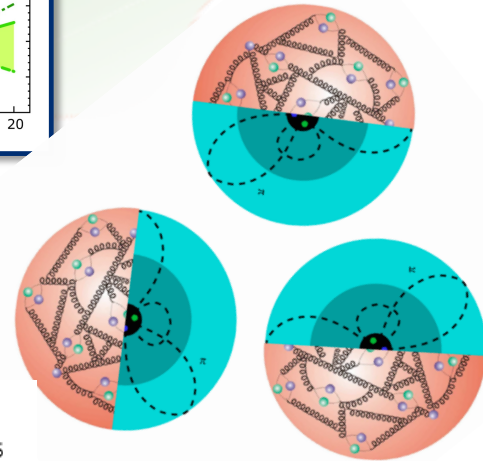
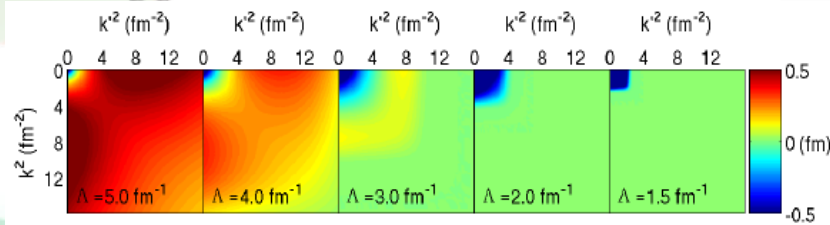
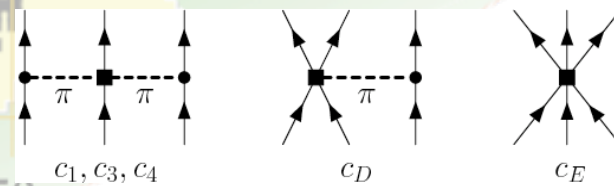
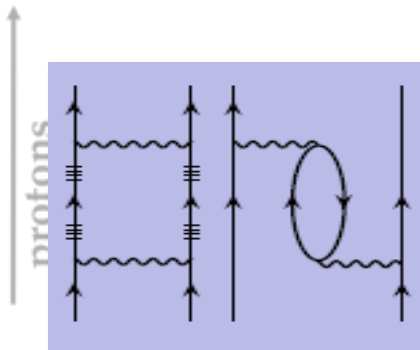
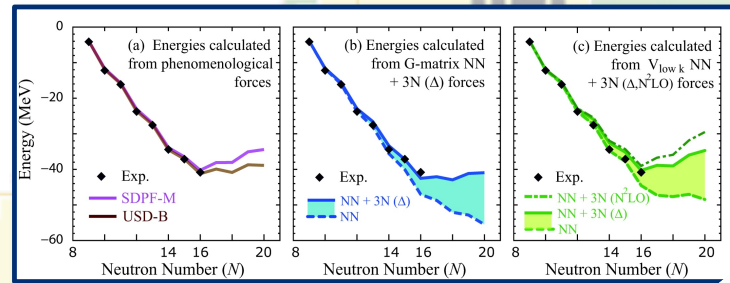


A Microscopic Approach to Shell Model

Jason D. Holt

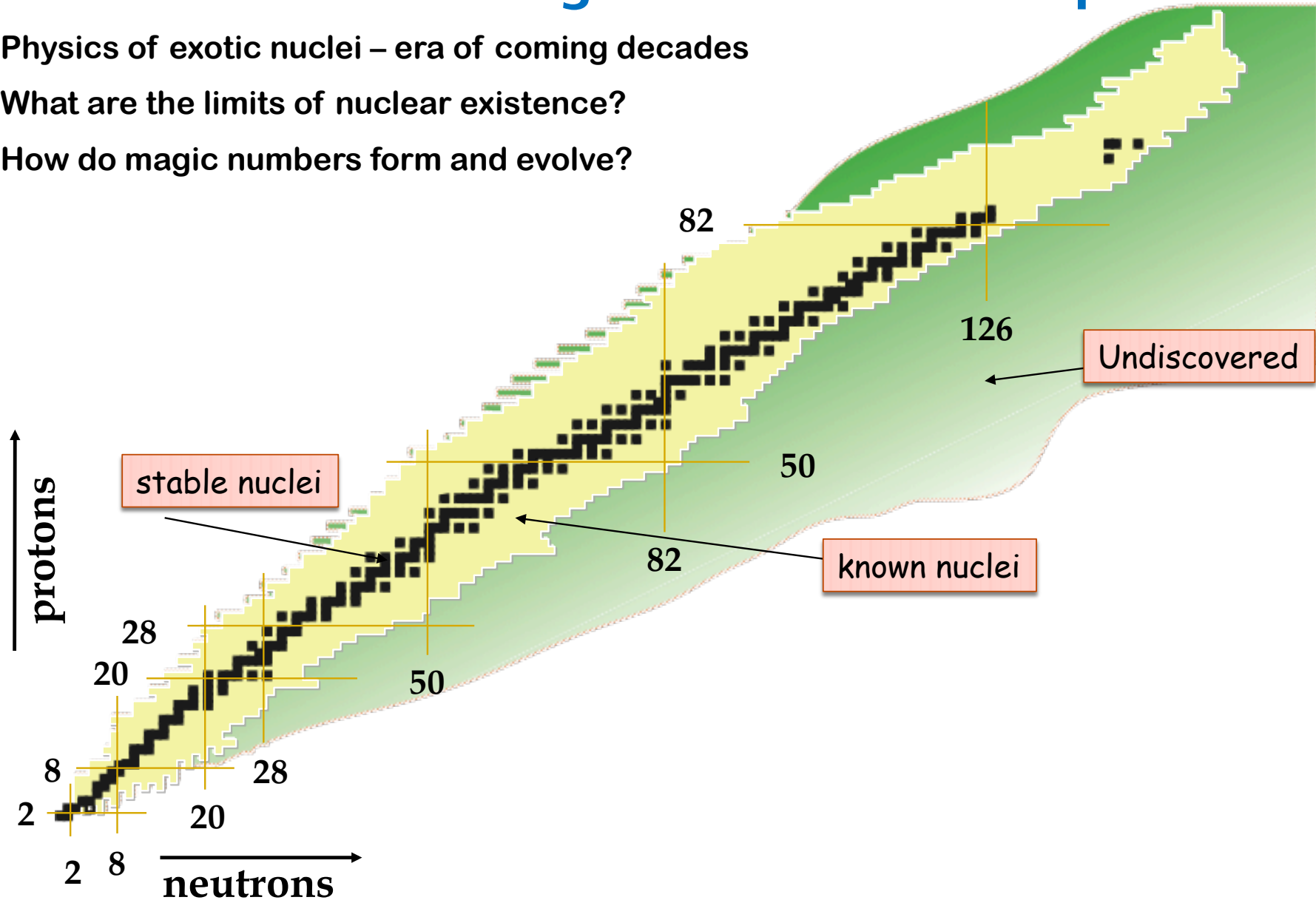


Drip Lines and Magic Numbers: The Evolving Nuclear Landscape

Physics of exotic nuclei – era of coming decades

What are the limits of nuclear existence?

How do magic numbers form and evolve?



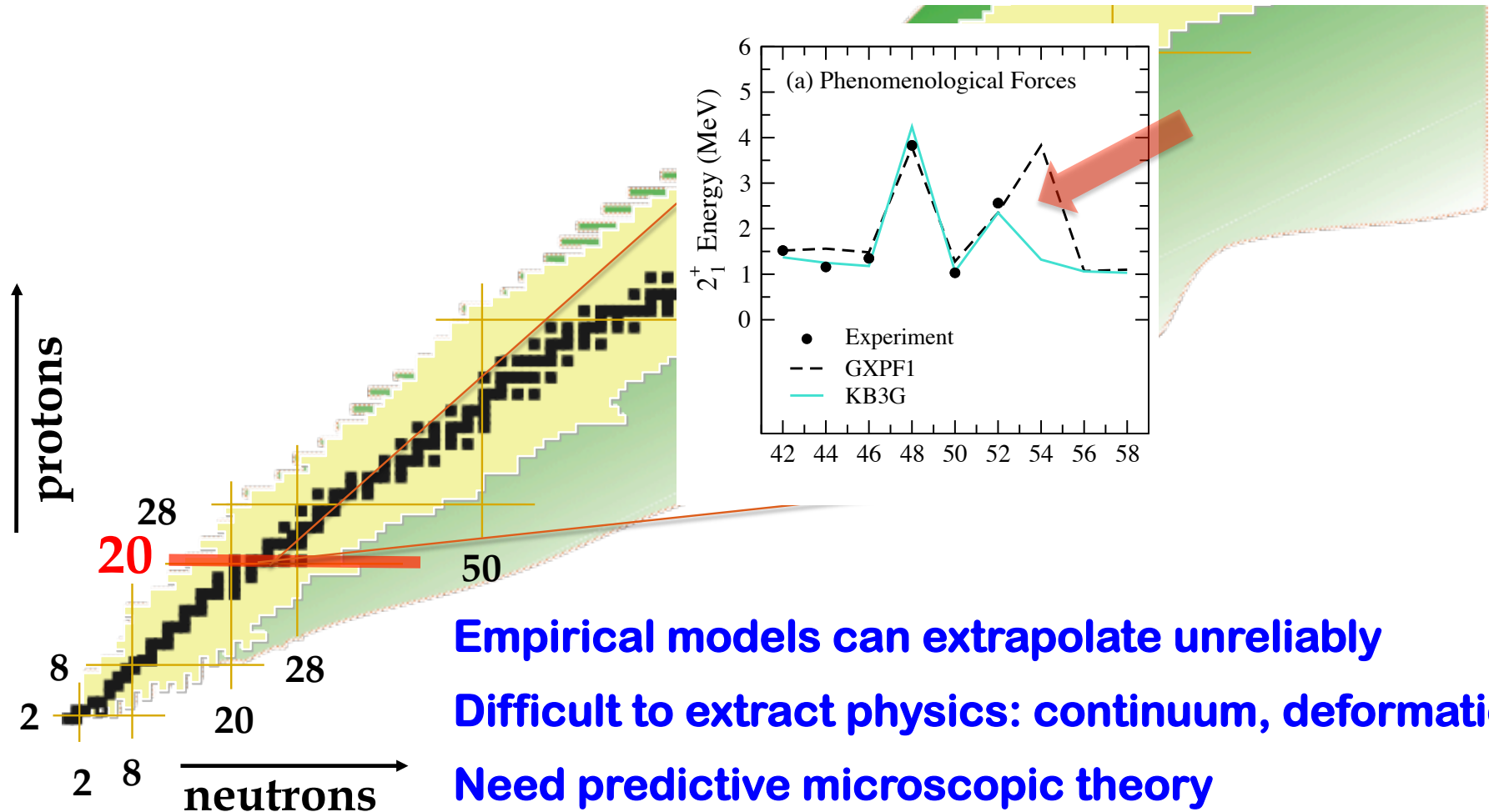
Drip Lines and Magic Numbers: The Evolving Nuclear Landscape

Physics of exotic nuclei – era of coming decades

What are the limits of nuclear existence?

How do magic numbers form and evolve?

N=54 magic number in calcium?



Empirical models can extrapolate unreliably

Difficult to extract physics: continuum, deformation

Need predictive microscopic theory

Drip Lines and Magic Numbers: The Evolving Nuclear Landscape

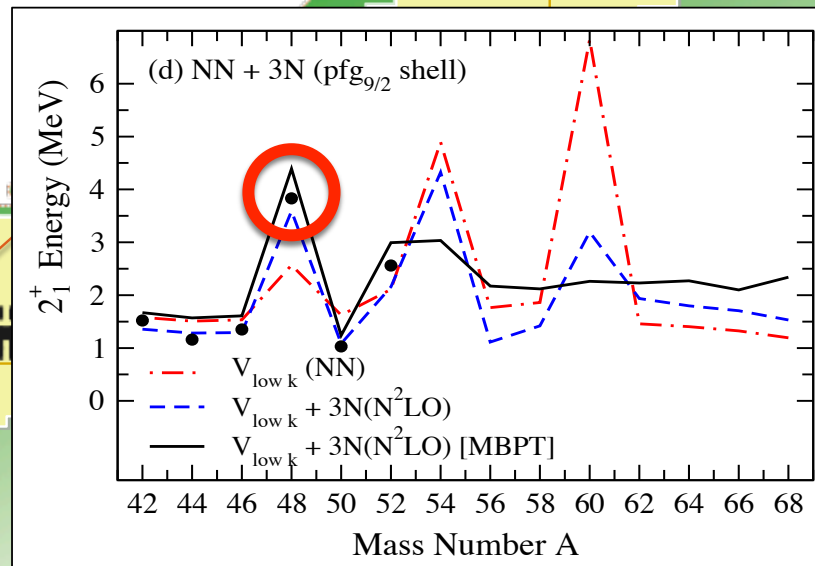
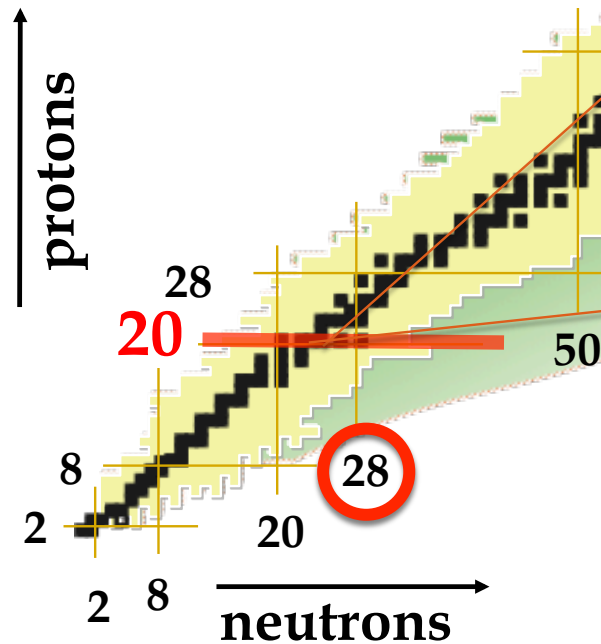
Physics of exotic nuclei – era of coming decades

What are the limits of nuclear existence?

How do magic numbers form and evolve?

N=28 magic number in calcium

Holt, Otsuka, Schwek,
Suzuki, arXiv:1009.5984



Provide a microscopic foundation for shell model

Familiarize with modern advances in microscopic theory

How these allow us to move towards predictive shell model

Drip Lines and Magic Numbers: The Evolving Nuclear Landscape

Physics of exotic nuclei – era of coming decades

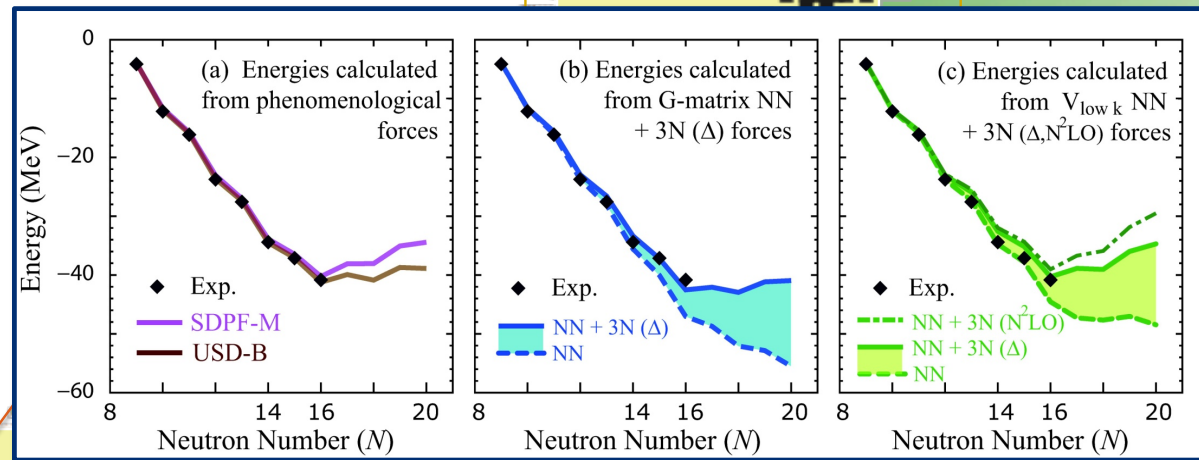
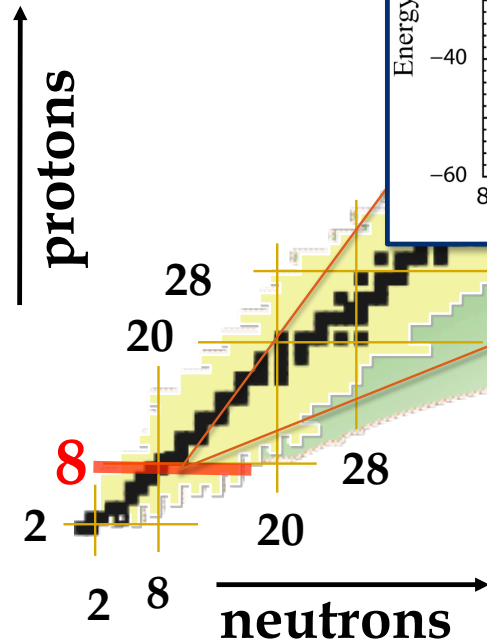
What are the limits of nuclear existence?

How do magic numbers form and evolve?

Otsuka, Suzuki, Holt, Schwenk, Akaishi, PRL (2010)

82

Heaviest oxygen isotope



Approaches to Nuclear Structure

“The first, the basic approach, is to study the elementary particles, their properties and mutual interaction. Thus one hopes to obtain knowledge of the nuclear forces. If the forces are known, one should, in principle, be able to calculate deductively the properties of individual nuclei. Only after this has been accomplished can one say that one completely understands nuclear structure...

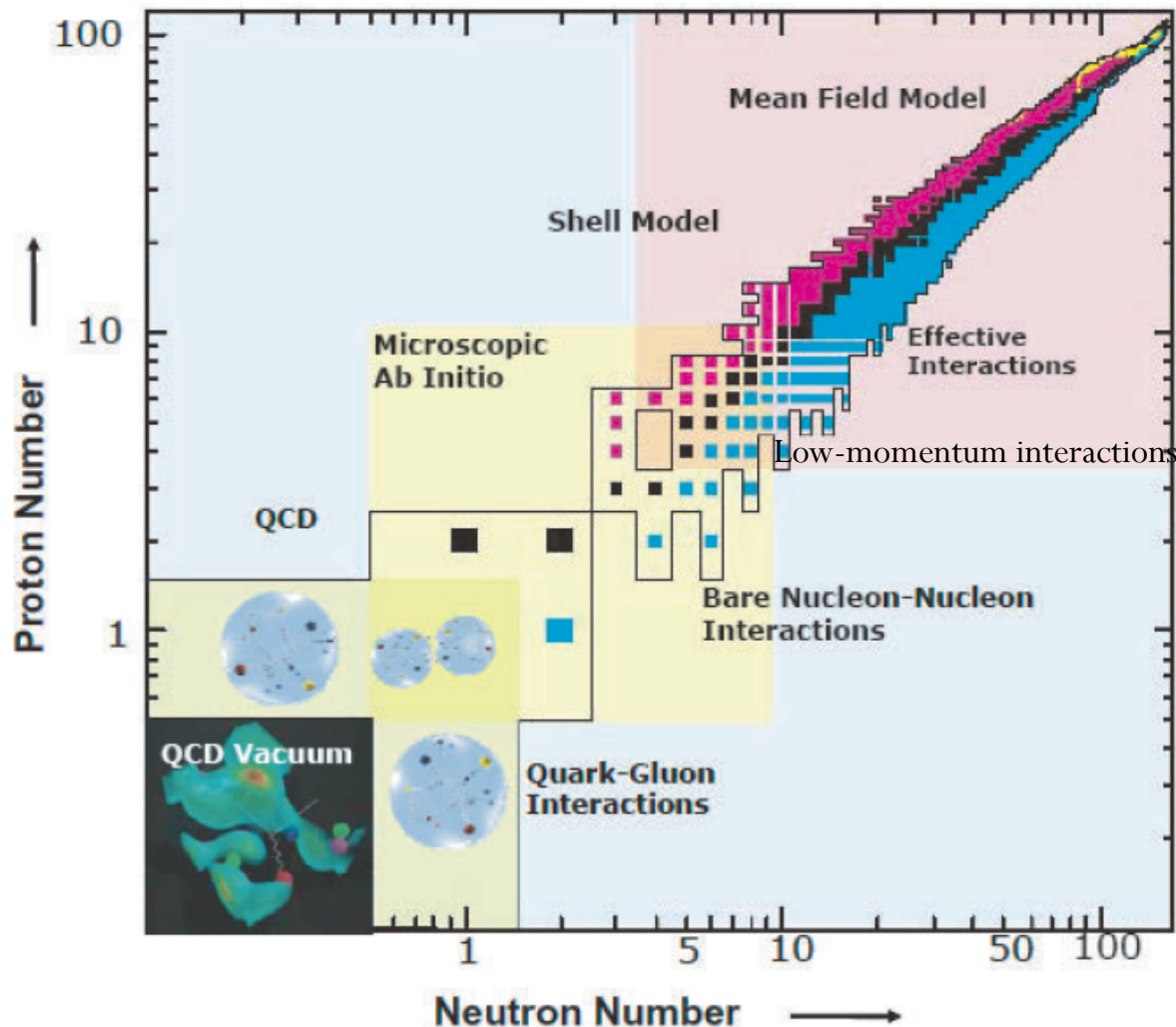
The other approach is that of the experimentalist and consists in obtaining by direct experimentation as many data as possible for individual nuclei. One hopes in this way to find regularities and correlations which give a clue to the structure of the nucleus... The shell model, although proposed by theoreticians, really corresponds to the experimentalist's approach.”

–M. Goepfert-Mayer, Nobel Lecture

Purpose of these lectures is to show how shell model can be based on the first approach!

The Challenge of Microscopic Nuclear Theory

To understand the properties of complex nuclei from elementary interactions



Two significant issues:

Interaction

Not well understood

Not obtainable from QCD

Too “hard” to be useful

Multiple scales

Many-body Problem

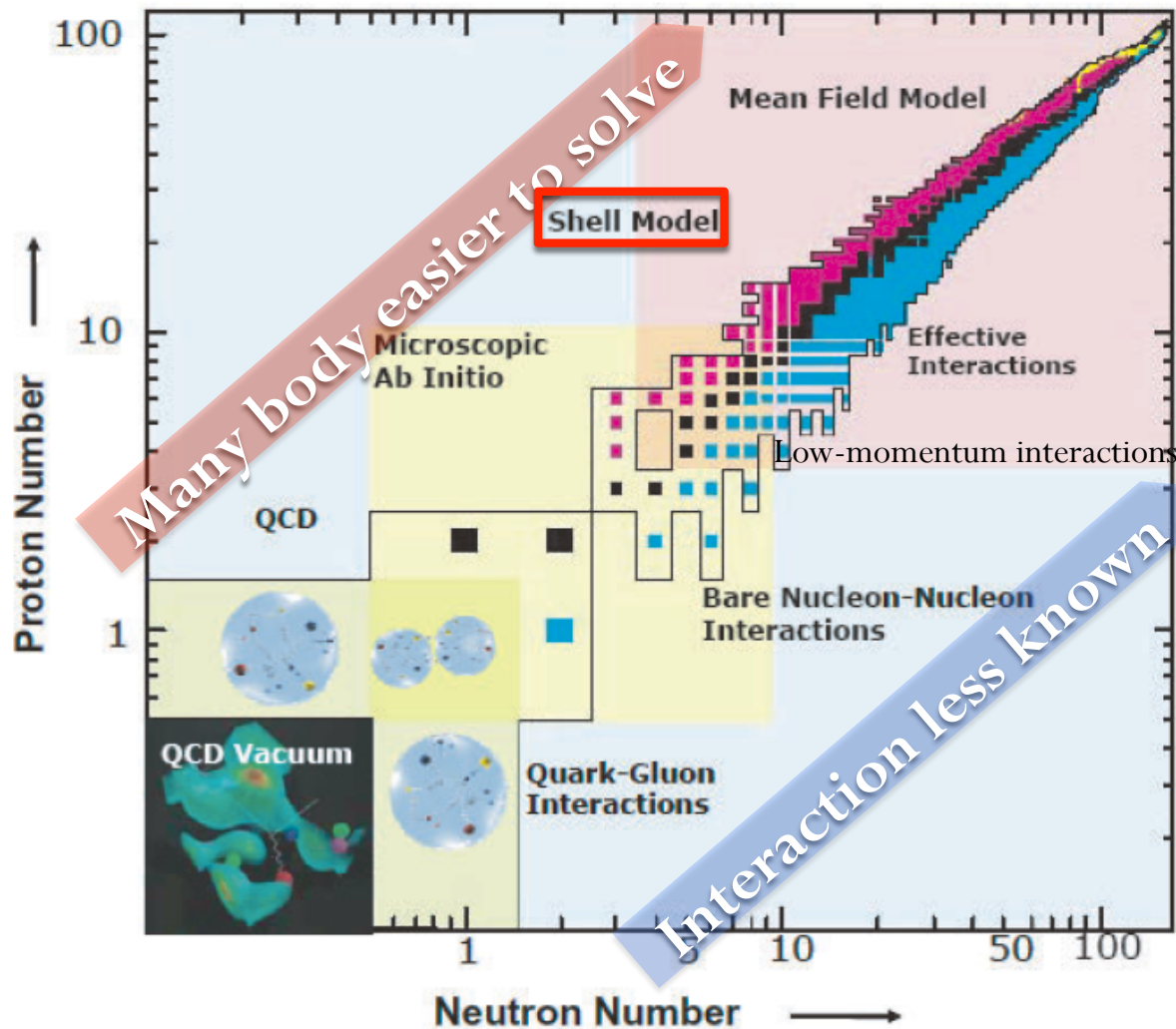
Not ‘exactly’ solvable above

$A \sim 16$ (*ab-initio*)

Here we focus on shell model

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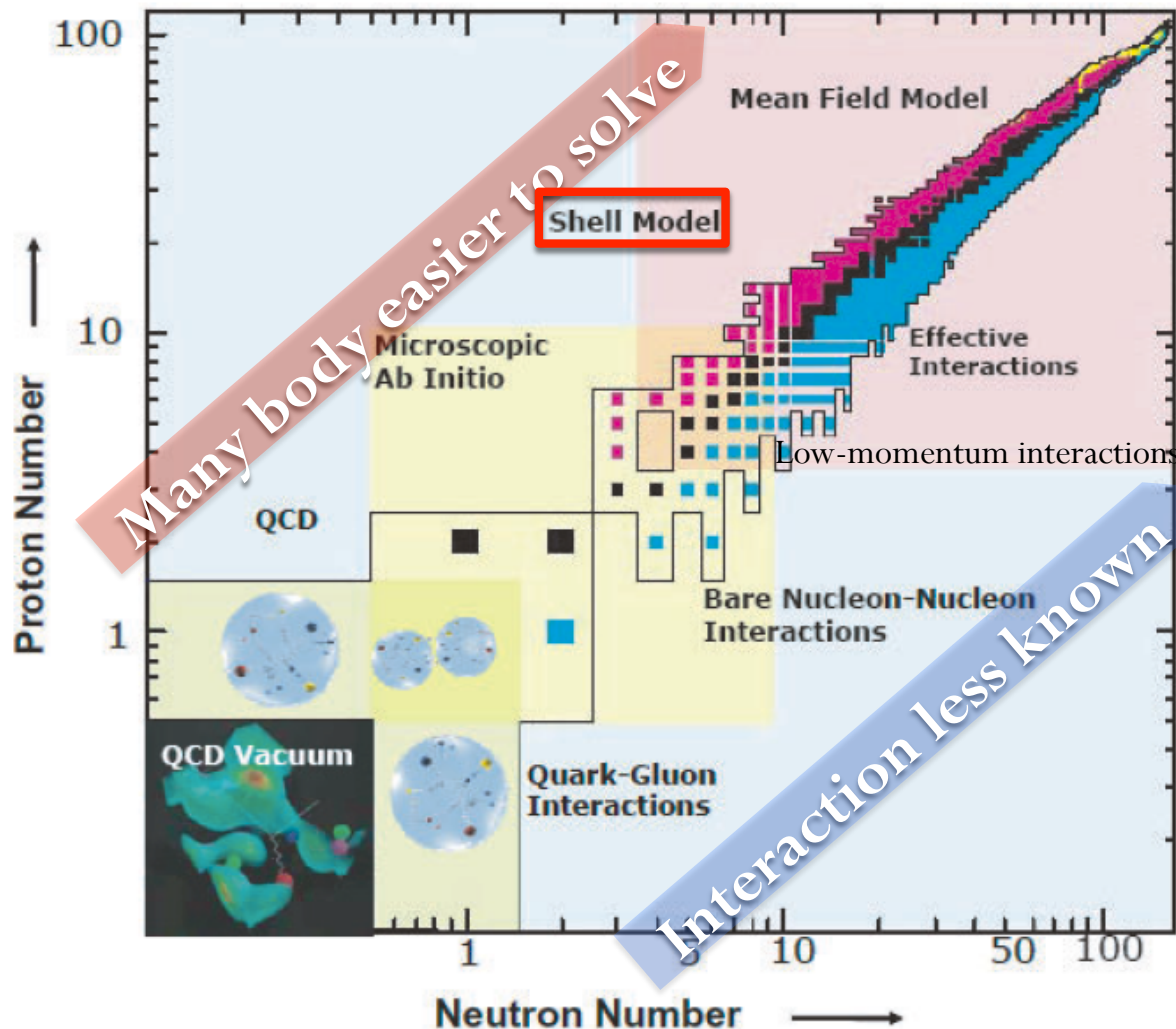
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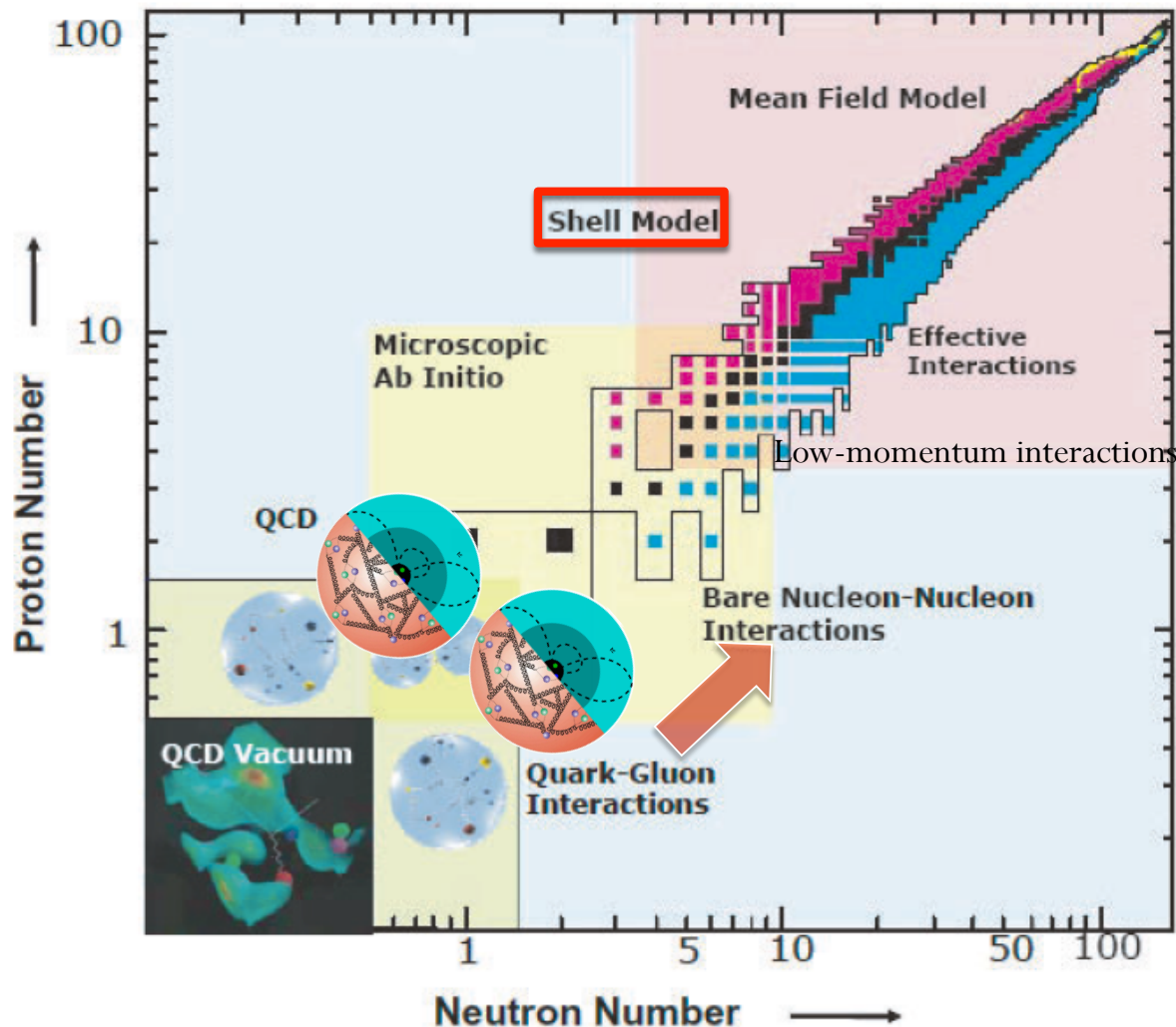
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How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow Solve many-body problem \rightarrow Predictions

The Challenge of Microscopic Nuclear Theory

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Nucleon-nucleon interaction

Some history

Anatomy of an NN interaction

Construction from QCD?

Ideas of Effective Field Theory

Chiral EFT for nuclear forces

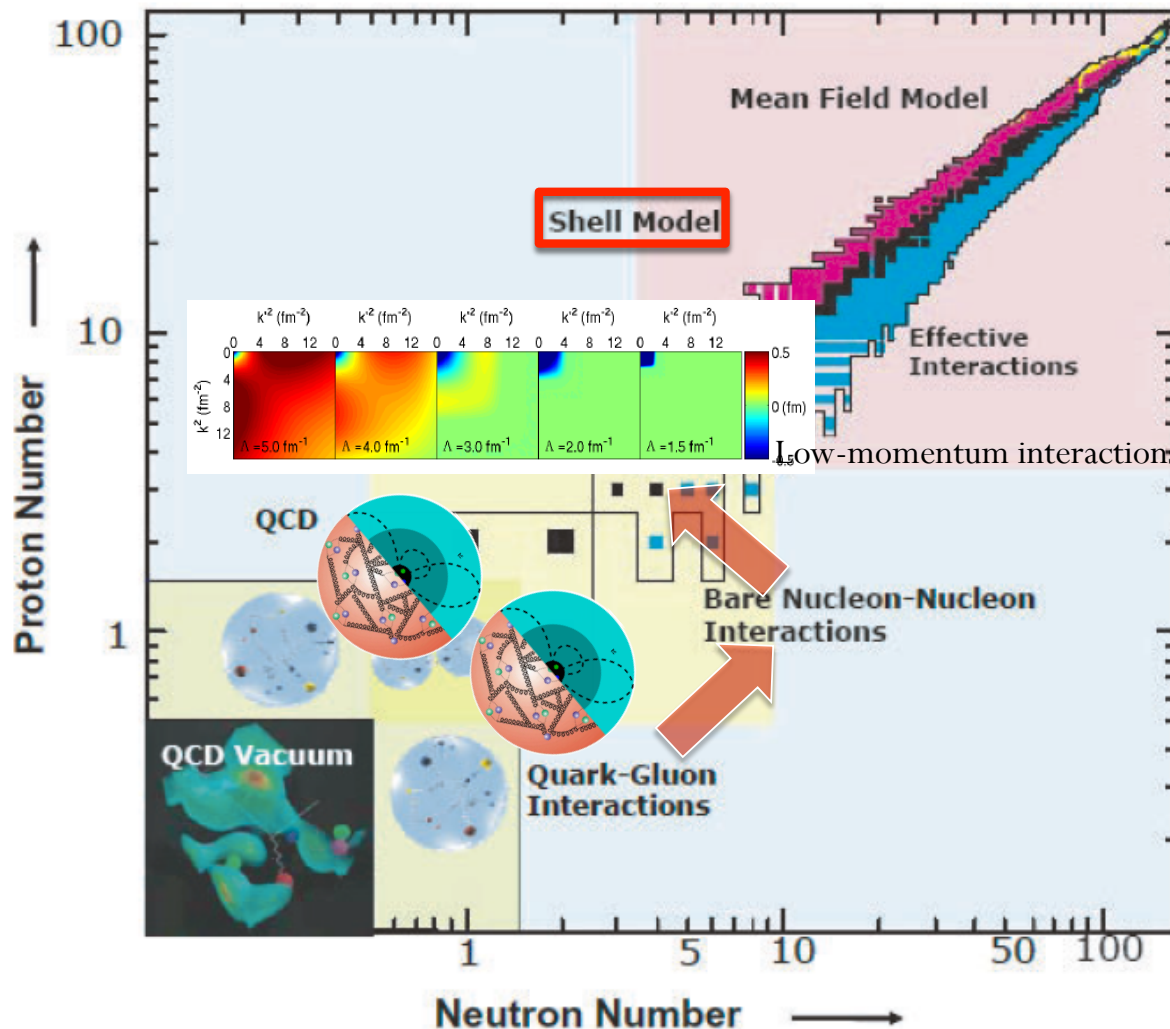
Constraint by data

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow Solve many-body problem \rightarrow Predictions

The Challenge of Microscopic Nuclear Theory

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Renormalizing NN Interactions

Basic ideas of RG

Calculating low-momentum interactions

Benefits of low cutoffs

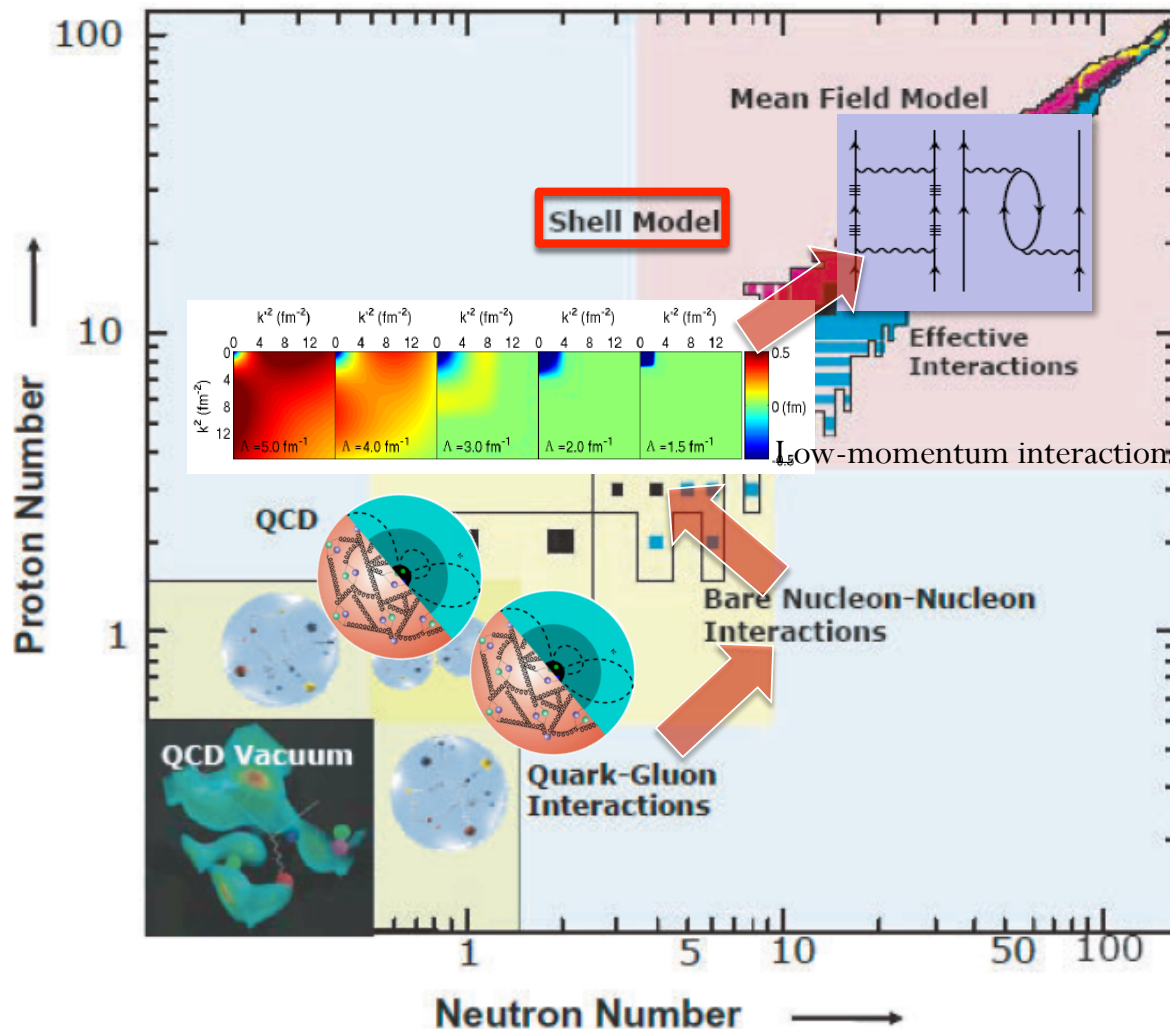
G-matrix renormalization

How will we approach this problem:

QCD → NN (3N) forces → Renormalize → Solve many-body problem → Predictions

The Challenge of Microscopic Nuclear Theory

To understand the properties of complex nuclei from elementary interactions



Microscopic Valence-Space Interactions

Model spaces

Many-body perturbation theory (MBPT)

Calculating effective interaction

Monopole part of interaction

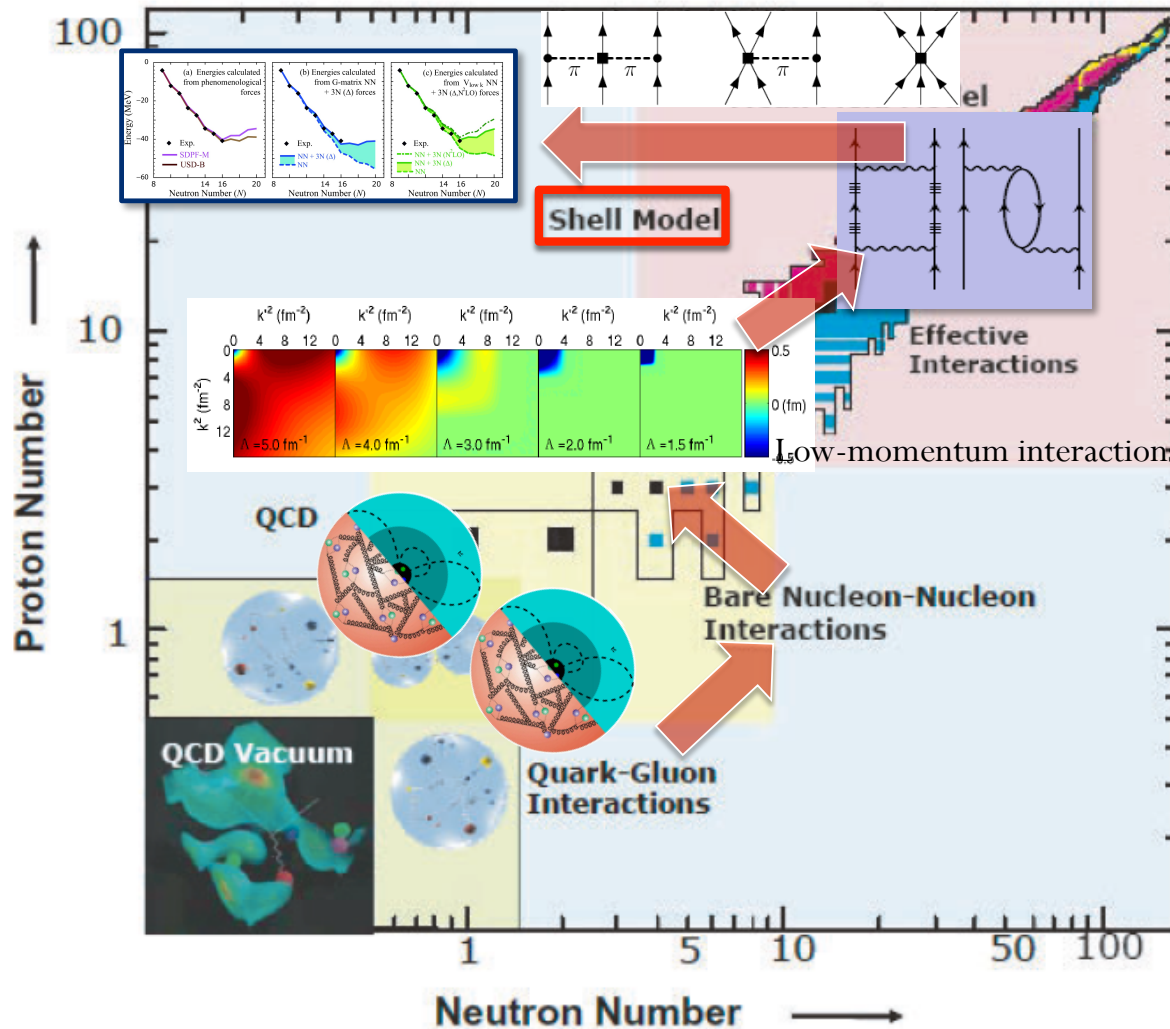
Deficiencies of this approach

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow Solve many-body problem \rightarrow Predictions

The Challenge of Microscopic Nuclear Theory

To understand the properties of complex nuclei from elementary interactions



Three-Nucleon Forces

Basic ideas – why do we need?

