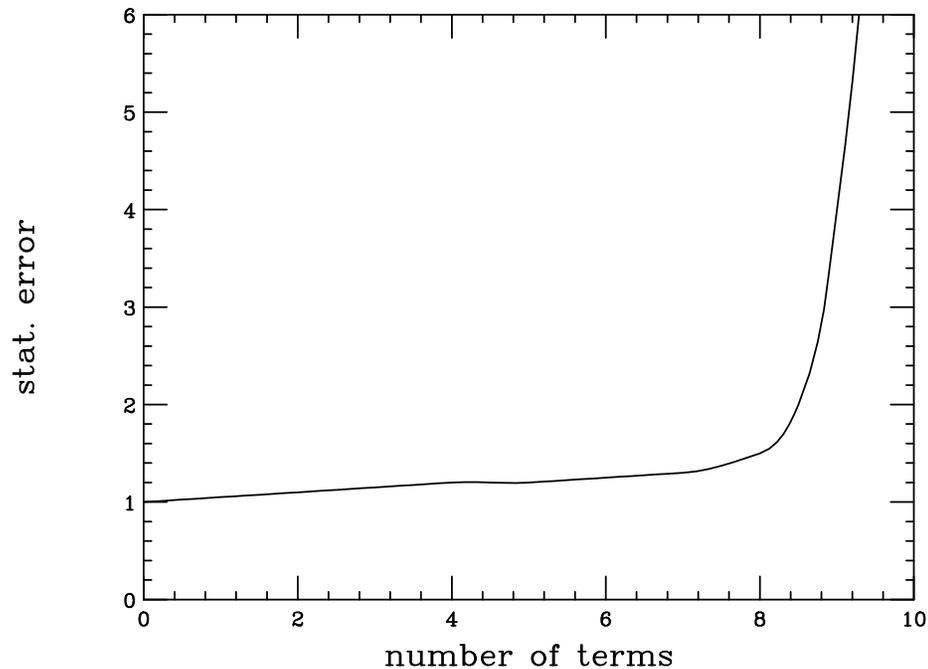
attempts to overcome model dependence

practical once $q_{max} \sim 3 \text{ fm}^{-1}$

expansion of $\rho(r)$ in complete basis

typically Fourier-Bessel

but: $\delta\rho(r)$ arbitrary (function of N) unless constrain $F(q > q_{max})$



to get significant $\delta\rho(r)$ must introduce physics constraint

FB misbehaves at $r \sim r_{max}$, detrimental for determination of R (see below)

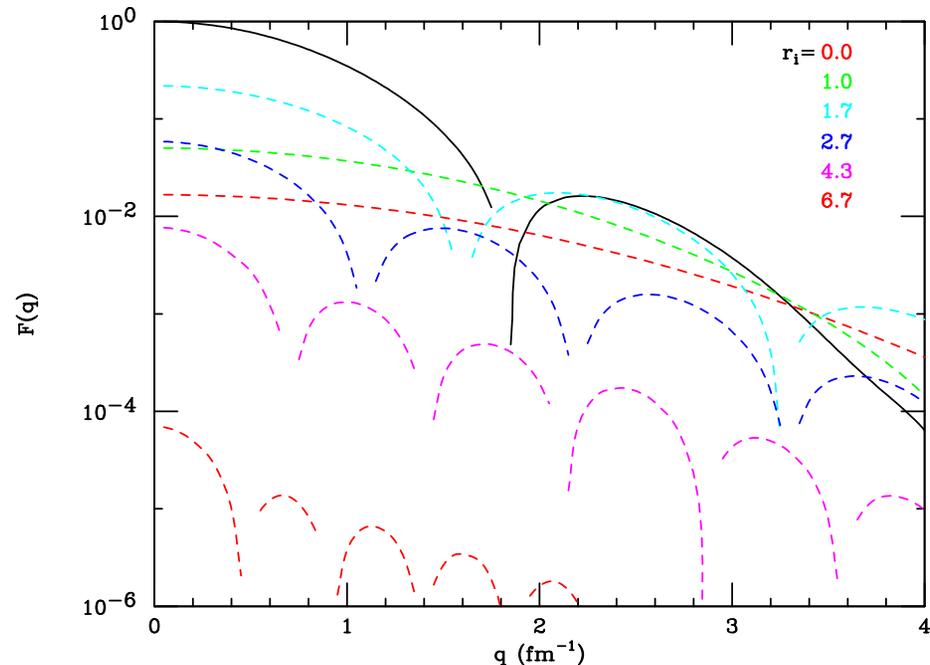
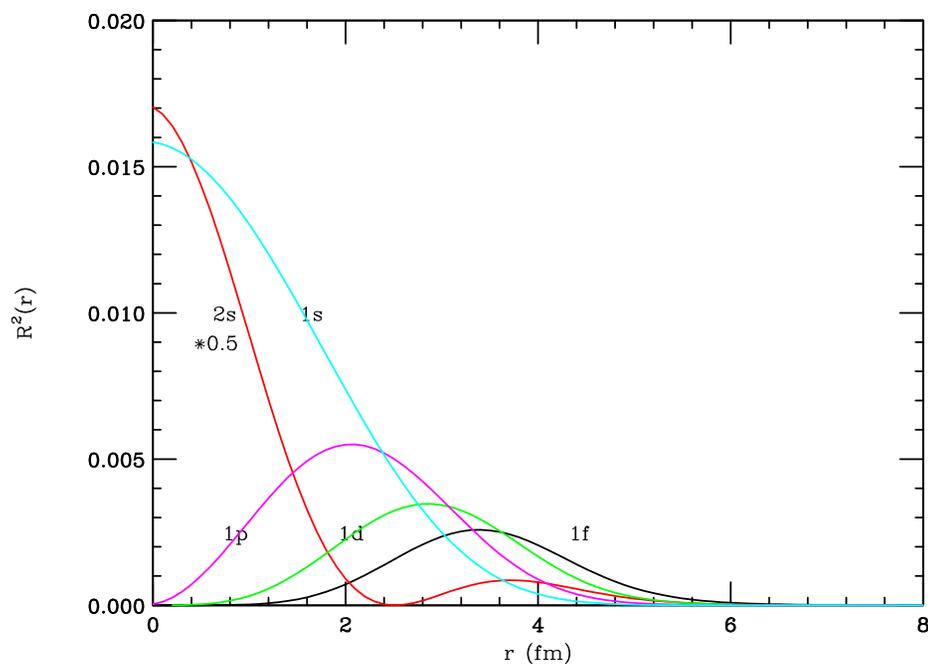
Physics constraint: minimal width Γ of peaks in ρ [3]

smoothly varying $R()$, peaks of minimal width Γ

property of solutions of Schrödinger equation: minimal curvature, $\langle E \rangle$

Parameterize ρ as sum of *many* gaussians SOG with width Γ , placed at radii r_i

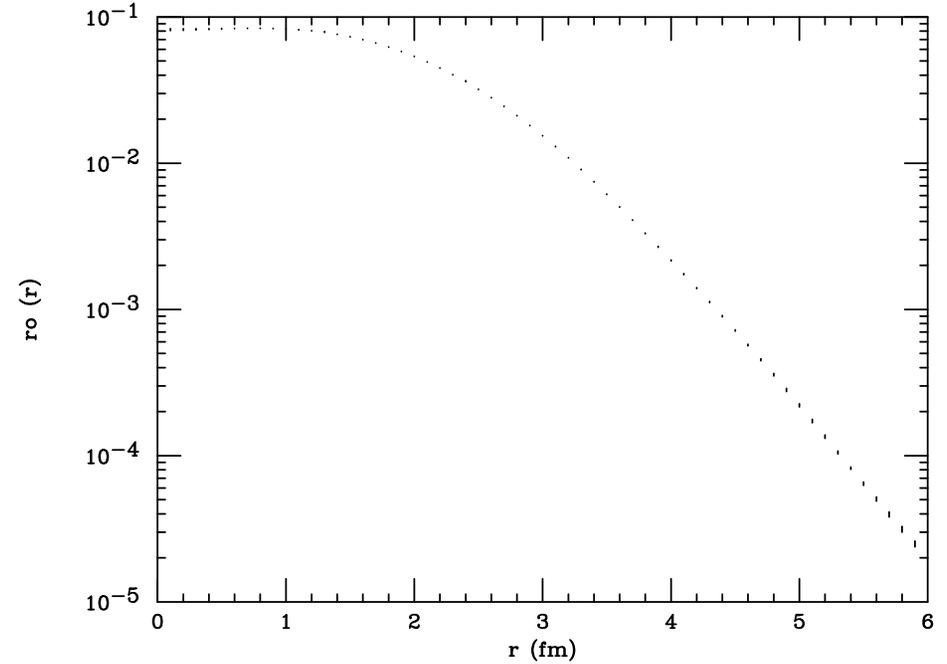
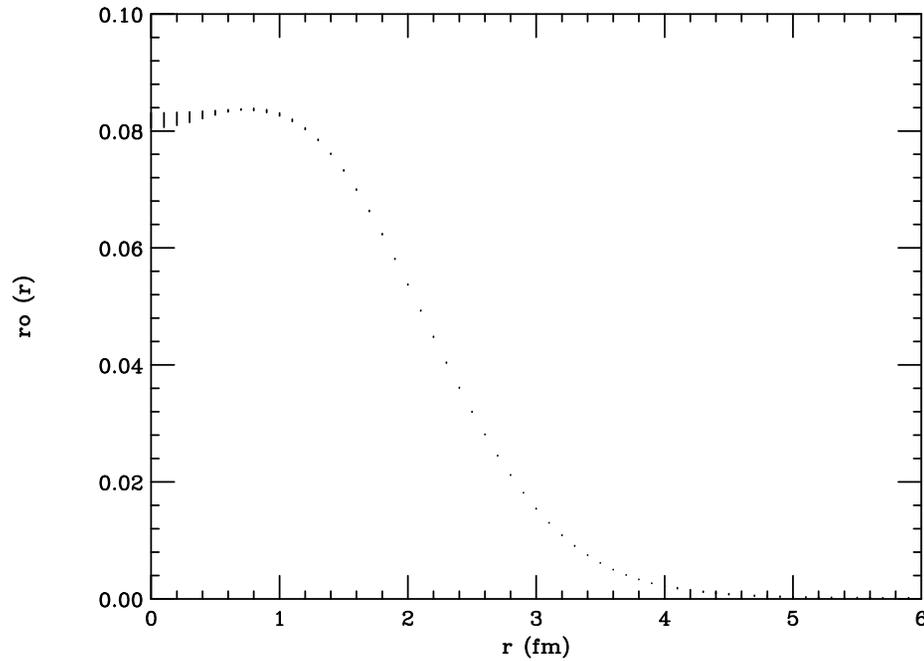
Example: ^{12}C



With physics constraint get significant $\delta\rho(r)$

note: large $r \leftrightarrow$ high-frequency oscillations in q

Get density with significant error bar: ^{12}C



Note (for later)

$\delta\rho$ in tail small (accurate extensive data)

displays typical exponential behavior

$$R(r) \sim e^{-\kappa r} / r$$

4. Eighties: most transparent extraction of ρ : DFT [4]

form factor $F(q)$ = Fourier transform of $\rho(r)$

density $\rho(r)$ = Fourier transform of $F(q)$

$$\rho(r) = \dots \int_0^\infty F(q) \frac{\sin(qr)}{qr} q^2 dq$$

Can exploit? Potential difficulties:

- $F(q)$ not available, only cross sections
Coulomb distortion!
- q_{max} not infinite