3N from chiral EFT

Implementing in shell model

Relation to monopoles

Predictions/Results

Density-dependent 3N

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow Solve many-body problem \rightarrow Predictions

Interaction Between Two Nucleons

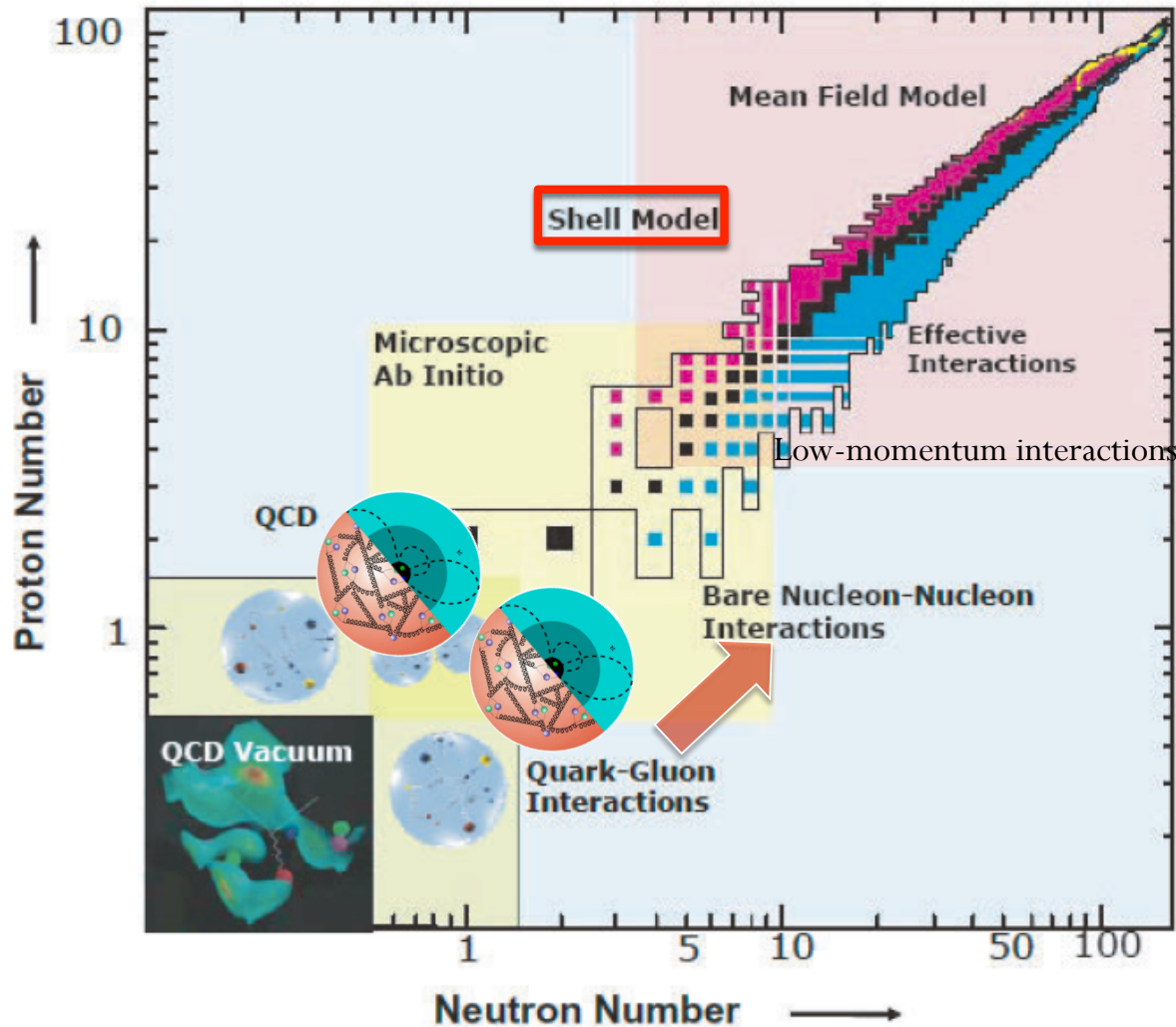
“In the past quarter century physicists have devoted a huge amount of experimentation and mental labor to this problem – probably more man-hours than have been given to any other scientific question in the history of mankind.”

–*H. Bethe*

So let's burn a few more man-hours of mental labor on this...

Part I: The Nucleon-Nucleon Interaction

To understand the properties of complex nuclei from elementary interactions



Nucleon-nucleon interaction

Some history

Anatomy of an NN interaction

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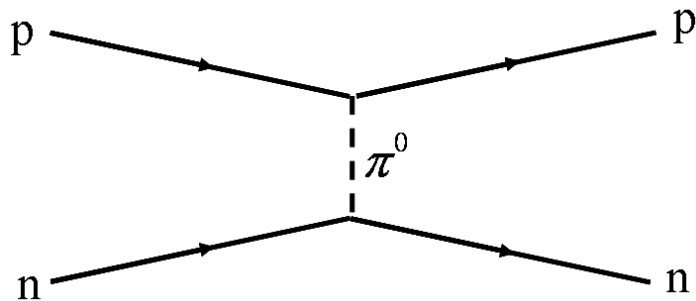
Fits to phase shifts

How will we approach this problem:

QCD \rightarrow NN (3N) forces \rightarrow Renormalize \rightarrow Solve many-body problem \rightarrow Predictions

Meson-Exchange Potentials: Yukawa

- First field-theoretical model of nucleon interaction proposed by **Yukawa** 1935
- Pion discovered 1947

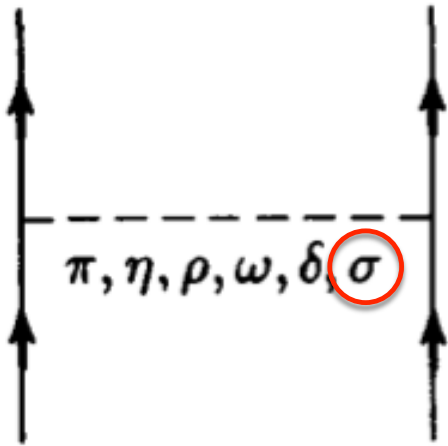


$$V(r) = -\frac{f_{\pi}^2}{m_{\pi}^2} \left\{ \sigma_1 \cdot \sigma_2 + C_T \left(1 + \frac{3}{m_{\alpha} r} + \frac{3}{(m_{\alpha} r)^2} \right) S_{12}(\hat{r}) \right\} \frac{e^{-m_{\pi} r}}{m_{\pi} r}$$

- **Attractive, long range, spin dependent, non-central (tensor) part**
- Successful in explaining scattering data, deuteron
- Advanced to multi-pion theories in 1950's – failed

One-Boson Exchange Potentials

- Heavy mesons discovered in 1950s – theories developed based on these
- Intermediate range – **attractive central, spin-orbit**



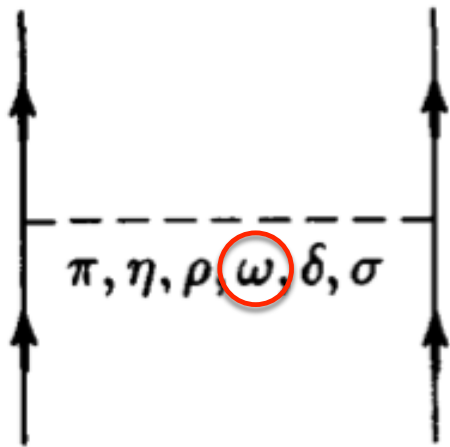
$$V^\sigma = g_{\sigma NN}^2 \frac{1}{\mathbf{k}^2 + m_\sigma^2} \left(-1 + \frac{\mathbf{q}^2}{2M_N^2} - \frac{\mathbf{k}^2}{8M_N^2} - \frac{\mathbf{L} \cdot \mathbf{S}}{2M_N^2} \right)$$

$$\vec{q}_i \equiv \vec{p}'_i - \vec{p}_i \quad \vec{k}_i \equiv \frac{1}{2}(\vec{p}'_i + \vec{p}_i)$$

Baryons	Mass (MeV)	Mesons	Mass (MeV)
p, n	938.926	π	138.03
Λ	1116.0	η	548.8
Σ	1197.3	σ	≈ 550.0
Δ	1232.0	ρ	770
Σ^*	1385.0	ω	782.6
		δ	983.0
		K	495.8
		K*	895.0

One-Boson Exchange Potentials

- Heavy mesons discovered in 1950s – theories developed based on these
- Short range; **repulsive central force, attractive spin-orbit**

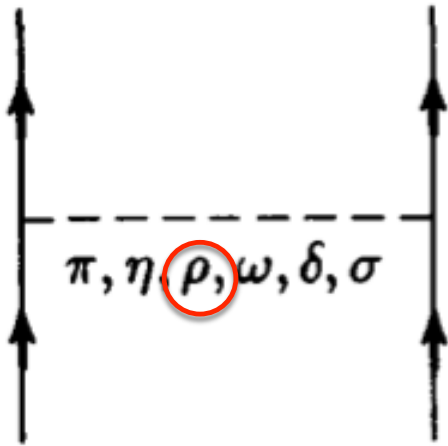


$$V^\omega = g_{\omega NN}^2 \frac{1}{\mathbf{k}^2 + m_\omega^2} \left(1 - 3 \frac{\mathbf{L} \cdot \mathbf{S}}{2M_N^2} \right)$$

Baryons	Mass (MeV)	Mesons	Mass (MeV)
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One-Boson Exchange Potentials

- Heavy mesons discovered in 1950s – theories developed based on these
- Short range; **tensor force opposite sign of pion exchange**



$$V^{\rho} = g_{\rho NN}^2 \frac{\mathbf{k}^2}{\mathbf{k}^2 + m_{\rho}^2} \left(-2\sigma_1\sigma_2 + S_{12}(\hat{\mathbf{k}}) \right) \tau_1\tau_2$$

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Parameterizing the NN Interaction

Starting from any NN-interaction

We first solve scattering Lipmann-Schwinger scattering T-matrix equation:

$$T_{ll'}^\alpha(kk'K) = V_{ll'}^\alpha(kk') + \frac{2}{\pi} \sum_{l''} \int_0^\infty dq q^2 V_{ll''}^\alpha(kq) \frac{1}{k^2 - q^2 + i\epsilon} T_{l''l'}^\alpha(qk'K).$$

Where

$$T_{ll'}^\alpha(k, k'; K) = \langle kK, lL; JST | T | k'K, l'L; JST \rangle$$

Parameterized in partial waves α – in **relative/center of mass frame (K,L)**

$$\tan \delta(p) = -pT(p, p)$$

Fully-on-shell T -matrix directly related to experimental data

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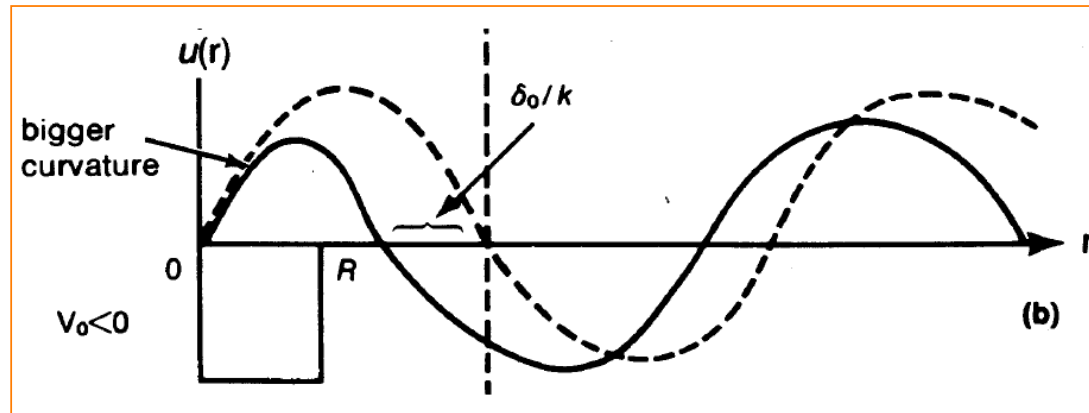
Fully-on-shell T -matrix directly related to experimental data

Note intermediate momentum allowed to infinity (but cutoff by regulators)

Coupling of low-to-high momentum in V

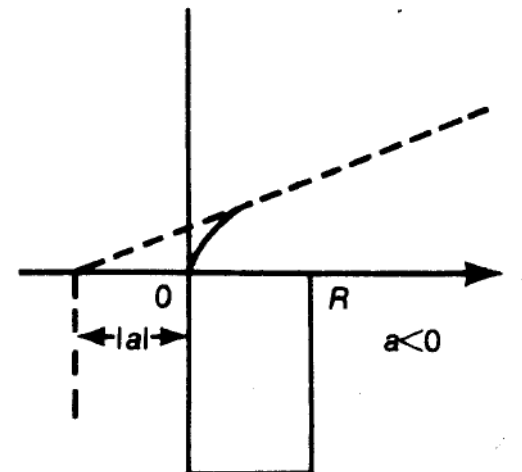
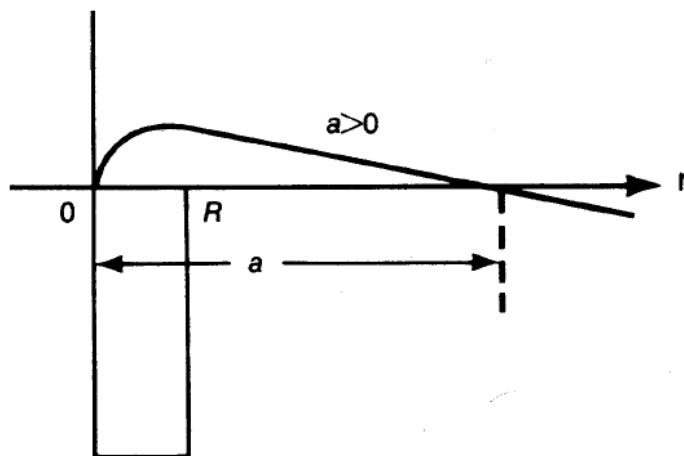
Constraining NN Scattering Phase Shifts

Phase shift is a function of relative momentum k ; Figure shows s -wave



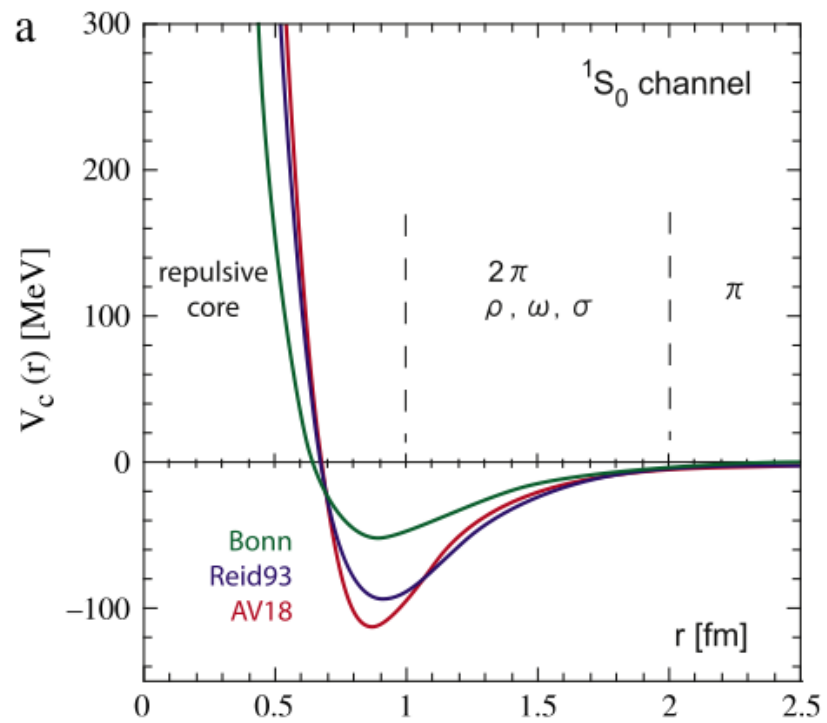
Scattering length:

$$k \cot \delta(k) \approx -\frac{1}{a}; \quad \sigma_{\text{tot}} \approx 4\pi a^2 \quad \text{for } k \rightarrow 0$$



Form of NN Interactions

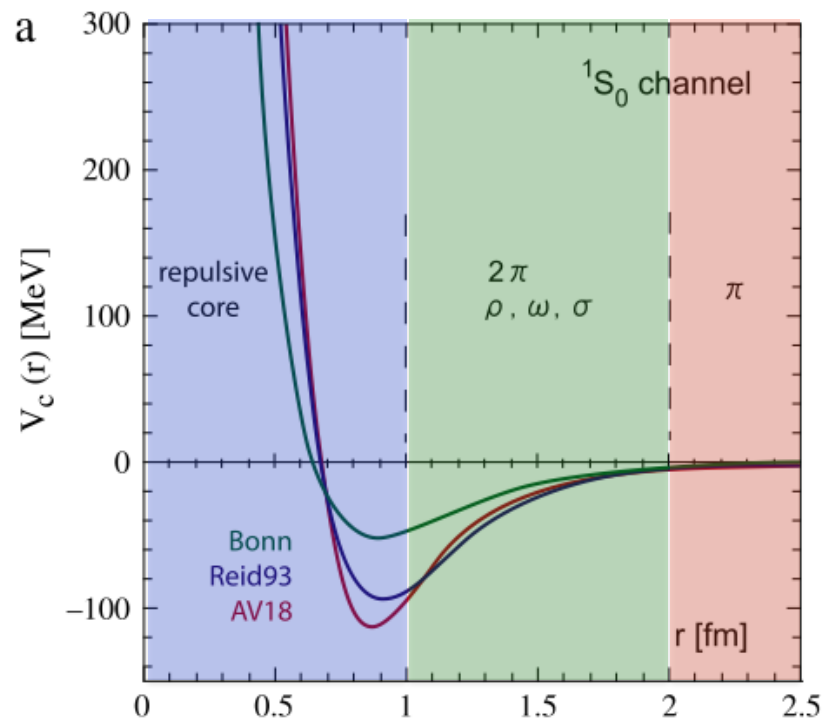
Textbook nuclear potentials in \mathbf{r} -space



Form of NN Interactions

Textbook nuclear potentials in \mathbf{r} -space

Hard core, **intermediate-range 2π** , **long-range 1π exchange (OPE)**



Form of NN Interactions

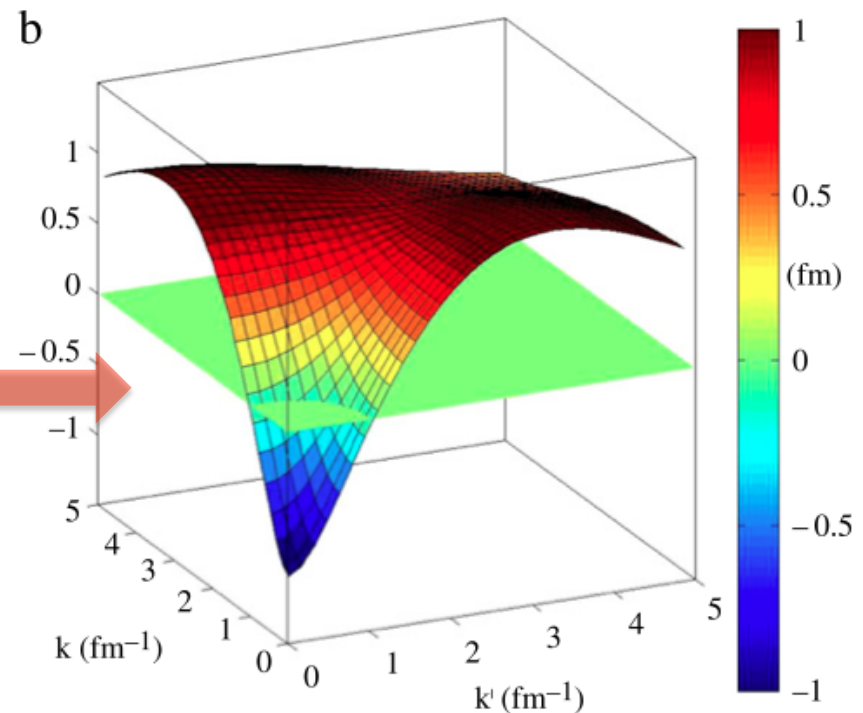
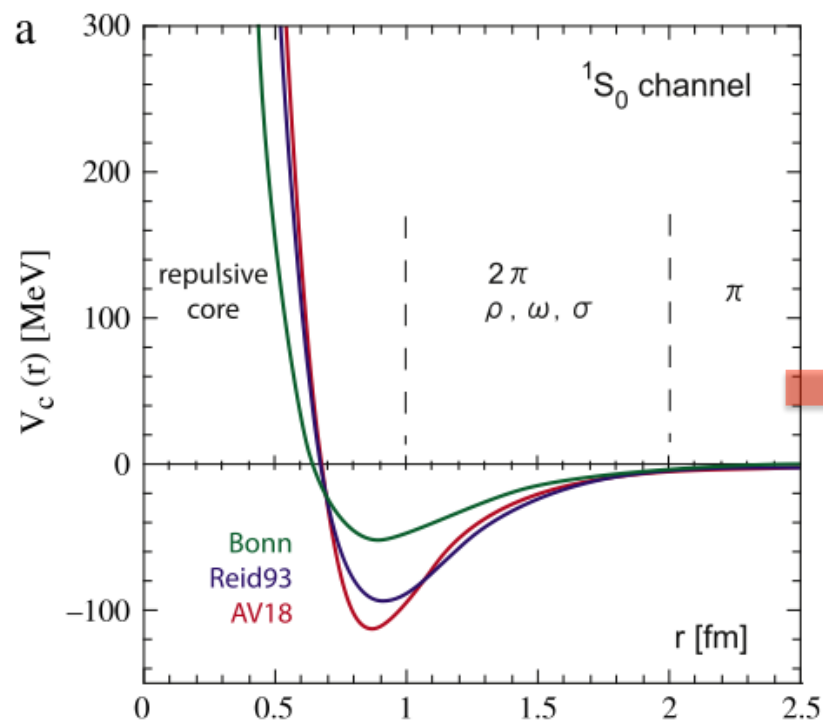
Textbook nuclear potentials in \mathbf{r} -space

- Hard core, intermediate-range 2π , long-range 1π exchange

Transform to momentum space via Fourier-Bessel Transformation

- **Strong high-momentum repulsion**, **low-momentum attraction**

$$V_l(k, k') = \frac{2}{\pi} \int_0^{\infty} r^2 dr j_l(kr) V(r) j_l(k'r)$$



Form of NN Interactions

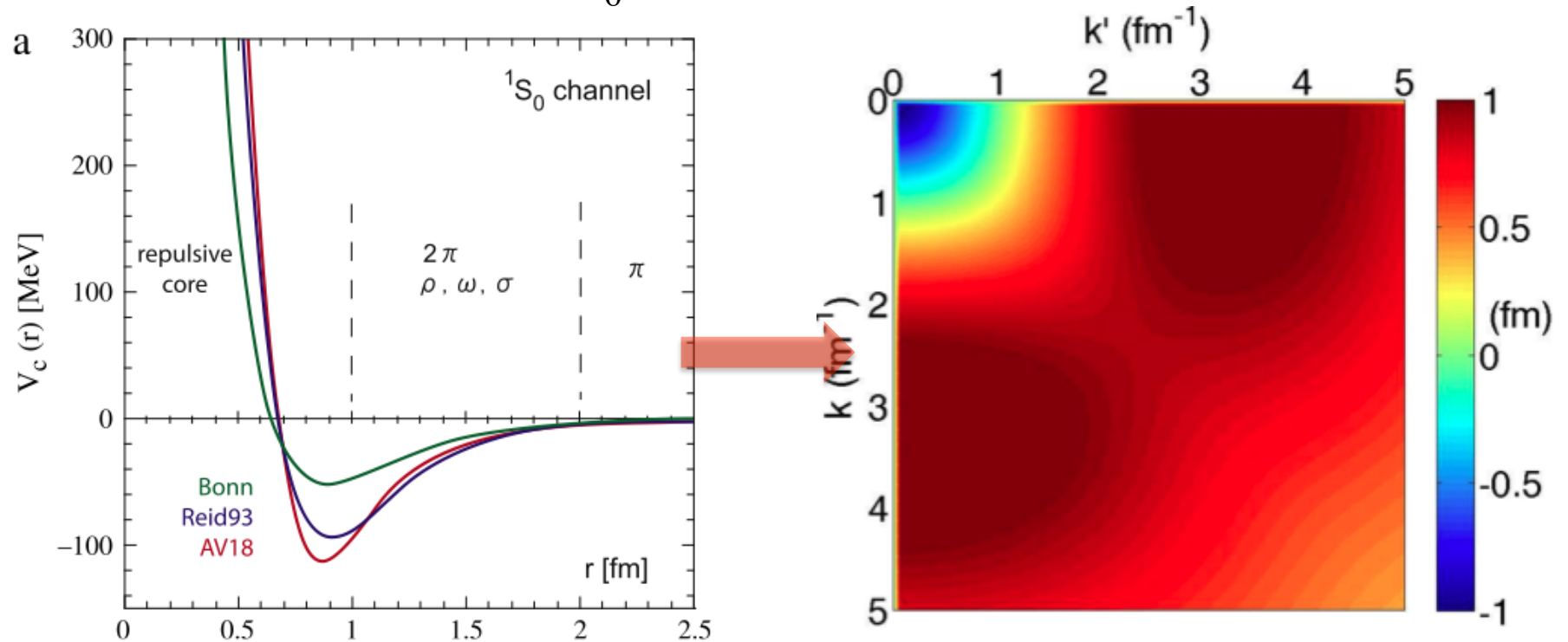
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Form of NN Interactions

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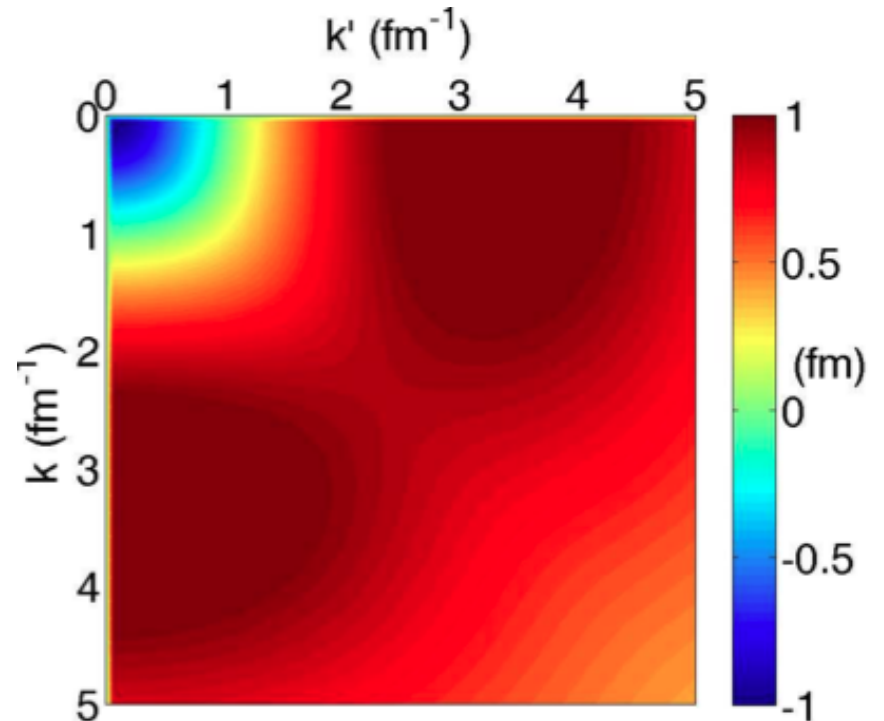
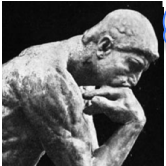
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$$V_l(k, k') = \frac{2}{\pi} \int_0^{\infty} r^2 dr j_l(kr) V(r) j_l(k'r)$$

Wait a minute... these potentials can't really go to extremely high energies; that would be QCD!

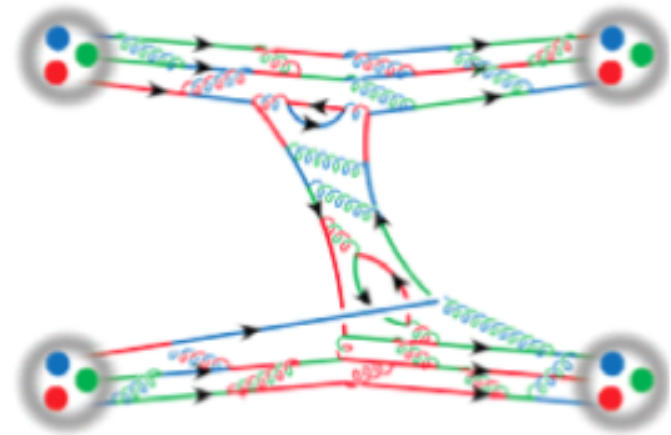
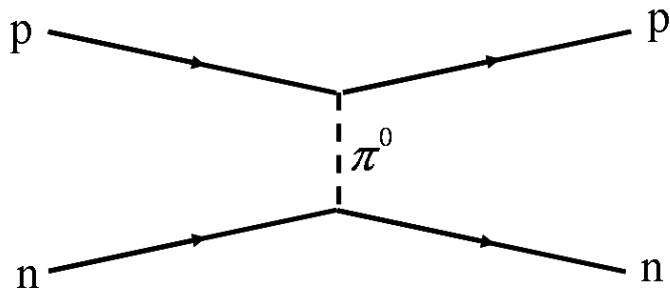


NN Interaction from QCD

Meson exchange described in QCD

Low-energy region non-perturbative – treat in the context of **Lattice QCD**

Directly from QCD Lagrangian, solve numerically on discretized space-time

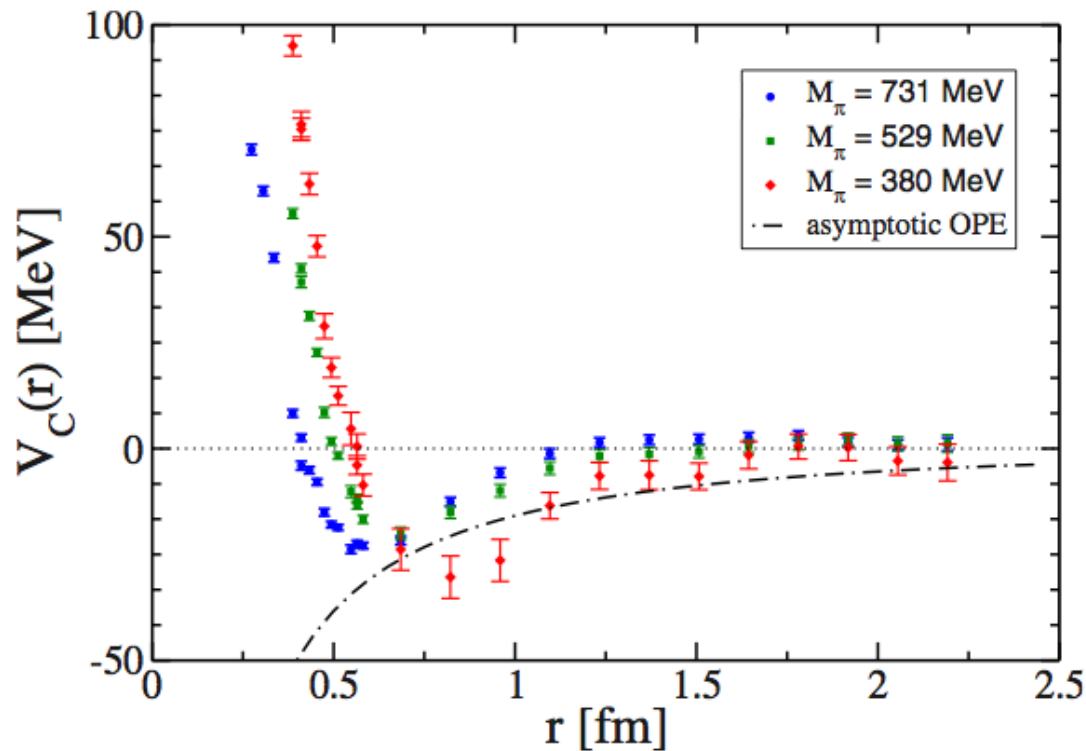


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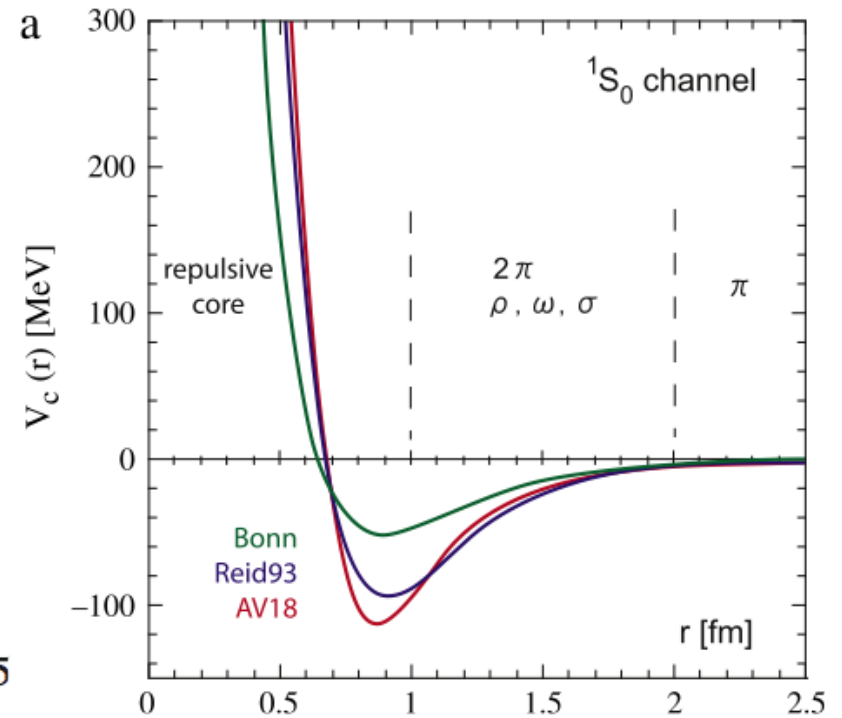
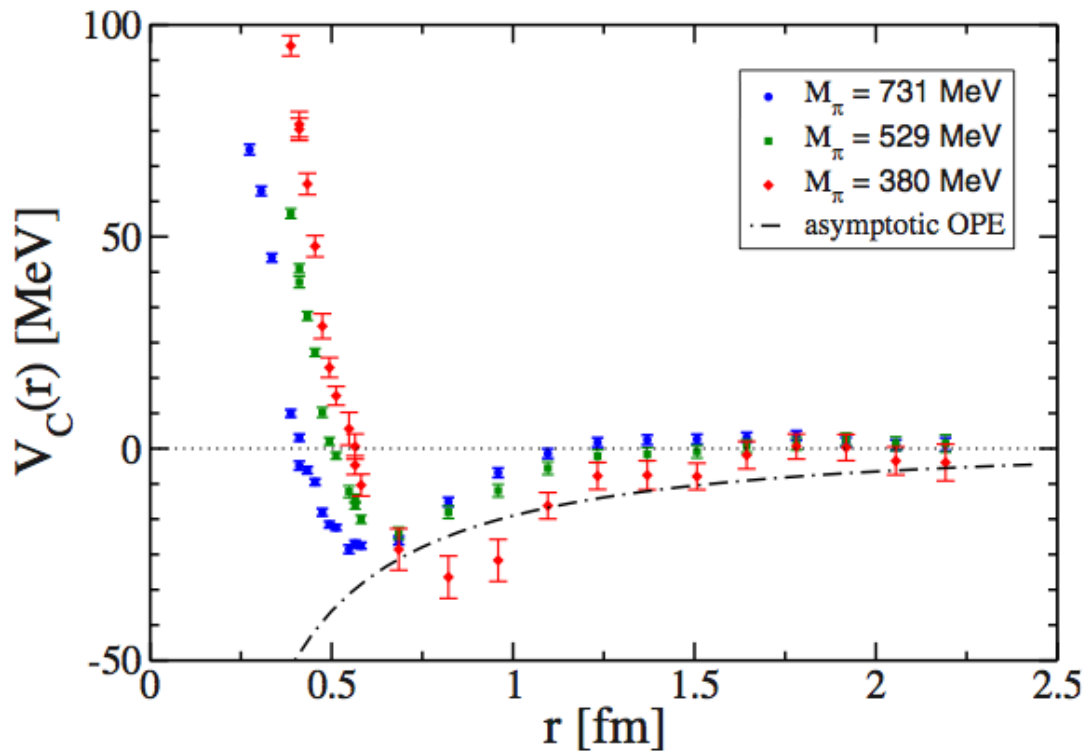
Lattice results give long-range OPE tail, hard core

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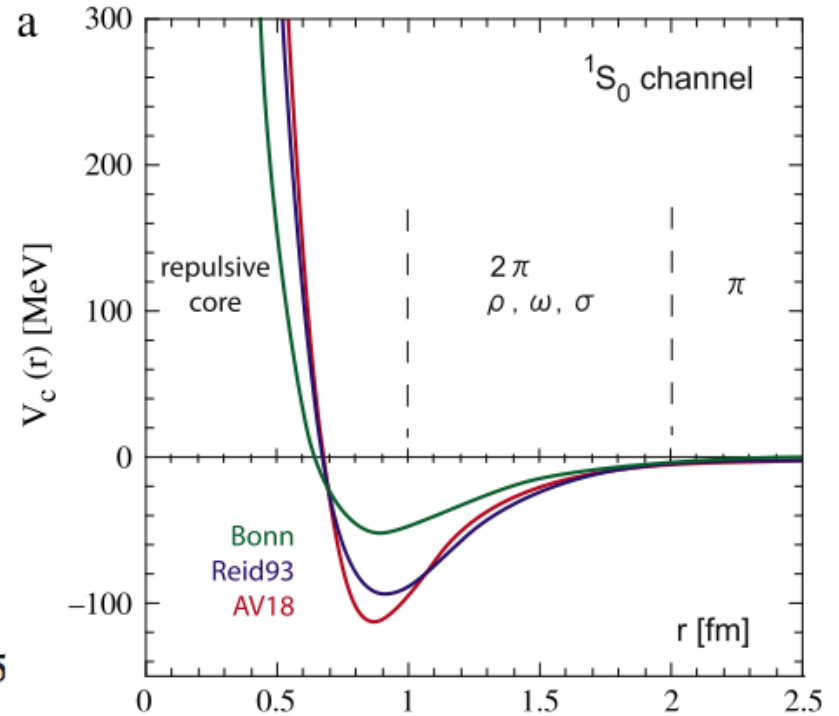
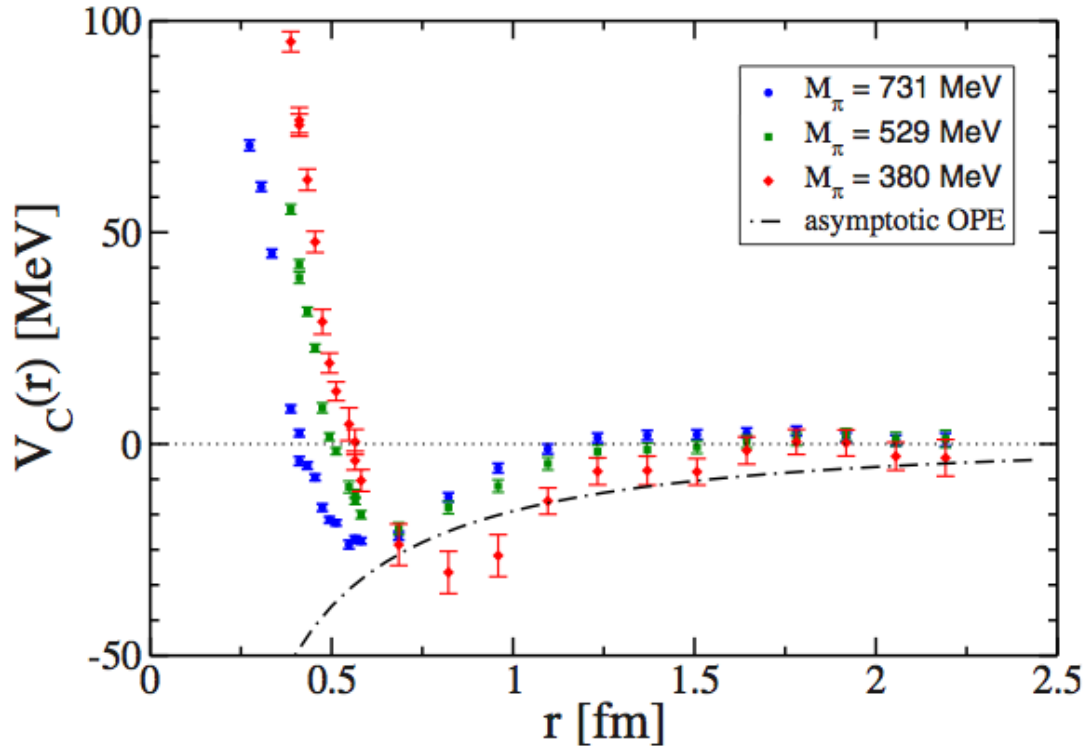
Lattice results give long-range OPE tail, hard core

Not yet to physical pion mass – work in progress – so we're done, right?

Unique NN Potential?

What does this tell us in our quest for an NN-potential?

Expected form seems to be confirmed by QCD



OBE Potentials: Summary/Problems

First generation (1960-1990): Paris, Reid, Bonn-A,B,C $\chi^2/\text{dof} \approx 2$

High precision potentials (1990s): ~ 40 parameters fit to NN data $\chi^2/\text{dof} \approx 1$

ArgonneV18, Reid93, Nijmegen, CD-Bonn

NN problem “solved” !!

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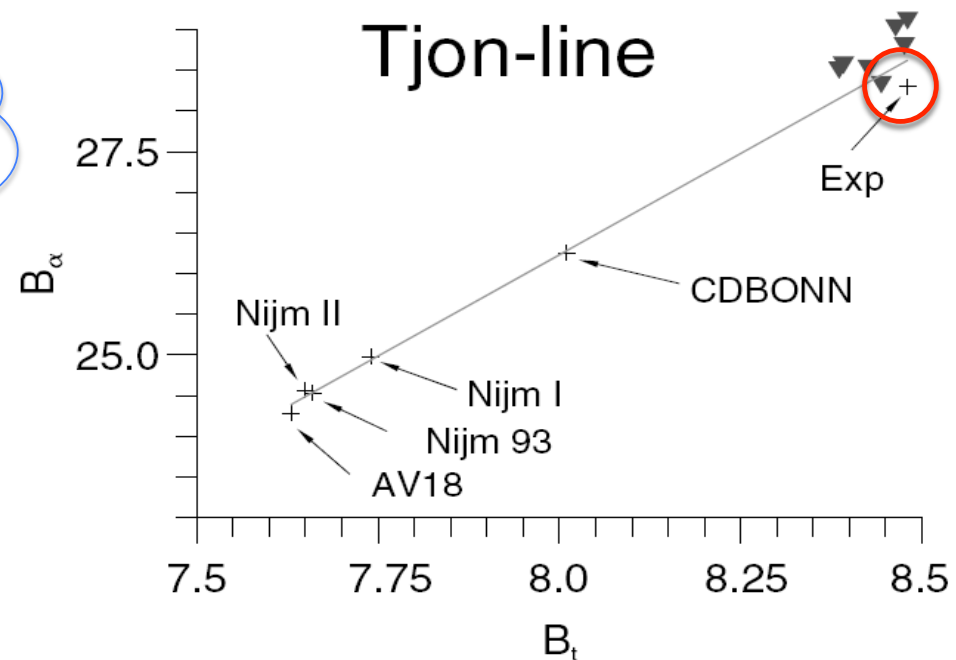
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Hmmm... what if we
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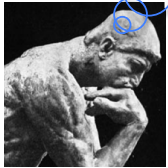
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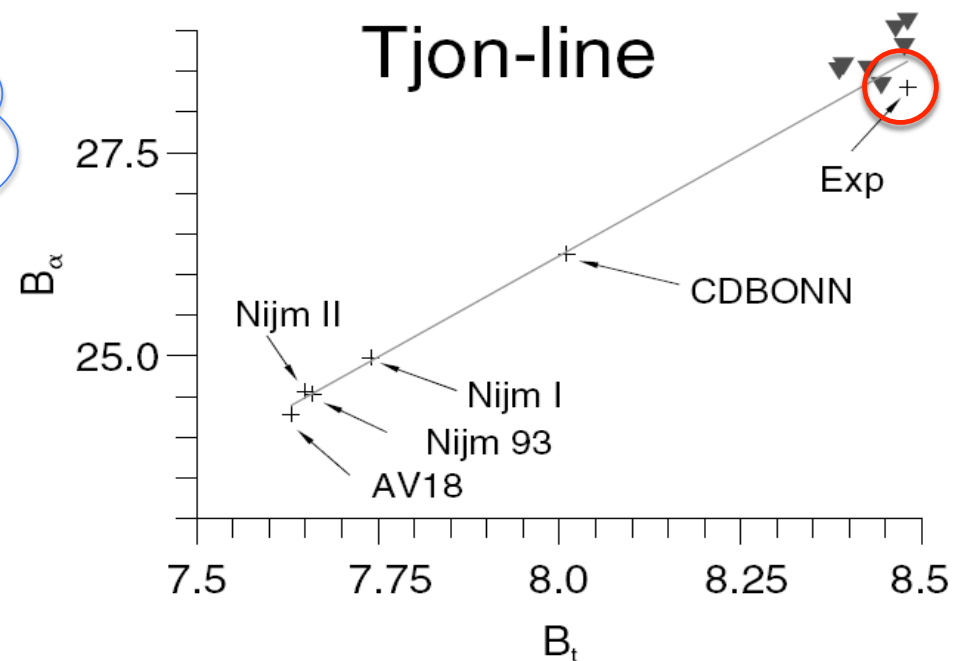
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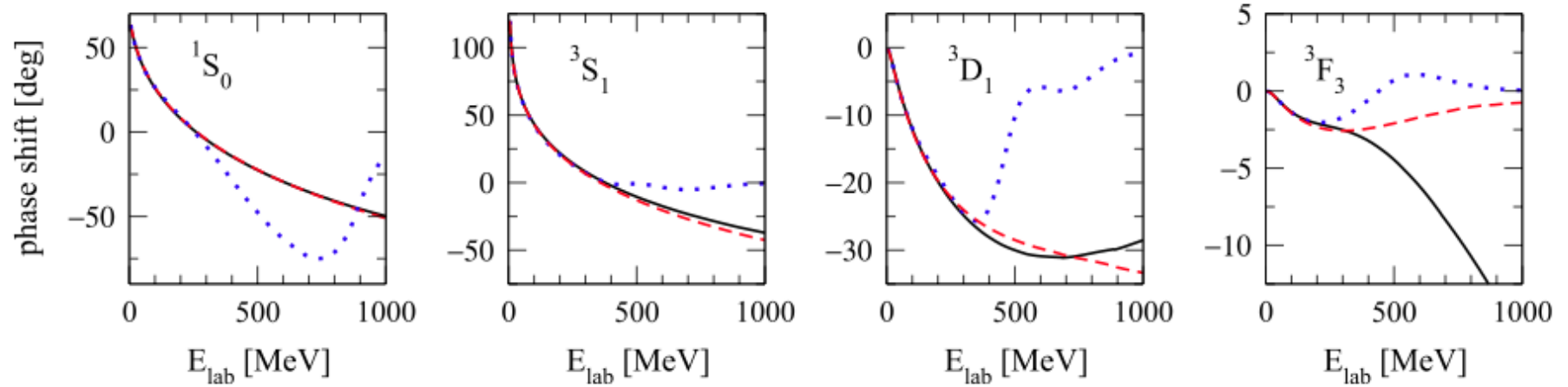
Many successes, but...

- 1) Difficult (impossible) to assign
- 2) 3N forces not consistent with NN forces
- 3) No clear connection to QCD
- 4) Clear **model dependence**...



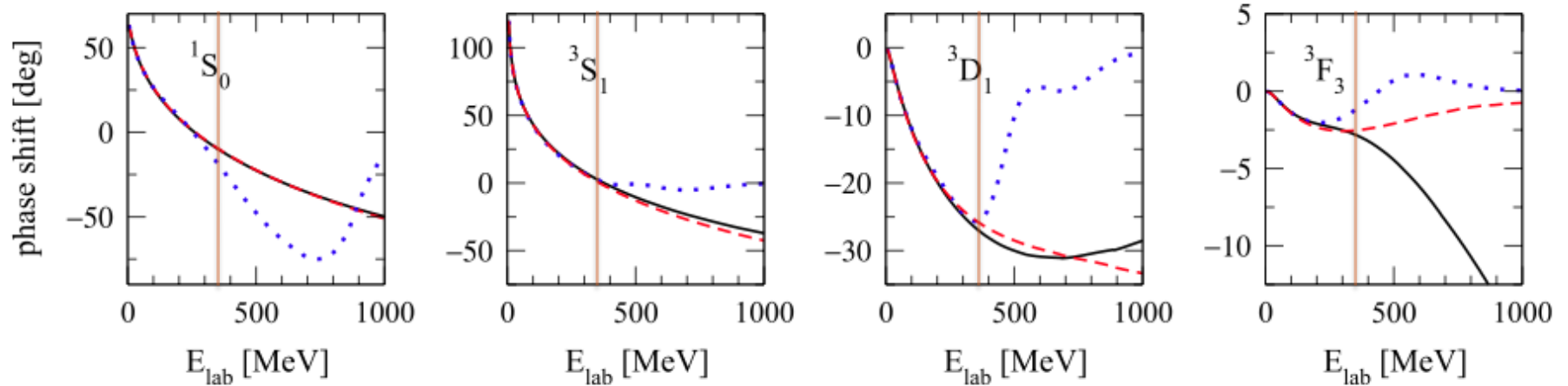
Meson-Exchange Potentials and Phase Shifts

Examples of phase shift reproduction by NN potentials

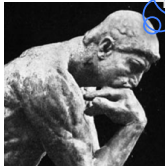


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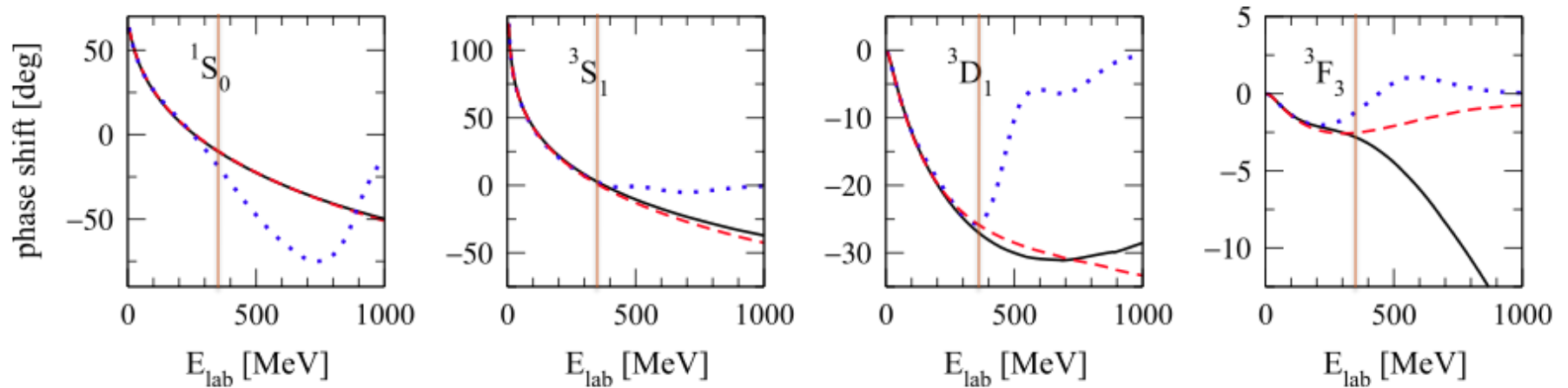
That's strange...
why do they only
agree to 350MeV?



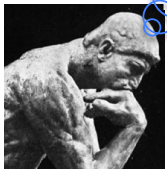
Remember, all have: $\chi^2/\text{dof} \approx 1$

Meson-Exchange Potentials and Phase Shifts

More model dependence: examples of phase shift reproduction by NN potentials



That's strange...
why do they only
agree to 350MeV?



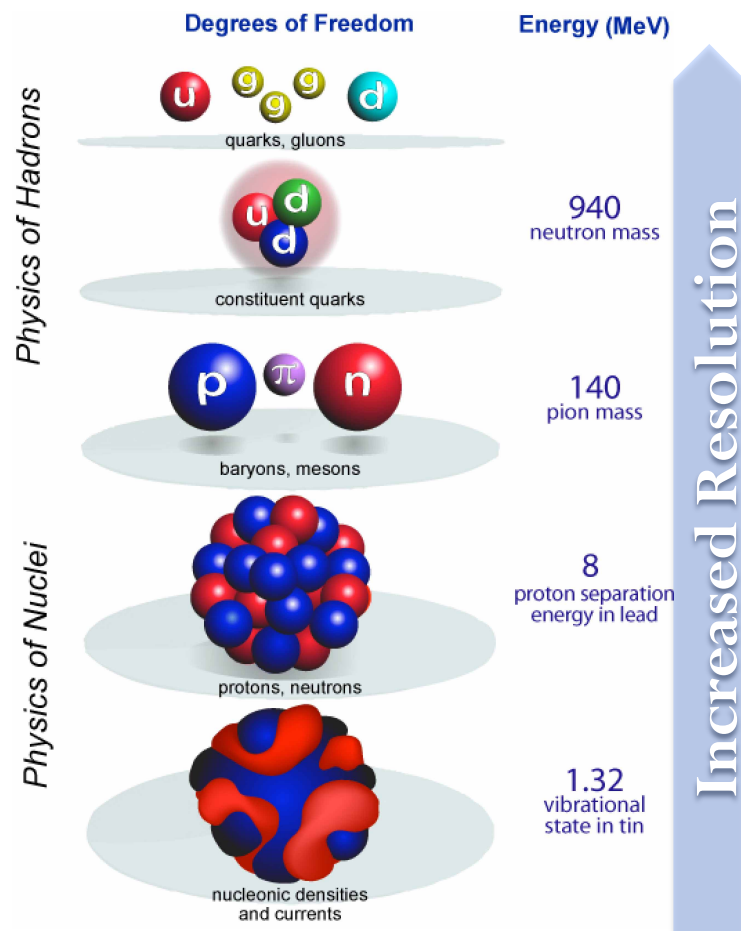
Agree well up to **pion-production threshold** $\sim 350\text{MeV}$

Data sparse – most models don't fit above this point - **unconstrained**

From QCD to Nuclear Interactions

How do we determine interactions between nucleons?

$$H(\Lambda) = T + V_{NN}(\Lambda) + V_{3N}(\Lambda) + V_{4N}(\Lambda) + \dots$$



Old view:

Multiple scales complicate life
No meaningful way to connect them

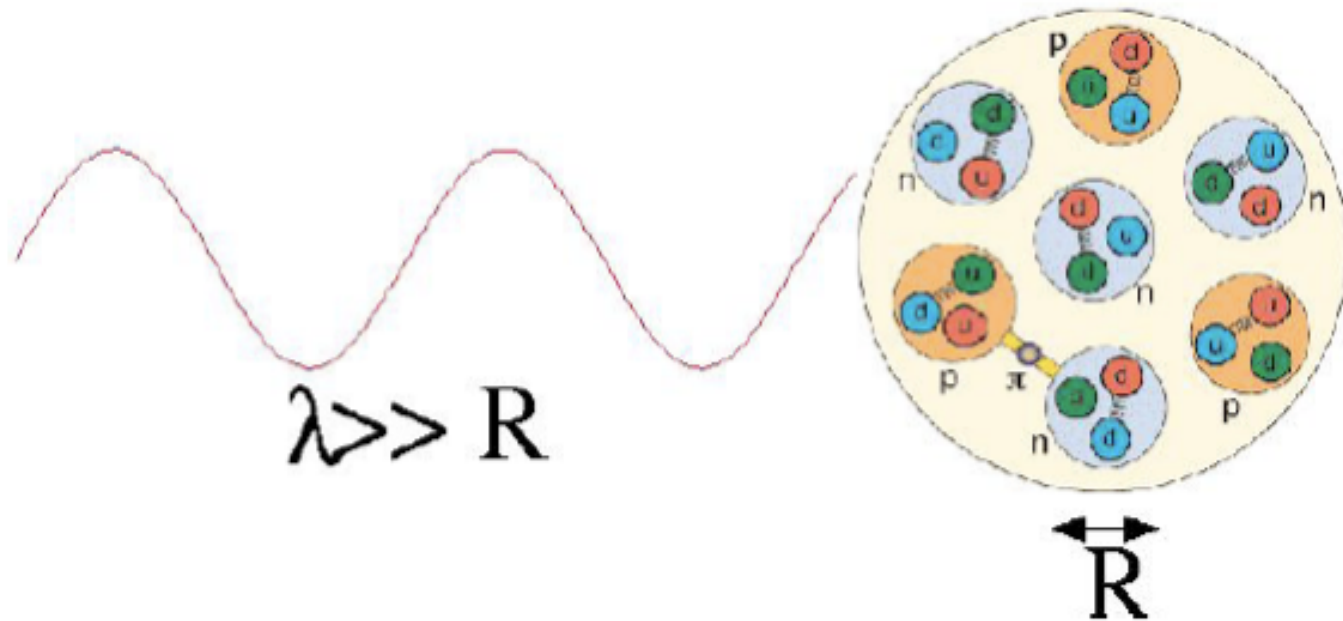
Modern view:

Ratio of scales – small parameters
Effective field theory at each scale
connected by RG

Choose convenient resolution scale

Ideas Behind Effective Theories

Resolution scales

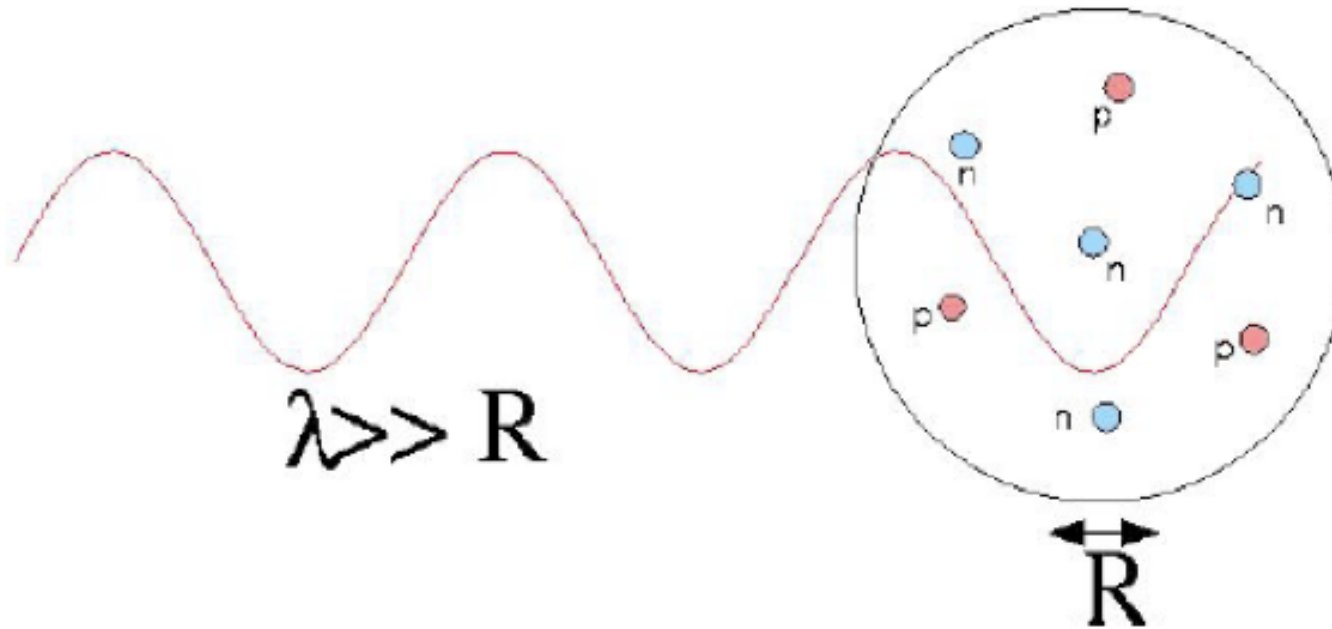


High energy probe resolves fine details

Need high-energy degrees of freedom

Ideas Behind Effective Theories

How do we determine interactions between nucleons?



Low-energy probe doesn't resolve such details

Don't need high-energy degrees of freedom – replace with something simpler

Use dof that are more convenient, but preserve low-energy observables

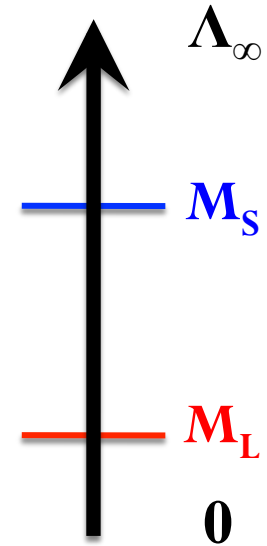
Ideas Behind Effective Theories

Underlying theory
with cutoff Λ_∞

$$V = V_L + V_S$$

Known **long-distance**
physics (like 1π -exchange)
with some scale M_L

Short-distance physics
(ρ, ω -exchange) with
Some scale M_S



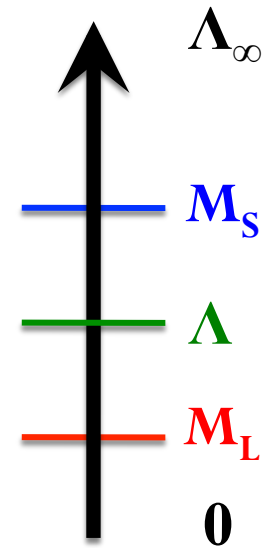
Ideas Behind Effective Theories

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Say we want a **low-energy effective theory** describing physics
up to some $M_L < \Lambda < M_S$

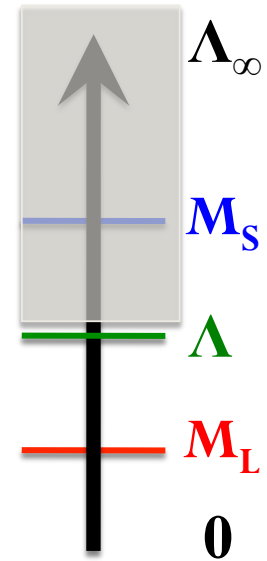
Ideas Behind Effective Theories

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Say we want a **low-energy effective theory** describing physics
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Integrate out states above Λ using **Renormalization Group (RG)**

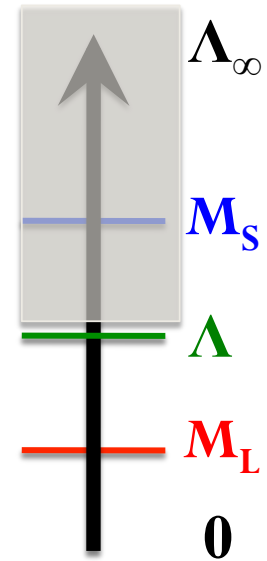
Ideas Behind Effective Theories

Underlying theory
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Short-distance physics
(ρ, ω -exchange) with
Some scale M_S



Say we want a **low-energy effective theory** describing physics
up to some $M_L < \Lambda < M_S$

Integrate out states above Λ using **Renormalization Group (RG)**

General form of effective theory: $V_{eff} = V_L + \delta V_{c.t.}(\Lambda)$

$$\delta V_{c.t.}(\Lambda) = C_0(\Lambda)\delta^3(\vec{r}) + C_2(\Lambda)\nabla^2\delta^3(\vec{r}) + \dots$$

Ideas Behind Effective Theories

General form of effective theory: $V_{eff} = V_L + \delta V_{c.t.}(\Lambda)$

$$\delta V_{c.t.}(\Lambda) = C_0(\Lambda) \delta^3(\vec{r}) + C_2(\Lambda) \nabla^2 \delta^3(\vec{r}) + \dots$$

Encodes effects of high-E
dof on low-energy observables

Universal; depends only
on symmetries

TWO choices:

Short distance structure of “true theory” captured in several numbers

- **Calculate from underlying theory**

When short-range physics is unknown or too complicated

- **Extract from low-energy data**

How do we apply these ideas to nuclear physics?

Chiral Effective Field Theory: Philosophy

“At each scale we have different degrees of freedom and different dynamics. Physics at a larger scale (largely) decouples from physics at a smaller scale... thus a theory at a larger scale remembers only finitely many parameters from the theories at smaller scales, and throws the rest of the details away.

More precisely, when we pass from a smaller scale to a larger scale, we average out irrelevant degrees of freedom... The general aim of the RG method is to explain how this decoupling takes place and why exactly information is transmitted from one scale to another through finitely many parameters.”

- *David Gross*

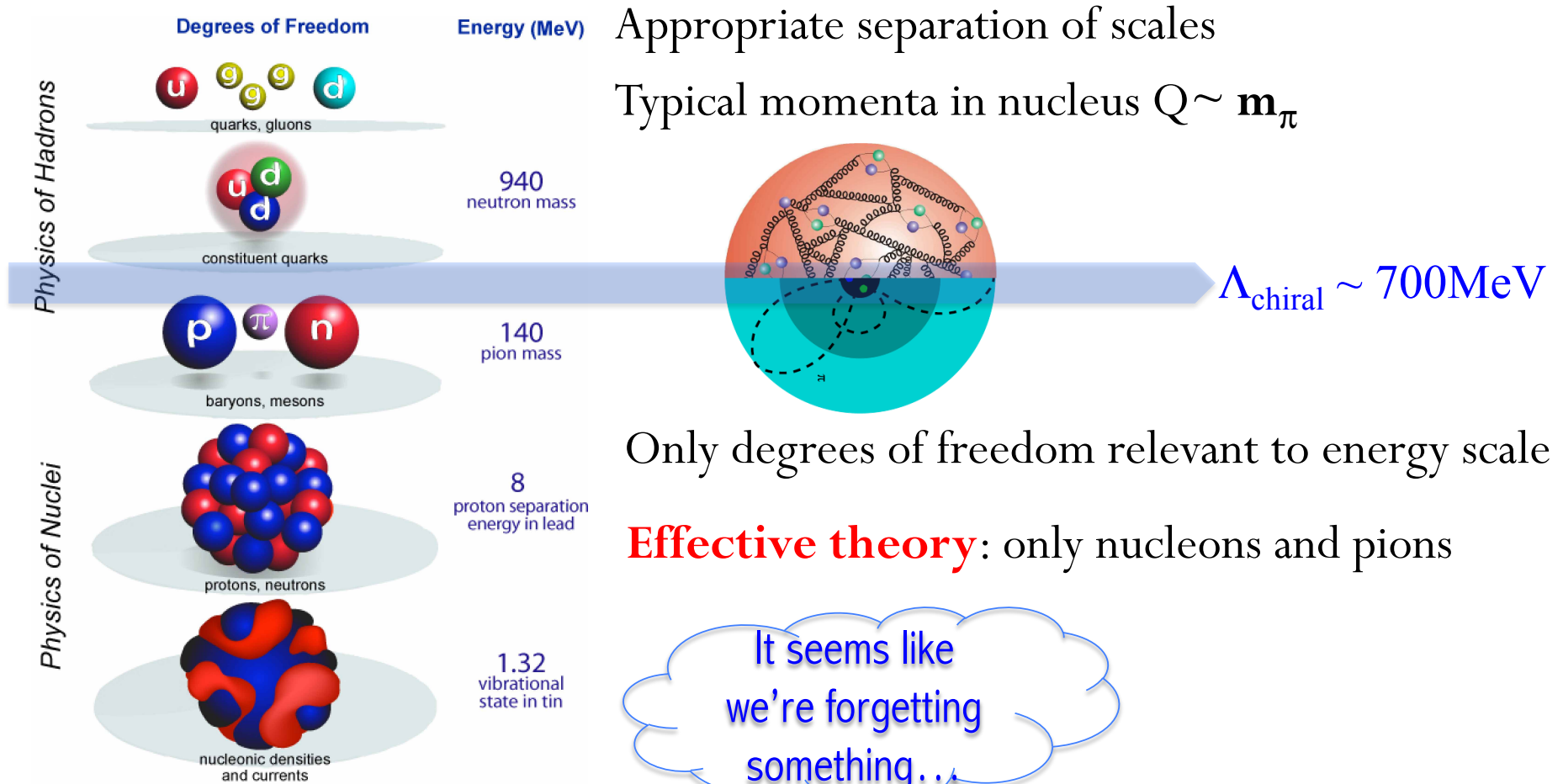
“The method in its most general form can.. be understood as a way to arrange in various theories that the degrees of freedom that you’re talking about are the relevant degrees of freedom for the problem at hand.”

- *Steven Weinberg*

5 Steps to constructing the theory

Separation of Scales in Nuclear Physics

Step I: Identify appropriate separation of scales, degrees of freedom



Chiral EFT Symmetries

Step II: Identify relevant symmetries of underlying theory QCD

1. $SU(3)$ color symmetry from QCD
(Nucleons and pions are color singlets)
2. Chiral symmetry: u and d quarks are almost massless
 - Left and right-handed (massless) quarks do not mix: $SU(2)_L \times SU(2)_R$ symmetry
 - Explicit symmetry breaking: u and d quarks have a small mass
 - Spontaneous breaking of chiral symmetry (no parity doublets observed in Nature)
 - $SU(2)_L \times SU(2)_R$ symmetry spontaneously broken to $SU(2)_V$
 - Pions are the Nambu-Goldstone bosons of spontaneously broken symmetry
 - Low-energy pion Lagrangian completely determined

Construct Lagrangian based on these symmetries

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\pi\pi} + \mathcal{L}_{N\pi} + \mathcal{L}_{NN}$$

Chiral EFT Lagrangian

Step III: Construct Lagrangian based on identified symmetries

Pion-pion Lagrangian: U is $SU(2)$ matrix parameterized by three pion fields

$$\mathcal{L}_{\pi}^{(0)} = \frac{F^2}{4} \langle \nabla^\mu U \nabla_\mu U^\dagger + \chi_+ \rangle,$$

Leading-order pion-nucleon

$$\mathcal{L}_{\pi N}^{(0)} = \bar{N}(i v \cdot D + \hat{g}_A u \cdot S)N,$$

Leading-order nucleon-nucleon (encodes unknown short-range physics)

$$\mathcal{L}_{NN}^{(0)} = -\frac{1}{2} C_S (\bar{N}N)(\bar{N}N) + 2C_T (\bar{N}SN) \cdot (\bar{N}SN)$$

EFT Power Counting

Step IV: Design an **organized** scheme to distinguish more from less important processes: **Power Counting**

Organize theory in powers of $\left(\frac{Q}{\Lambda_\chi}\right)$ where $Q \sim m_\pi$, typical momentum in nucleus

Expansion only valid for small expansion parameter, *i.e.*, low momentum

Irreducible time-ordered diagram has order $\left(\frac{Q}{\Lambda_\chi}\right)^v$, where

$$v = -4 + 2N + 2L + \sum_i V_i \Delta_i \quad \Delta_i = d_i + \frac{1}{2} n_i - 2 \quad \text{“Chiral dimension”}$$

N = Number of nucleons

L = Number of pion loops

V_i = Number of vertices of type i

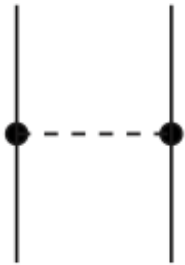
d = Number of derivatives or insertions of

n = Number of nucleon field operators m_π

Chiral EFT: Lowest Order (LO)

Step V: Calculate Feynmann diagrams to the desired accuracy

Leading order ($\nu = 0$)



One-pion exchange

$$V_{NN}^{(0)} = - \frac{g_A^2}{4F_\pi^2} \frac{\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q}}{\vec{q}^2 + M_\pi^2} \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2$$

$$\vec{q}_i \equiv \vec{p}'_i - \vec{p}_i$$

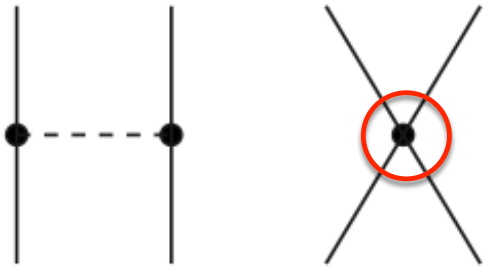
$$g_A = 1.26$$

$$F_\pi = 92.4 \text{ MeV}$$

Chiral EFT: Lowest Order (LO)

Step V: Calculate Feynmann diagrams to the desired accuracy

Leading order ($\nu = 0$)



One-pion exchange
NN contact interaction

$$V_{NN}^{(0)} = -\frac{g_A^2}{4F_\pi^2} \frac{\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q}}{\vec{q}^2 + M_\pi^2} \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2 + \textcircled{C_S} + \textcircled{C_T} \vec{\sigma}_1 \cdot \vec{\sigma}_2$$

$$\vec{q}_i \equiv \vec{p}'_i - \vec{p}_i$$

$$g_A = 1.26$$

$$F_\pi = 92.4 \text{ MeV}$$

Two **low-energy constants (LECs)**: C_S, C_T

Chiral EFT

Step V: Calculate Feynmann diagrams to the desired accuracy

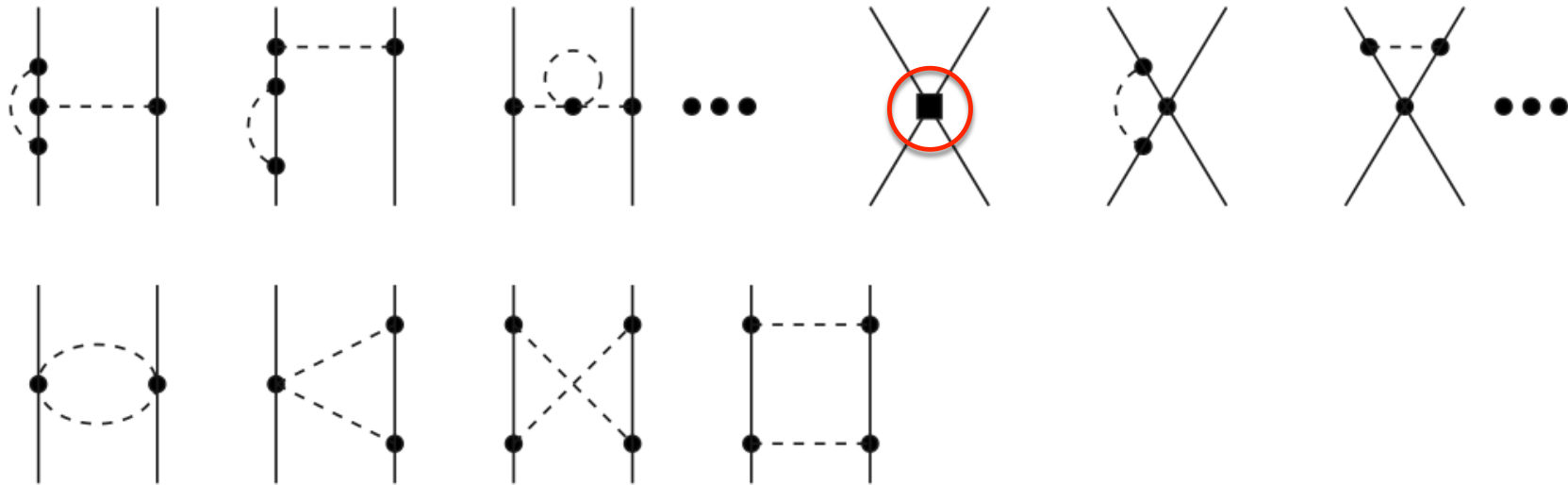
Question: What will $\nu = 1$ look like?