Advantage of electrons

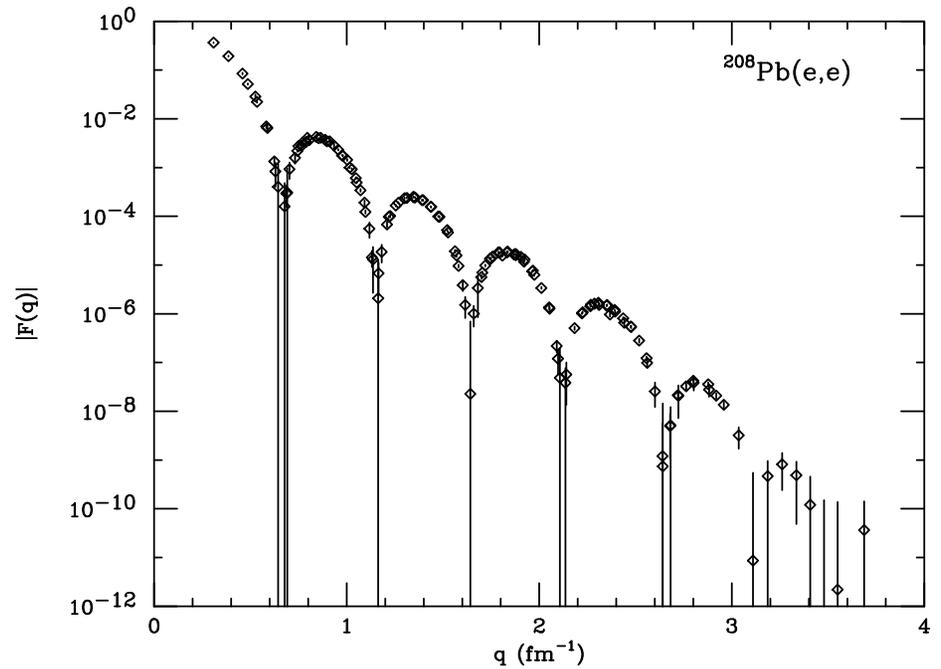
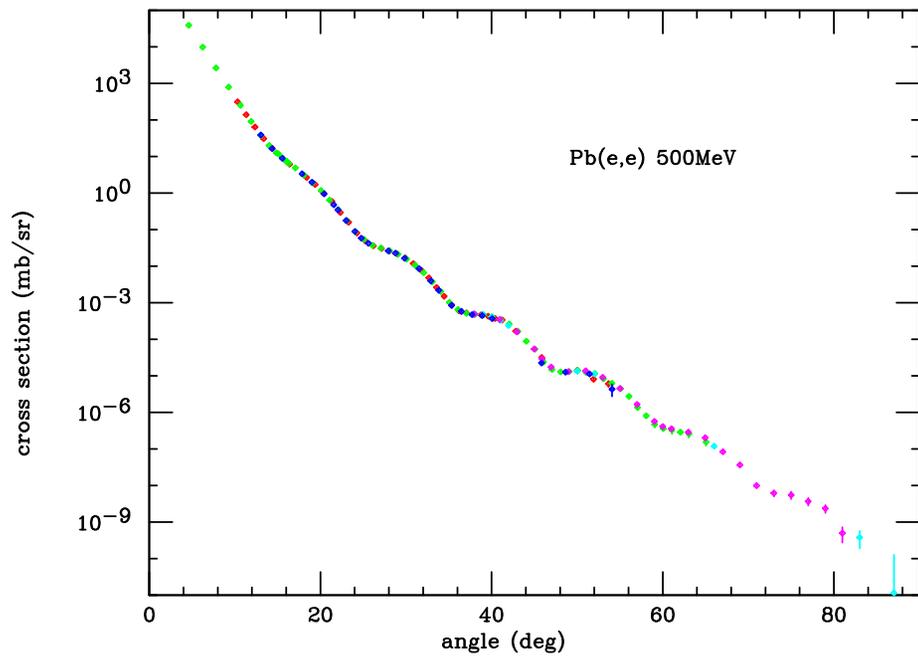
Coulomb distortion relatively small

only function of Z (known)

can remove to get $F(q) \pm \delta F(q)$ via DW code

$$F_{exp}^2 = F_{model}^2 \frac{\sigma_{exp}}{\sigma_{model}}, \quad \delta F_{exp}^2 = \delta \sigma_{exp} / \sigma_{point}$$

Example: cross sections for Pb



experimental $F(q)$ contain same information as σ

To deal with finite q_{max} : study \int as function of upper integration limit

$$\rho(r, q_{max}) = \dots \int_0^{q_{max}} F(q) \frac{\sin(qr)}{qr} q^2 dq$$

$\rho(r, q_{max})$ = damped oscillatory function of q_{max}
 extrema at every zero of $F(q)$, $\sin(qr)$
 in limit $q_{max} = \infty$ equals $\rho(r)$

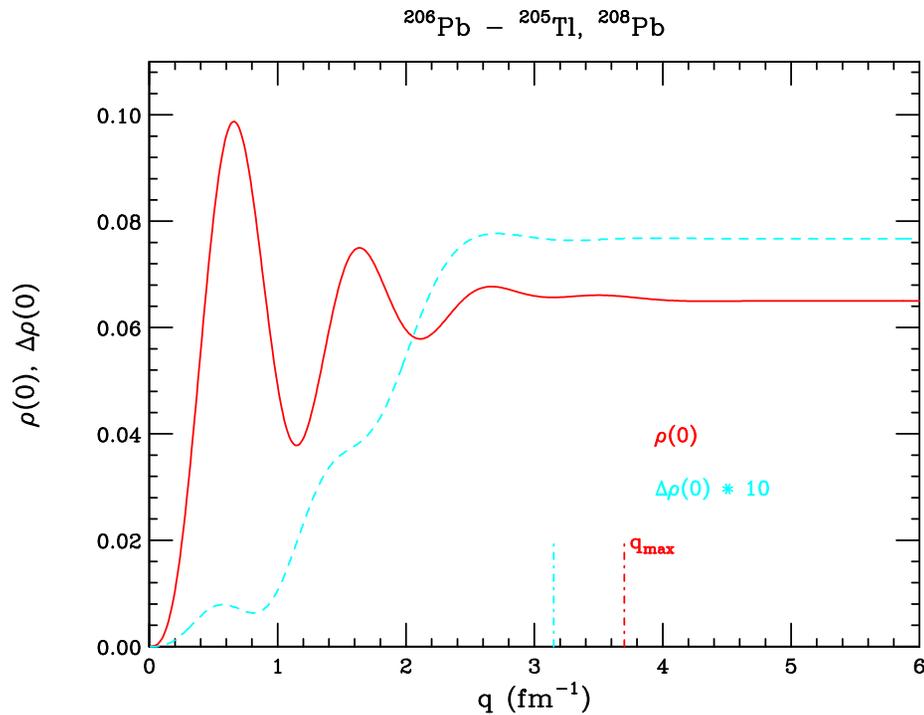
difference last minimum/maximum = uncertainty

→ model-independent estimate of $\delta\rho$ → benchmark for $\delta\rho$

Illustration: $r = 0$ (most difficult case)

$$\rho(0, q_{max}) = \dots \int_0^{q_{max}} F(q) q^2 dq$$

Example: Pb (red curve)



difference last mini/maxi $< 1\%$

Major assets of Direct Fourier Transform

bias-free determination of $\delta\rho$, transparent relation data - ρ , benchmark $\delta\rho$

$F(q)$ most useful to get insight, only in the end need solution Dirac eq.

Data of low q_{max} for exotic nuclei: DFT not possible

must rely on model densities

must aim at *integral* properties such as rms , c , z , ...

Discuss first: rms-radius $R = \langle r^2 \rangle^{1/2}$

the integral quantity of interest

most often quoted

can compare to radii from electronic/muonic atoms

needed to convert isotope shifts to radii

particularly relevant for unstable nuclei

discuss partly in PWBA (more transparent), in practice use DWBA

R strongly correlated with half-density radius c , discuss only R

Determination of R is deceptively easy:

$$F(q) = 1 - q^2 \langle r^2 \rangle / 6 + q^4 \langle r^4 \rangle / 120 - \dots$$

with $F(q)$ obtained as in DFT

apparently model-independent way to get R

± everybody uses it (at least conceptually)

but it does not work!!

Insights (in part not so recent [5])

1. Expansion of $F(q)$ in q^2 has very small convergence radius, $\sim 6/R^2$
2. $F(q)$ does not correspond to a density, Fourier transform diverges
3. Higher moments always contribute

if go to small enough q_{max} :

then finite size effect too small
measure only the "1" in $1 - q^2 R^2/6$

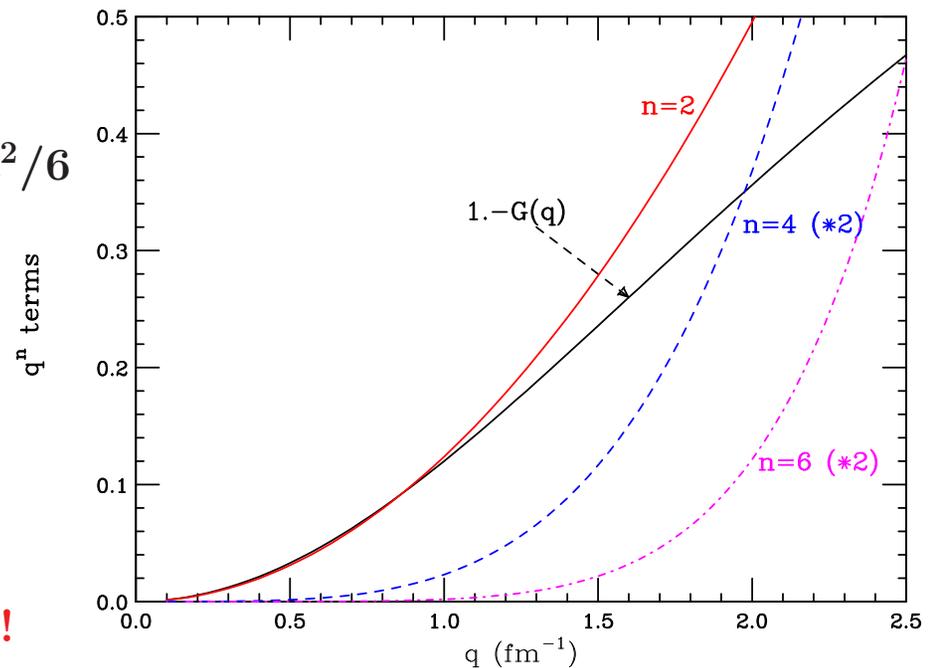
if q_{max} is larger for small δR :

then $\langle r^4 \rangle$ important
but $\delta \langle r^4 \rangle$ affects δR

get smaller $\delta \langle r^4 \rangle$ through higher q_{max}
get same problem with $\langle r^6 \rangle$,

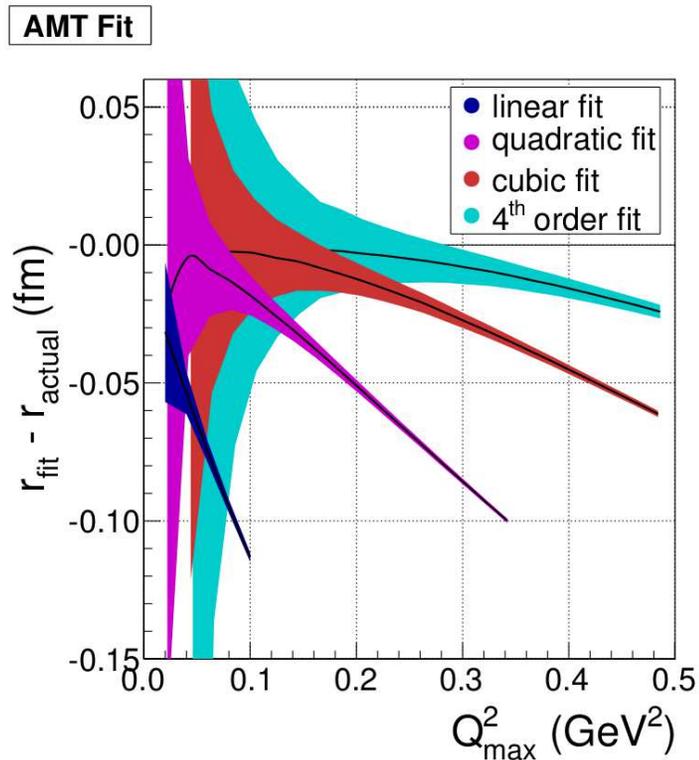
known since a decade for p, but ignored!

Problems with proton rms-radius have illustrated this point



Quantitative demonstration: Kraus *et al.* [6]

generate $F(q)$ -data from known density, extract R from fit of $F(q)$
find systematic differences input- $R \leftrightarrow$ output- R



Expansion in powers of q^2 useless!