Answer: No contribution at this order

Chiral EFT: NLO

Step V: Calculate Feynmann diagrams to the desired accuracy

Next-to-leading order ($\nu = 2$)



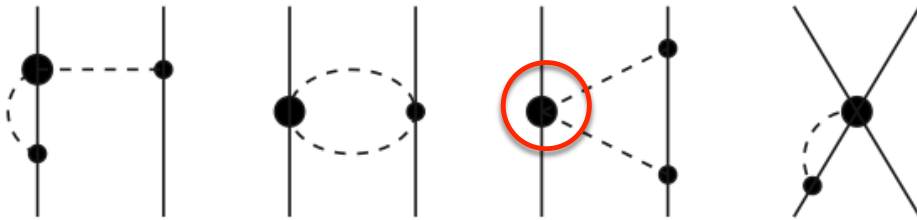
Higher order contact interaction: 7 new LECs, spin-orbit

$$\begin{aligned}
 &+ \mathcal{C}_1 \vec{q}^2 + \mathcal{C}_2 \vec{k}^2 + (\mathcal{C}_3 \vec{q}^2 + \mathcal{C}_4 \vec{k}^2) \vec{\sigma}_1 \cdot \vec{\sigma}_2 \\
 &+ i \mathcal{C}_5 \frac{1}{2} (\vec{\sigma}_1 + \vec{\sigma}_2) \cdot \vec{q} \times \vec{k} + \mathcal{C}_6 \vec{q} \cdot \vec{\sigma}_1 \vec{q} \cdot \vec{\sigma}_2 \\
 &+ \mathcal{C}_7 \vec{k} \cdot \vec{\sigma}_1 \vec{k} \cdot \vec{\sigma}_2,
 \end{aligned}$$

Chiral EFT: N²LO

Step V: Calculate Feynmann diagrams to the desired accuracy

Next-to-next-to-leading order ($\nu = 3$)



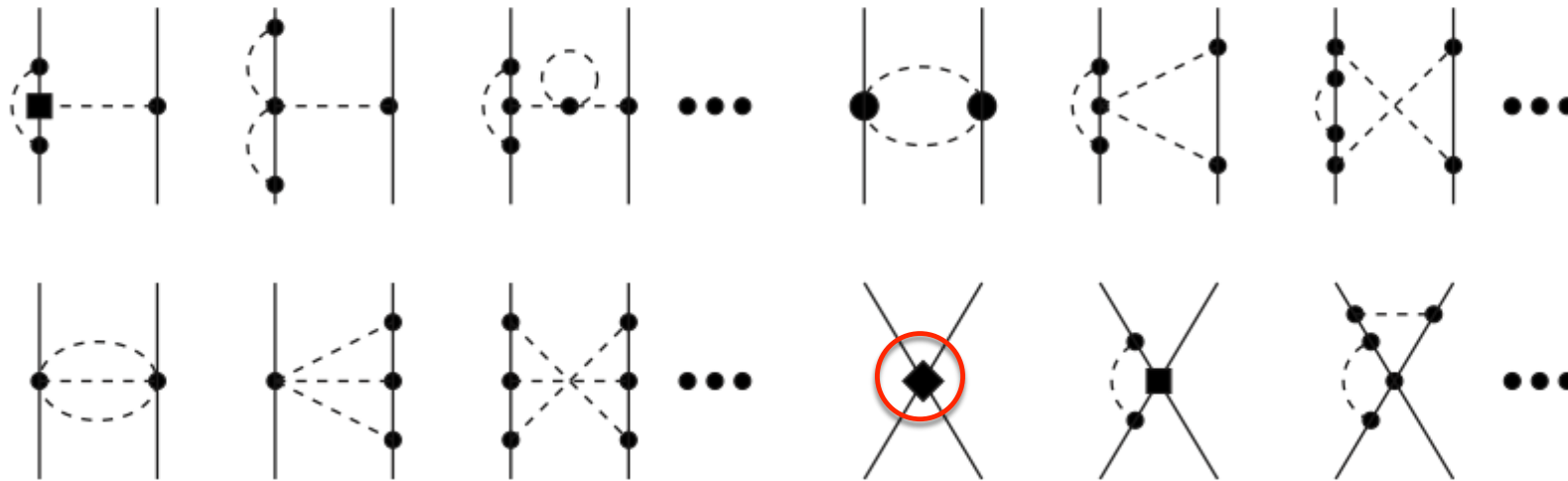
3 new LECs from $\pi\pi NN$ vertex

$$\begin{aligned}
 V_{NN}^{(3)} = & -\frac{3g_A^2}{16\pi F_\pi^4} [2M_\pi^2 (2c_1 - c_3) - c_3 \vec{q}^2] \\
 & \times (2M_\pi^2 + \vec{q}^2) A^{\tilde{\Lambda}}(q) - \frac{g_A^2 c_4}{32\pi F_\pi^4} \tau_1 \cdot \tau_2 (4M_\pi^2 \\
 & + q^2) A^{\tilde{\Lambda}}(q) (\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q} - \vec{q}^2 \vec{\sigma}_1 \cdot \vec{\sigma}_2),
 \end{aligned}$$

Chiral EFT: N³LO

Step V: Calculate Feynmann diagrams to the desired accuracy

Next-to-next-to-next-to-leading order ($\nu = 4$)



Higher order contact interaction: 15 new LECs

Regularization of Chiral potentials

Remember: constructing potential involves solving **L-S equation**

All NN potentials cutoff loop momenta at some value $> 1\text{GeV}$

Impose exponential regulator, Λ , in Chiral EFT potentials – **not in integral**

$$\hat{T}(\vec{p}', \vec{p}) = \hat{V}(\vec{p}', \vec{p}) + \int d^3 p'' \hat{V}(\vec{p}', \vec{p}'') \frac{M}{p^2 - p''^2 + i\epsilon} \hat{T}(\vec{p}'', \vec{p})$$

$$\hat{V}(\vec{p}', \vec{p}) \longmapsto \hat{V}(\vec{p}', \vec{p}) e^{-(p'/\Lambda)^{2n}} e^{-(p/\Lambda)^{2n}}$$

LECs will depend on regularization approach and Λ

Infinitely many ways to do this

Infinitely many chiral potentials

Indeed, many on the market – some fit well to phase shifts, others not

Chiral EFT: Resulting fits to Phase shifts

Systematic improvement of chiral EFT potentials fit to phase shifts

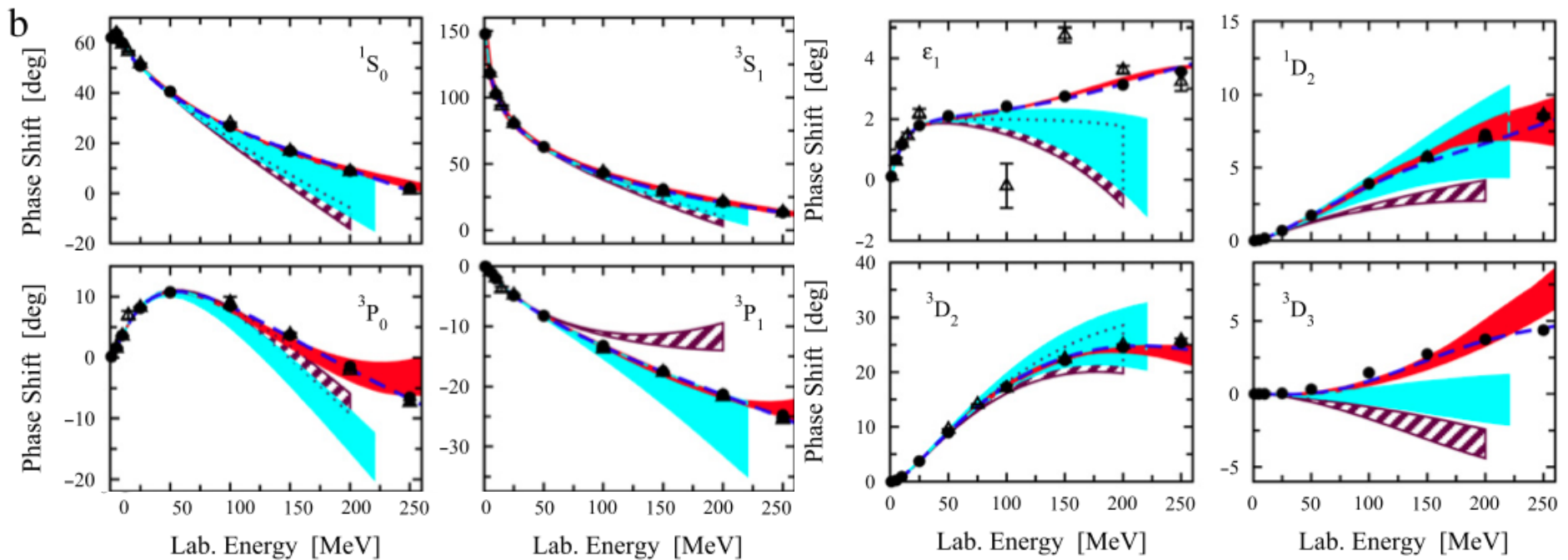
Cutoff variation – information about missing physics

NLO: dashed band  9 Parameters

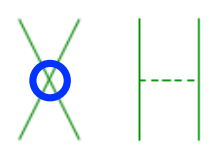


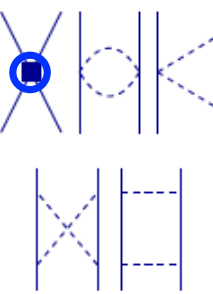


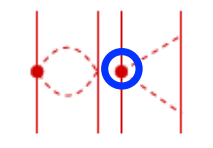
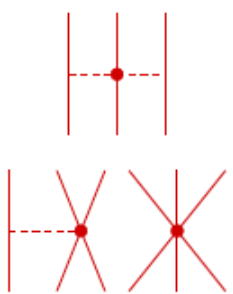

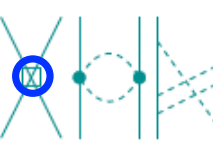

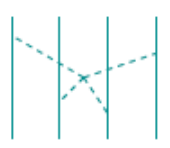
N²LO: light band  12 Parameters

N³LO: dark band  27 Parameters

Generally decreasing error and increasing accuracy – not entirely...



Chiral Effective Field Theory: Nuclear Forces

	2N forces	3N forces	4N forces
LO			
NLO			
N ² LO			
N ³ LO			
	+ ...	+ ...	+ ...

Meson exchange potentials were an admirable effort

Using ideas of **effective field theory**:

Nucleons interact via pion exchanges and contact interactions

Lower momentum

Systematic – can assign error

Connected to QCD

Hierarchy: $V_{NN} > V_{3N} > \dots$

Consistent treatment of NN, 3N, ... electroweak operators

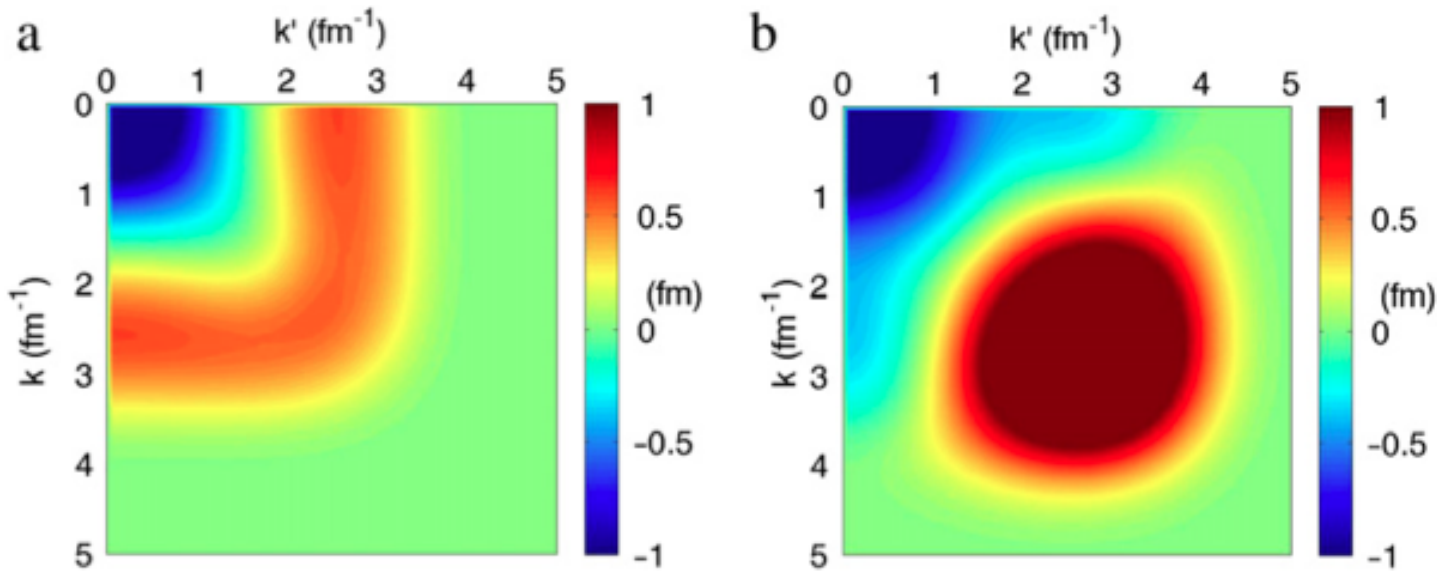
Couplings fit to experiment once

Weinberg, van Kolck, Kaplan, Savage, Wise, Epelbaum, Kaiser, Meissner, ...

Chiral NN Potentials

Two chiral potentials with regulators of 500MeV and 600MeV

Still low-to-high momentum coupling: poor convergence, non perturbative, etc.



How do these compare to the potential you drew?

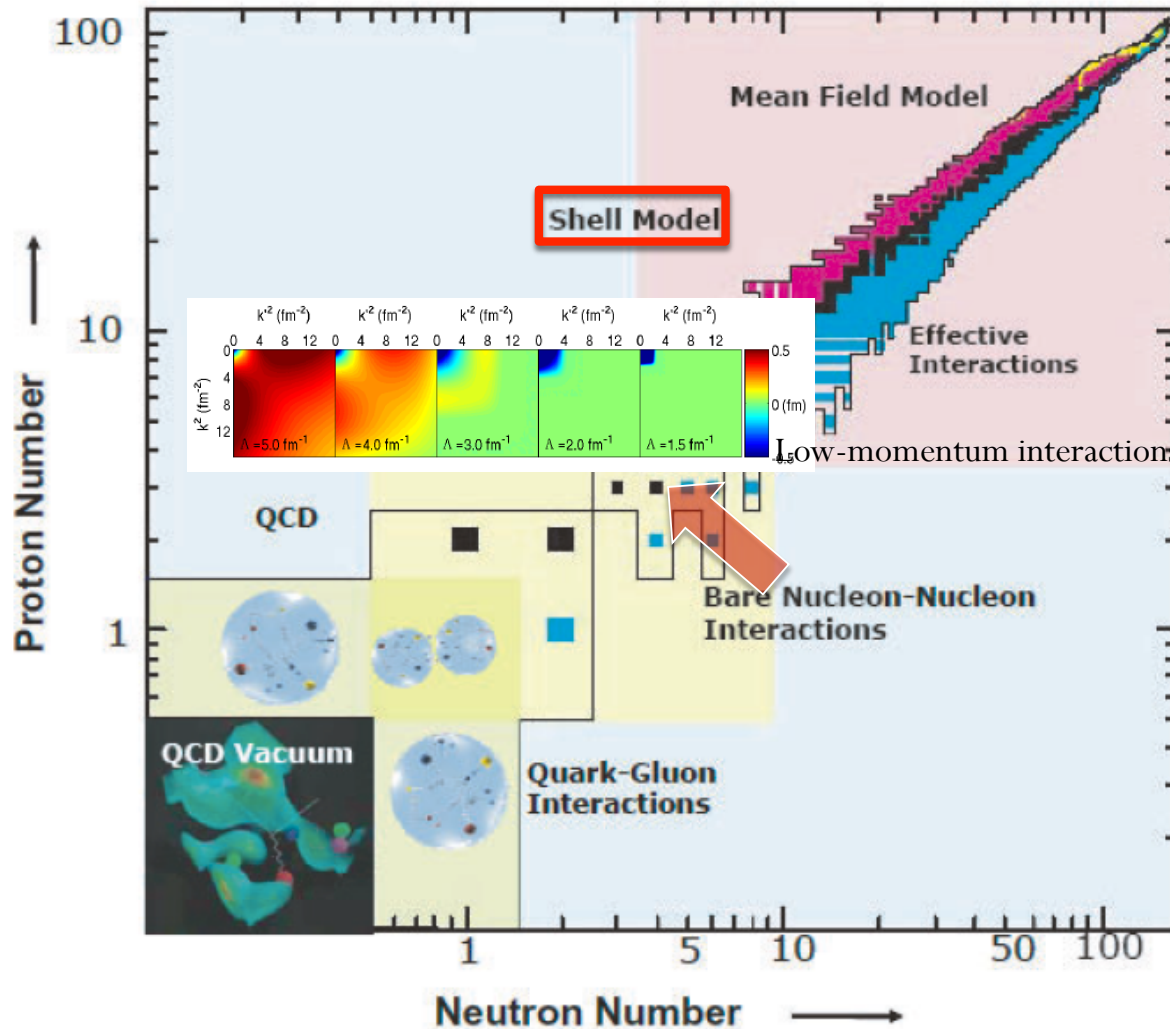
Lesson: Infinitely many phase-shift equivalent potentials

$$E_n = \langle \Psi_n | H | \Psi_n \rangle = (\langle \Psi_n | U^\dagger) U H U^\dagger (U | \Psi_n \rangle) = \langle \tilde{\Psi}_n | \tilde{H} | \tilde{\Psi}_n \rangle$$

NN interaction not observable Low-to-high momentum makes life difficult for low-energy nuclear theorists

Part II: RG and Low-Momentum Interactions

To understand the properties of complex nuclei from elementary interactions



Renormalizing NN Interactions

Basic ideas of RG

Calculating low-momentum interactions

Benefits of low cutoffs

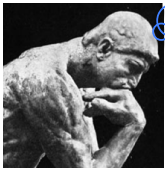
G-matrix renormalization

How will we approach this problem:

QCD → NN (3N) forces → Renormalize → Solve many-body problem → Predictions

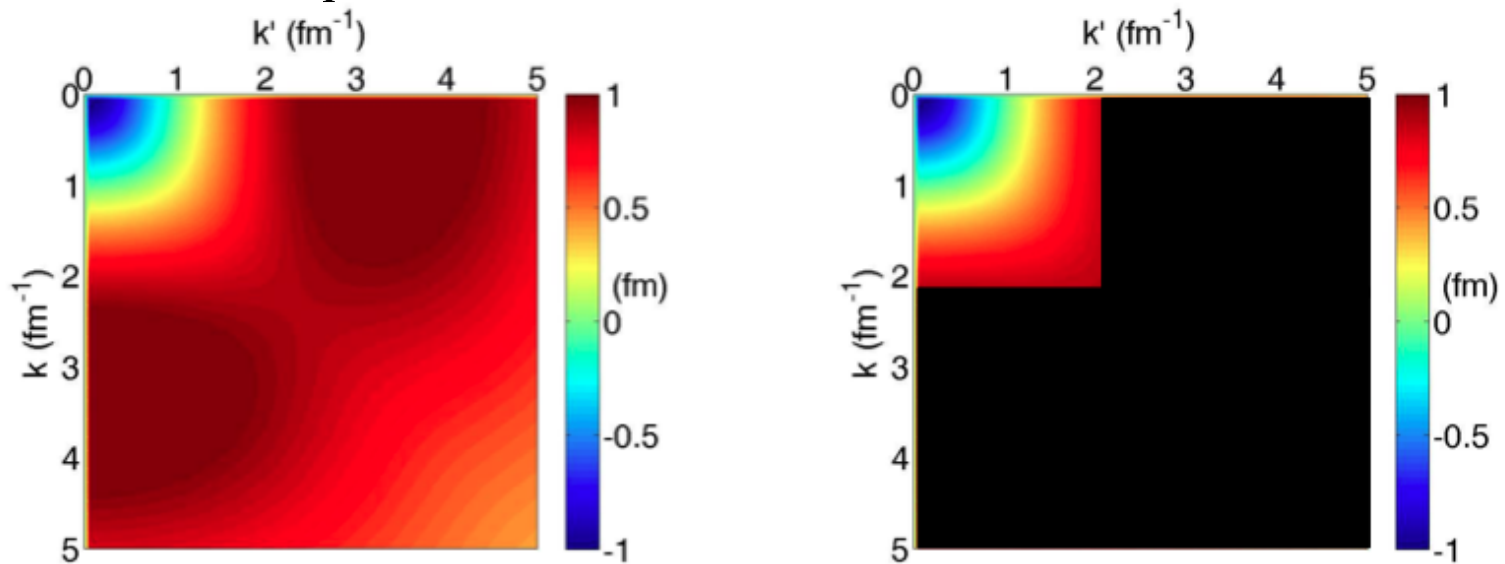
Renormalization of Meson-Exchange Potentials

Ok, high momentum is a pain. I wonder what would happen to low-energy observables...



Low-to-high momentum makes life difficult for low-energy nuclear theorists

Can we just make a sharp cut and see if it works?



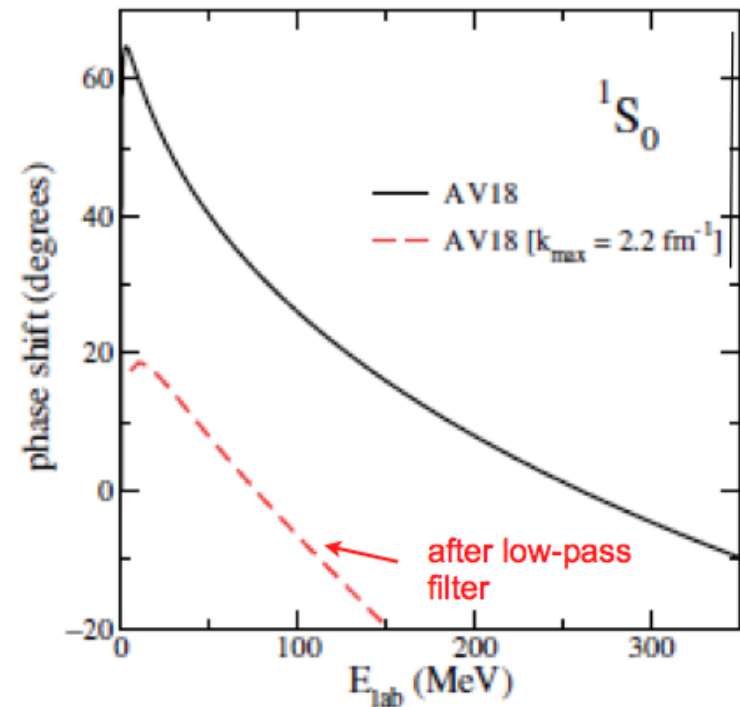
$$V_{filter}(k', k) \equiv 0 \quad k, k' > 2.2 \text{ fm}^{-1}$$

Renormalization of Meson-Exchange Potentials

Can we just make a sharp cut without renormalizing?

Low-energy physics is not correct

Lesson: Must ensure low-energy physics is preserved



Low and high k are coupled by quantum fluctuations (virtual states)

$$\langle k|V|k'\rangle + \sum_{q=0}^{\Lambda} \frac{\langle k|V|q\rangle\langle q|V|k'\rangle}{\epsilon_{k'} - \epsilon_q} + \sum_{q=\Lambda}^{\infty} \frac{\langle k|V|q\rangle\langle q|V|k'\rangle}{\epsilon_{k'} - \epsilon_q}$$

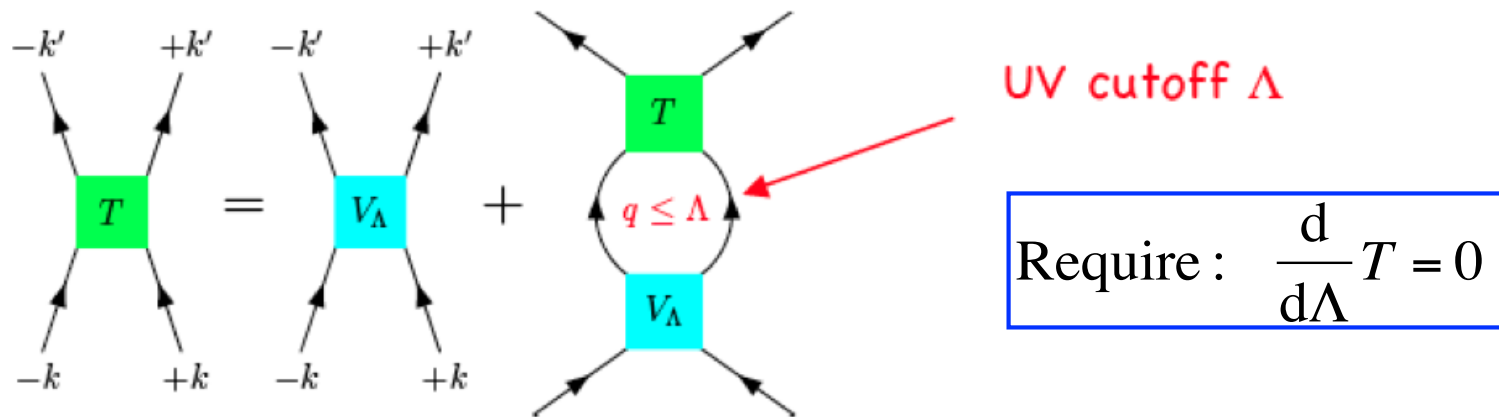
Can't simply drop **high q** without changing low k observables.

Renormalization of Meson-Exchange Potentials

To do properly: from T -matrix equation, define **low-momentum** equation:

$$T(k', k; k^2) = V_{\text{NN}}(k', k) + \frac{2}{\pi} \mathcal{P} \int_0^{\Lambda_\infty} \frac{V_{\text{NN}}(k', p) T(p, k; k^2)}{k^2 - p^2} p^2 dp,$$

$$= V_{\text{low } k}^\Lambda(k', k) + \frac{2}{\pi} \mathcal{P} \int_0^\Lambda \frac{V_{\text{low } k}^\Lambda(k', p) T(p, k; k^2)}{k^2 - p^2} p^2 dp,$$

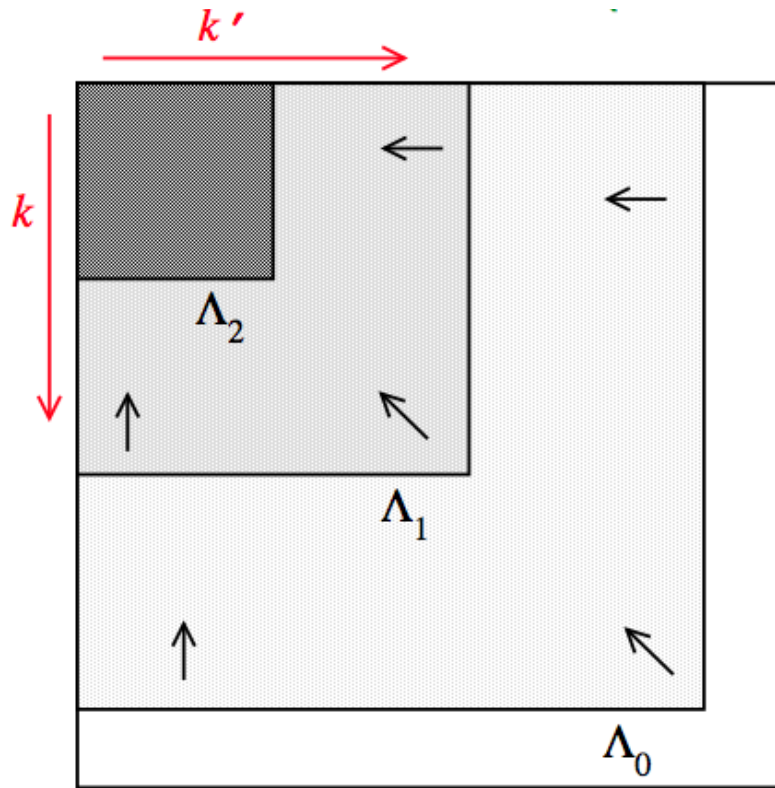


Leads to **renormalization group** equation for low-momentum interaction

$$\frac{d}{d\Lambda} V_{\text{low } k}^\Lambda(k', k) = \frac{2}{\pi} \frac{V_{\text{low } k}^\Lambda(k', \Lambda) T^\Lambda(\Lambda, k; \Lambda^2)}{1 - (k/\Lambda)^2}$$

Renormalization of Meson-Exchange Potentials

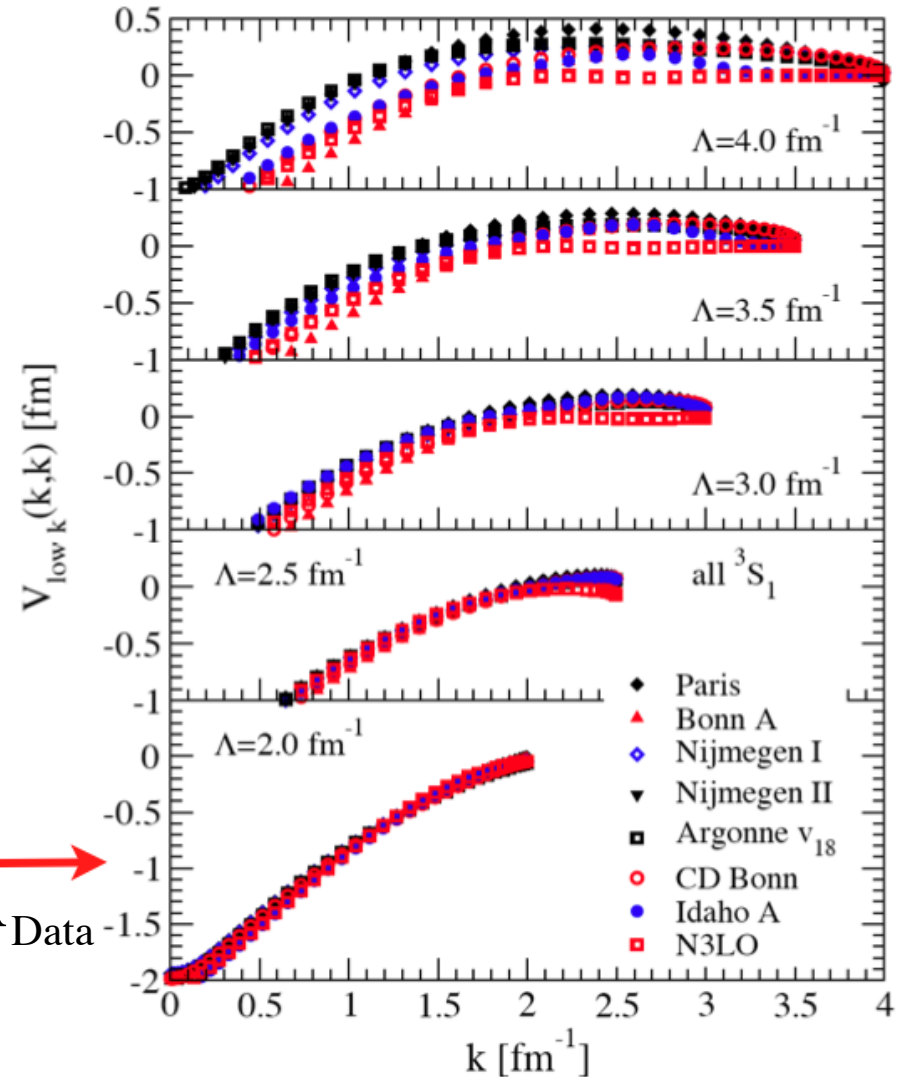
Run cutoff to lower values – decouples high-momentum modes



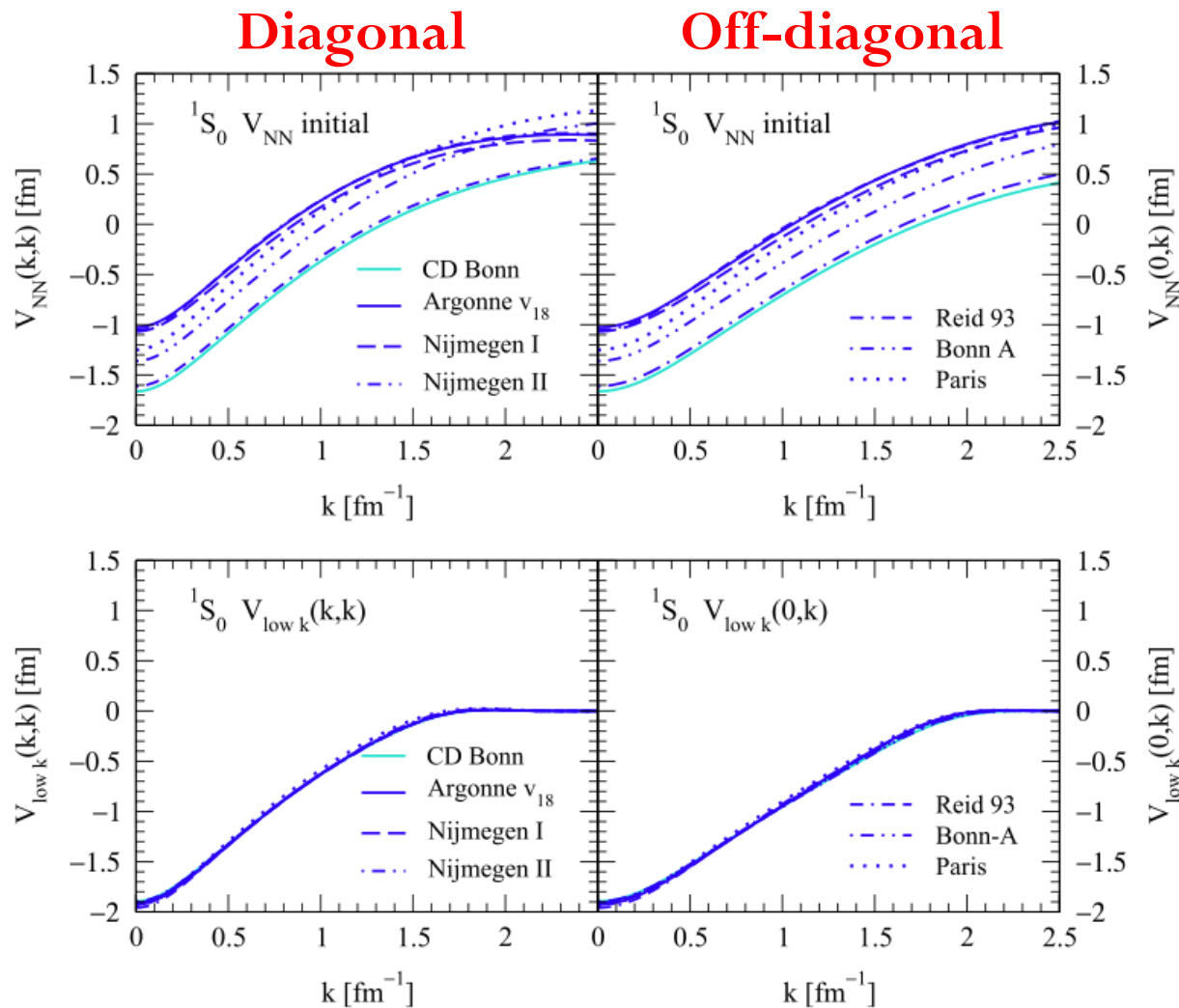
Start from some “bare” NN potential at high cutoff

Universality at low momentum

$$\Lambda \approx \Lambda_{\text{Data}}$$



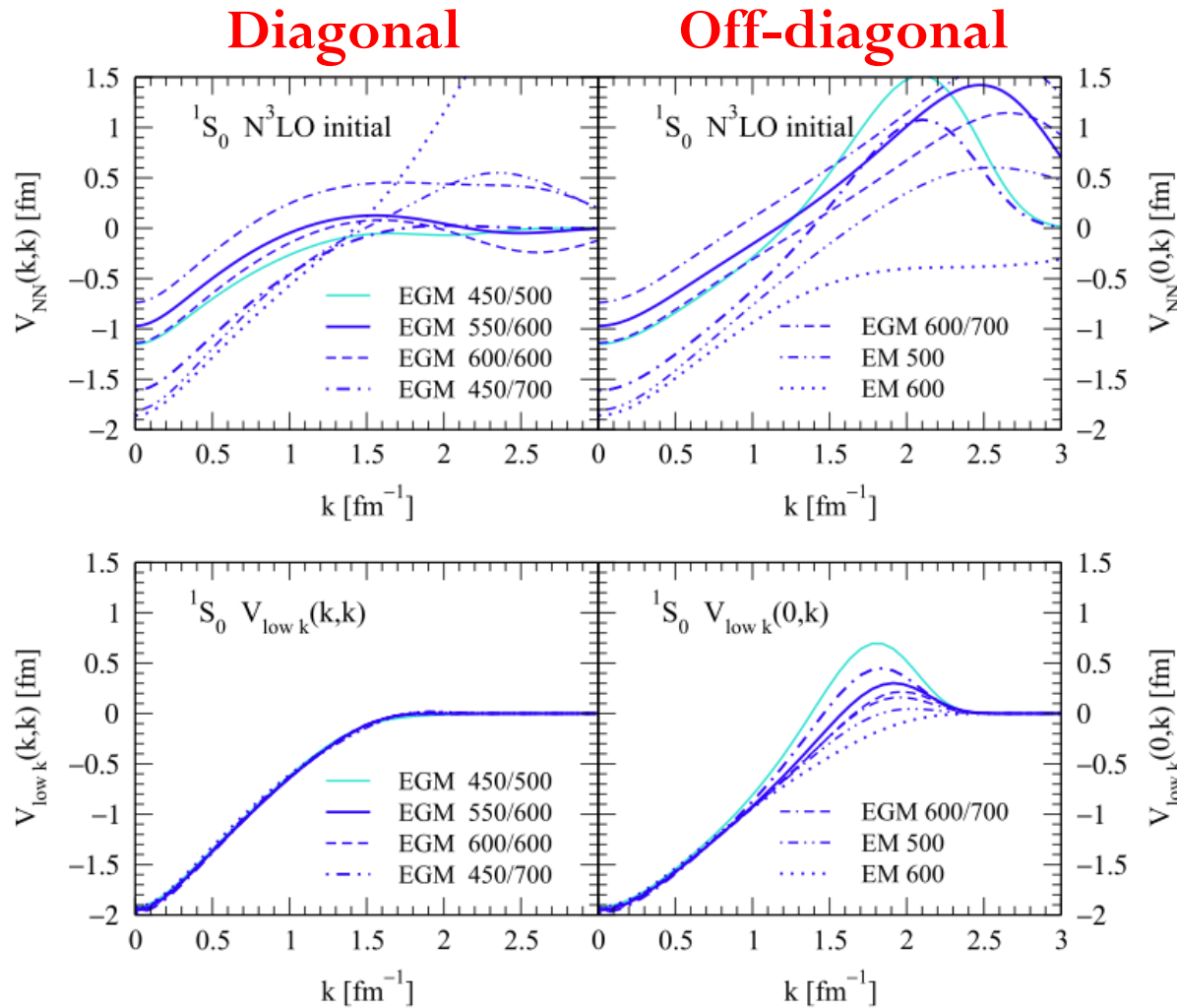
Renormalization of Meson-Exchange Potentials



These are all our favorite OBE NN potentials...

These are all our favorite OBE NN potentials...
at low momentum

Renormalization of Chiral EFT Potentials

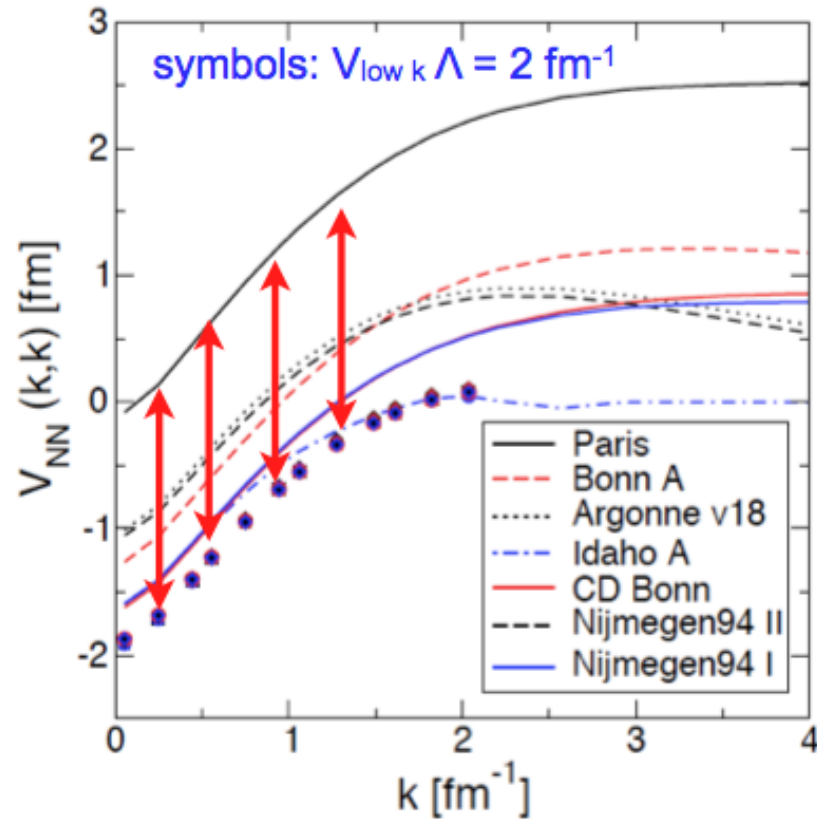
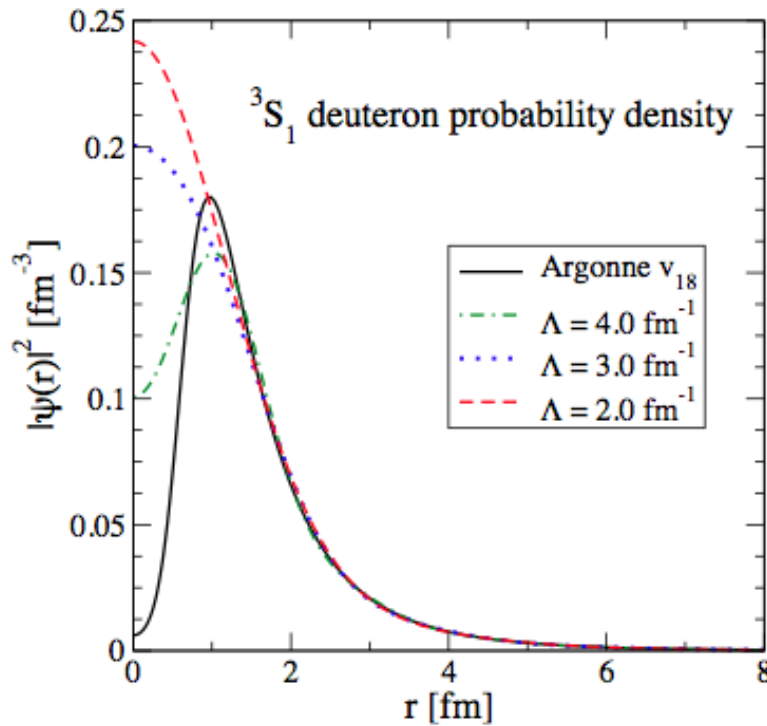


These are all our favorite Chiral EFT NN potentials...

These are all our favorite Chiral EFT NN potentials...
at low momentum

Differences remain in off-diagonal matrix elements
Sensitive to agreement for phase shifts (not all fit perfectly)

Renormalization of NN Potentials



$$V_{\text{eff}} = V_L + \delta V_{c.t.}(\Lambda)$$

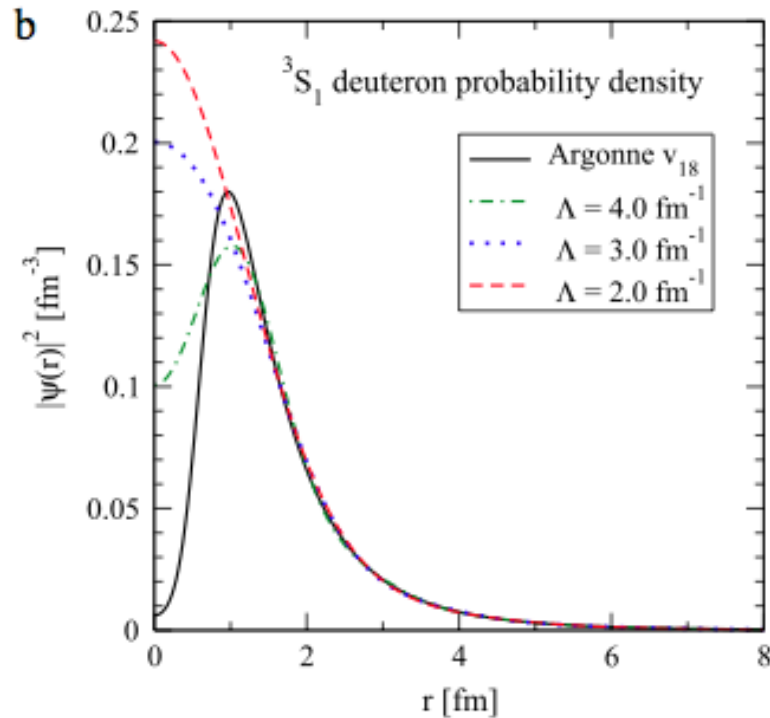
Long-range tail of deuteron wavefunction preserved

Main effect is shift in momentum space – delta function

Removes hard core!

Renormalization of Nuclear Interactions

Short-distance behaviour of the deuteron – striking difference between potentials



But what's **really**
going on at short
distances?



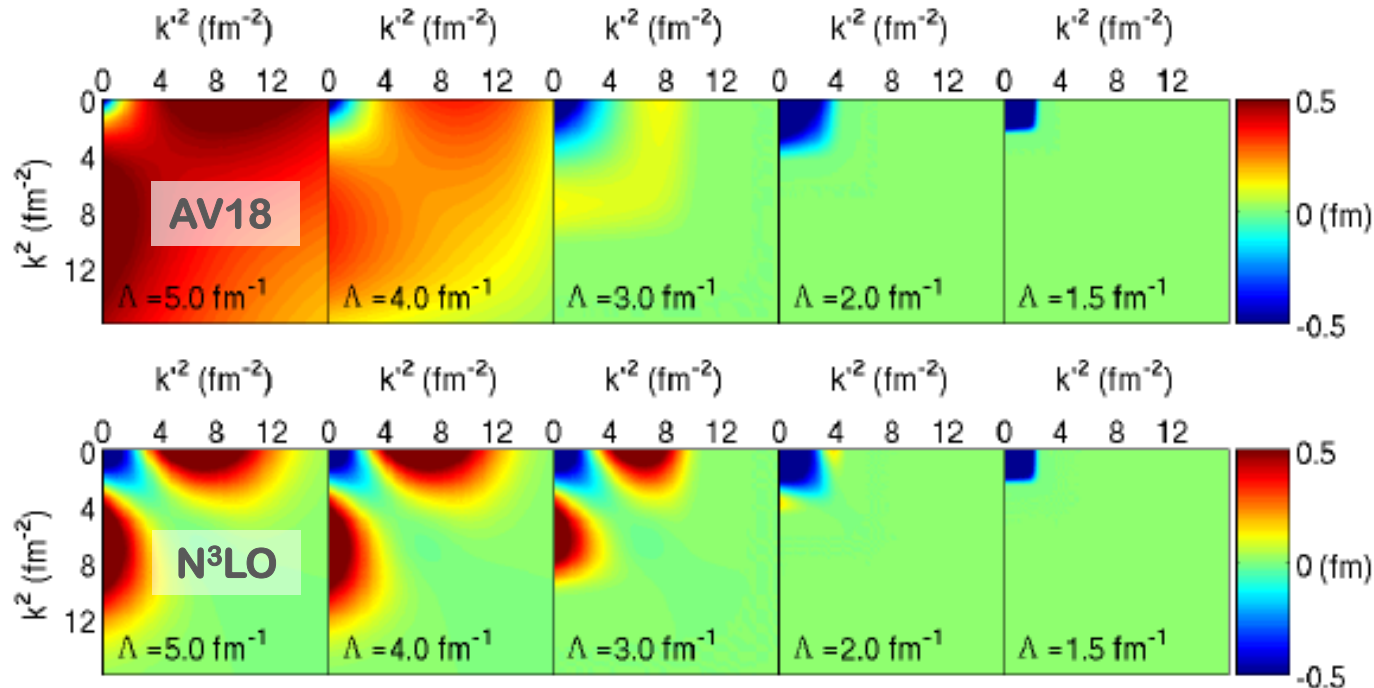
- A) Argonne is correct: Short range repulsion prohibits nucleons from
- B) Vlowk is correct: the nucleons really will overlap in space
- C) Some superposition of these
- D) It doesn't matter

Renormalization of Nuclear Interactions

$$H(\Lambda) = T + V_{\text{NN}}(\Lambda) + V_{\text{3N}}(\Lambda) + V_{\text{4N}}(\Lambda) + \dots$$

Evolve momentum resolution scale of chiral interactions from initial Λ_χ
 Remove coupling to high momenta, low-energy physics unchanged

Bogner, Kuo, Schwenk, Furnstahl



Universal at
 low-momentum

$V_{\text{low } k}(\Lambda)$: lower cutoffs advantageous for nuclear structure calculations

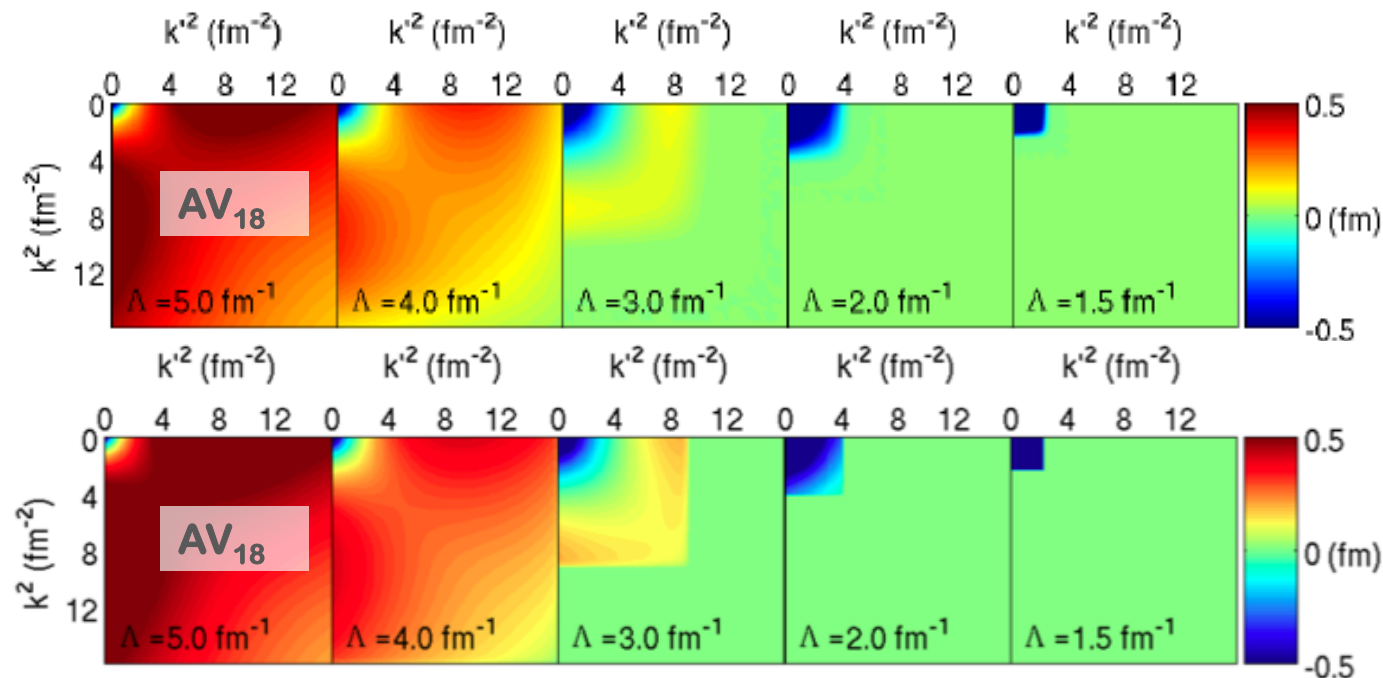
Smooth vs. Sharp Cutoffs

$$H(\Lambda) = T + V_{\text{NN}}(\Lambda) + V_{\text{3N}}(\Lambda) + V_{\text{4N}}(\Lambda) + \dots$$

Can have sharp as well as smooth cutoffs (codes only do sharp)

Remove coupling to high momenta, low-energy physics unchanged

Bogner, Kuo, Schwenk, Furnstahl



Similar but not exact same results – will be differences in calculations

Benefits of Lower Cutoffs

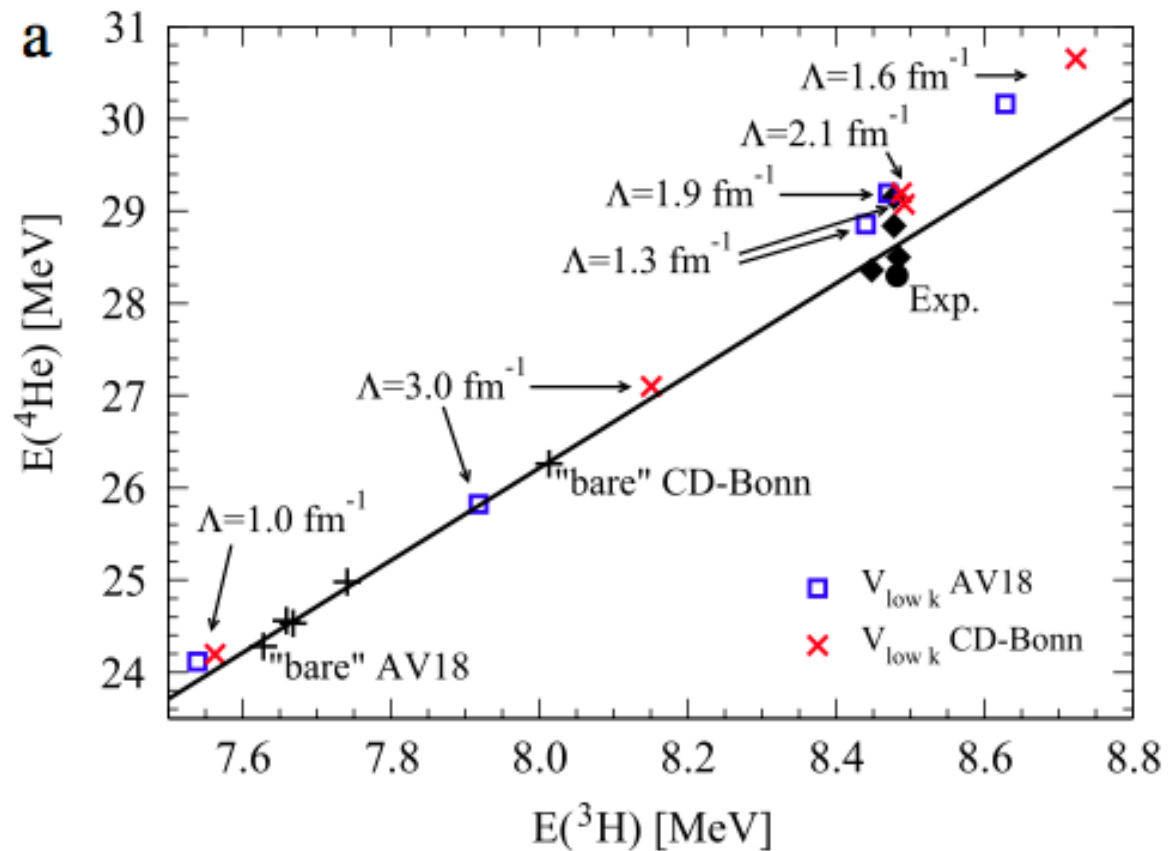
Also use cutoff dependence to assess missing physics: return to Tjon line

Varying cutoff moves along line

Never breaks off to experiment

Lesson: Variation in physical observables with cutoff denotes missing physics beyond NN

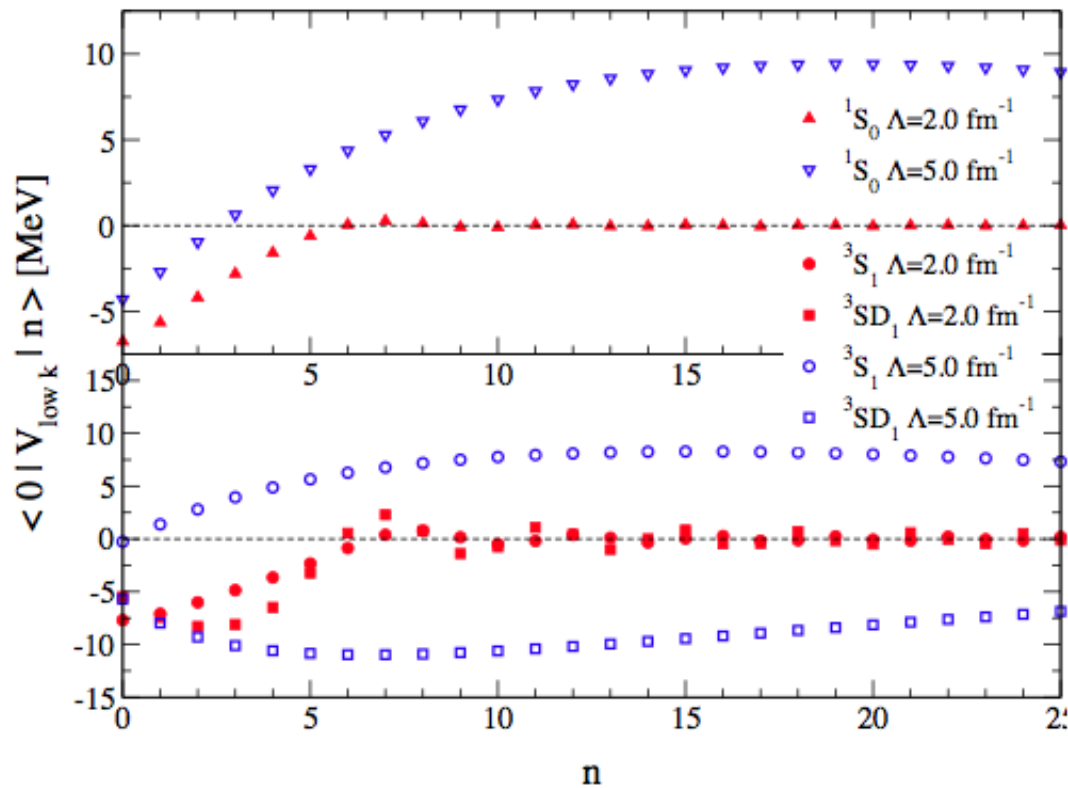
Tool not a parameter!



Benefits of Lower Cutoffs

Removes coupling from low-to-high harmonic oscillator states

Expect to speed convergence in HO basis

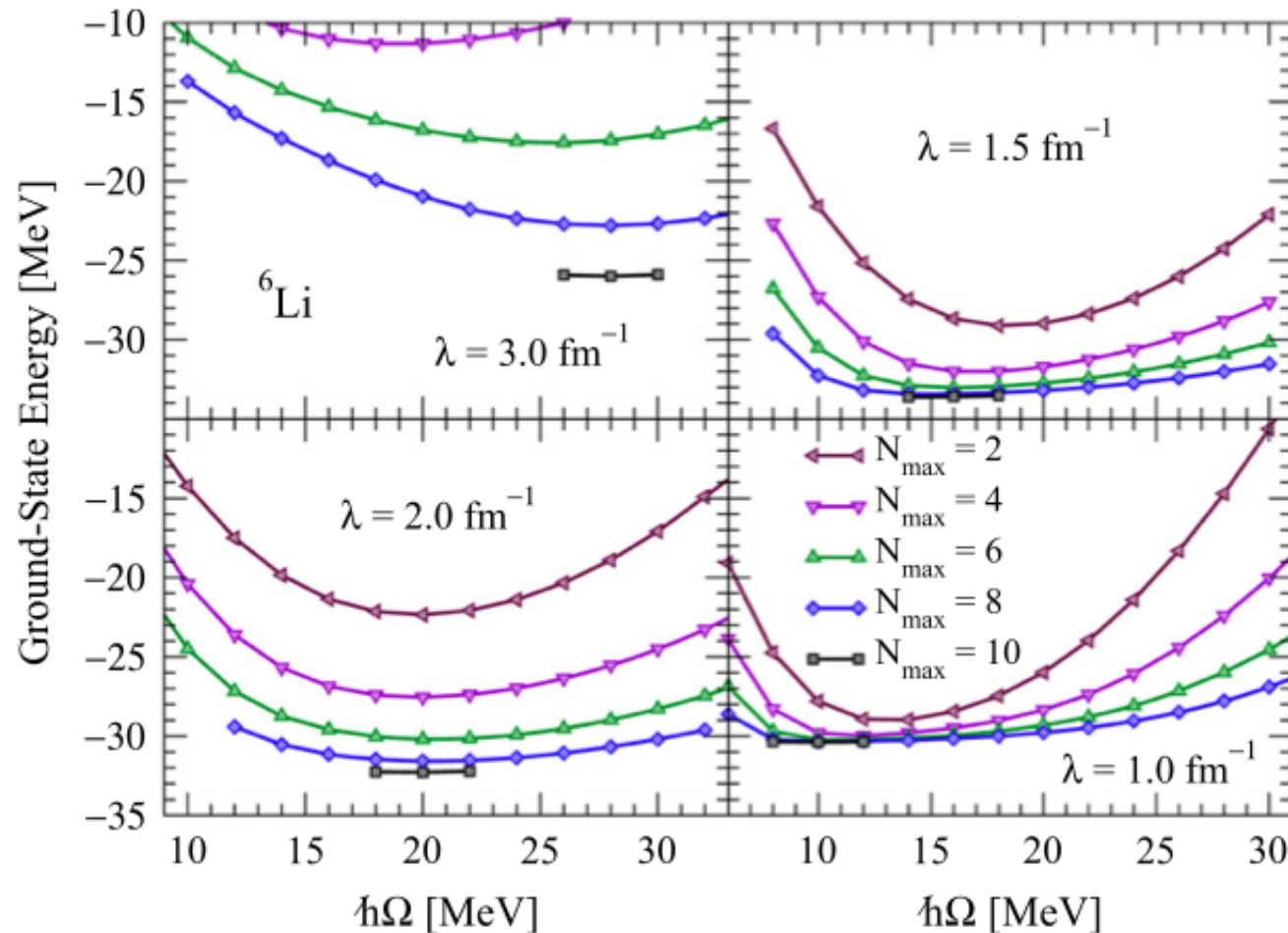


Benefits of Lower Cutoffs

Exactly what happens in no-core shell model calculations

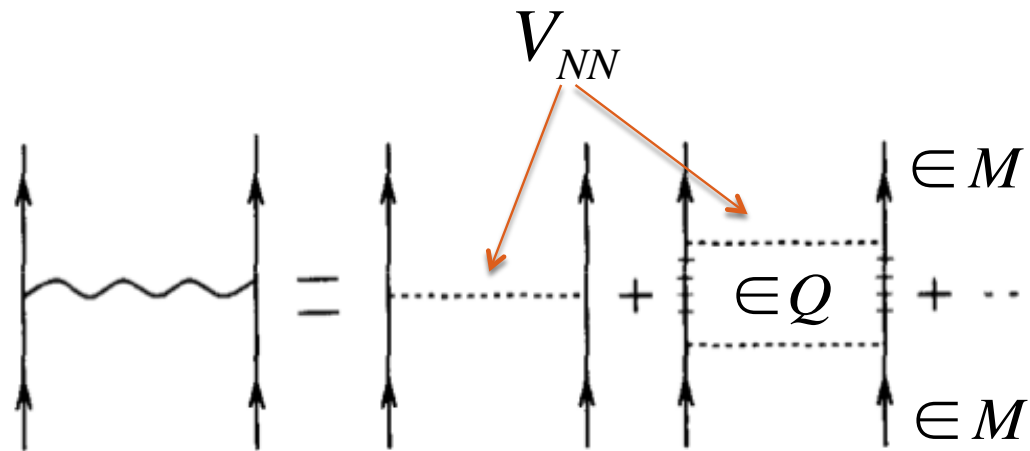
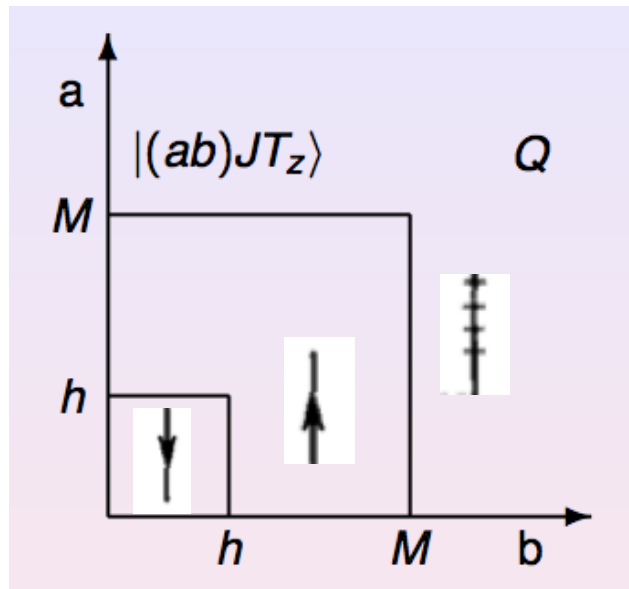
Probably equally helpful in normal shell model calculations

Come back to this later...



G-matrix Renormalization

Standard method for softening interaction in nuclear structure for decades:



Infinite summation of ladder diagrams

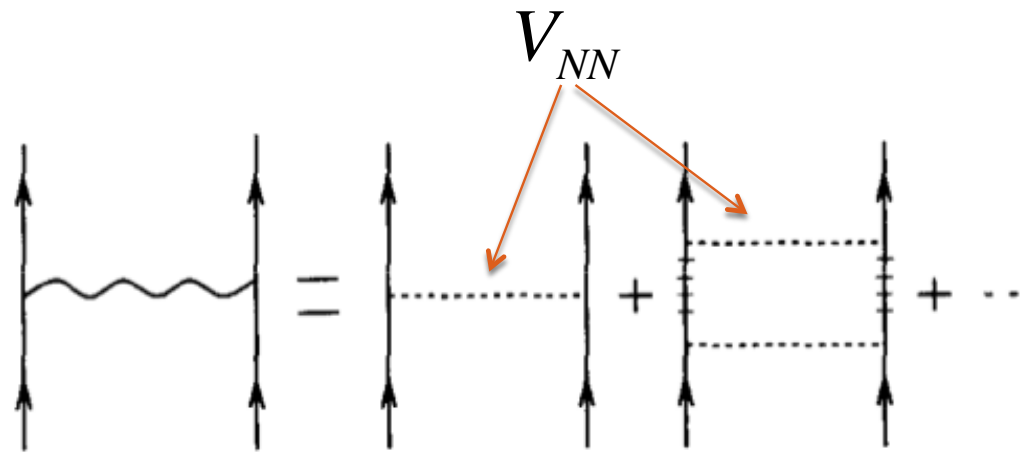
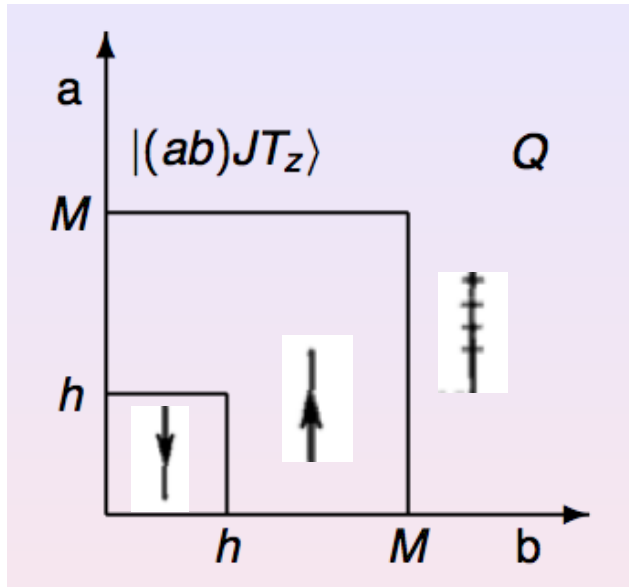
Need two model spaces:

- 1) **M** space in which we will want to calculate (excitations allowed in M)
- 2) Large space **Q** in which particle excitations are allowed

To avoid double counting, can't overlap – **matrix elements depend on M**

G-matrix Renormalization

Standard method for softening interaction in nuclear structure for decades:



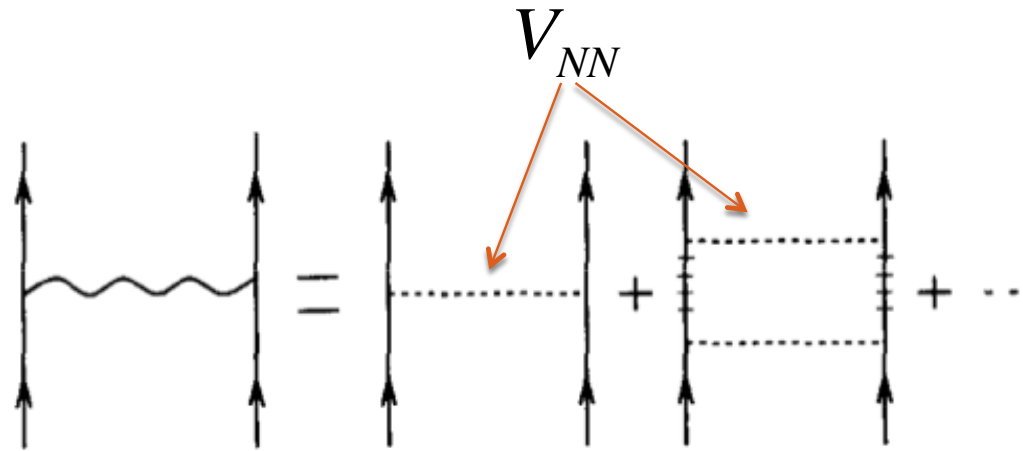
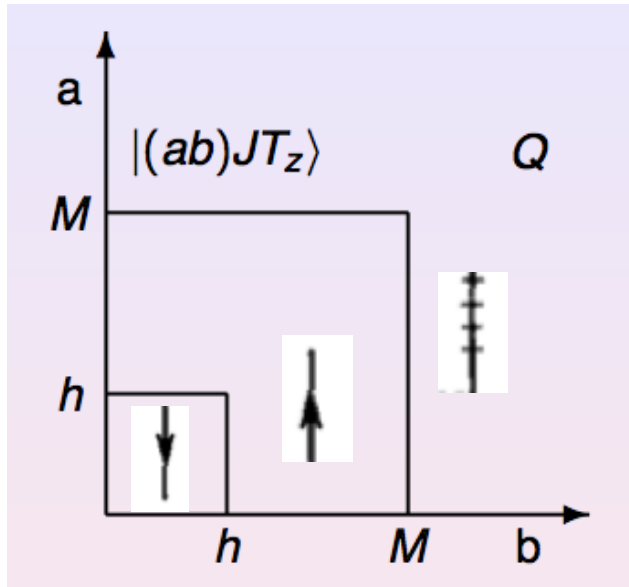
$$G_{ijkl}(\omega) = V_{ijkl} + \sum_{mn \in Q} V_{ijmn} \frac{Q}{\omega - \varepsilon_m - \varepsilon_n} G_{mnkl}(\omega)$$

Iterative procedure

Dependence on arbitrary starting energy!

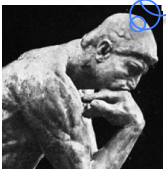
G-matrix Renormalization

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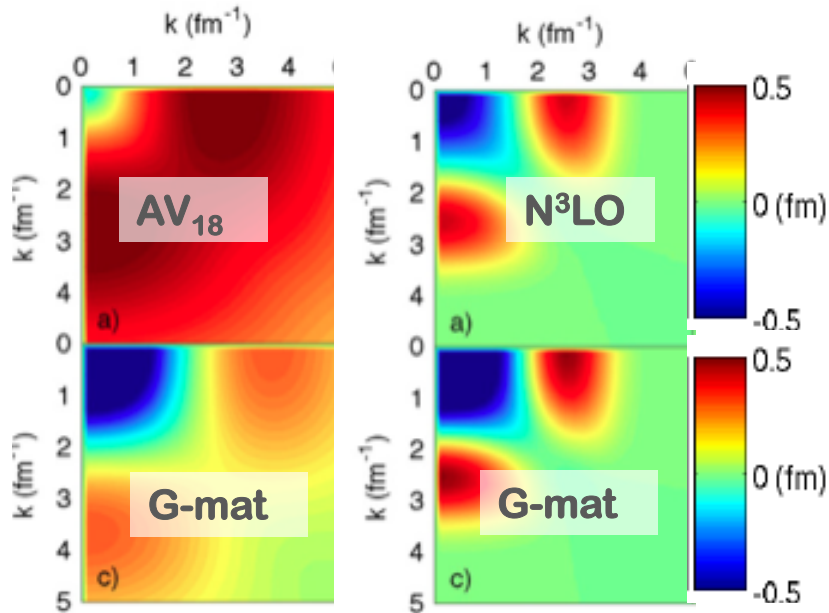
$$G_{ijkl}(\omega) = V_{ijkl} + \sum_{mn \in Q} V_{ijmn} \frac{Q}{\omega - \varepsilon_m - \varepsilon_n} G_{mnkl}(\omega)$$

What happens as we keep increasing M?



G-matrix Renormalization

Results of **G-matrix** renormalization:



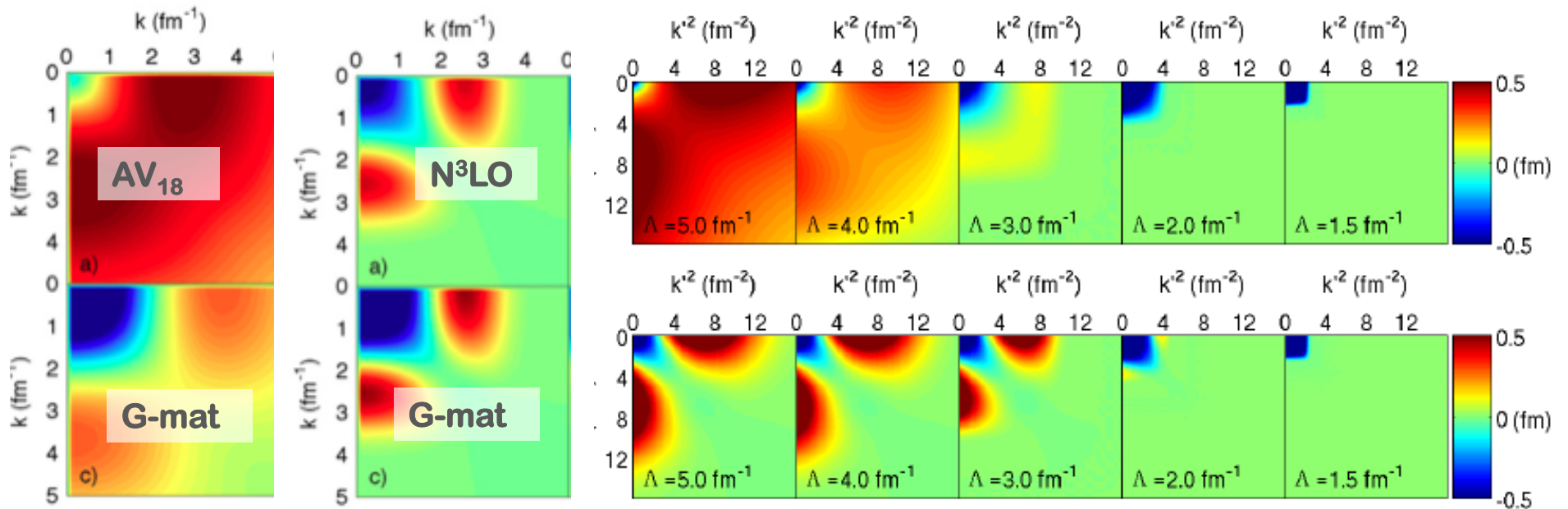
Removes some diagonal high-momentum components

Still large low-to-high coupling in both interactions

No indication of universality

G-matrix Renormalization

Results of **G-matrix** renormalization:



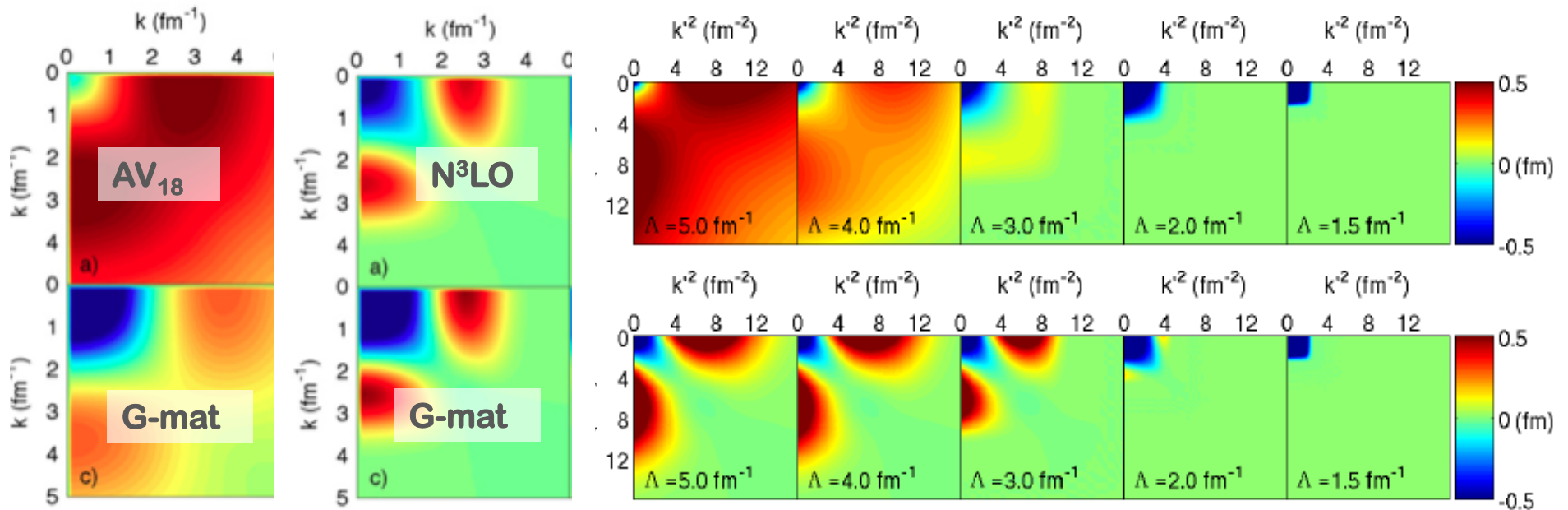
Removes some diagonal high-momentum components

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Summary

Low-momentum interactions can be constructed from any V_{NN} via RG



Low-to-high momentum coupling not desirable in low-energy nuclear physics

Evolve to low-momentum while preserving low-energy physics

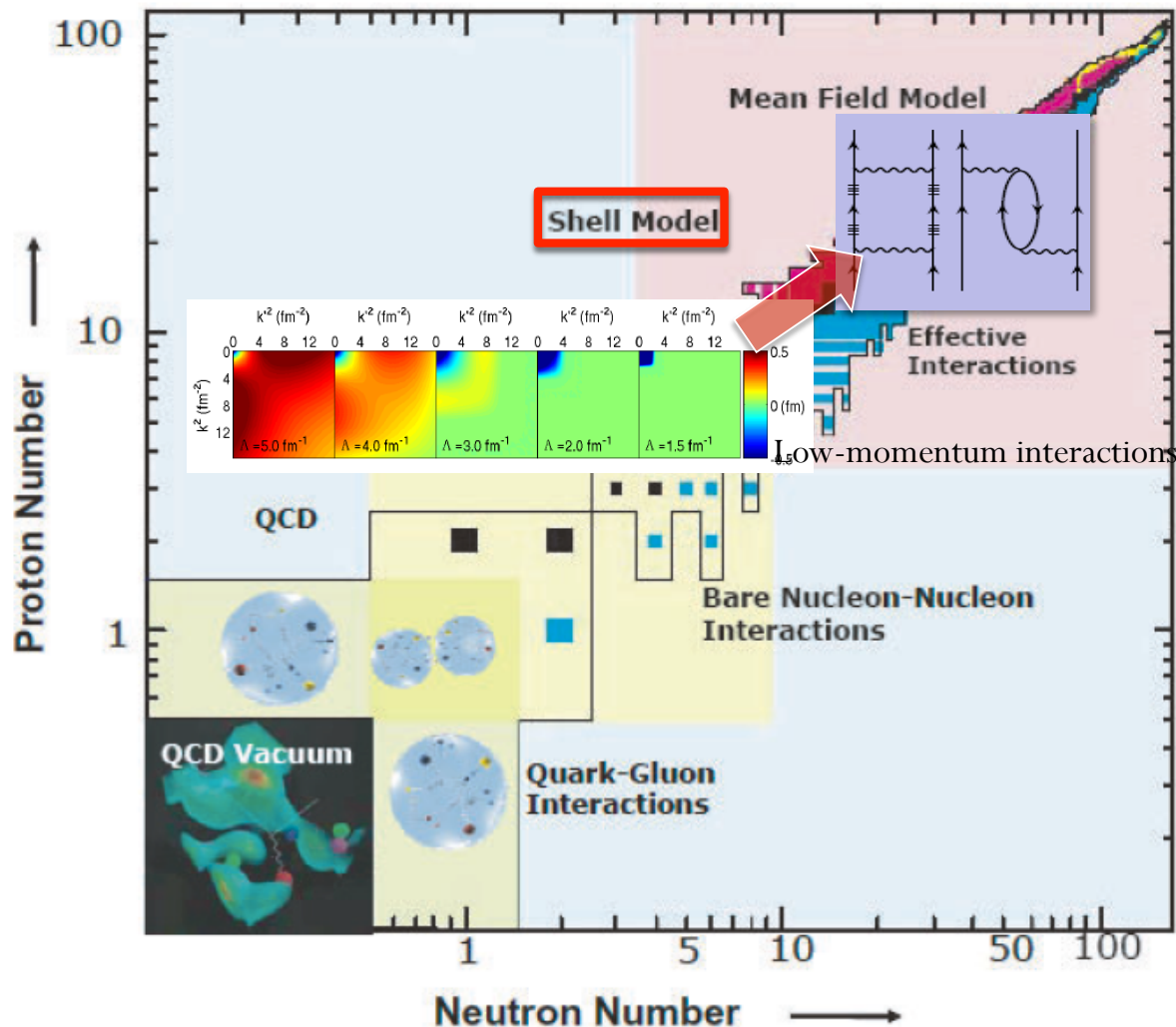
Universality attained near cutoff of data

Low-momentum cutoffs remove low-to-high harmonic oscillator couplings

Cutoff variation assesses missing physics at the level of interactions: tool not a parameter

Part III: Many-Body Perturbation Theory

To understand the properties of complex nuclei from elementary interactions



Microscopic Valence-Space Interactions

Model spaces

Many-body perturbation theory (MBPT)

Calculating effective interaction

Monopole part of interaction

Deficiencies of this approach

How will we approach this problem:

QCD → NN (3N) forces → **Renormalize → Solve many-body problem** → Predictions

Solving the Many-Body Problem

Matrix elements now given in momentum space, partial waves

$$\langle kK, lL; JST | V | k' K, l' L; JS' T \rangle$$

To go to finite nuclei begin from Hamiltonian

$$H|\psi\rangle = (T + V)|\psi\rangle = E|\psi\rangle$$

Assume many particles in the nucleus generate a **mean field** U :

U a one-body potential simple to solve (typically Harmonic Oscillator)

$$H = H_0 + H_1 \quad H_0 = T + U \quad H_1 = V - U$$

So transform from momentum space to **Harmonic Oscillator Basis**

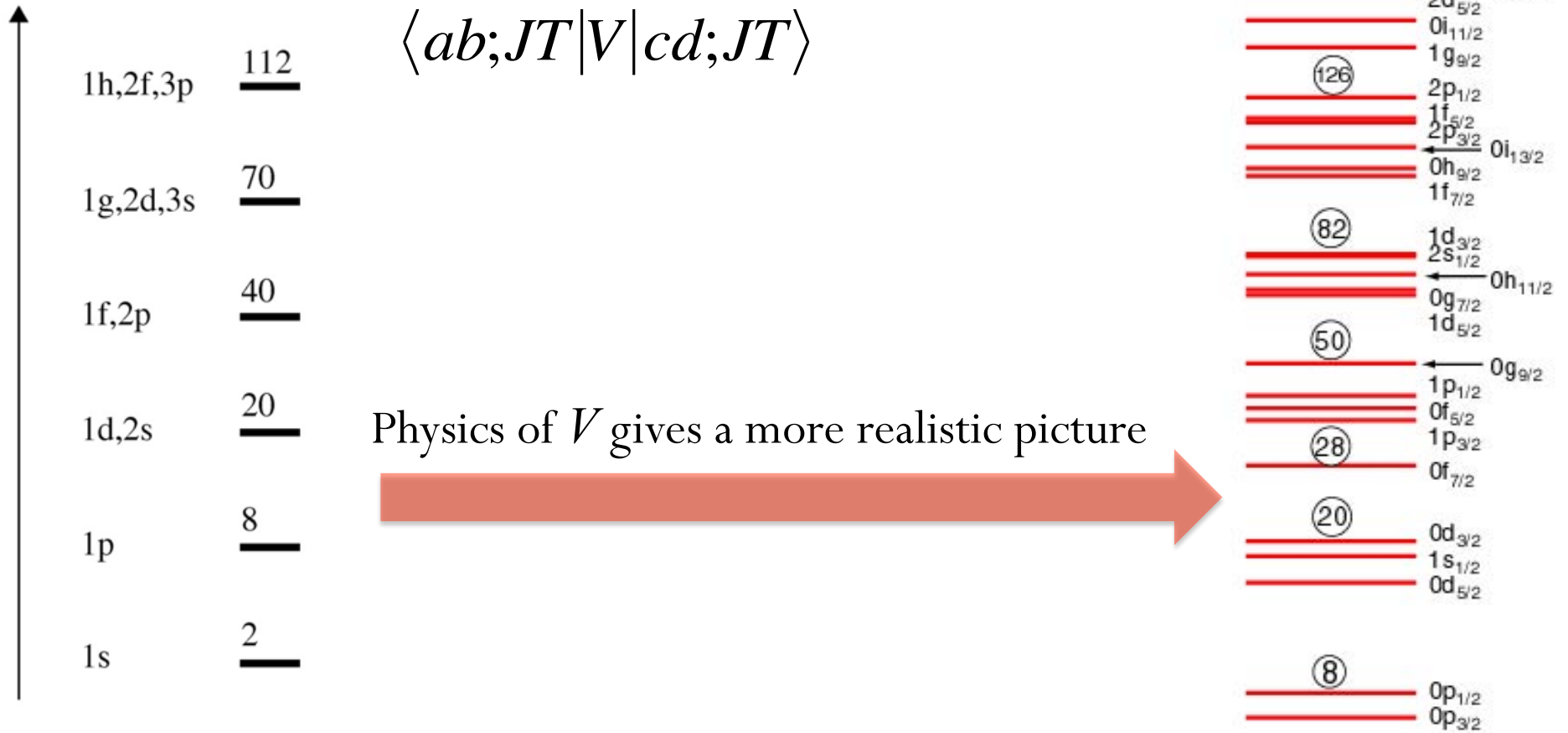
$$|nl, NL; JST\rangle = \int k^2 dk K^2 dK R_{nl}(\sqrt{2}\alpha k) R_{NL}(\sqrt{1/2}\alpha K) |kl, KL; JST\rangle$$

One more (ugly) transformation from center-of-mass to lab frame:

$$\langle ab; JT | V | cd; JT \rangle$$

Solving the Nuclear Many-Body Problem

Matrix elements now given between degenerate HO levels



Non-degeneracy of levels must come from theory

Problem: Can't solve Schrodinger equation in full Hilbert space

Possible with approximations only in light nuclei (*ab initio*)