Consequence: need model densities/physics to (implicitly) fix higher moments

for resulting problems (model dependence) see below

Important consideration: which q 's are important?

most often ignored

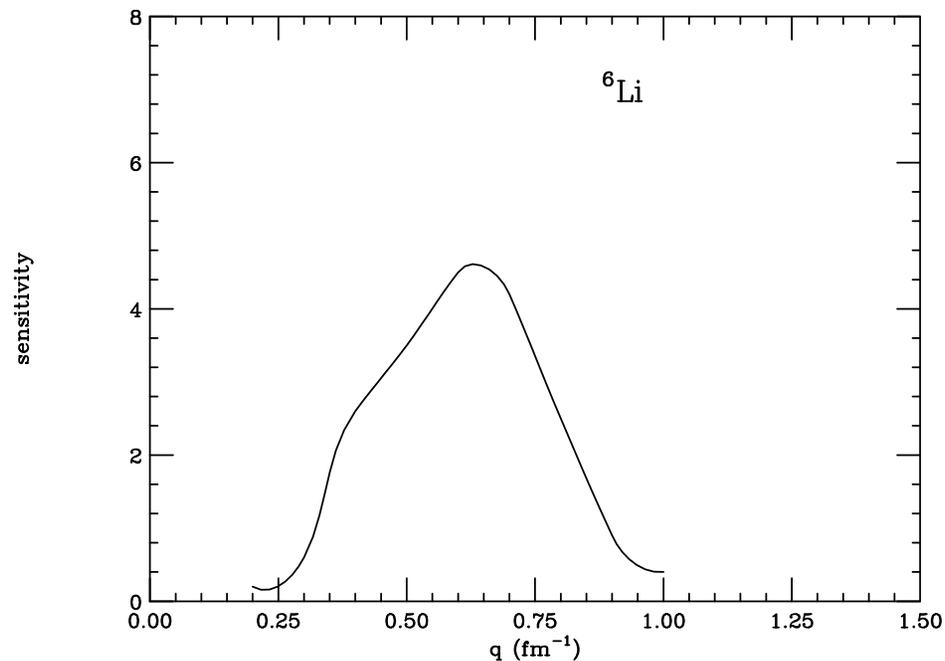
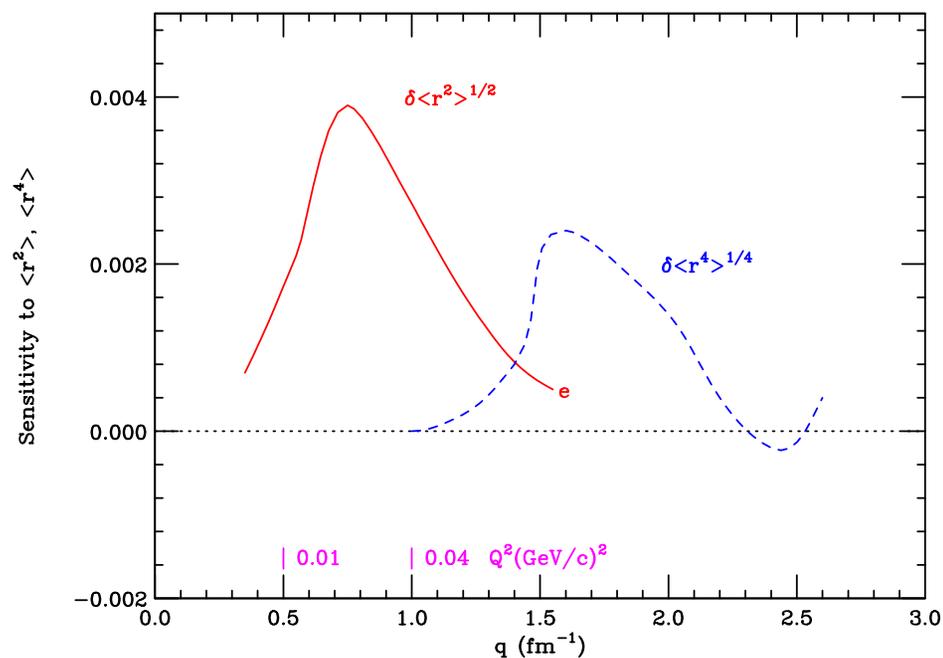
standard answer "small q " not good enough

Quantitative study: notch test [7]

change data by (say) 1% in narrow region around q_0

refit data

plot changes of R as function of q_0

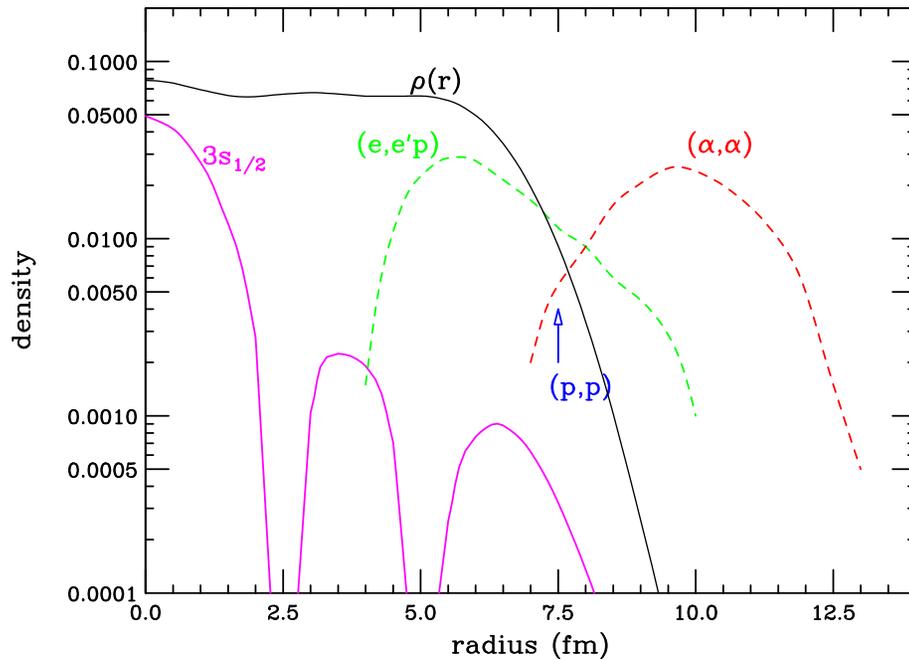


For nuclei heavier than proton:

maximum sensitivity to R at q where $F(q) \sim 0.7 - 0.8$

Important consideration: which r do matter?

mainly relevant for comparison with *other* probes aiming at ρ 's and R 's
all too often ignored



(e,e) = only probe sensitive to *all* r

strongly interacting probes = strongly absorbed probes
only sensitive to very large r
can never determine an *rms*-radius!
but are regularly used to do so

Not everything simple with (e,e) either: R dominated by large r , $\int ..r^4 dr$

Importance of large- r tail [8]

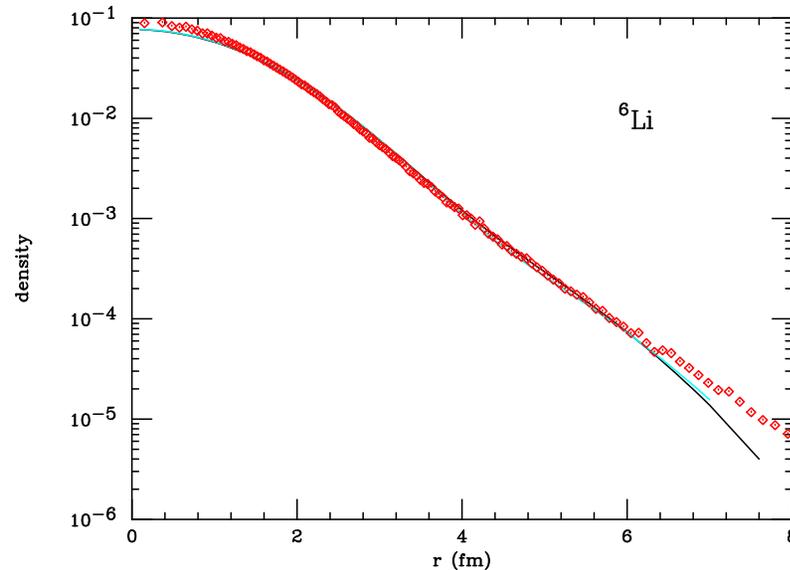
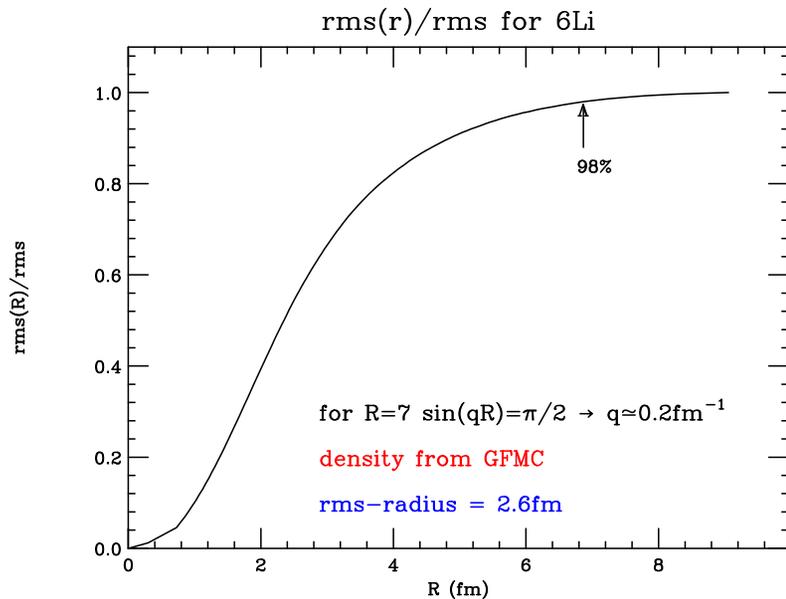
most often ignored in analyses using models
perhaps even worse in 'model-independent' analyses (see above)
parameterizations allow for too much flexibility of ρ at large r

Large- r particularly important for rms-radius and low SE!

Example: ${}^6\text{Li}$

study $[\int_0^{r_{cut}} \rho(r) r^4 dr / \int_0^\infty \rho(r) r^4 dr]^{1/2}$ as function of cutoff r_{cut}

note: rms-radius of ${}^6\text{Li} = 2.6\text{fm}$

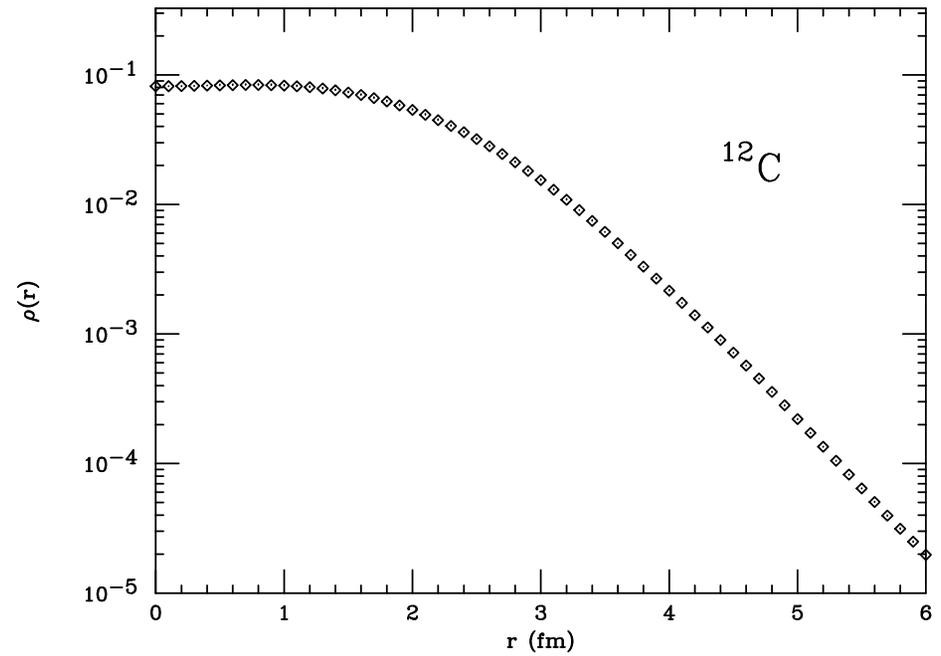
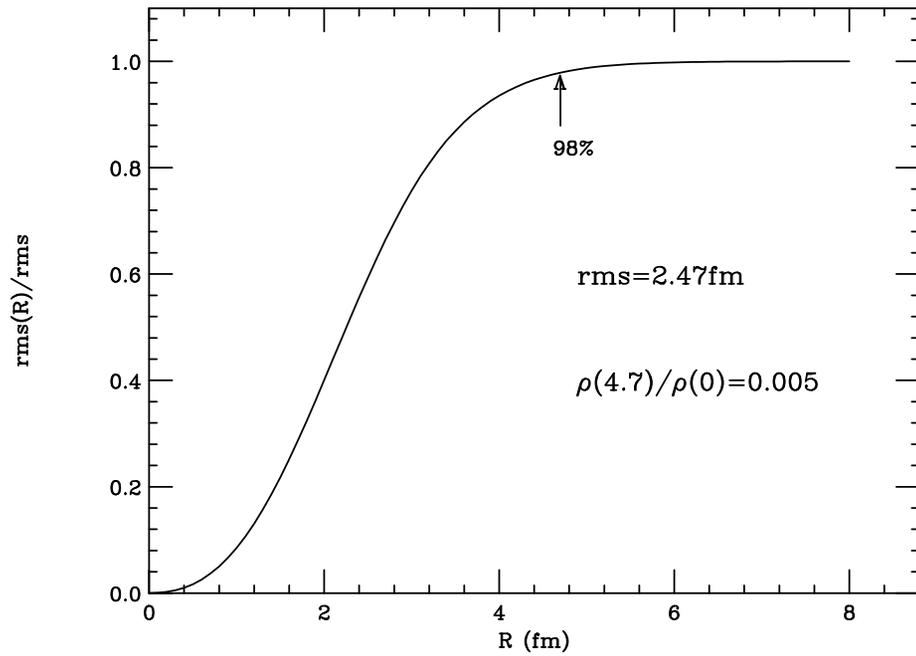