Solving the Nuclear Many-Body Problem

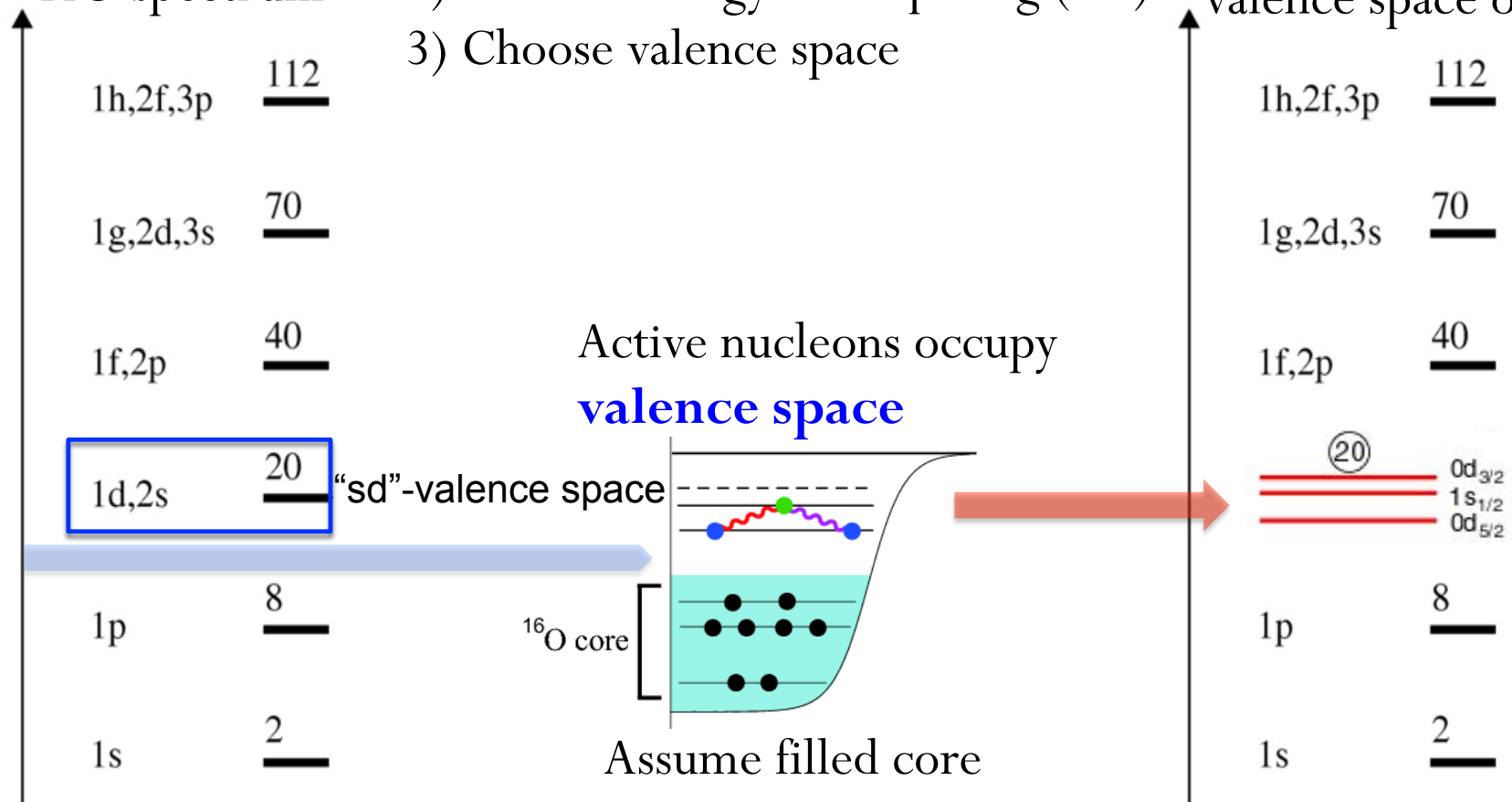
Nuclei understood as many-body system starting from closed shell, add nucleons

Recipe

Unperturbed

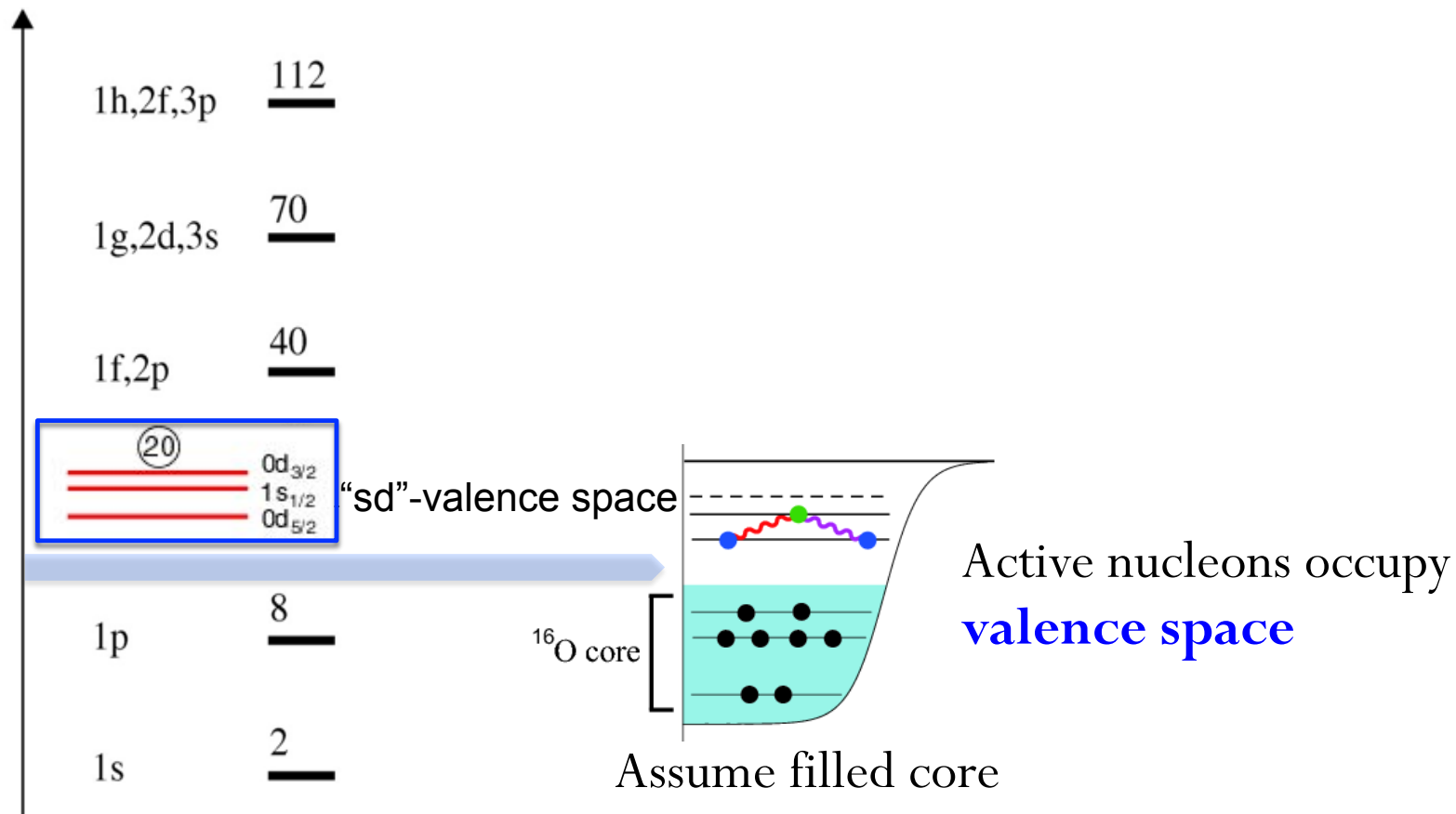
- 1) Choose appropriate core
- 2) Fix HO energy level spacing ($\hbar\omega$)
- 3) Choose valence space

Removes degeneracy in valence space only



Solving the Nuclear Many-Body Problem

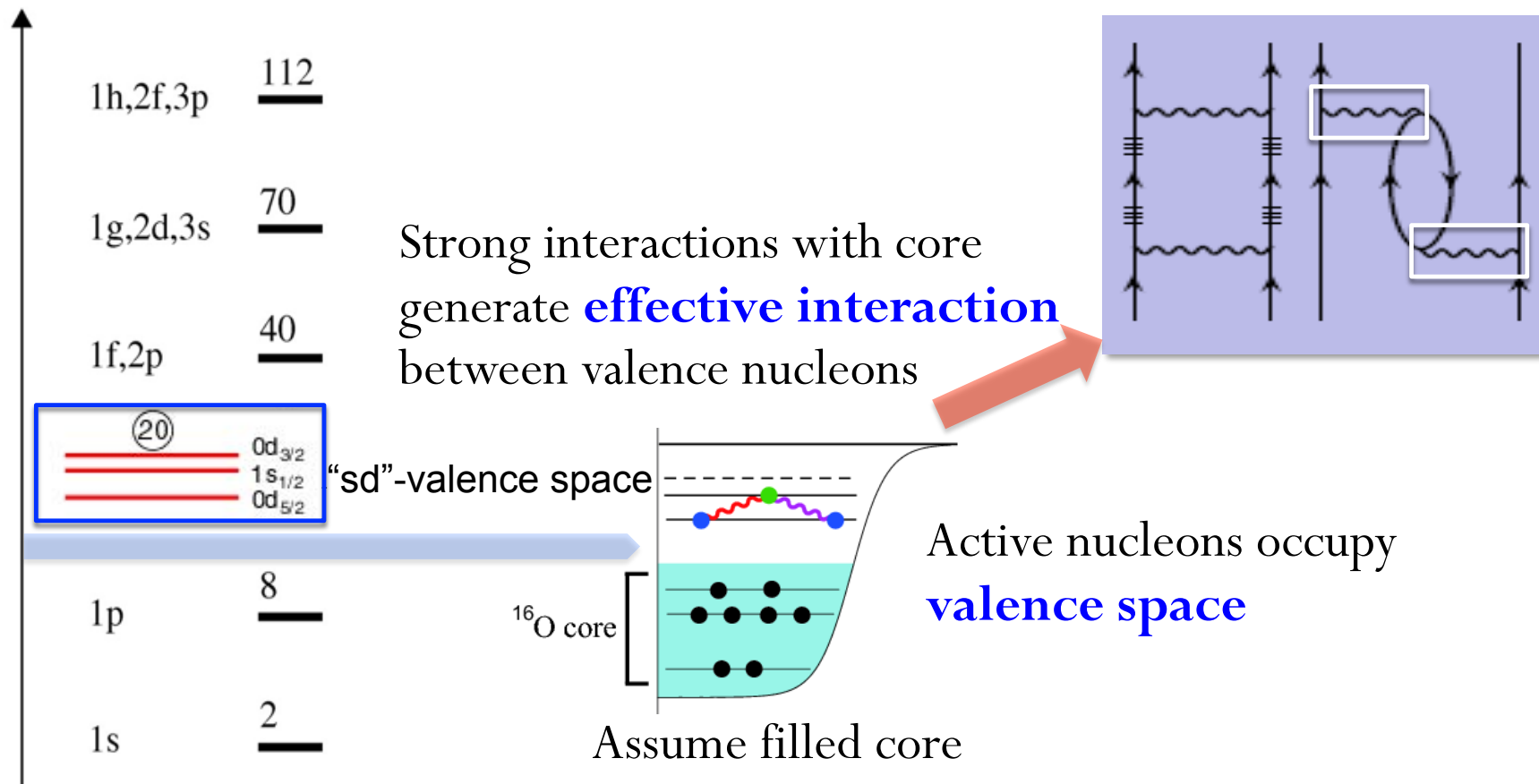
Now have interaction and energies of valence space orbitals from original V
This alone does not reproduce experimental data



Solving the Nuclear Many-Body Problem

Now have interaction and energies of valence space orbitals from original V
 This alone does not reproduce experimental data

Hjorth-Jensen, Kuo, Osnes (1995)

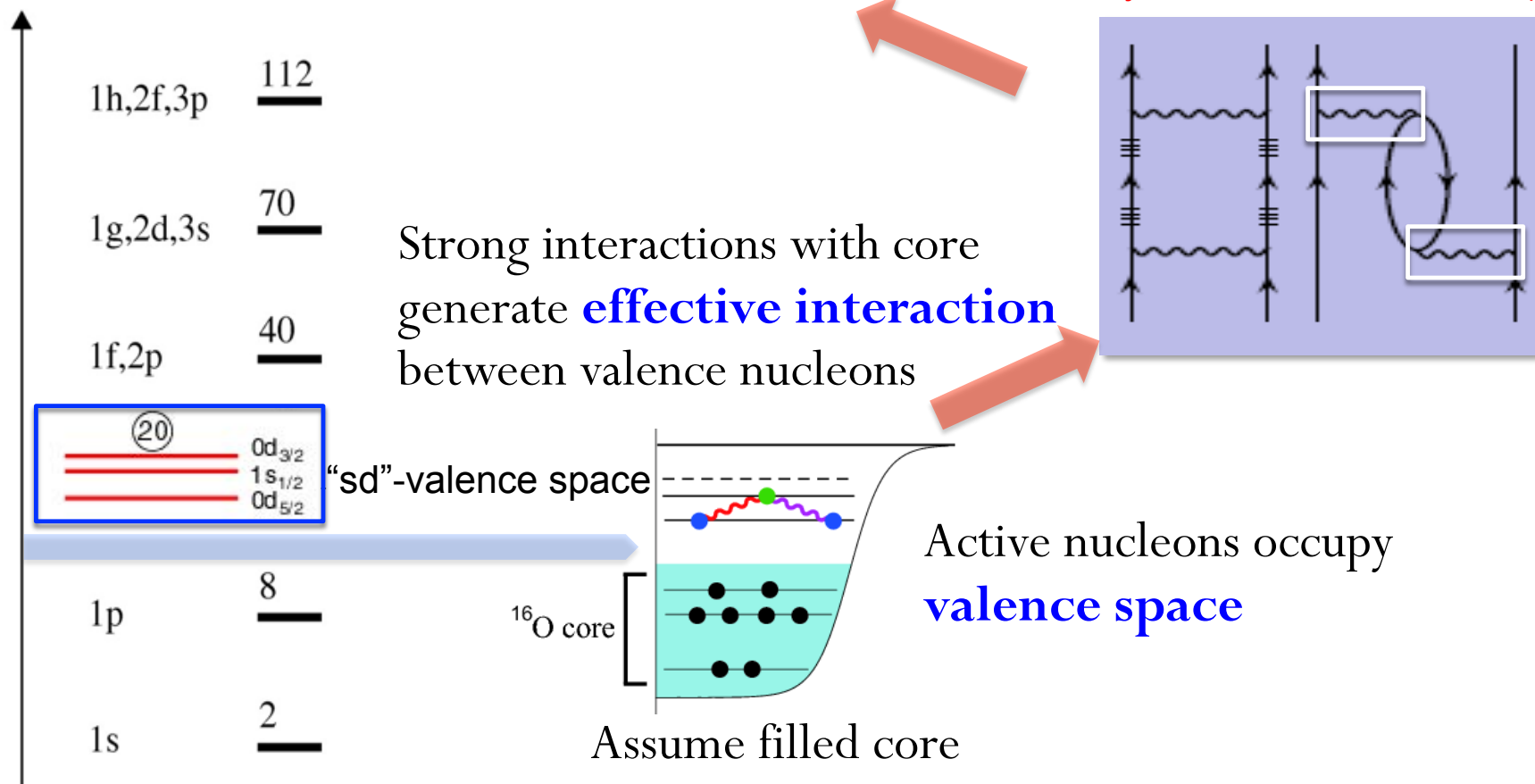


Solving the Nuclear Many-Body Problem

Now have interaction and energies of valence space orbitals from original V
 This alone does not reproduce experimental data

Effective two-body matrix elements
Single-particle energies (SPEs)

Hjorth-Jensen, Kuo, Osnes (1995)



Many-Body Perturbation Theory

How do we calculate valence space interactions and SPEs??

Define operator P that projects onto the model space

$$P = \sum_{i=1}^D |\psi_i\rangle\langle\psi_i| \quad Q = \sum_{i=1+D}^{\infty} |\psi_i\rangle\langle\psi_i|$$

With relations:

$$PQ = 0 \quad P^2 = P \quad Q^2 = Q \quad P + Q = 1$$

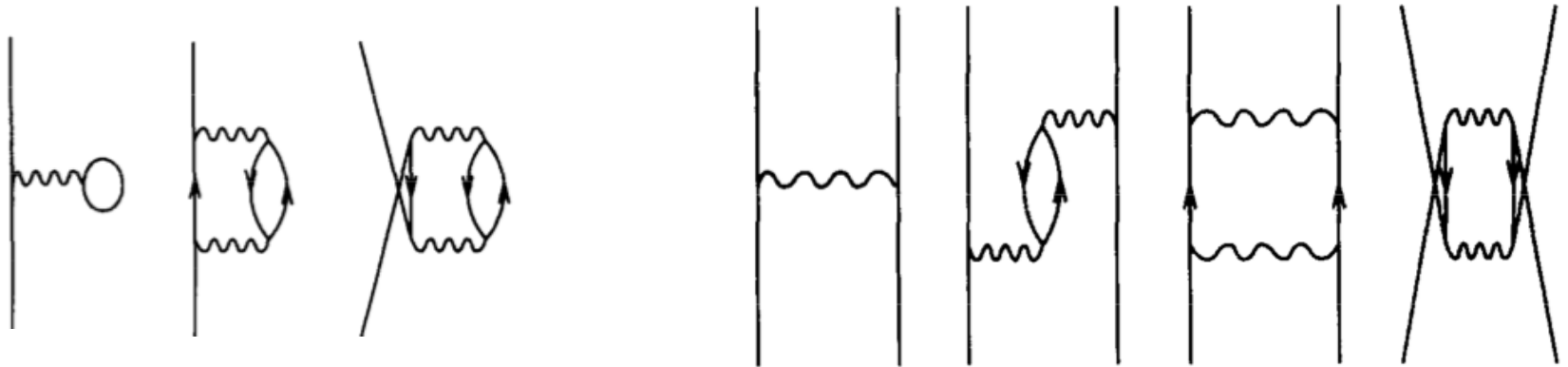
Project full Schrodinger equation into model space eqn that's easy to solve:

$$PH_{eff}P\psi = EP\psi; \quad H_{eff} = H_0 + V_{eff}$$

Need to construct V_{eff}

Many-Body Perturbation Theory

To construct the effective interaction, define \hat{Q} -box = sum of all possible topologically distinct diagrams which are **irreducible** and **valence linked**:



1-body Q-box to 2nd order

2-body Q-box to 2nd order

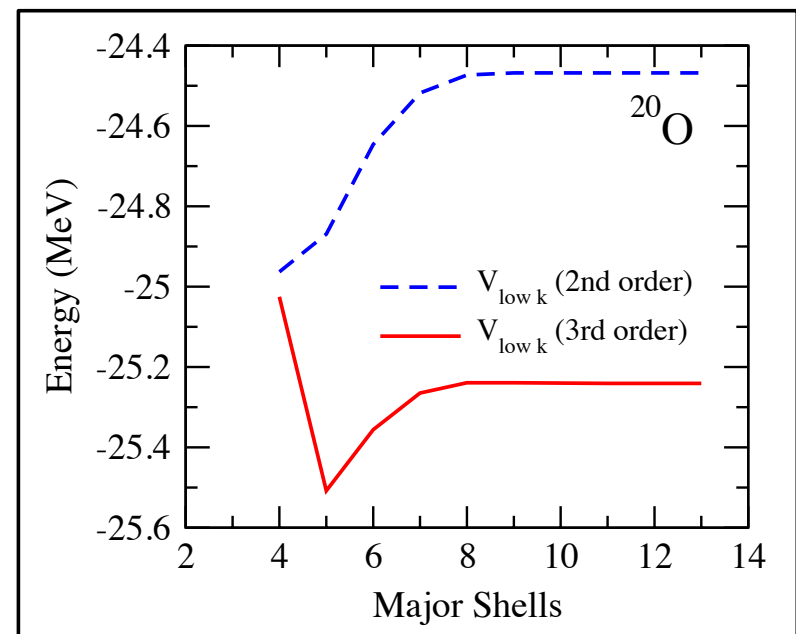
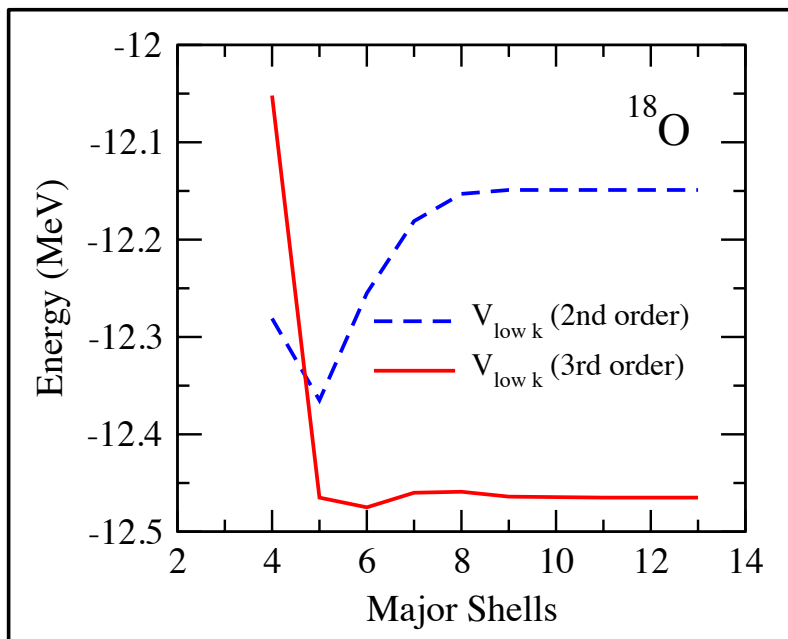
Single-particle energies can be calculated from one-body part
Traditionally taken from experimental one-particle spectrum or empirical values

Calculation Details

Convergence in terms of Harmonic Oscillator basis size

NN matrix elements derived from:

- Chiral N³LO (Machleidt, 500 MeV) using smooth-regulator $V_{\text{low } k}$
- 3rd-order in perturbation theory
- 13 major shells for intermediate state configurations (converged)

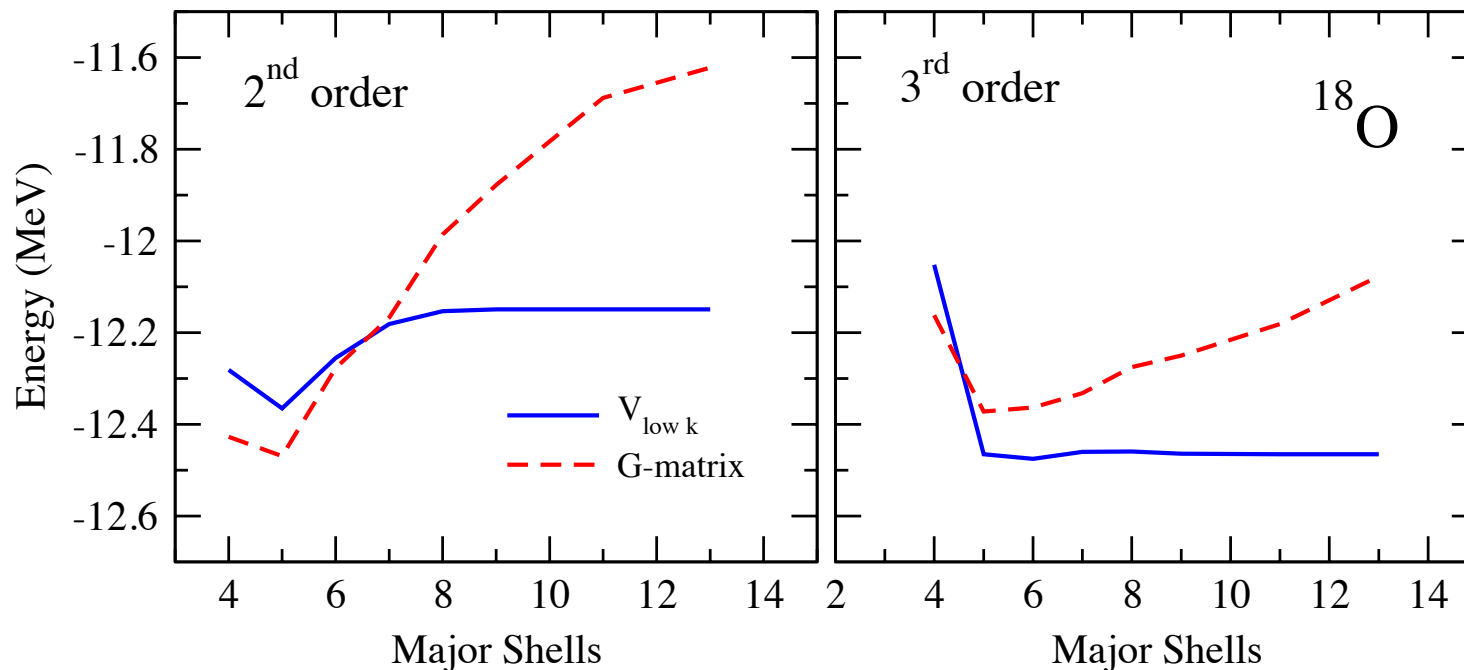


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Monopole Part of Valence-Space Interactions

Microscopic MBPT – effective interaction in chosen model space

Works near closed shells: deteriorates beyond this

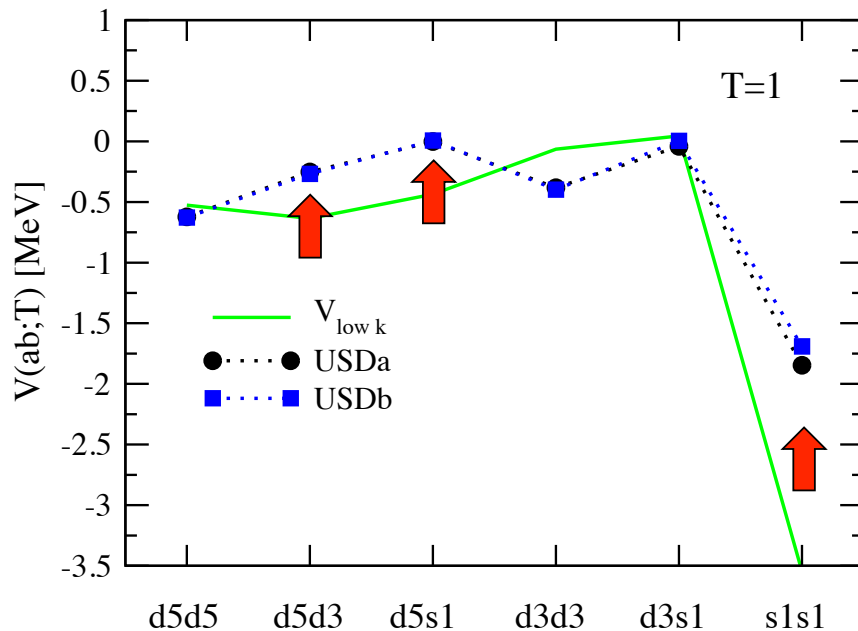
Deficiencies improved adjusting particular two-body matrix elements

Monopoles:
Angular average of interaction

$$V_{ab}^T = \frac{\sum_J (2J+1) V_{abab}^{JT}}{\sum_J (2J+1)}$$

Determines interaction of orbit a with b : evolution of orbital energies

$$\Delta \epsilon_a = V_{ab} n_b$$



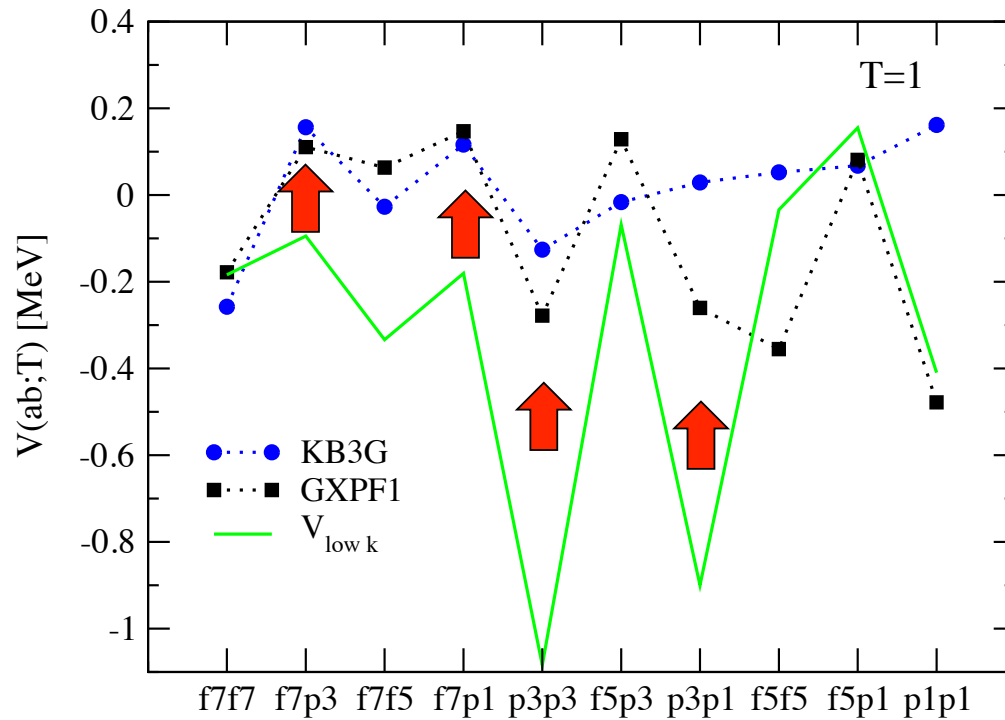
Microscopic **low-momentum** interactions

Phenomenological **USD** interactions

Clear shifts in **low-lying orbitals**:

- T=1 repulsive shift

Phenomenological vs. Microscopic



Compare monopoles from:

Microscopic **low-momentum** interactions

Phenomenological **KB3G, GXPF1** interactions

Shifts in **low-lying orbitals**:

- T=1 repulsive shift

Limits of Nuclear Existence: Oxygen Anomaly

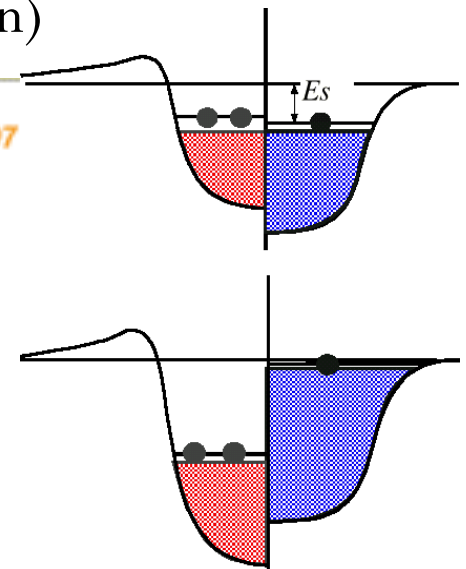
Where is the nuclear dripline?

Limits defined as last isotope with positive neutron separation energy

- Nucleons “drip” out of nucleus

Neutron dripline experimentally established to $Z=8$ (Oxygen)

²⁸ Si	²⁹ Si	³⁰ Si	³¹ Si	³² Si	³³ Si	³⁴ Si	³⁵ Si	³⁶ Si	³⁷ Si	³⁸ Si	³⁹ Si	⁴⁰ Si	⁴¹ Si	⁴² Si	⁴³ Si	⁴⁴ Si	2007
²⁷ Al	²⁸ Al	²⁹ Al	³⁰ Al	³¹ Al	³² Al	³³ Al	³⁴ Al	³⁵ Al	³⁶ Al	³⁷ Al	³⁸ Al	³⁹ Al	⁴⁰ Al	⁴¹ Al	⁴² Al	⁴³ Al	
²⁶ Mg	²⁷ Mg	²⁸ Mg	²⁹ Mg	³⁰ Mg	³¹ Mg	³² Mg	³³ Mg	³⁴ Mg	³⁵ Mg	³⁶ Mg	³⁷ Mg	³⁸ Mg	⁴⁰ Mg				
⁵⁷ Na	²⁶ Na	²⁷ Na	²⁸ Na	²⁹ Na	³⁰ Na	³¹ Na	³² Na	³³ Na	³⁴ Na	³⁵ Na		³⁷ Na	2002				
²⁴ Ne	²⁵ Ne	²⁶ Ne	²⁷ Ne	²⁸ Ne	²⁹ Ne	³⁰ Ne	³¹ Ne	³² Ne		³⁴ Ne	2002						
²³ F	²⁴ F	²⁵ F	²⁶ F	²⁷ F		²⁹ F		³¹ F	1999								
²² O	²³ O	²⁴ O	1970														
²¹ N	²² N	²³ N															
²⁰ C		²² C															



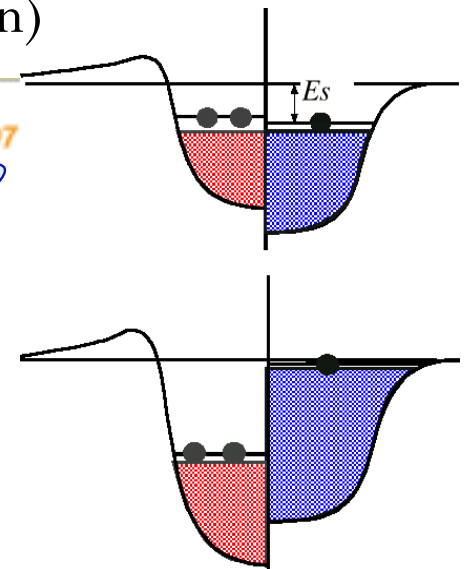
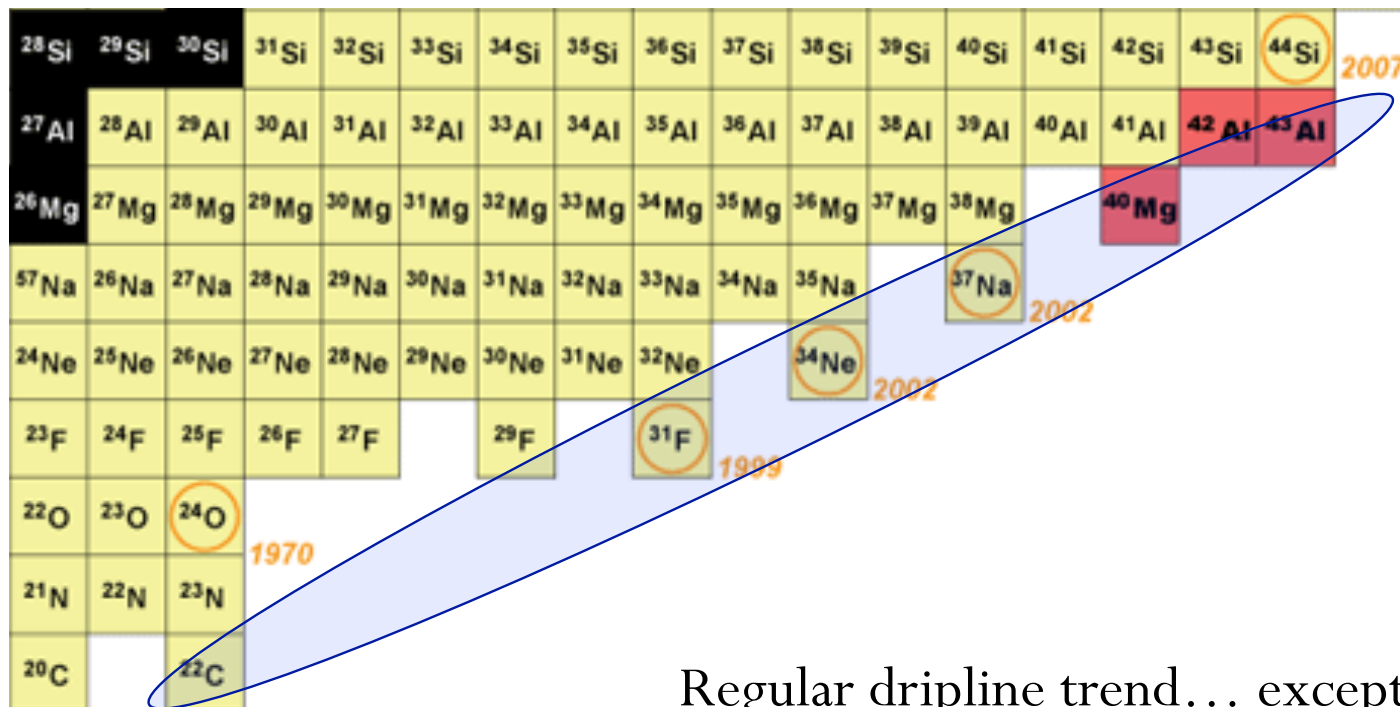
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Where is the nuclear dripline?