if want R to $<2\%$ must know ρ at $r > 7\text{fm}$ where $\rho/\rho(0) \sim 10^{-4}!!$

Only a problem for ${}^6\text{Li}$ with long tail?

similar, though less pronounced for ${}^{12}\text{C}$

SE=16MeV!

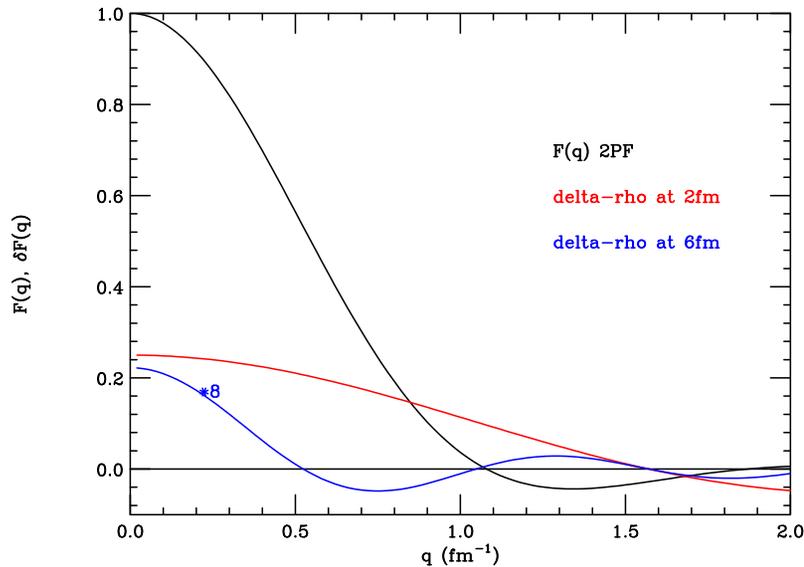


Qualitative demonstration large $r \leftrightarrow$ low q

density = 2-parameter Fermi, 1/2-density radius $c = 4\text{fm}$

add local change of $\rho(2\text{fm})$, or

add local change of $\rho(6\text{fm})$, same rms-radius change



large r produce curvature of $F(q)$ at very low q

affects extrapolated slope $F(q = 0)$ which yields rms-radius

Consequence: must be VERY careful about $\rho(\text{large } r)$!

parameterization \longrightarrow implicit assumption on $\rho(\text{large } r)$

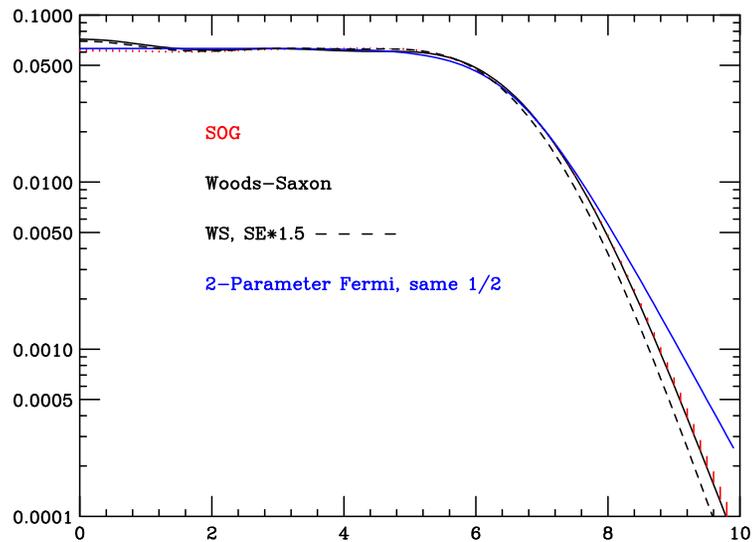
Example: compare tails for *Pb*

experimental SOG density

Woods-Saxon density with exp. separation energy

Woods-Saxon with increased separation energy

2-parameter Fermi with same 1/2-density radius



tail of 2PF leads to 2% larger rms-radius

Lesson for analysis using model- ρ 's:

must constrain large- r behavior using physics knowledge

fall-off of $\rho(r)$ dominated by least-bound protons

What know on $\rho(r)$ at large r ?

dominated by least-bound proton, radial wave function = Whittaker function

$$R(r) \sim W_{-\eta, l+1/2}(2\kappa r)/r \text{ with } \kappa^2 = 2\mu E/\hbar^2, \eta = (Z-1)e^2/\hbar\sqrt{\mu/2E}$$

depends only on removal energy

presumably available from accurate mass-measurements

easy to calculate

overall normalization (asymptotic normalization) generally not known \rightarrow shape

Shape automatically given if calculate $R(r)$ using Woods-Saxon potential

solution of Schrödinger equation for sensible potential

\rightarrow physical shape of asymptotic $R(r)$

Note: Harmonic oscillator has *wrong* asymptotic shape

has been used often in past

has caused many problems (*e.g.* Darmstadt rms-radii)

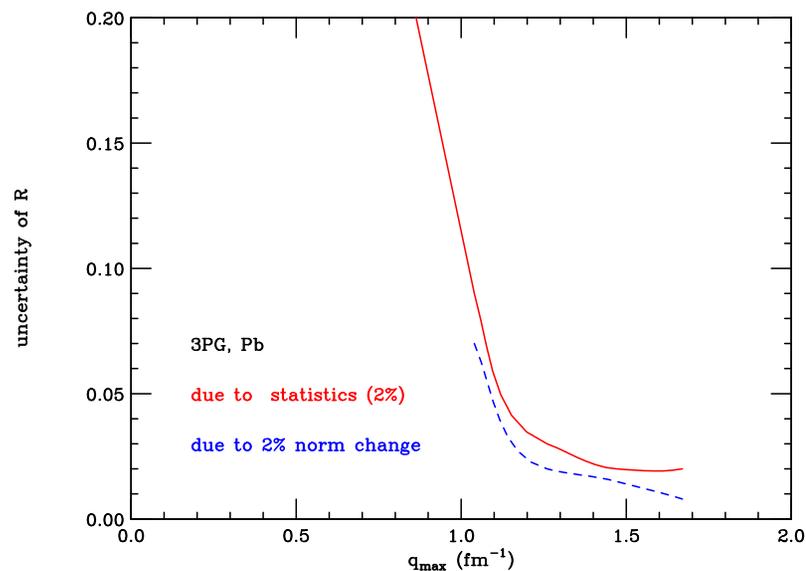
keep away from it

Important to use physics information on large- r behavior !

Should have been done in past analyses of data for stable nuclei

Essential insight for determination of R: role of normalization

generate cross sections for Pb using 3PG, $\pm 2\%$
in realistic q -range
analyze as function of q_{max} using 3PG
study effect of normalization change



Absolute normalization of cross sections very important

must make great effort to determine

Replace absolute normalization by floating of data?

traditional: people float data, as absolute norm difficult to determine
main purpose: get good-looking χ^2

Problems:

1. Ignores $\sim 50\%$ of effort of experimentalists

often great effort made to get *absolute* normalization, don't want to throw away

2. When floating *extrapolate* via fit to $q = 0$

become very sensitive to q -dependent systematic errors

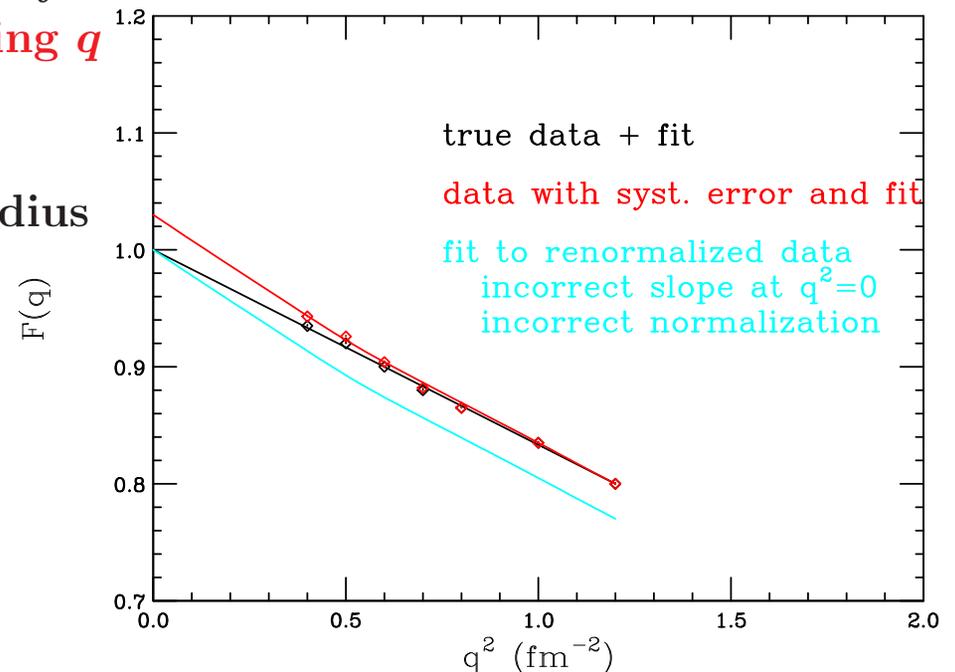
errors clearly increasing with decreasing q

otherwise data taken at lower q

extrapolation increases effect

affects in particular extracted *rms*-radius

→ High premium on *absolute* σ 's



Alternative look at accuracy needed

can estimate from sensitivity-curves discussed above
no need for detailed quantitative study

maximum sensitivity to R at q where $F(q) \sim 0.75$

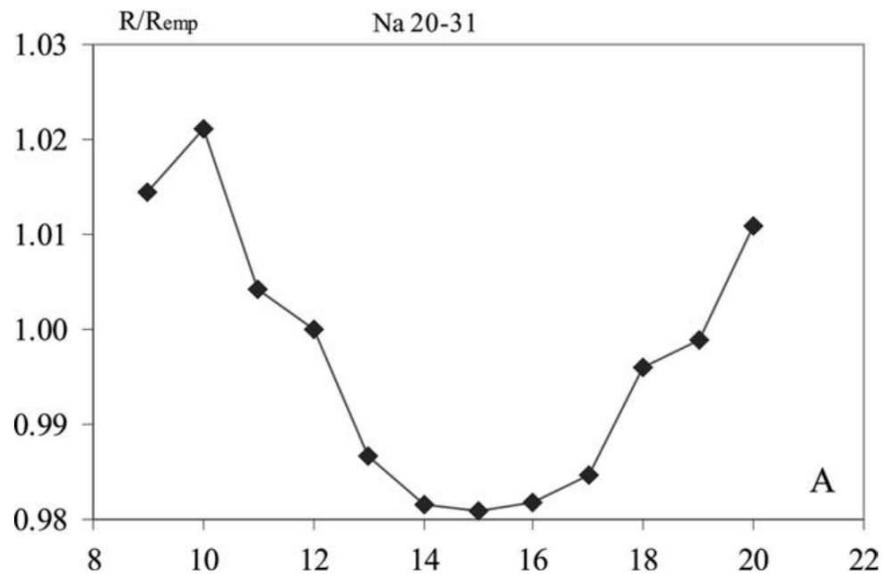
→ finite-size effect in $\sigma \sim 50\%$

If measure σ to $\pm 1\%$

measure R^2 to $\sim 2\%$

measure R to 1%

then can see evolution of R of interest



Absolute σ 's to $\pm 1\%$ for stored ions??

already difficult enough for stable isotopes
already there dominating experimental error
presumably not possible

For stored ions

measure cross section *ratios* to stable isotope
 \pm straightforward as these much more abundant
modest increase of time

Requirement

accurate (relative) luminosity monitor
same response for different isotopes
should be feasible

- elastic scattering at very low q (recoils near 90°)
- or exploiting atomic electrons (?), nuclear Bremsstrahlung

But remember

results only as good as result for reference nucleus
(re-)analyze reference nucleus data respecting above insights

Conditio sine qua non: precise *ratio*-measurements!

Special difficulty for unstable isotopes: halo, for small SE

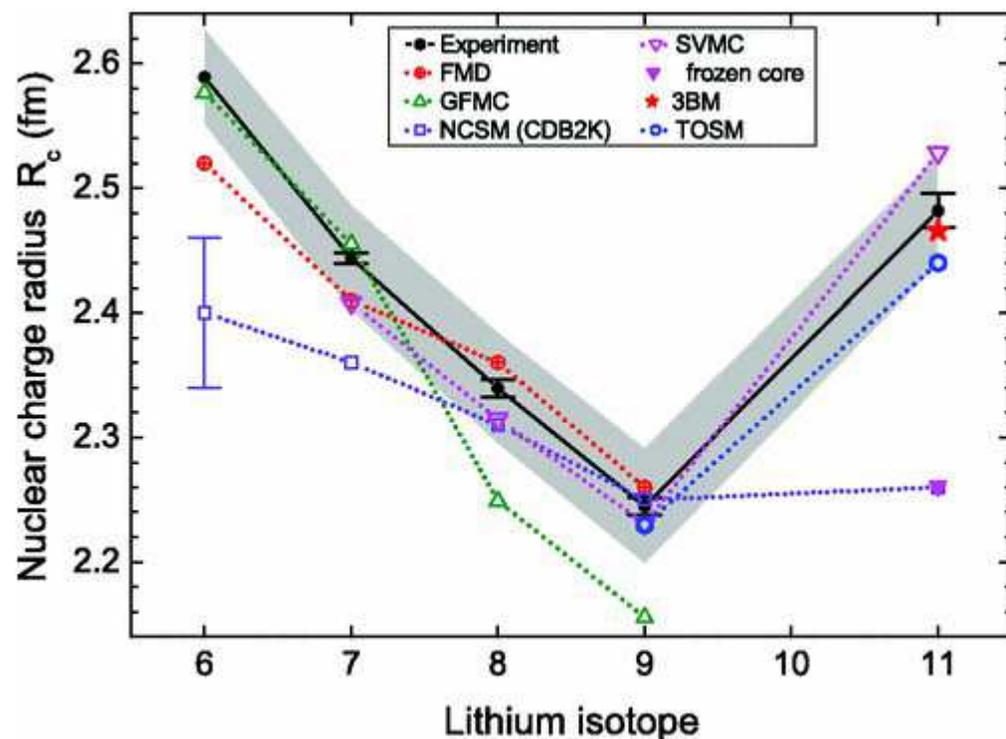
Example: isotope shifts for Li

measured by laser spectroscopy with stored ions

Nörtershäuser *et al.*[9]

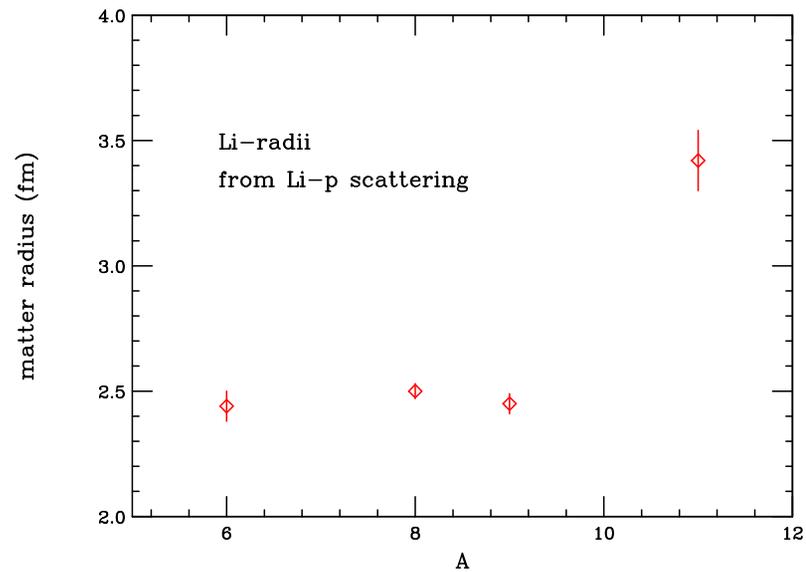
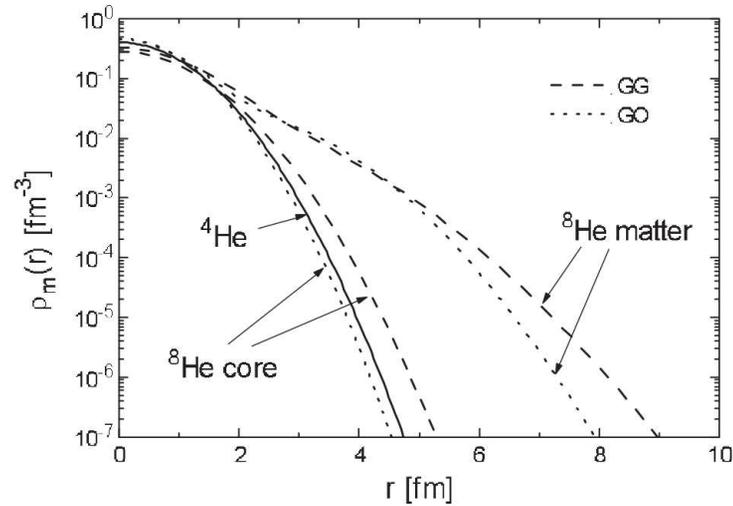
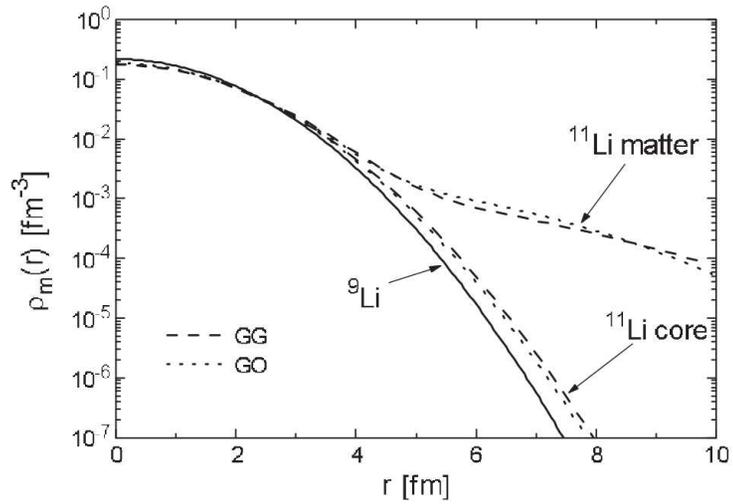
^{11}Li = Borromean nucleus ($2n$, ^{10}Li unbound)

$2n$ -separation energy only 369KeV



Jump in R due to large- r tail from outermost proton

Extreme example of tail-importance: matter radii [10]



Densities from GeV/N p-nucleus scattering in inverse kinematics

How to proceed in presence of such a halo?

standard analysis with model for ρ not promising
density has too 'complicated' a shape
large- r tail affects R way too much

More promising approach

tail due to *one* shell with low separation energy
core in general more 'normal'

Parameterize ρ as model+tail

model = standard 2PF, 2PG, ...

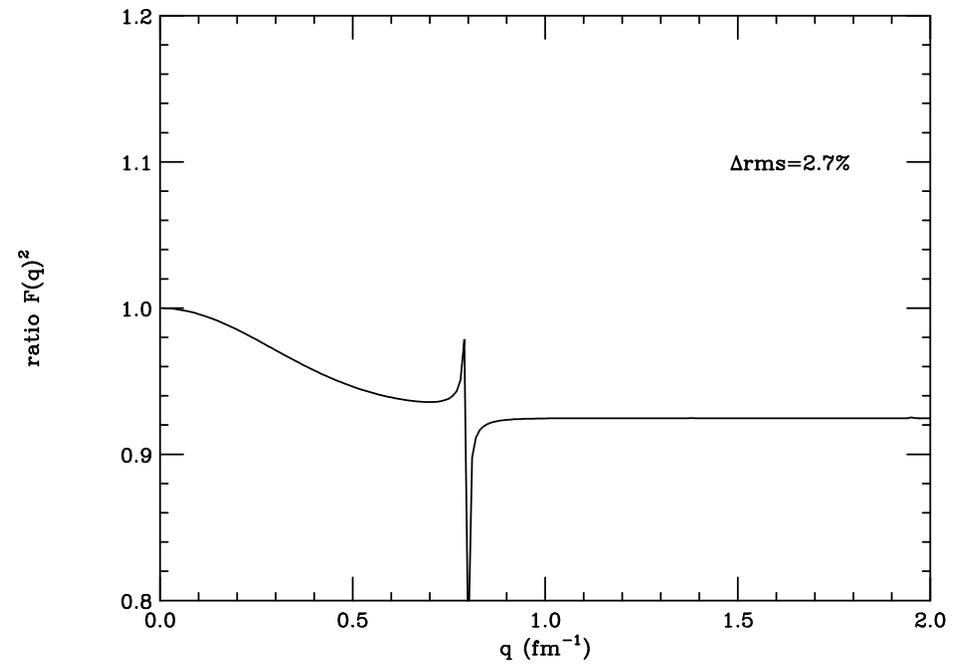
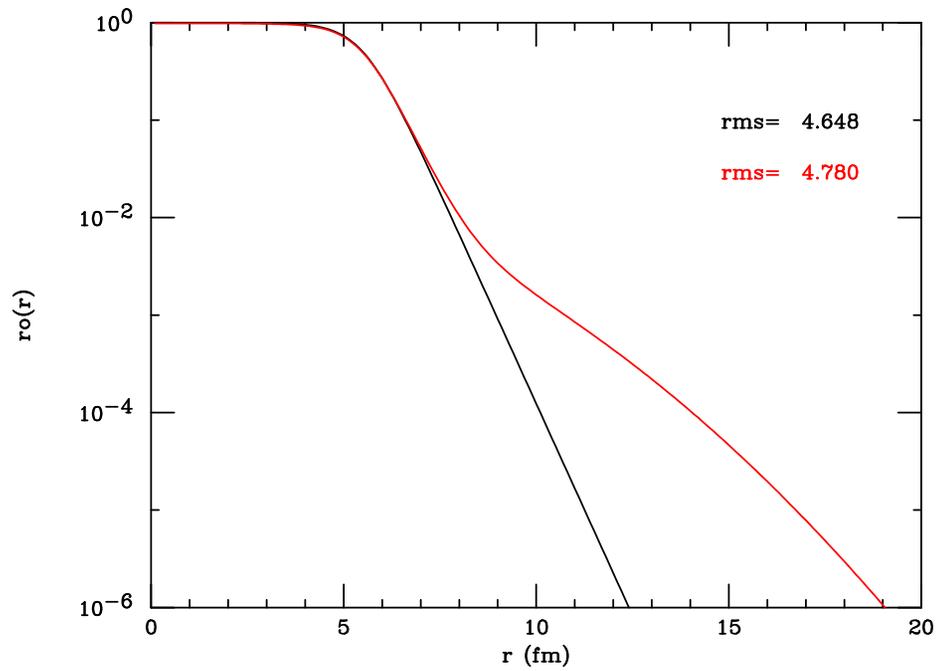
tail = calculated in WS potential using experimental separation energy