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Neutron dripline experimentally established to $Z=8$ (Oxygen)



Regular dripline trend... except oxygen

Adding one proton binds 6 additional neutrons

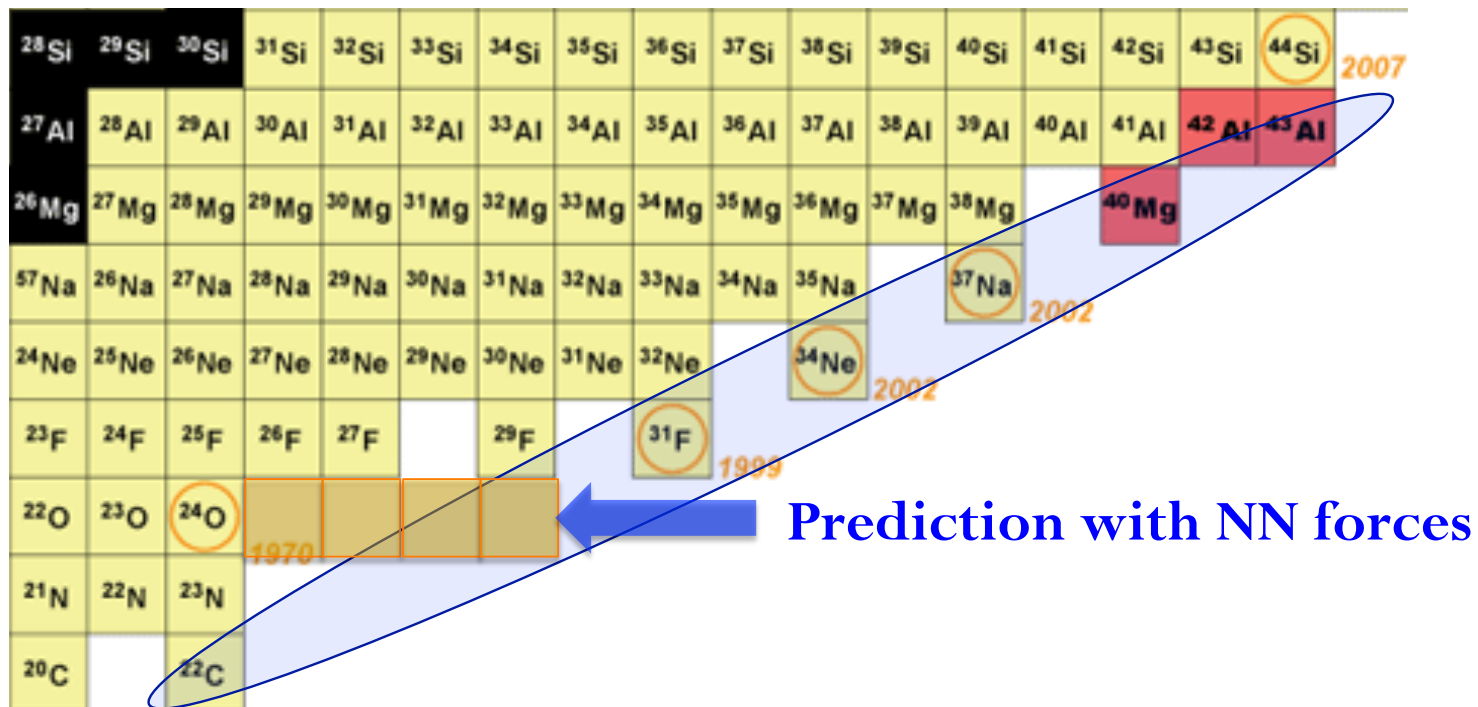
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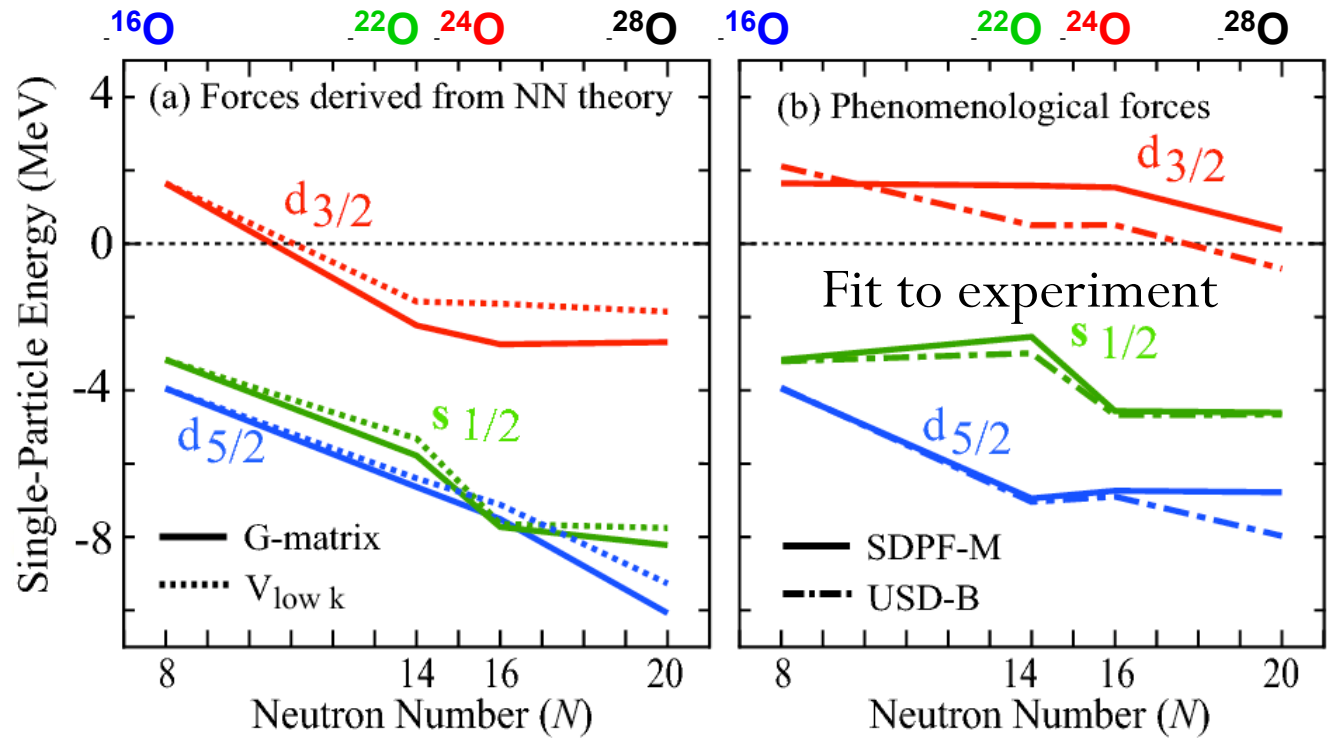
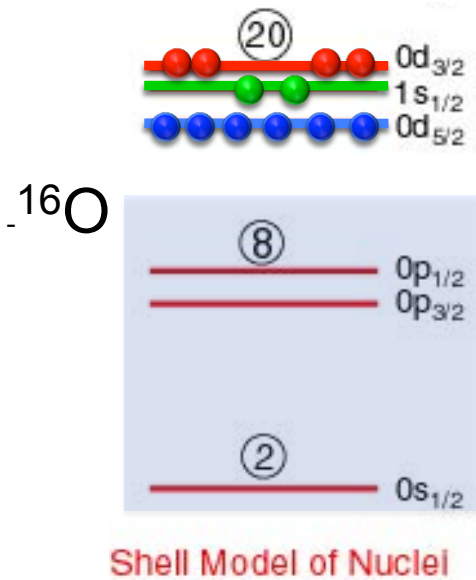


Microscopic picture: **NN-forces too attractive**

Incorrect prediction of dripline

Physics in Oxygen Isotopes

Calculate evolution of sd -orbital energies from interactions



Microscopic NN Theories

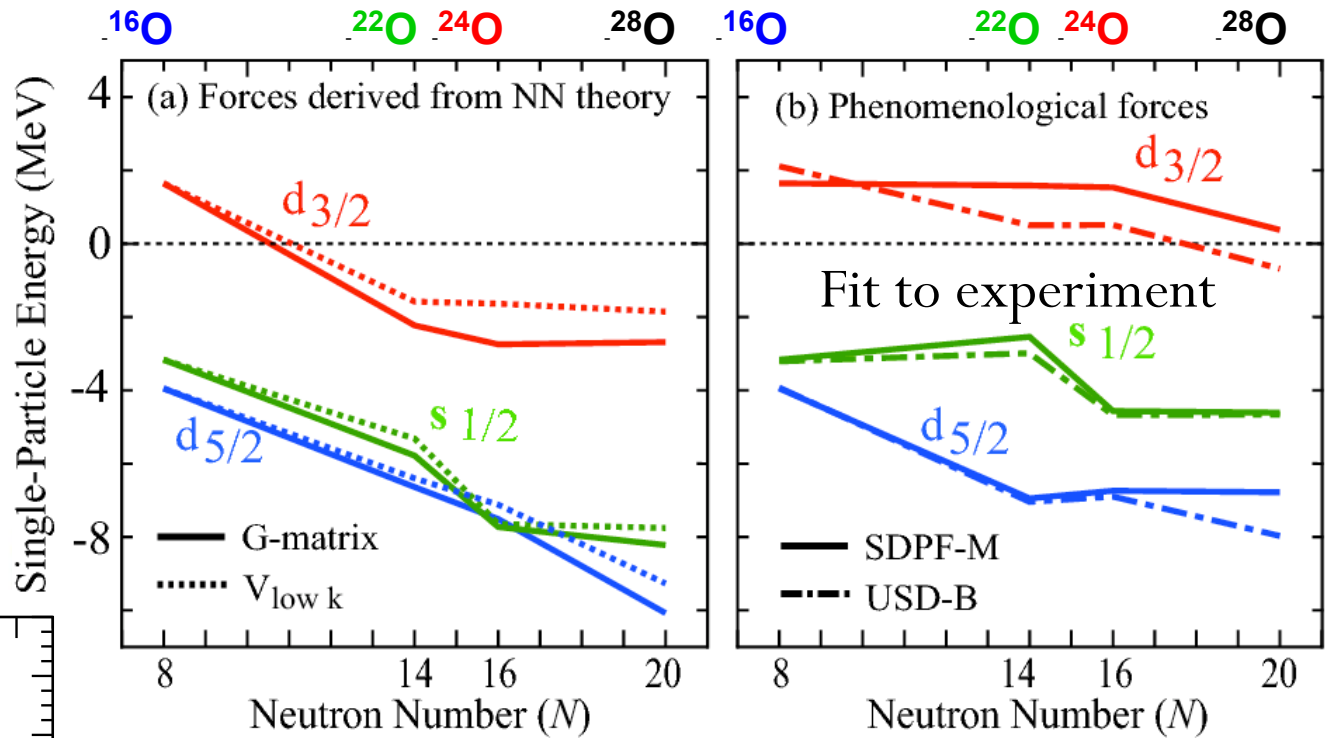
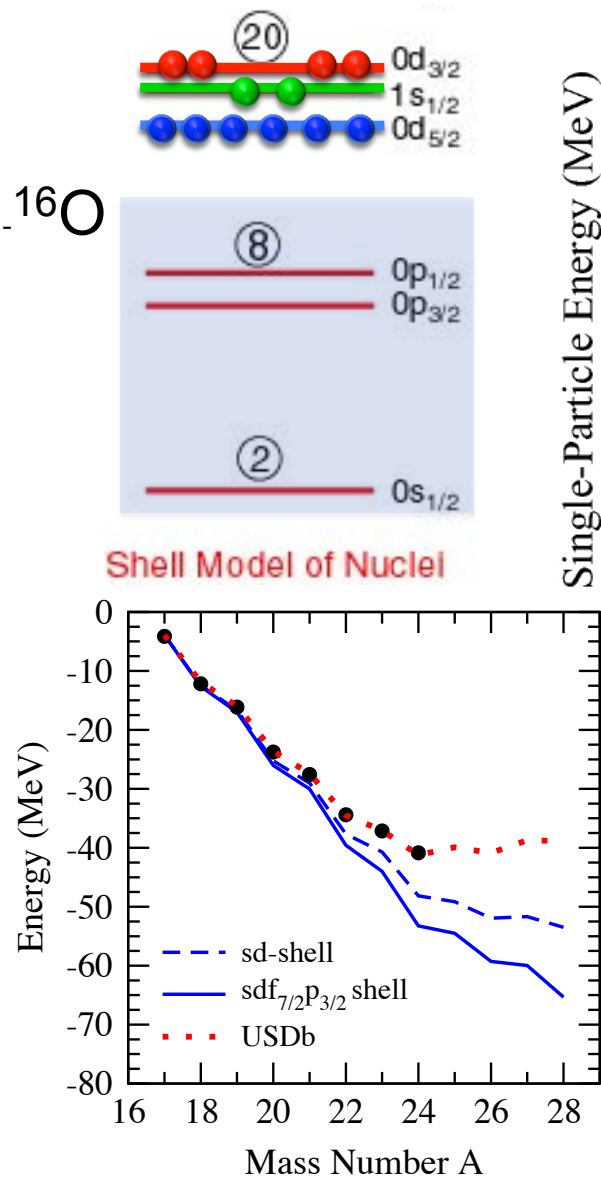
$d_{3/2}$ orbit bound to ^{28}O

Phenomenological Models

$d_{3/2}$ orbit unbound

Physics in Oxygen Isotopes

Calculate evolution of sd -orbital energies from interactions



Microscopic NN Theories

$d_{3/2}$ orbit bound to ^{28}O
Dripline at ^{28}O

Phenomenological Models

$d_{3/2}$ orbit unbound
Dripline at ^{24}O

Oxygen anomaly unexplained with NN forces

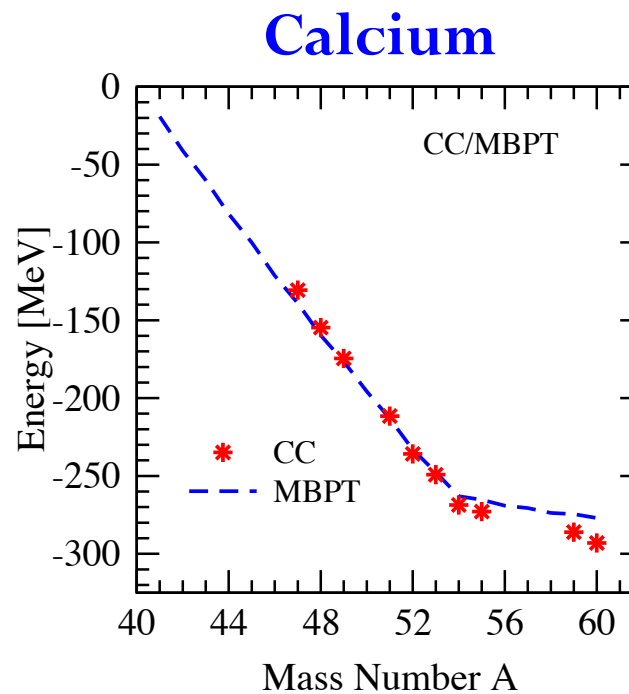
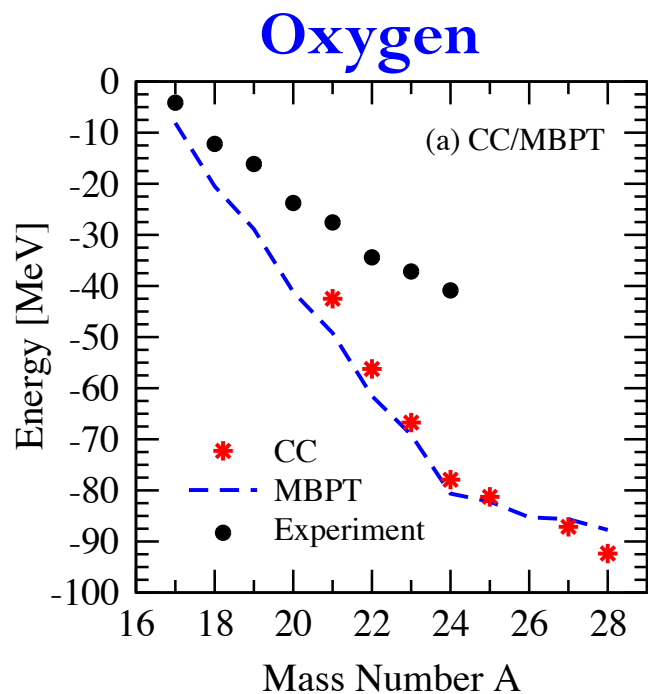
Origin of monopole shifts: Neglected 3N forces

-- **Zuker (2003)**

Comparison to Coupled Cluster

Many-body method insufficient?

Benchmark against *ab-initio* Coupled Cluster at NN-only level



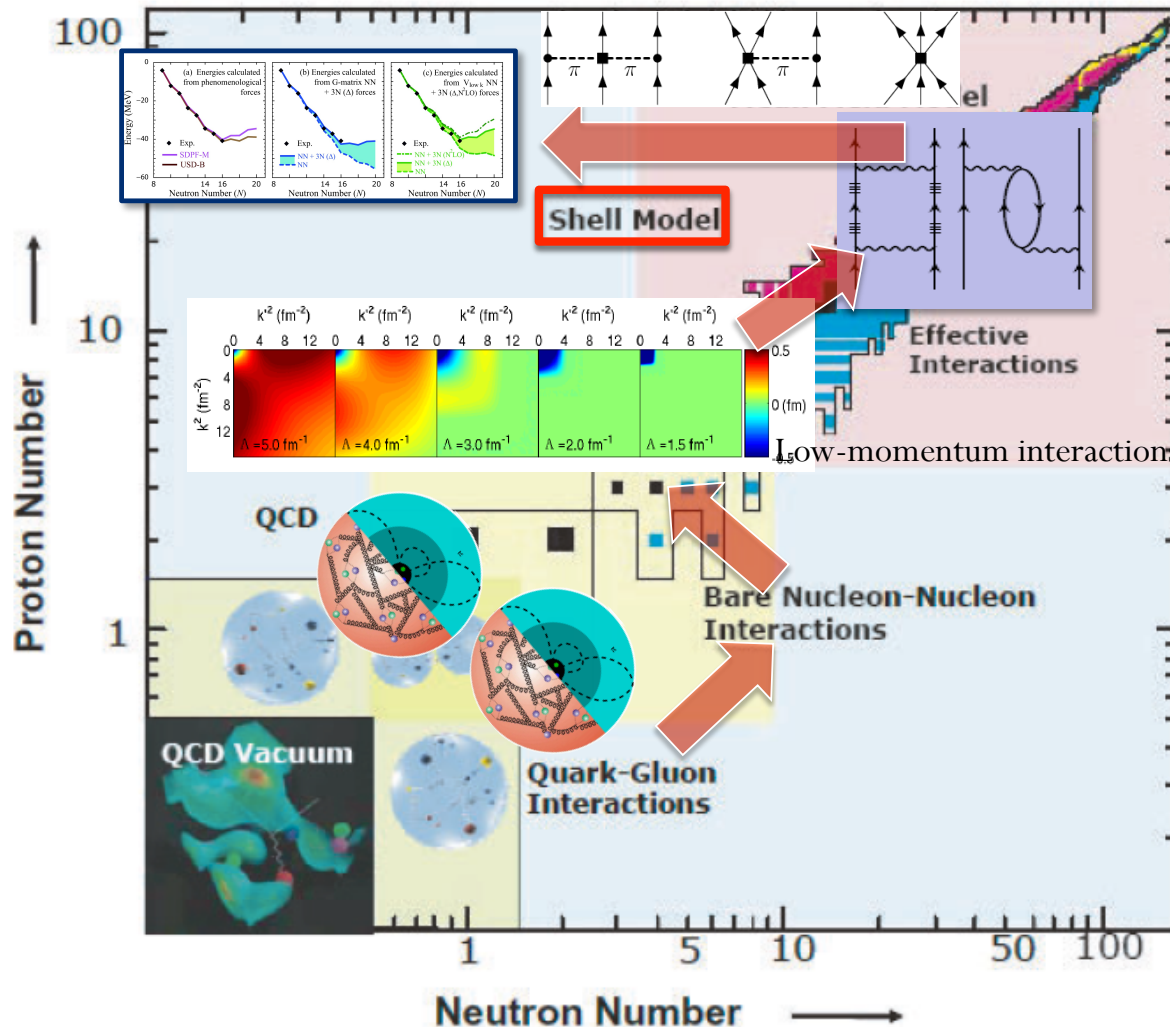
SPEs: one-particle attached CC energies in ^{17}O and ^{41}Ca

Small difference in many-body methods

Include **3N forces** to improve agreement with experiment

The Challenge of Microscopic Nuclear Theory

To understand the properties of complex nuclei from elementary interactions



Three-Nucleon Forces

Basic ideas – why do we need?

3N from chiral EFT

Implementing in shell model

Relation to monopoles

Predictions/Results

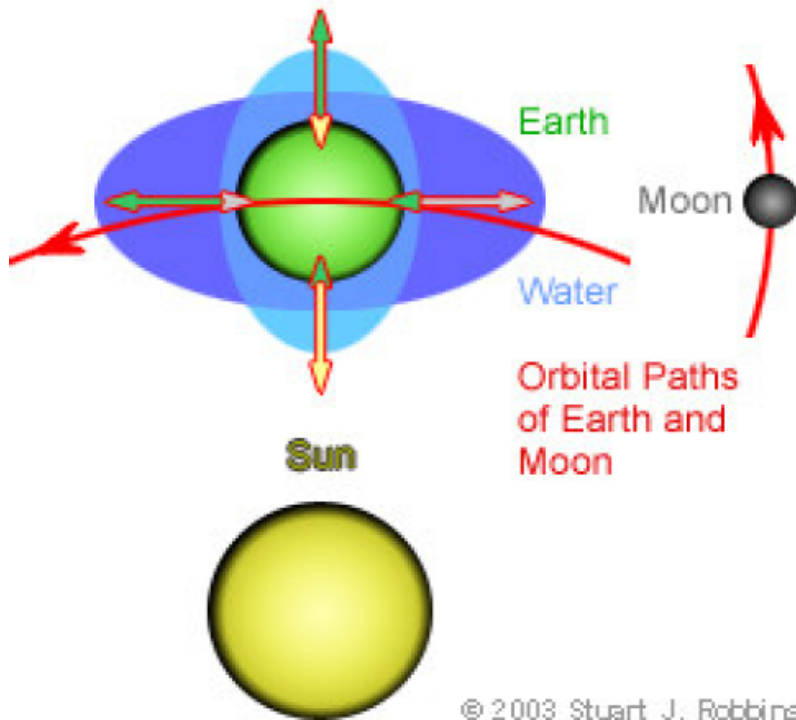
Density-dependent 3N

How will we approach this problem:

QCD → NN (3N) forces → Renormalize → Solve many-body problem → Predictions

Why Three-Body Forces?

Tidal Bulges from Moon and Sun



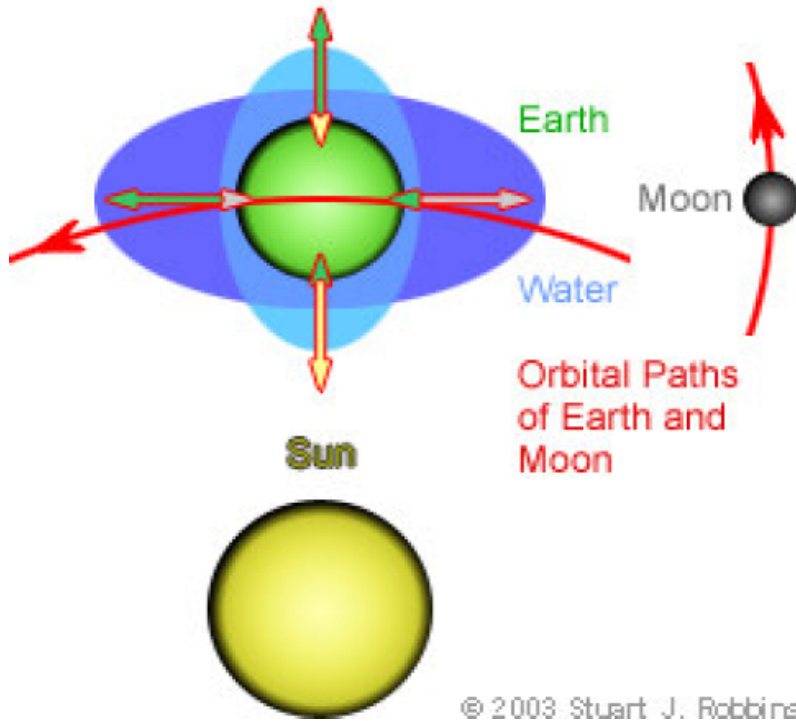
Earth not point particle

Experiences tidal forces from sun *and* moon

Lead to 3-body forces in E-M-S system

Why Three-Body Forces?

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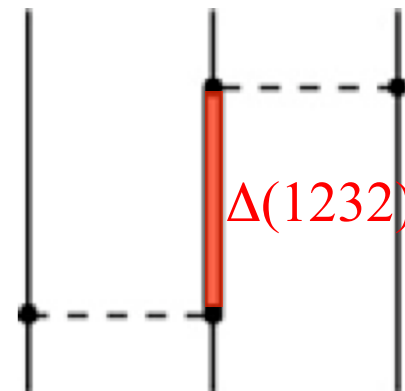
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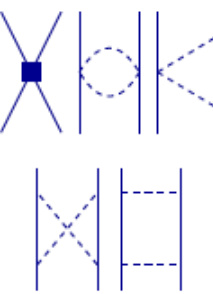


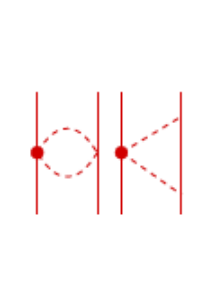
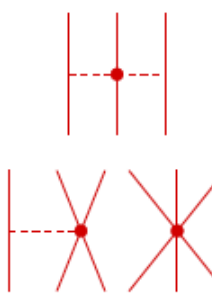

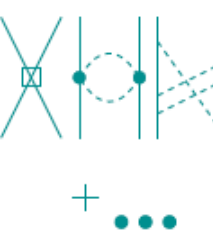
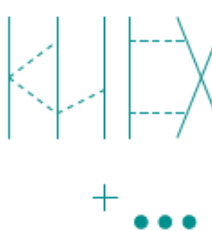

Nucleons are composite particles

Can be excited to resonances



Leads to non-negligible effects

Chiral Effective Field Theory: Summary

	2N forces	3N forces	4N forces
LO			
NLO			
N ² LO			
N ³ LO	 + ...	 + ...	 + ...

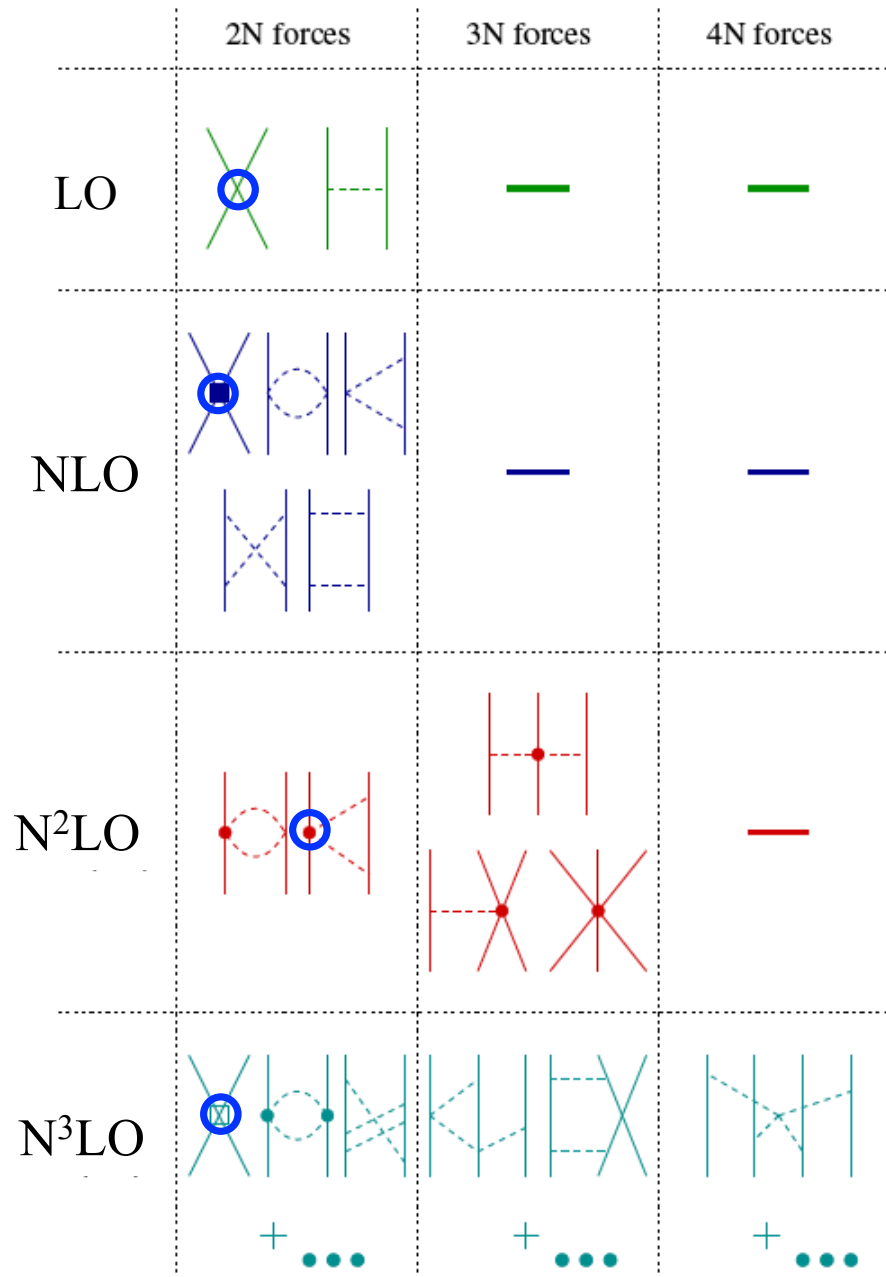
Nucleons interact via pion exchanges and contact interactions

Hierarchy: $V_{\text{NN}} > V_{\text{3N}} > \dots$

Consistent treatment of NN, 3N, ... electroweak operators

Weinberg, van Kolck, Kaplan, Savage, Wise, Epelbaum, Kaiser, Meissner, ...

Chiral Effective Field Theory: Nuclear Forces



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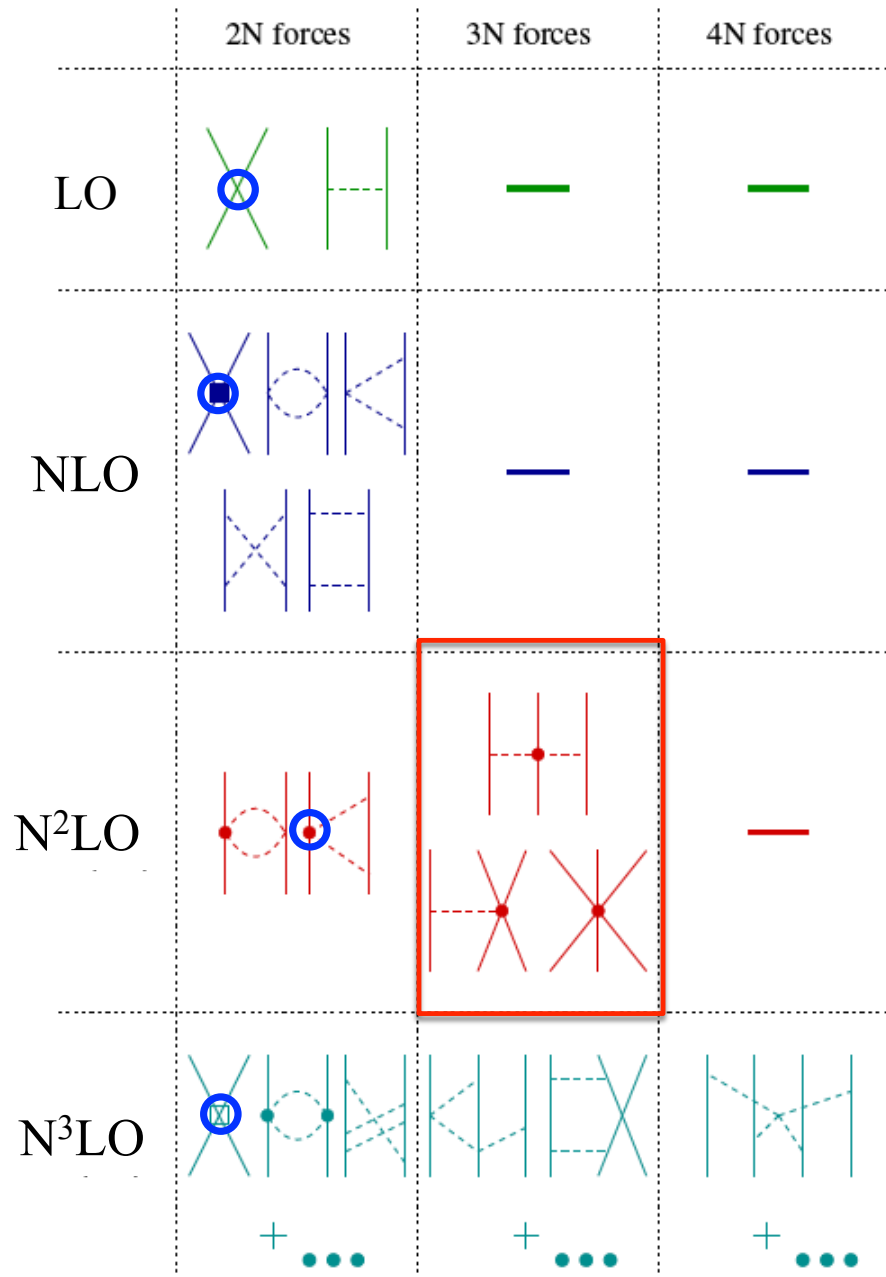
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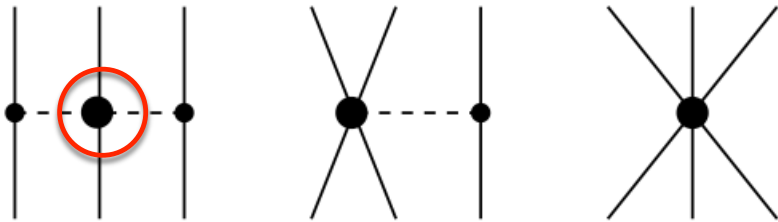
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Chiral EFT: N²LO

First non-vanishing 3N contributions

Next-to-next-to-leading order ($\nu = 3$)



$$\vec{q}_i \equiv \vec{p}'_i - \vec{p}_i$$

$$g_A = 1.26$$

$$F_\pi = 92.4 \text{ MeV}$$

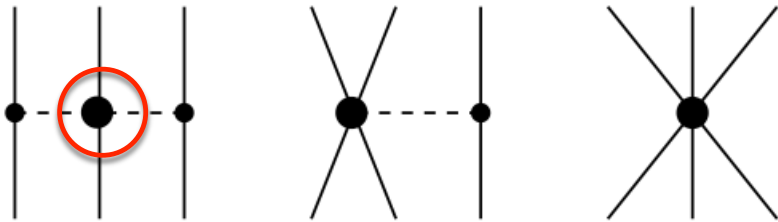
$$V_{2\pi}^{(3)} = \sum_{i \neq j \neq k} \frac{1}{2} \left(\frac{g_A}{2F_\pi} \right)^2 \frac{(\vec{\sigma}_i \cdot \vec{q}_i) (\vec{\sigma}_j \cdot \vec{q}_j)}{(\vec{q}_i^2 + M_\pi^2) (\vec{q}_j^2 + M_\pi^2)} F_{ijk}^{\alpha\beta} \tau_i^\alpha \tau_j^\beta$$

$$F_{ijk}^{\alpha\beta} = \delta^{\alpha\beta} \left(-\frac{4c_1 M_\pi^2}{F_\pi^2} + \frac{2c_3}{F_\pi^2} \vec{q}_i \cdot \vec{q}_j \right) + \sum_\gamma \frac{c_4}{F_\pi^2} \varepsilon^{\alpha\beta\gamma} \tau_k^\gamma \vec{\sigma}_k \cdot (\vec{q}_i \times \vec{q}_j)$$

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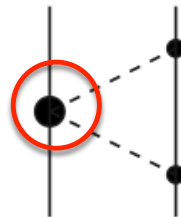
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$$V_{2\pi}^{(3)} = \sum_{i \neq j \neq k} \frac{1}{2} \left(\frac{g_A}{2F_\pi} \right)^2 \frac{(\vec{\sigma}_i \cdot \vec{q}_i) (\vec{\sigma}_j \cdot \vec{q}_j)}{(\vec{q}_i^2 + M_\pi^2) (\vec{q}_j^2 + M_\pi^2)} F_{ijk}^{\alpha\beta} \tau_i^\alpha \tau_j^\beta$$

$$F_{ijk}^{\alpha\beta} = \delta^{\alpha\beta} \left(-\frac{4c_1 M_\pi^2}{F_\pi^2} + \frac{2c_3}{F_\pi^2} \vec{q}_i \cdot \vec{q}_j \right) + \sum_\gamma \frac{c_4}{F_\pi^2} \varepsilon^{\alpha\beta\gamma} \tau_k^\gamma \vec{\sigma}_k \cdot (\vec{q}_i \times \vec{q}_j)$$

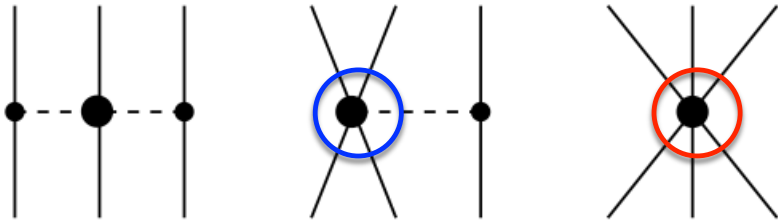
3 LECs – determined from NN fit



Chiral EFT: N²LO

First non-vanishing 3N contributions

Next-to-next-to-leading order ($\nu = 3$)



$$\vec{q}_i \equiv \vec{p}'_i - \vec{p}_i$$

$$g_A = 1.26$$

$$F_\pi = 92.4 \text{ MeV}$$

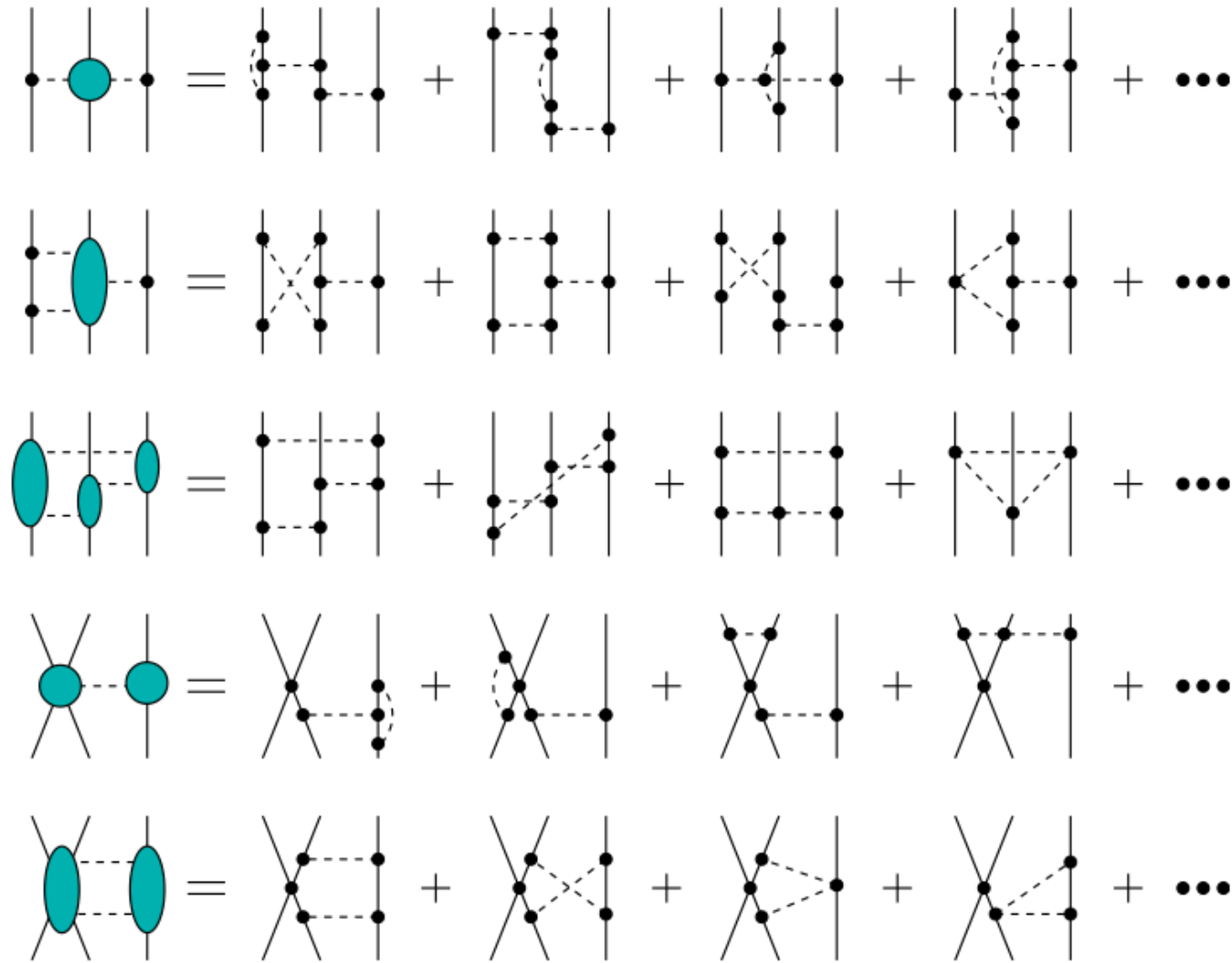
$$V_{\text{l}\pi, \text{cont}}^{(3)} = - \sum_{i \neq j \neq k} \left(\frac{g_A}{8F_\pi} \right)^2 \textcircled{D} \frac{(\vec{\sigma}_j \cdot \vec{q}_j)}{(\vec{q}_j^2 + M_\pi^2)} (\tau_i \cdot \tau_j) (\vec{\sigma}_i \cdot \vec{\sigma}_j)$$

$$V_{\text{cont}}^{(3)} = \frac{1}{2} \sum_{j \neq k} \textcircled{E} (\tau_j \cdot \tau_k)$$

Two new unconstrained couplings D, E: what should we fit to?

Chiral EFT: N³LO

Next-to-next-to-next-to-leading order ($\nu = 4$)



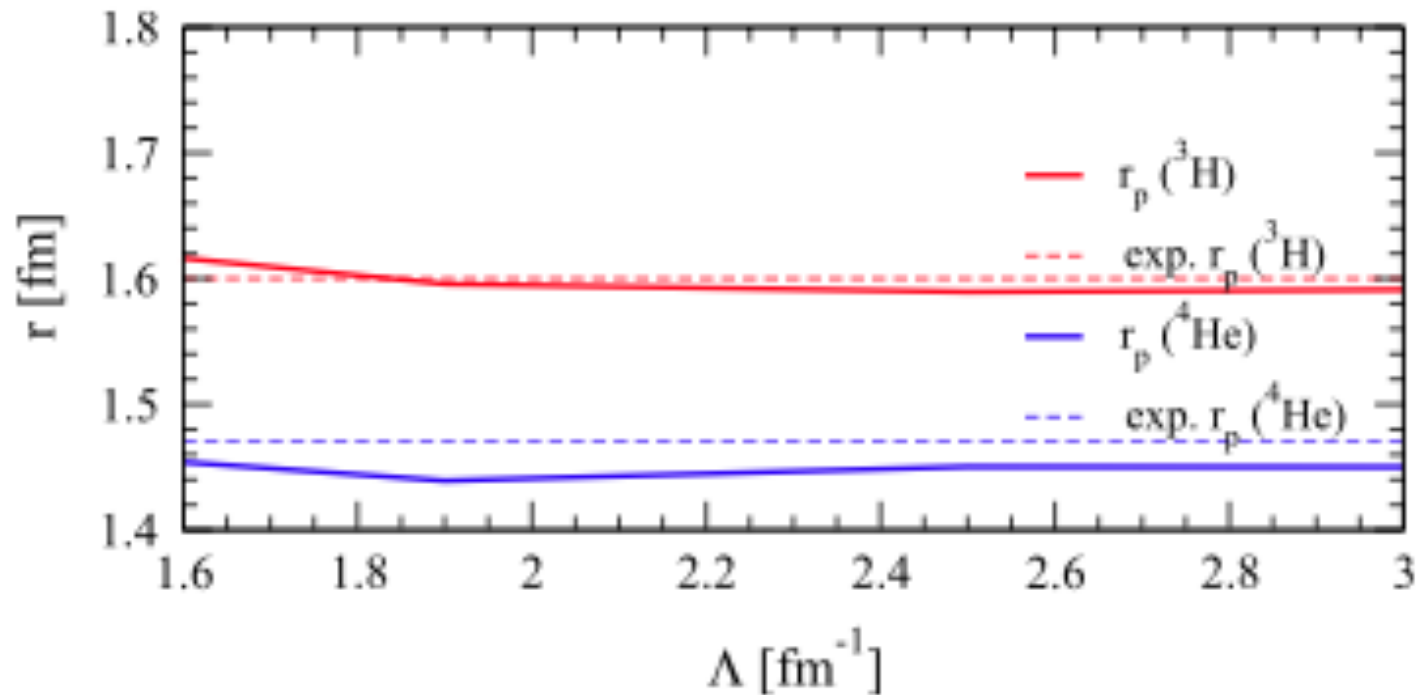
Good news: **no new constants**

Bad news: it's not obvious?