Example

add tail to Pb density

similar to findings from above example Li

compute change in cross section



Find (as expected)

main effect of halo at *very* low q
 need accurate data at these low q 's
 feasible from ratio measurements
 main issue: precision, *not* luminosity

Determination of z

Lessons from the "early days"

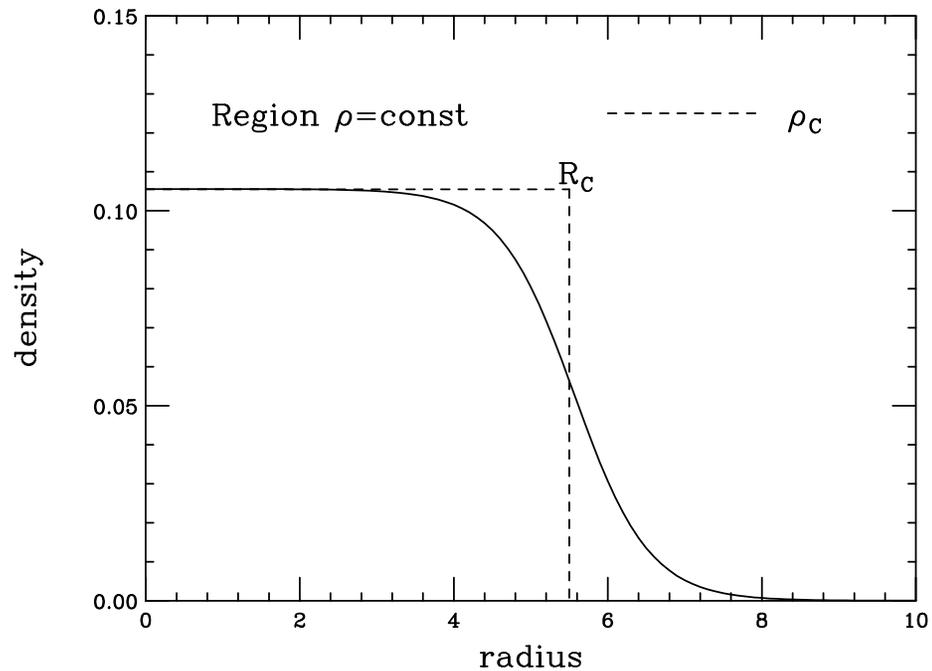
z obtained from $F(q)$ in 2. diffraction maximum

Understanding in PWBA:

consider $\rho(r)$ as result of folding *

fold uniform density ρ_c with gaussian $G(r-r')$ of width z

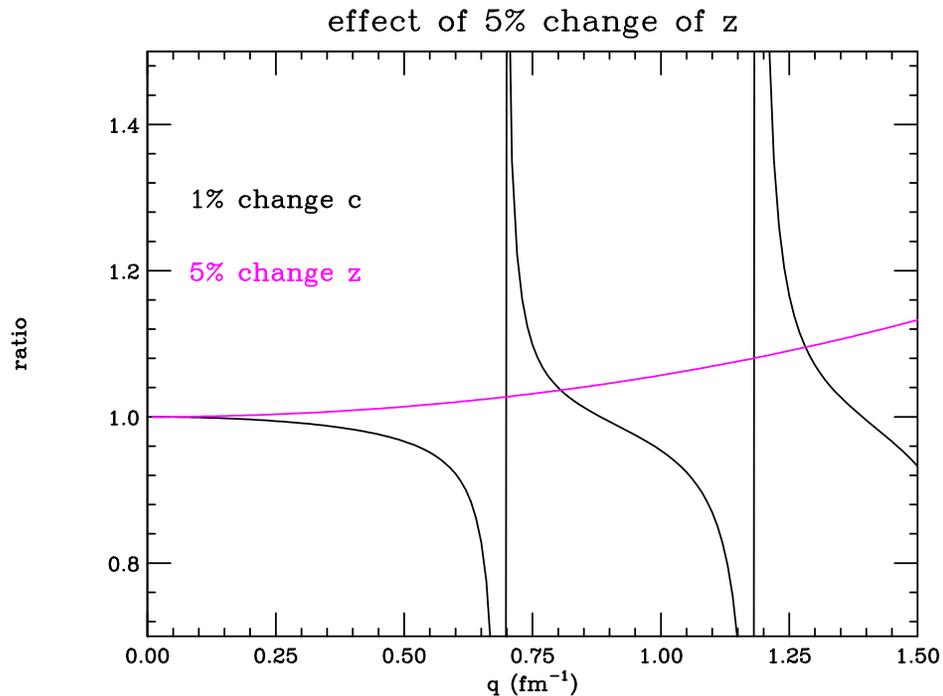
$$\rho(r) = \rho_c(r) * G(r)$$



Result for Fourier transform FT

$$F(q) = FT(\rho_c) \cdot FT(G)$$

Effect of $z \sim$ multiplication of $F_c(q)$ with FT of G i.e. $e^{-q^2 z^2}$



region of 2. diffraction maximum sensitive to z
shape of $F(q)$ for Δc and Δz different

but: effect not large

third diffraction maxi would be better

Particular interest in z for unstable nuclei

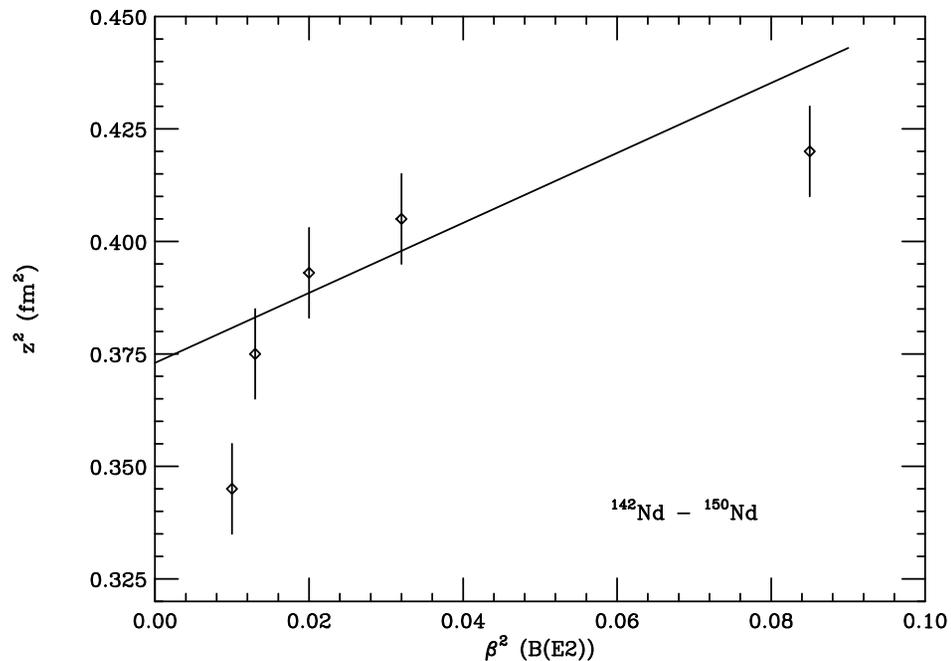
larger for halo-nuclei?

large for deformed nuclei

$$z_{eff}^2 = z_0^2 + \frac{3}{4\pi^3}\beta^2 c^2 + O(\beta^3)$$

Particularly useful for series of isotopes

Example: Nd-isotopes [1]



could be best way to get at deformation, is only way for I=0 nuclei

note: need rather *precise* z 's

Challenge beyond L: energy-resolution

near 1. minimum, inelastic F's peak

have typically several 10% of probability

must separate elastic/inelastic scattering

exotic nuclei typically far from magic

low excitation energies

also needed for data on transition form factors

Problem: long targets

required to get high L

makes small $\Delta E/E$ difficult

Advantage of fixed-target experiments

resolution needed $\sim E^*$

$\Delta E/E \sim 10$ times easier

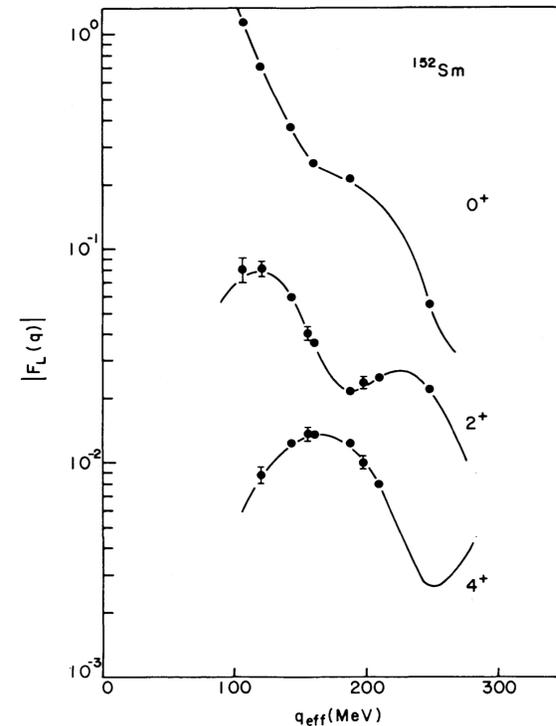
than for collider experiments

Most realistic solution

magnetic spectrometer

deflection \perp to scattering plane

decouples coordinate and E



Additional considerations → precise z

a priori correlation $c \leftrightarrow z$ not too strong
nevertheless: more precise c (or R) allows for more precise z
independent knowledge on R -type observable → smaller δz

For stable (reference)-isotope

add muonic X-ray information
available for *many* nuclei
yields *very* precise moments of ρ
for larger Z also yield higher moments of type $\langle r^4 \rangle$, $\langle r^6 \rangle$
(Barrett moments, to be precise)

For precise isotope shifts of R -type observables

include also electronic isotope shifts
occasionally easier to measure than (e,e) cross sections
particularly for exotic nuclei
need to know e-wave function at origin
 $\Psi(r = 0)$ calibrated using stable isotopes

General rule: add as much independent R -type info as possible

also helps to mitigate problems with absolute normalization of σ

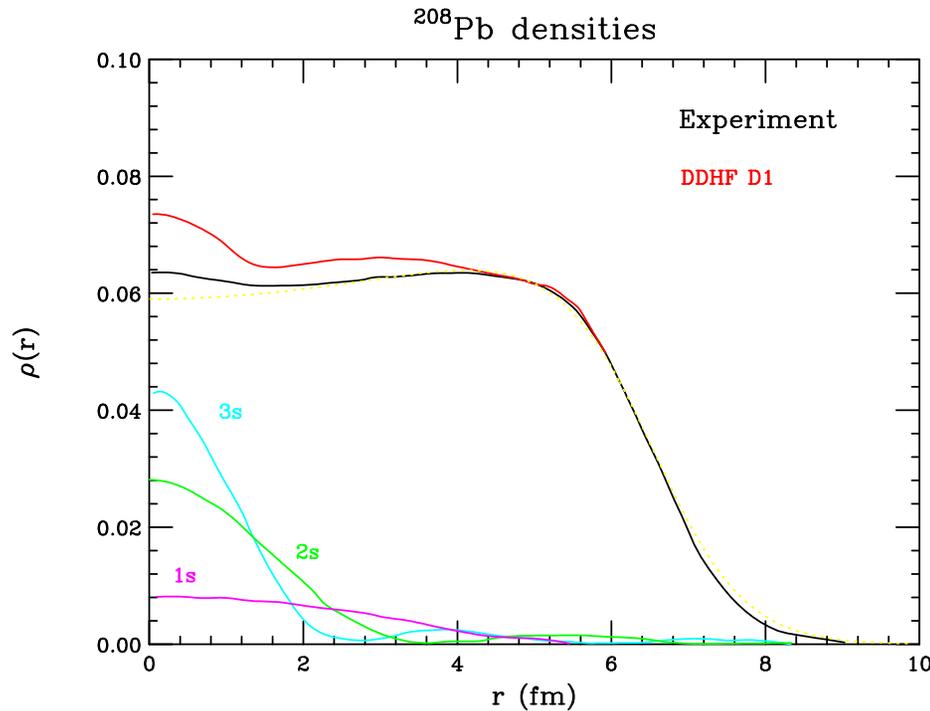
Central depression w

large for heavy nuclei away from valley? Bubble nuclei??
predicted for super-heavy, proton-rich nuclei

Lesson from 'early days'

not clear if w describes central depression
biggest effect of $(1 + wr^2/c^2)$ occurs in tail of $\rho(r)$ at *large* r

Example: ^{208}Pb : $w = 0.33 \rightarrow$ depression?? Not there in modelindependent ρ [11]



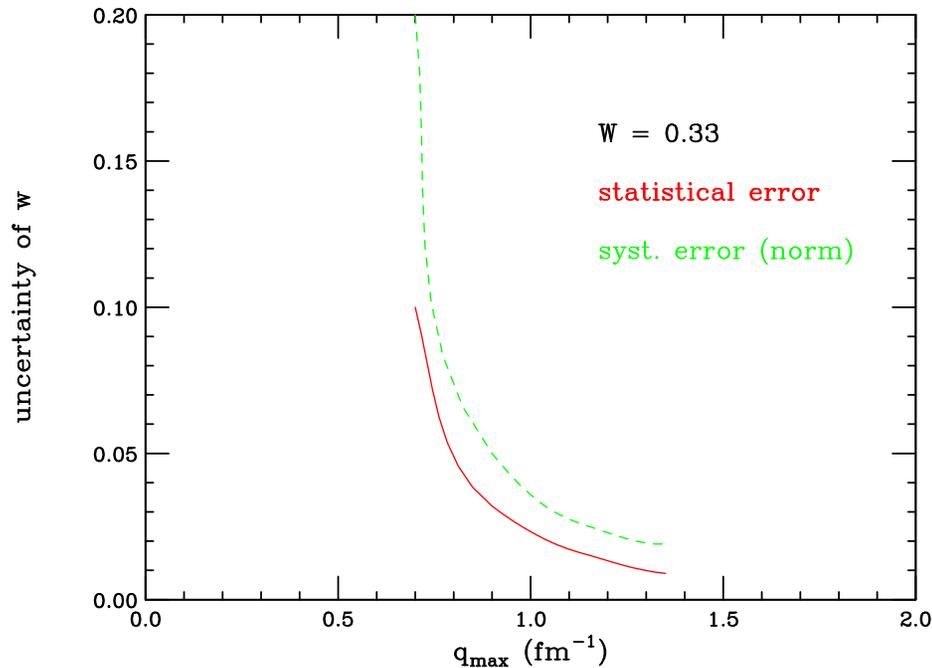
w essentially compensates too fast fall-off of Gaussian distribution

Sensitivity to w with 3PG (phase-shift calculation)

use data with 2% stat. errors, up to q_{max}

fit with 3PG

derive uncertainty of w , syst. error due to norm change



w with precise data to $q = 1.5 \text{ fm}^{-1}$ measurable

but: represents depression, or shape $\rho(r)$ at large r ?

My conclusion: w mainly used to affect large- r tail, *not* central depression

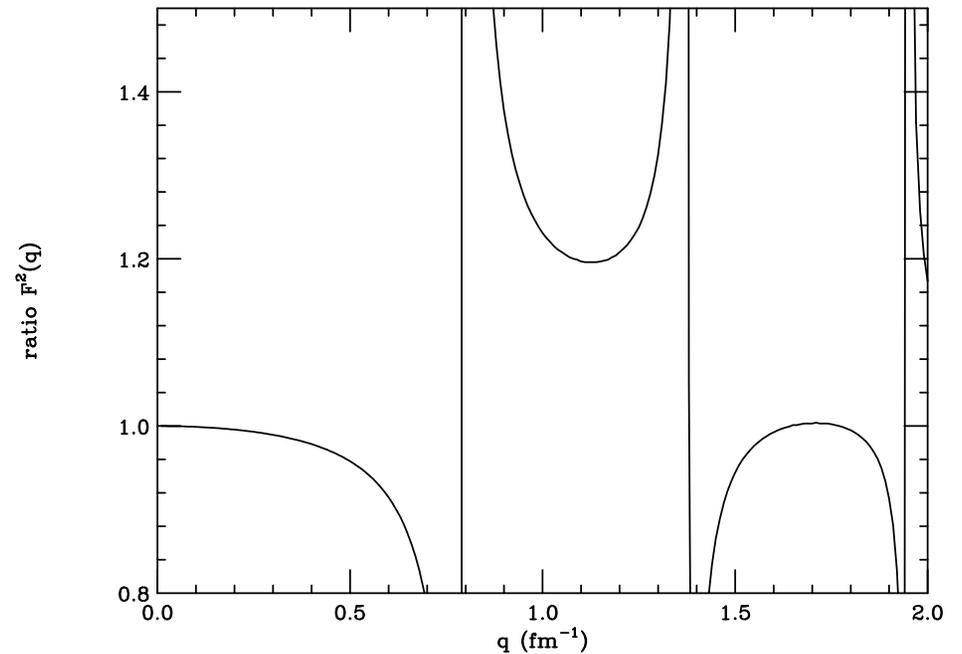
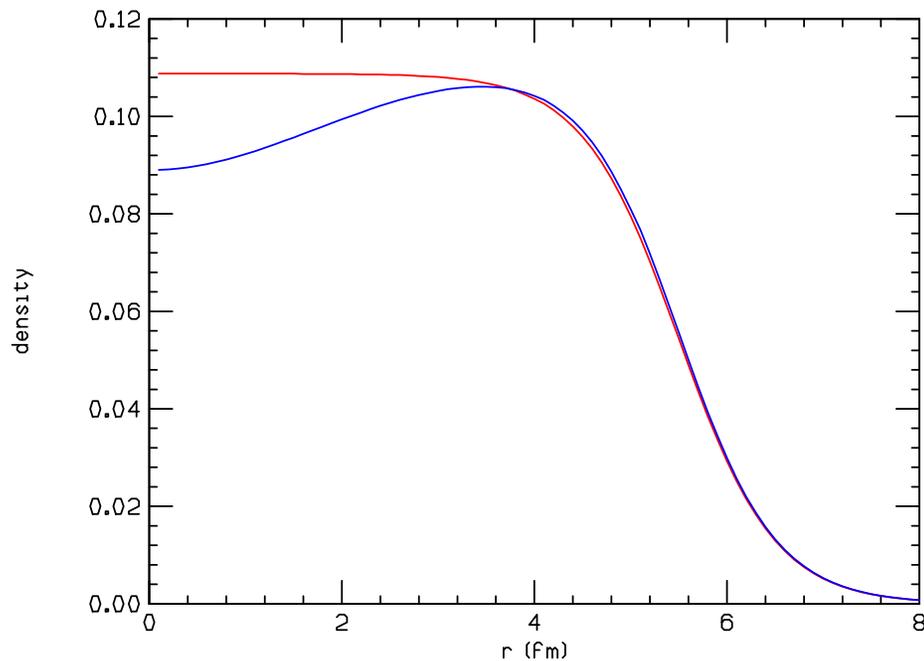
→ keep away from $w \sim$ central depression

w affects *both* small and large r

Better parameterization of eventual central dip:

$$\begin{aligned}\rho(r) &= 2PF - ae^{-r^2/b^2} && \text{for } r \leq c \\ &= 2PF && \text{for } r \geq c\end{aligned}$$

Result for $\rho(r)$ and F^2 -ratio in PWBA



Fairly characteristic change in F^2

increase of F^2 in 2. maximum, little change in third maxi
comparatively little change below first mini, above second mini

Measurable if can reach third maxi of F^2

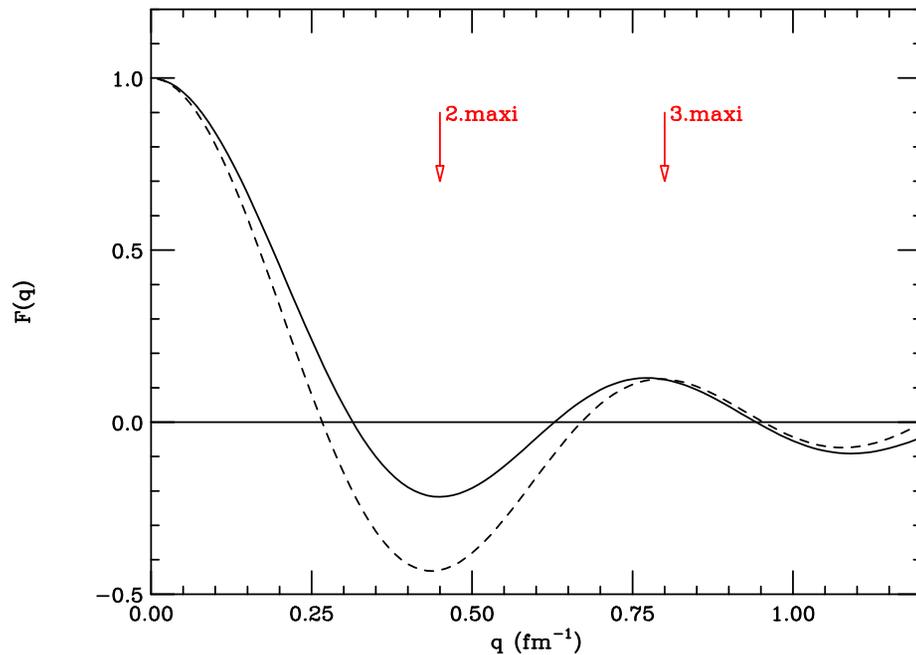
Qualitative understanding from q -dependent term in integrand for $F(q)$

compare for $r_1 = 0.3r_0$

$$\sin(qr_0)/qr_0$$

$$0.8 \sin(qr_0)/qr_0 + 0.2 \sin(qr_1)/qr_1$$

20% dip, 1/3 radial size



Confirms: main signal of central dip = enhancement in second maxi of $F(q)$

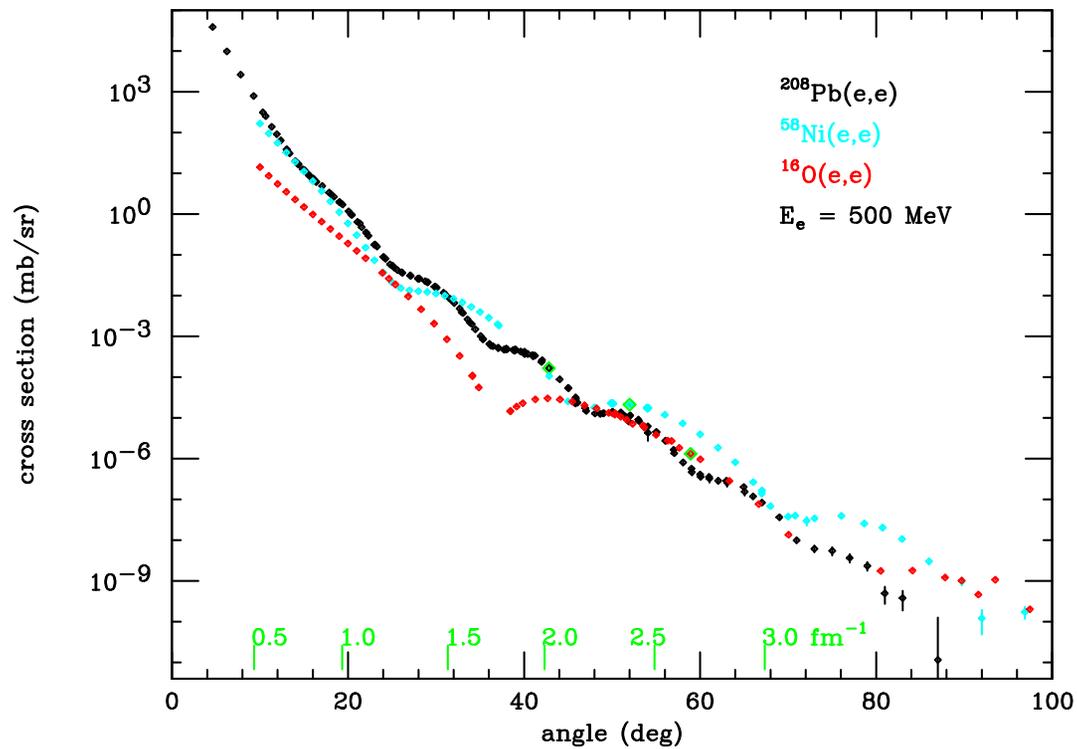
should be within reach of experiment

Light vs. heavy nuclei

above studies mainly performed for $A=208$
how about lighter nuclei?

Fall-off of $d\sigma/d\Omega$

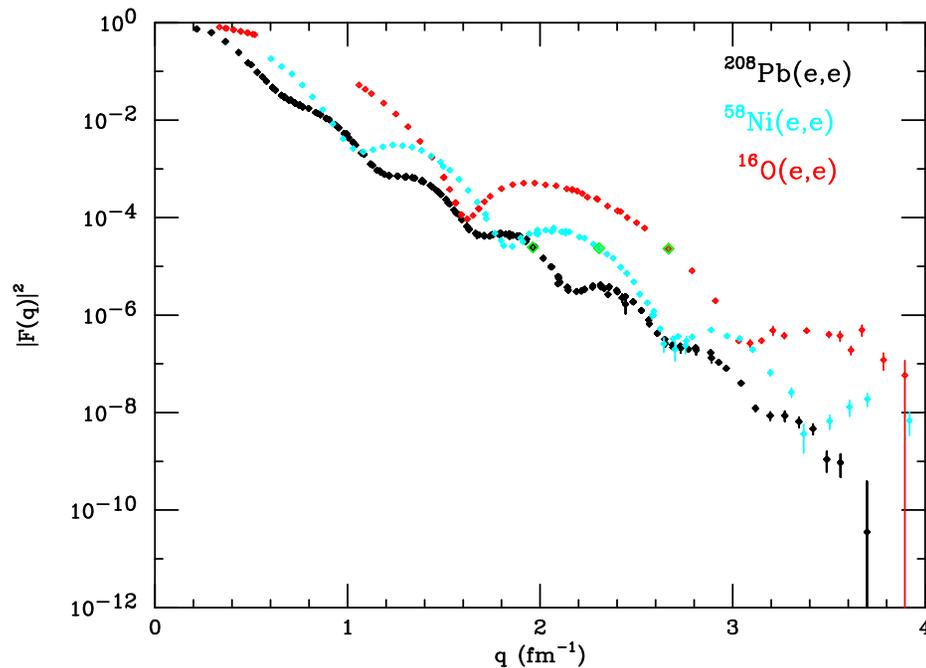
quite similar for different A
somewhat slower for small A



Relevant for accuracy of $\rho(r)$:

smallest $F(q)$ measured, not q or number of diffraction maxima

$$\rho(r, q_{max}) = \int_0^{q_{max}} F(q) \frac{\sin(qr)}{qr} q^2 dq$$



For small A

need larger q to get same, small $F(q)$

need larger q for same completeness error \rightarrow somewhat lower σ

determination of c , z , 'w' somewhat more difficult

Need/can determine further integral quantities?

can determine *more* than c , z , dip?

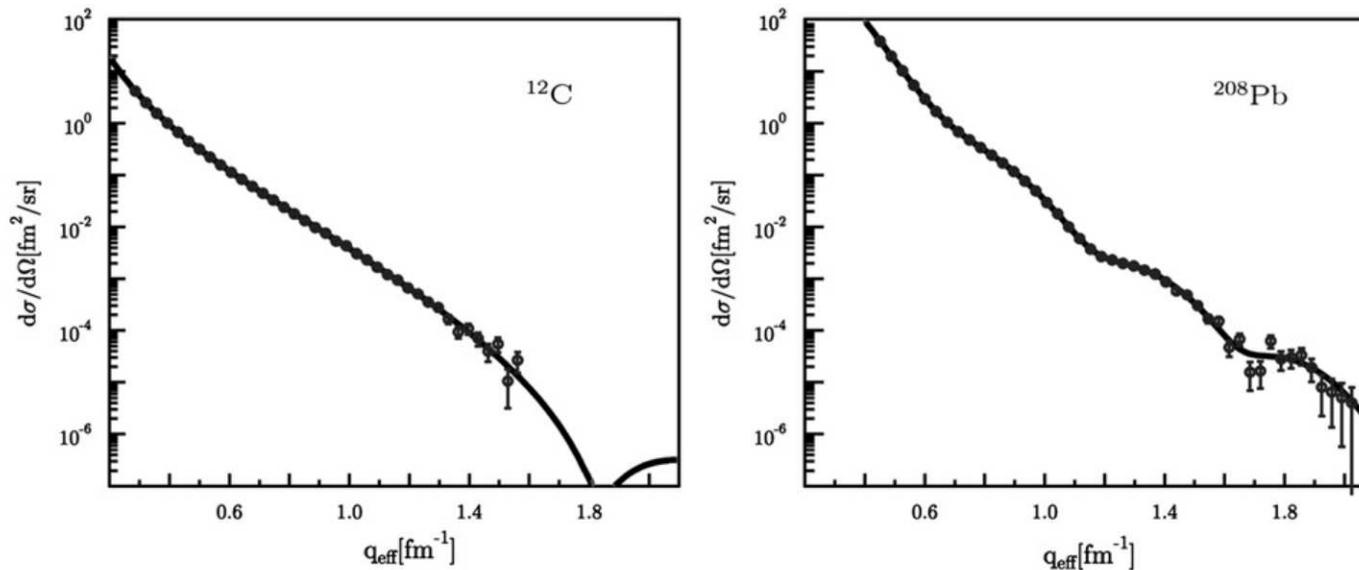
depends strongly on achievable luminosity L

depends on quality of e-spectrometer

evaluation beyond scope of present study

Use result of study for ELISE (Antonov *et al.* [12])

assume $L = 10^{28}$, 100msr, 4 weeks



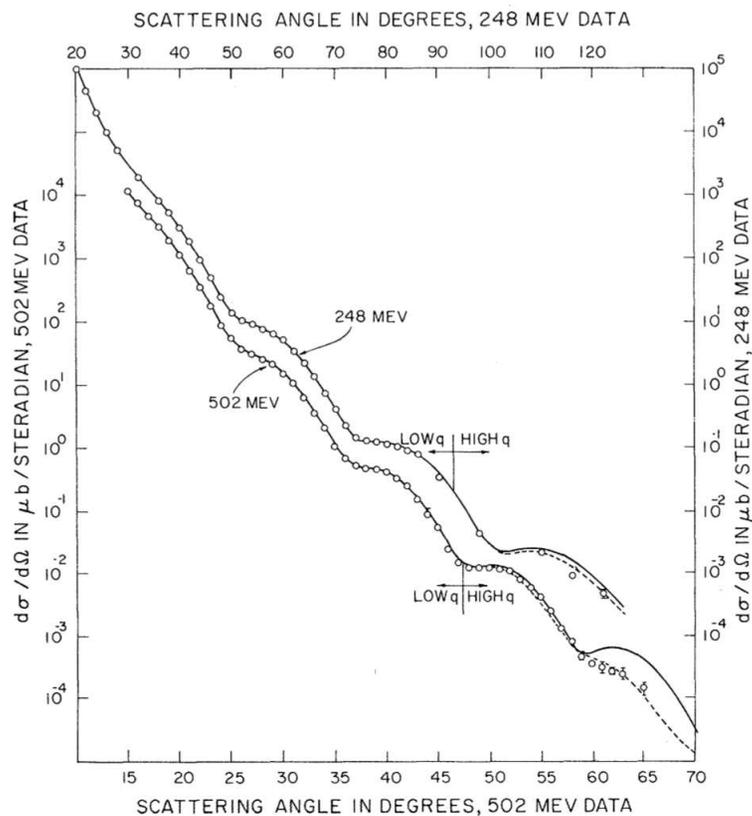
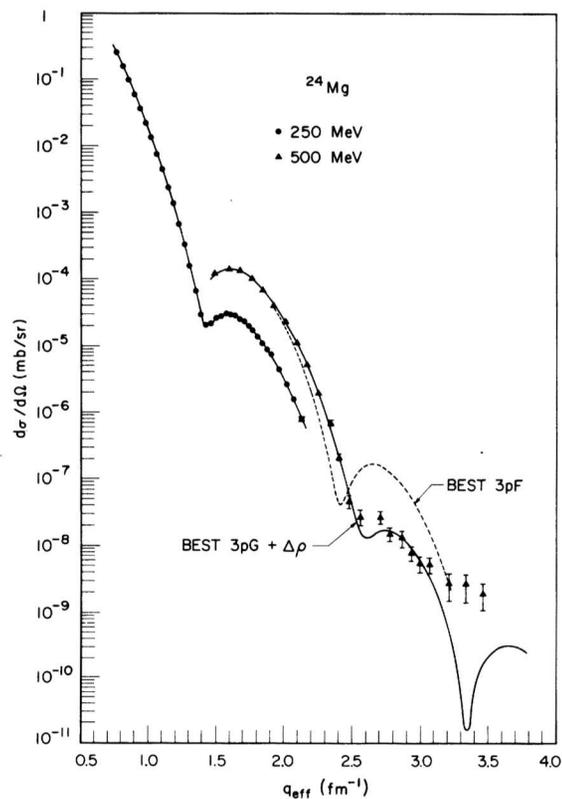
$q_{\text{max}} \sim 1.5(2.0) \text{ fm}^{-1}$ seems to be limit for $A=12(208)$

Implications for quantities that can be determined

experience from past fits to stable isotope data:

3PF, 3PG typically flexible enough for $q_{max} \sim 2 fm^{-1}$

Examples: Mg, Pb [13, 14]



Data from ETIC-type collider can be described using 3pF, 3pG, ...

For exotic nuclei can determine R, z , central dip, large- r tail

(e,e) and trapped ions: conclusions

great interest + potential

Limitations due to achievable luminosity

interpretation of data with models for $\rho(r)$
extraction of integral properties: R , z , dip, halo
look feasible given present estimates for L

Precise results for *rms*-radius R

need precise (absolute) cross sections
only achievable via *ratio* measurements to stable isotopes
requires precise monitors for (relative) luminosity

For precise extraction of R

use realistic large- r tail of ρ (not properly done in past)
reanalyze data for (stable) reference nucleus

For nuclei far from valley

must expect halo's, *i.e.* long larger tail
parameterize as halo ($R(r)$ from WS potential) + model-density
emphasizes interest in measuring SE via precise masses, (e,e'p)

References

- [1] J. Heisenberg, J.S. McCarthy, I. Sick, and M.R. Yearian. *Nucl. Phys. A*, 164:340, 1971.
- [2] F. Lenz. *Z. Phys.*, 222:491, 1969.
- [3] I. Sick. *Nucl. Phys. A*, 218:509, 1974.
- [4] I. Sick. *Lecture Notes in Physics*, 236:137, 1985.
- [5] I. Sick. *Phys. Lett. B*, 576:62, 2003.
- [6] E. Kraus, K.E. Mesick, A. White, R. Gilman, and S. Strauch. *Phys. Rev. C*, 90:045206, 2014.
- [7] I. Sick. *J. Phys. Chem. Ref. Data*, 44:031213, 2015.
- [8] I. Sick and D. Trautmann. *Phys. Rev. C*, 89:012201(R), 2014.
- [9] W. Nörtershäuser, T. Neff, R. Sanchez, and I. Sick. *Phys. Rev. C*, 84:024307, 2011.
- [10] P. Egelhof et al. *Eur. Phys. J. A*, 15:27, 2002.
- [11] B. Frois et al. *Phys. Rev. Lett.*, 38:152, 1977.
- [12] A.N. Antonov et al. *Nucl. Instr. Meth. A*, 637:60, 2011.
- [13] G.C. Li, I. Sick, R.R. Whitney, and M.R. Yearian. *Nucl. Phys. A*, 162:583, 1971.
- [14] J. Heisenberg, R. Hofstadter, J.S. McCarthy, I. Sick, B.C. Clark, R. Herman, and D.G. Ravenhall. *Phys. Rev. Lett.*, 23:1402, 1969.