Cutoff Variation with 3N Forces

Use cutoff variation to assess missing physics in few body systems

Radii of triton and alpha particle calculated from NN+3N forces

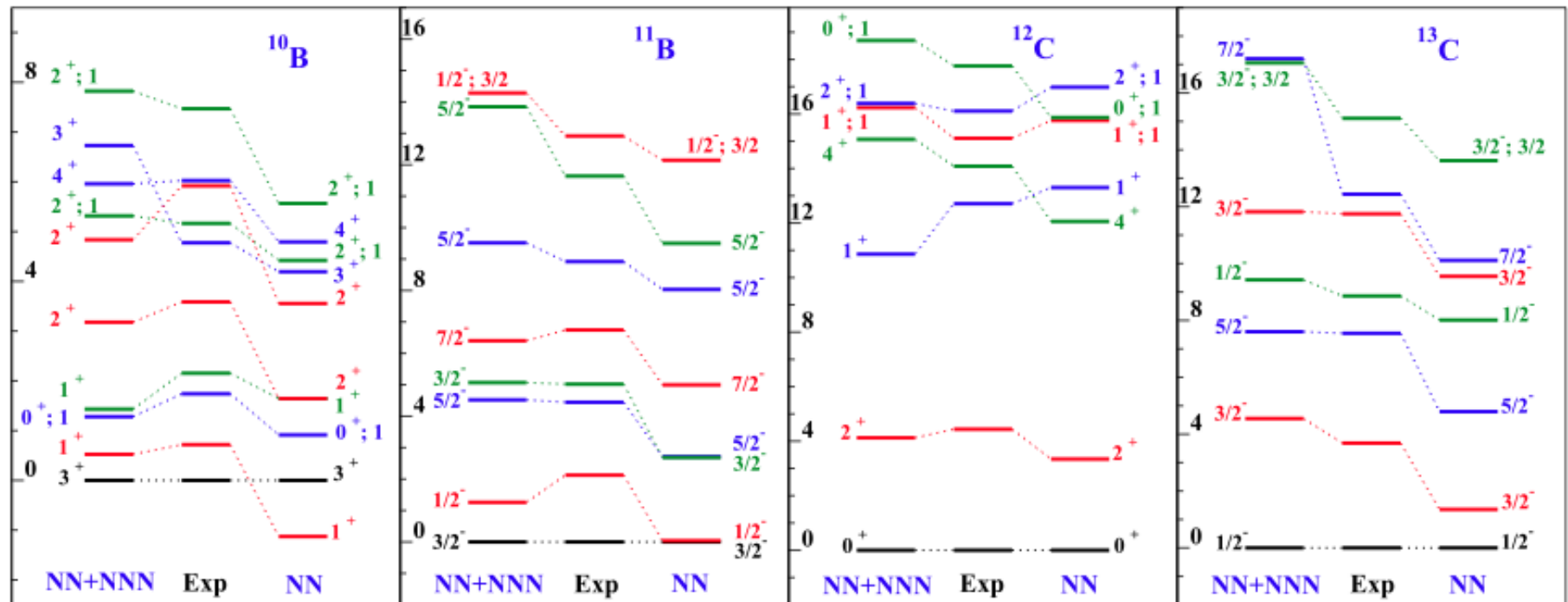


Clearly minimal cutoff variation

Chiral Three-Body Forces in Light Nuclei

Importance of chiral 3N forces established in light nuclei $A \leq 12$

Converged No-core shell model Navratil et al., 2007

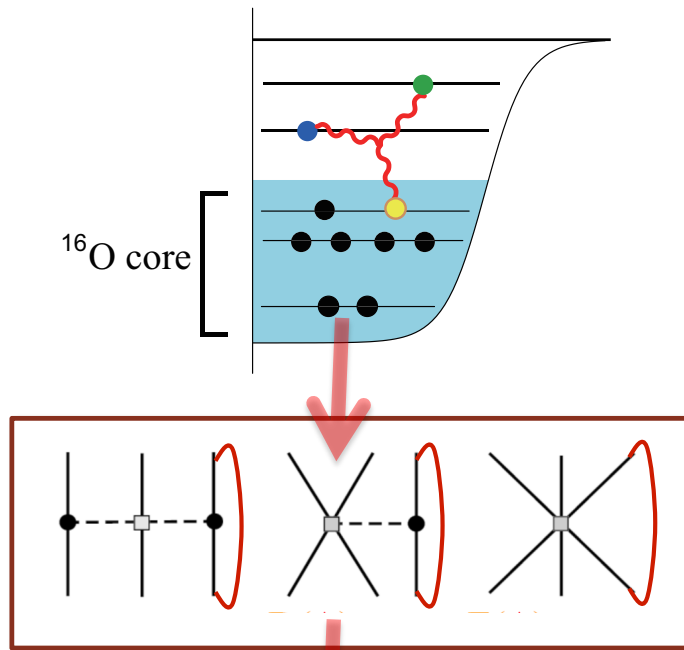


They work! What about medium-mass and exotic nuclei?

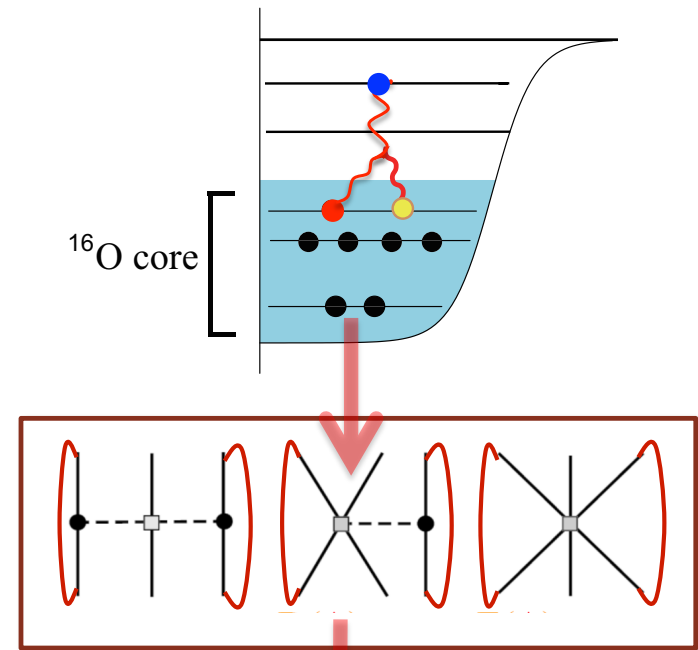
3N Forces for Valence-Shell Theories

Normal-ordered 3N: contribution to valence neutron interactions

Effective two-body



Effective one-body



$$\langle ab | V_{3N,\text{eff}} | a'b' \rangle = \sum_{\alpha=\text{core}} \langle \alpha ab | V_{3N} | \alpha a'b' \rangle$$

$$\langle a | V_{3N,\text{eff}} | a' \rangle = \frac{1}{2} \sum_{\alpha\beta=\text{core}} \langle \alpha\beta a | V_{3N} | \alpha\beta a' \rangle$$

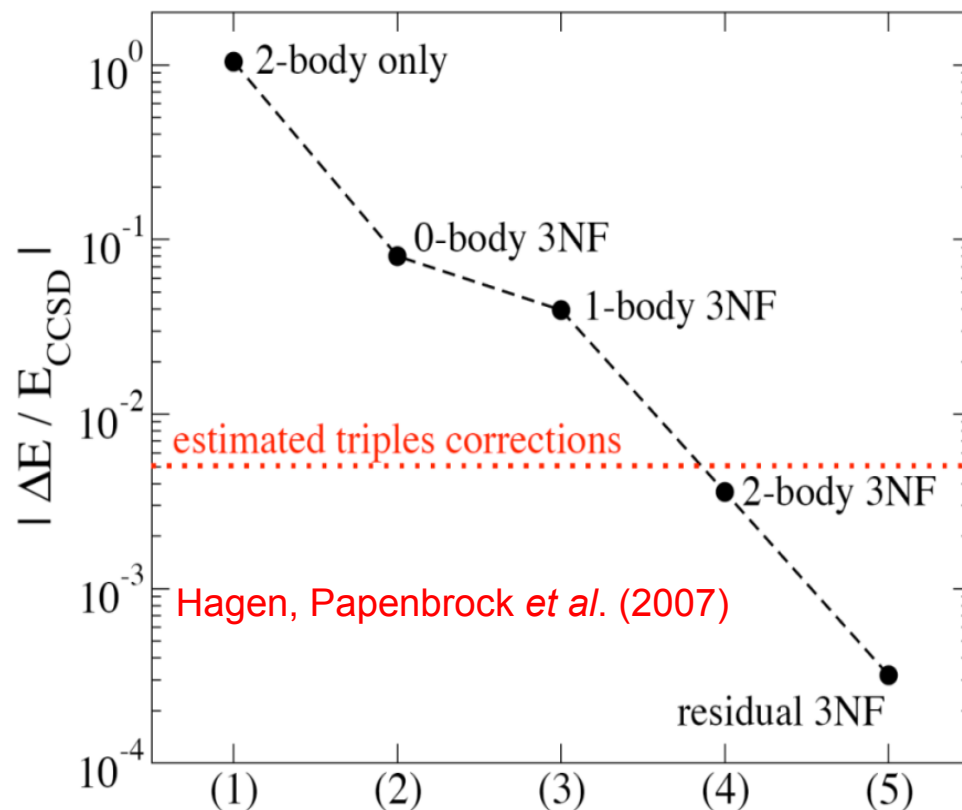
Combine with microscopic NN: eliminate empirical adjustments

3N Forces for Valence-Shell Theories

Effects of residual 3N between 3 valence nucleons?

Normal-ordered 3N: microscopic contributions to inputs for CI Hamiltonian

Effects of residual 3N between 3 valence nucleons?



Coupled-Cluster theory with 3N:
benchmark of ${}^4\text{He}$

0- 1- and 2-body of 3NF dominate

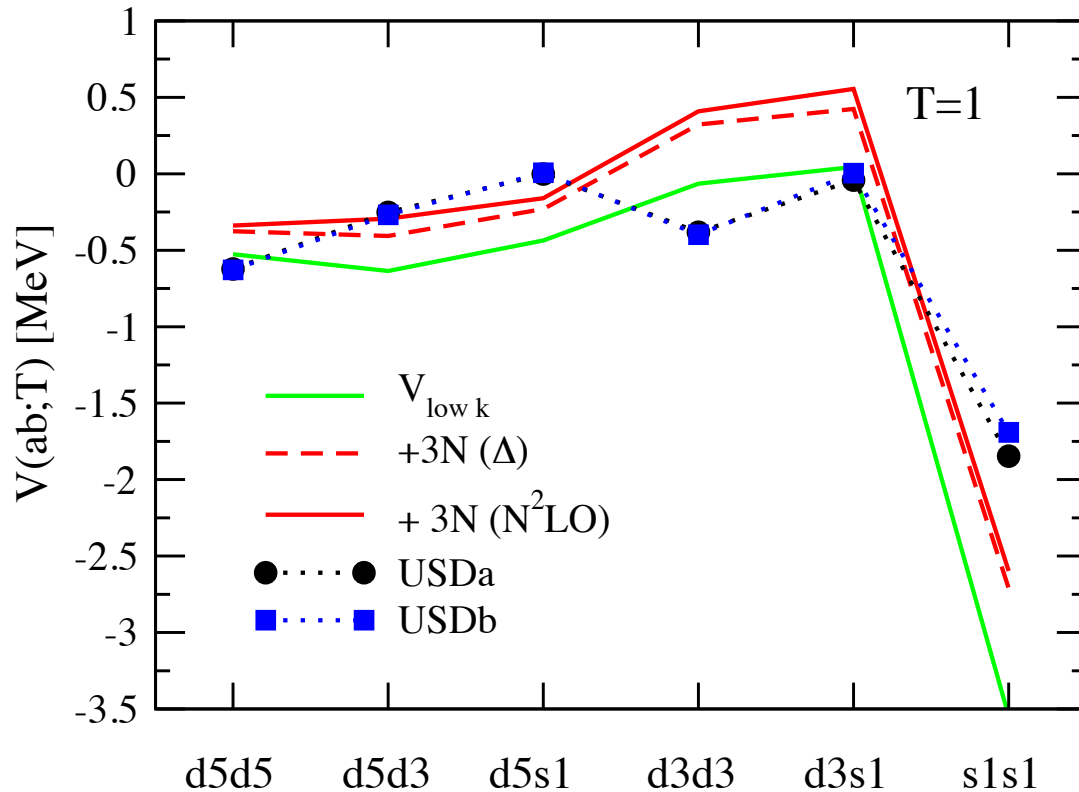
Residual 3N can be neglected

Work on ${}^{16}\text{O}$ in progress

Approximated residual 3N by summing over valence nucleon

– Nucleus-dependent: effect small, not negligible by ${}^{24}\text{O}$

Two-body 3N: Monopoles in *sd*-shell



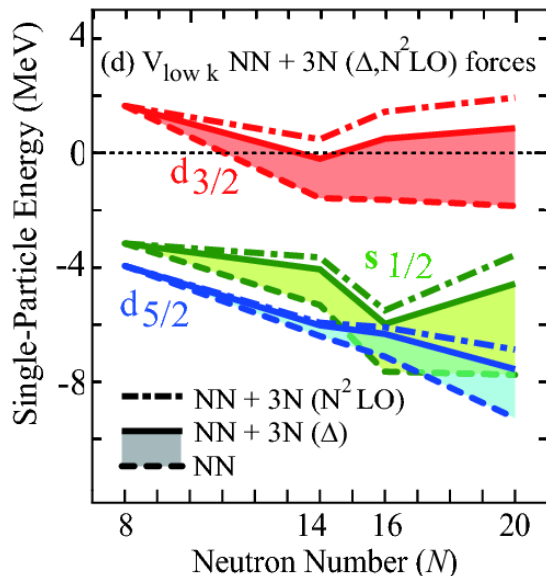
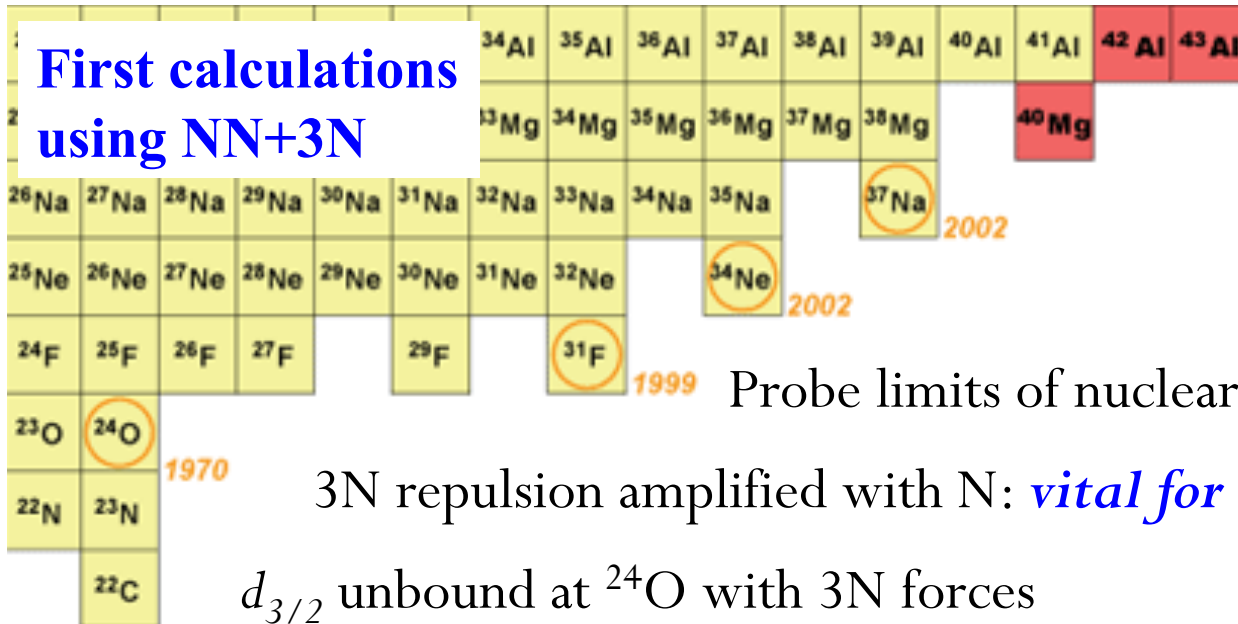
Dominant effect from **one- Δ** – as expected from cutoff variation

3N forces produce clear repulsive shift in monopoles

First calculations to show missing monopole strength due to neglected 3N

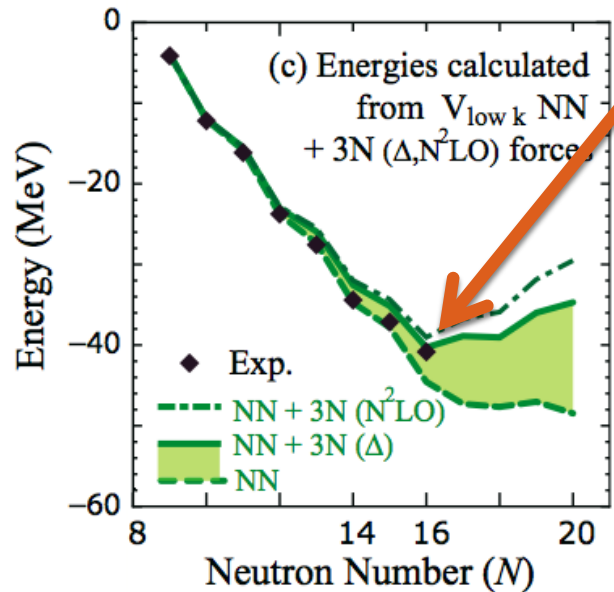
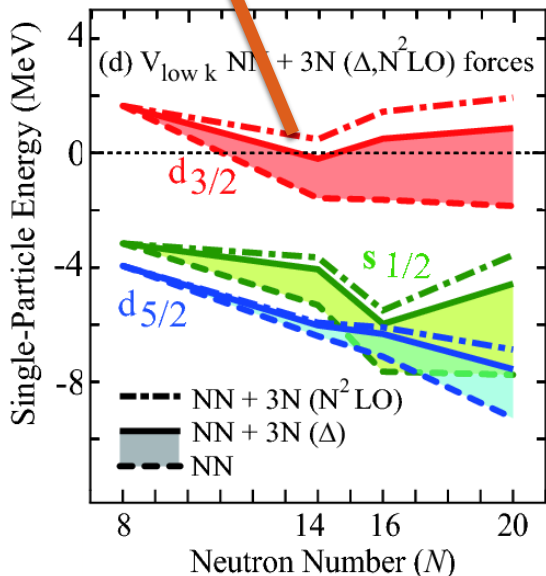
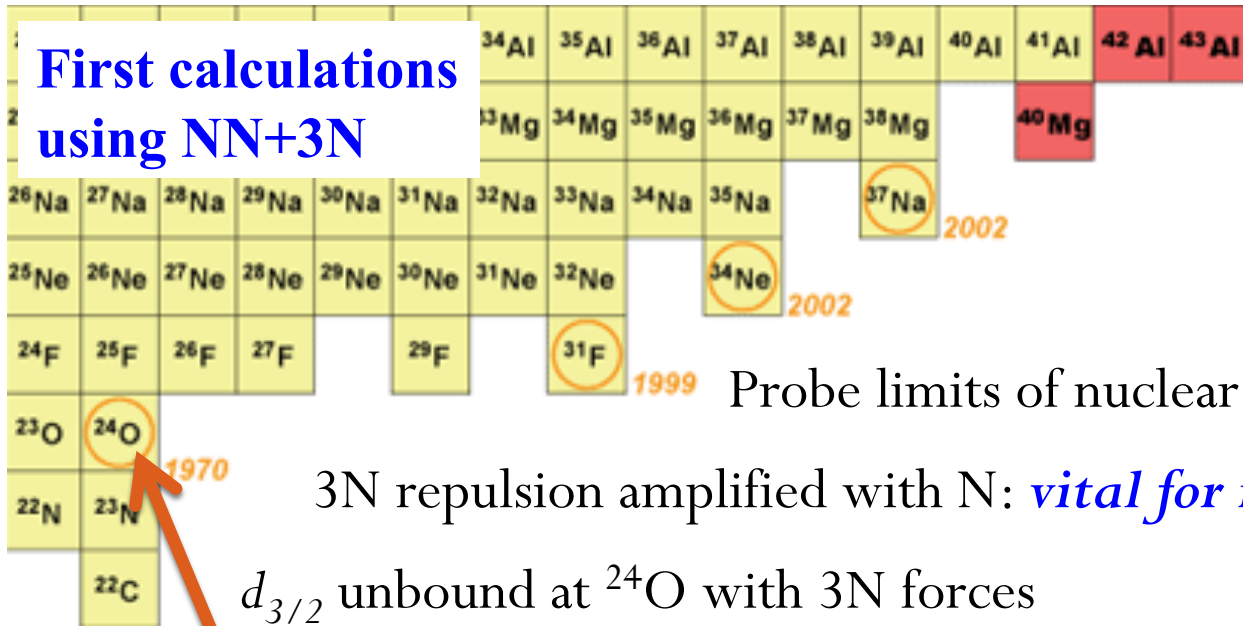
Future: Improved treatment of high-lying orbits

Oxygen Anomaly



Otsuka, Suzuki, Holt, Schwenk, Akaishi, PRL (2010)

Oxygen Anomaly



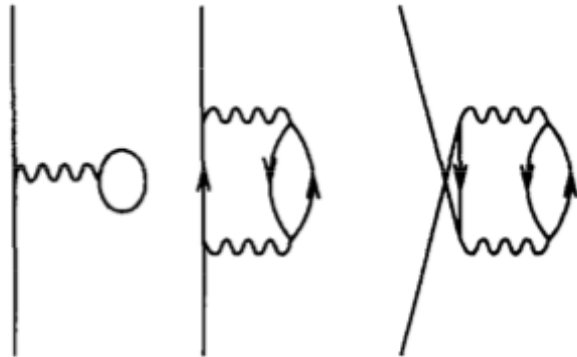
Isotopes unbound beyond ^{24}O

First microscopic explanation of oxygen anomaly

Otsuka, Suzuki, Holt, Schwenk, Akaishi, PRL (2010)

One-Body 3N: Single Particle Energies

NN-only microscopic SPEs yield poor results – rely on empirical adjustments

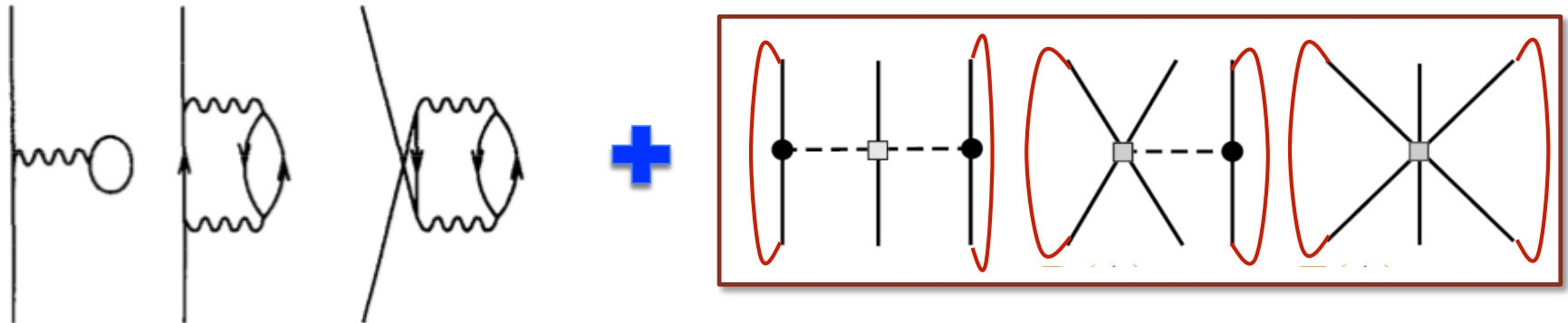


sd-shell: SPEs much too bound, unreasonable splitting

Orbit	“Exp”	USD b	$T+V_{NN}$
$d_{5/2}$	-4.14	-3.93	-5.43
$s_{1/2}$	-3.27	-3.21	-5.32
$d_{3/2}$	0.944	2.11	-0.97

One-Body 3N: Single Particle Energies

NN-only microscopic SPEs yield poor results – rely on empirical adjustments



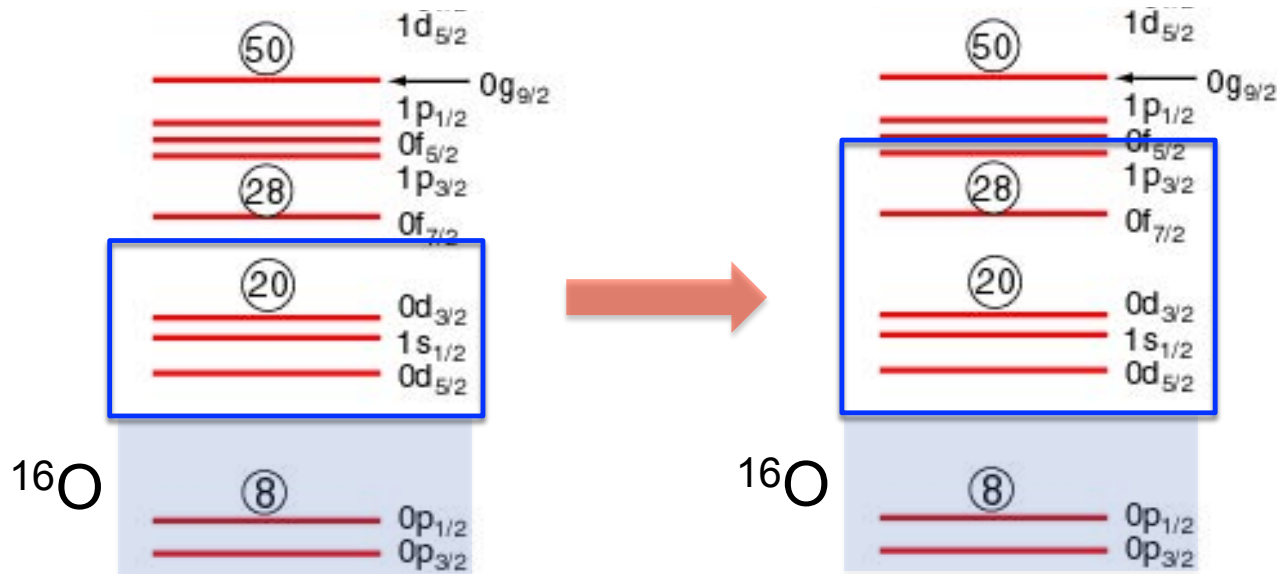
sd-shell: SPEs much too bound, unreasonable splitting

3N forces: additional repulsion – reasonable values!

Orbit	USD b	$T+V_{NN}+V_{3N}$
$d_{5/2}$	-3.93	-3.82
$s_{1/2}$	-3.21	-2.14
$d_{3/2}$	2.11	2.01

One-Body 3N: Single Particle Energies

Effects of correlations beyond one major oscillator shell:



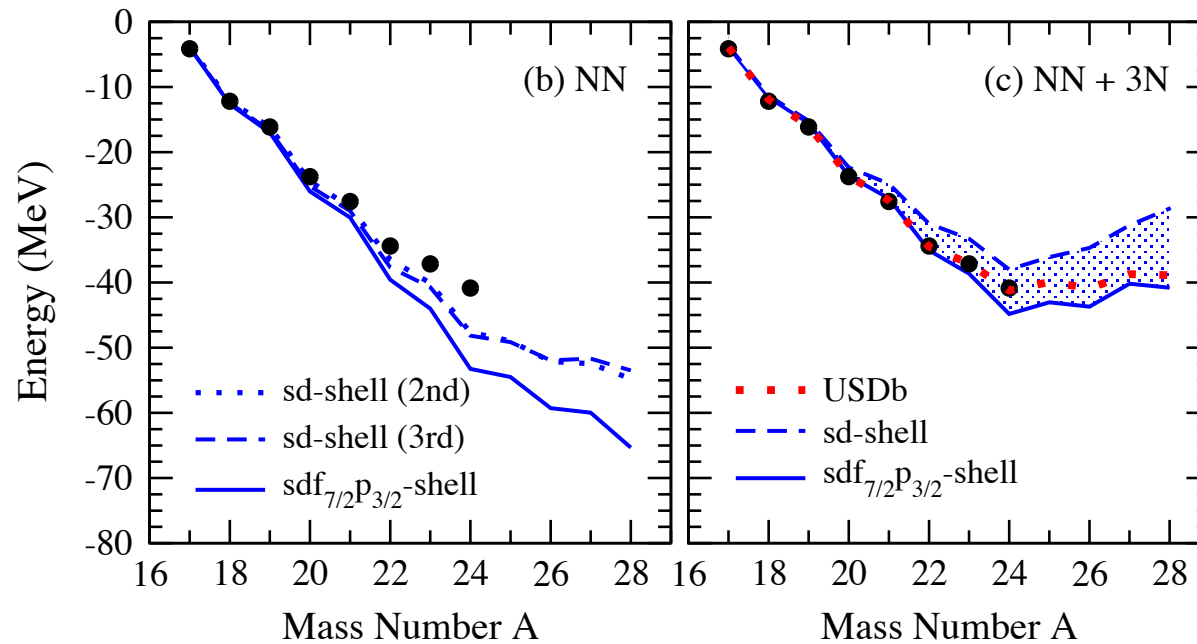
Orbit	USD _b	$T+V_{NN}+V_{3N}$	SDPF-M	$T+V_{NN}+V_{3N}$
$d_{5/2}$	-3.93	-3.82	-3.95	-3.75
$s_{1/2}$	-3.21	-2.14	-3.16	-2.10
$d_{3/2}$	2.11	2.01	1.65	2.13
$f_{7/2}$			3.10	2.96
$p_{3/2}$			3.10	4.82

Fully microscopic framework and extended valence space

Fully-Microscopic Calculations

Interaction and self-consistent SPEs from NN+3N

Empirical SPEs for NN-only



Holt, Schwenk, arXiv:1108.2680

NN-only: dripline at ^{28}O

NN+3N: dripline at ^{24}O

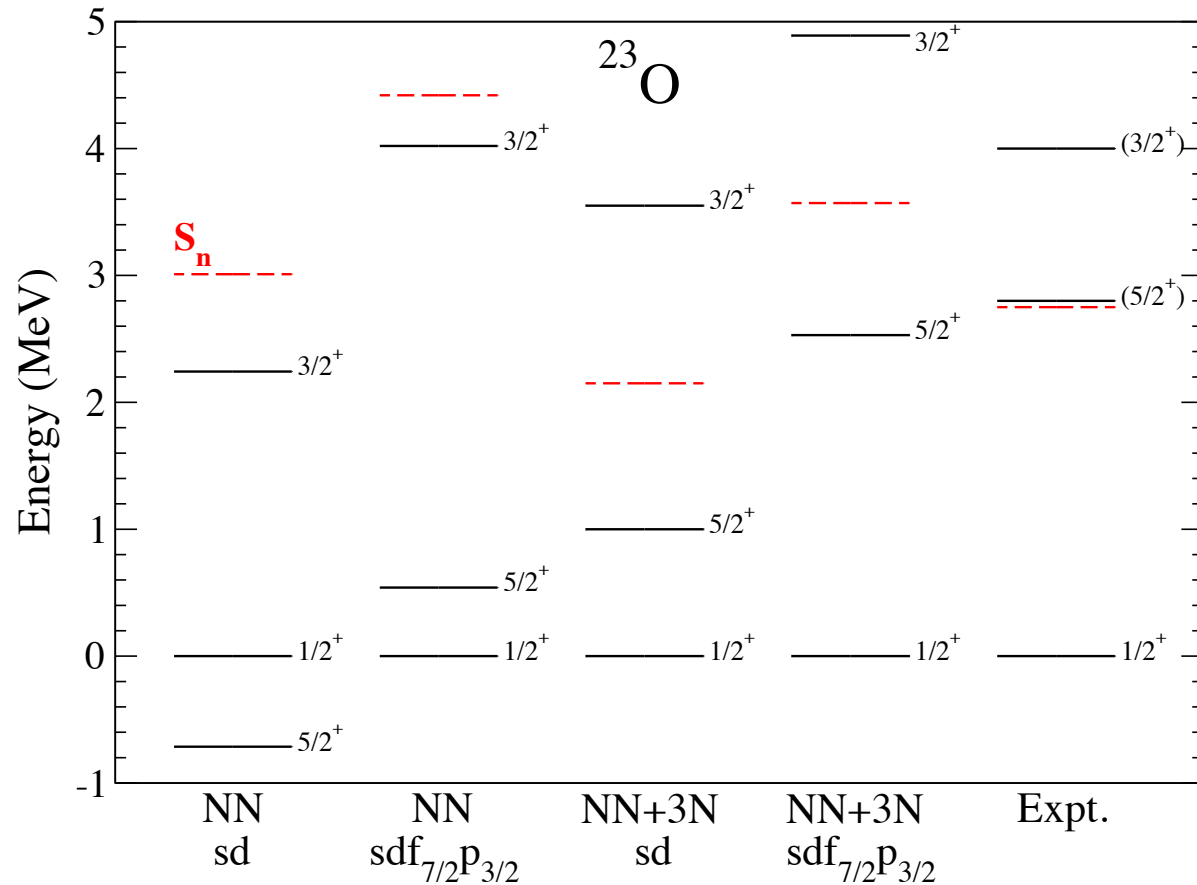
sd -shell results underbound; improved in $sdf_{7/2} p_{3/2}$

Continuum: $\sim 300\text{keV}$ more binding beyond ^{24}O (from CC)

Impact on Spectra: ^{23}O

Neutron-rich oxygen spectra with NN+3N

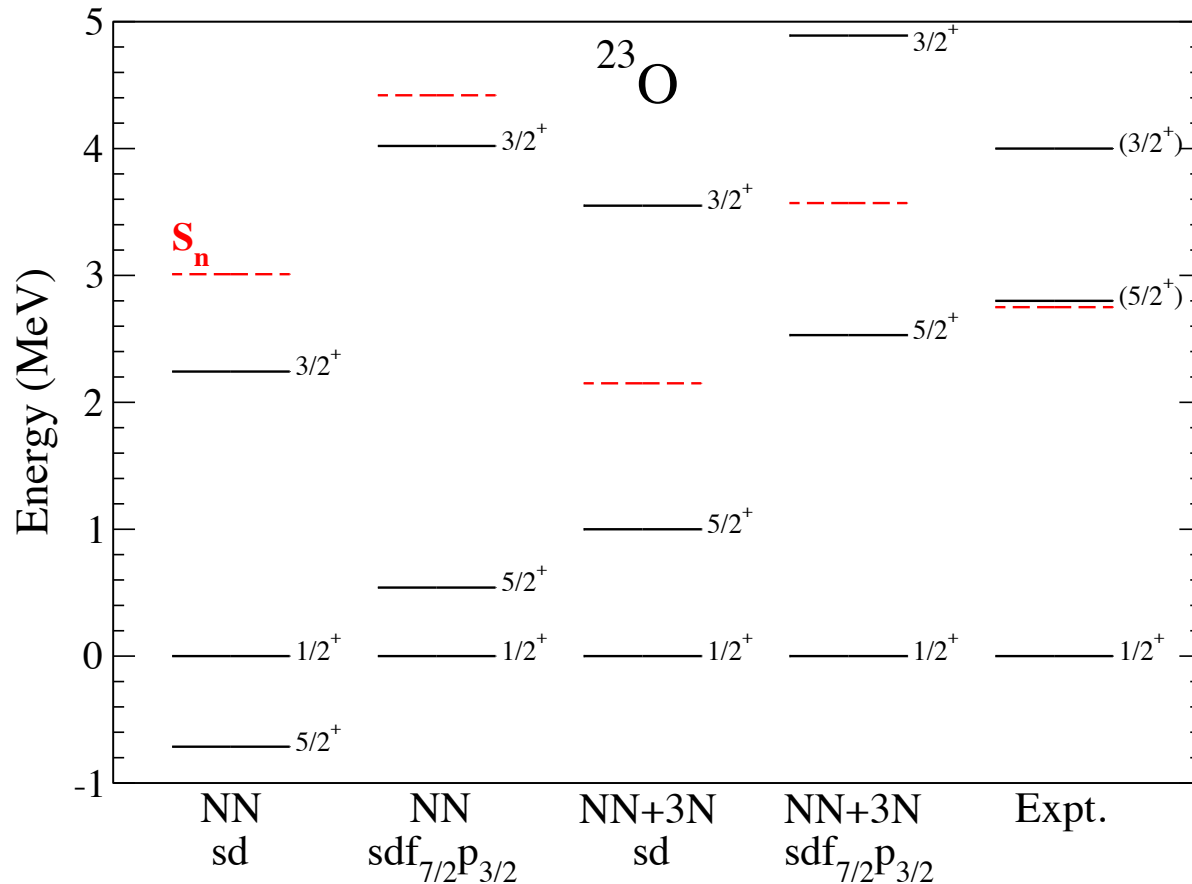
$5/2^+$, $3/2^+$ indicate position of $d_{5/2}$ and $d_{3/2}$ orbits



Impact on Spectra: ^{23}O

Neutron-rich oxygen spectra with NN+3N

$5/2^+$, $3/2^+$ indicate position of $d_{5/2}$ and $d_{3/2}$ orbits



sd-shell NN-only

Wrong ground state!

$5/2^+$ too low

$3/2^+$ bound

Microscopic NN+3N

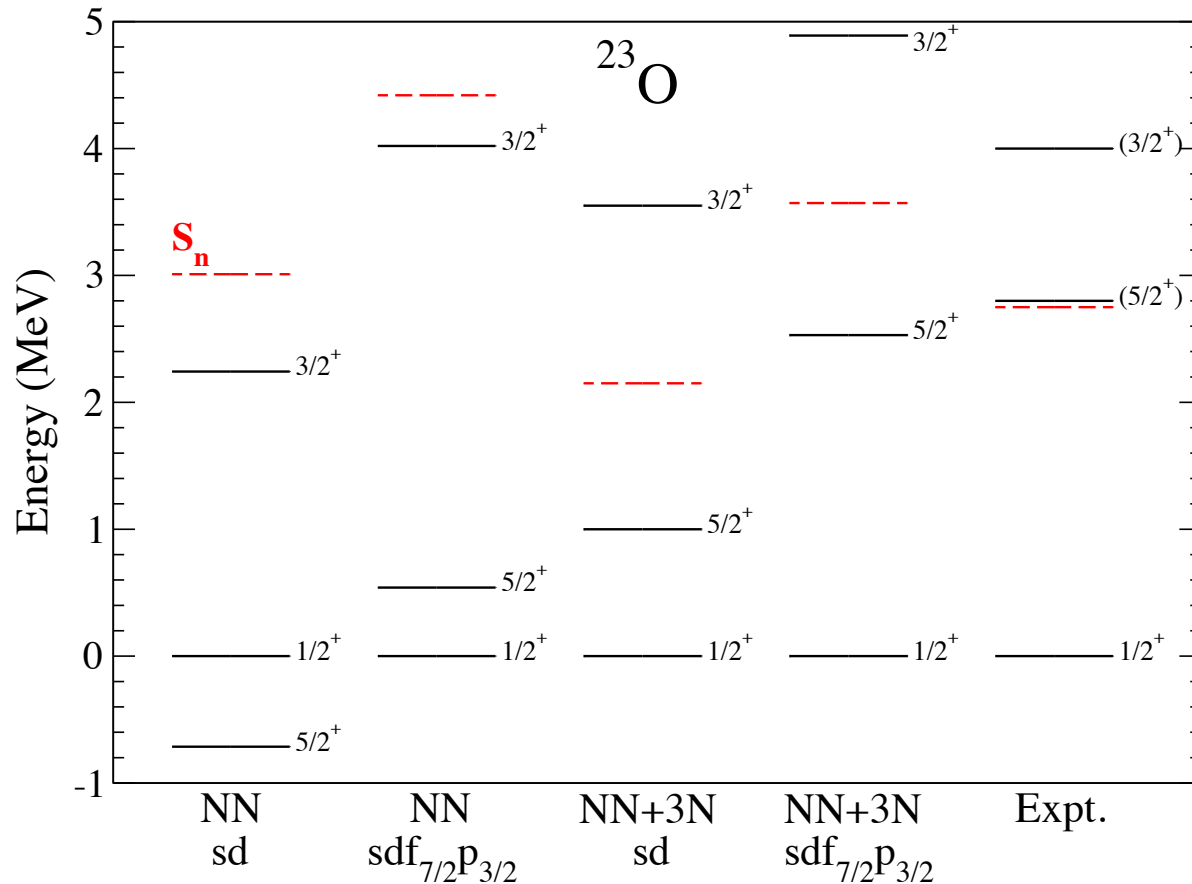
Great improvements in extended valence space!

Holt, Schwenk, arXiv:1108.2680

Impact on Spectra: ^{23}O

Neutron-rich oxygen spectra with NN+3N

$5/2^+$, $3/2^+$ indicate position of $d_{5/2}$ and $d_{3/2}$ orbits



sd-shell NN-only

Wrong ground state!

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Microscopic NN+3N

Great improvements in extended valence space!

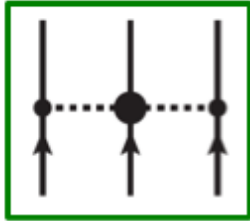
Holt, Schwenk, arXiv:1108.2680

Coupled Cluster spectrum reasonably close to extended space results

Continuum effectively lowers $3/2^+$ - vital for $^{24-28}\text{O}$ Hagen et al., arXiv:1202.2839

In-medium NN interactions

JWH, N. Kaiser, W. Weise, PRC (2009)

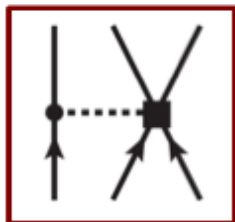
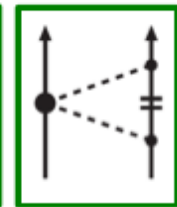
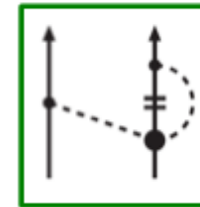
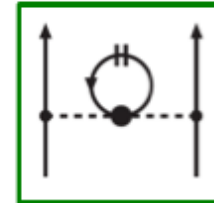


$$V_{3N}^{(2\pi)} = \sum_{i \neq j \neq k} \frac{g_A^2}{8f_\pi^4} \frac{\vec{\sigma}_i \cdot \vec{q}_i \vec{\sigma}_j \cdot \vec{q}_j}{(\vec{q}_i^2 + m_\pi^2)(\vec{q}_j^2 + m_\pi^2)} F_{ijk}^{\alpha\beta} \tau_i^\alpha \tau_j^\beta$$

$$F_{ijk}^{\alpha\beta} = \delta^{\alpha\beta} (-4c_1 m_\pi^2 + 2c_3 \vec{q}_i \cdot \vec{q}_j) + c_4 \epsilon^{\alpha\beta\gamma} \tau_k^\gamma \vec{\sigma}_k \cdot (\vec{q}_i \times \vec{q}_j)$$

$$N^3\text{LO} : c_1 = -0.81, c_3 = -3.2, c_4 = 5.4 \text{ [GeV}^{-1}\text{]}$$

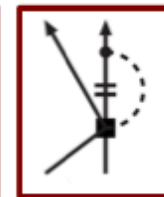
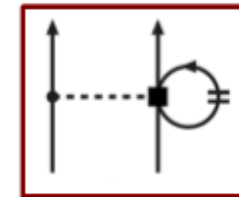
$$V_{\text{low-k}}(2.1) : c_1 = -0.76, c_3 = -4.78, c_4 = 3.96 \text{ [GeV}^{-1}\text{]}$$



$$V_{3N}^{(1\pi)} = - \sum_{i \neq j \neq k} \frac{g_{ACD}}{8f_\pi^4 \Lambda_\chi} \frac{\vec{\sigma}_j \cdot \vec{q}_j}{\vec{q}_j^2 + m_\pi^2} \vec{\sigma}_i \cdot \vec{q}_j \vec{\tau}_i \cdot \vec{\tau}_j$$

$$c_D(N^3\text{LO}) = -0.2$$

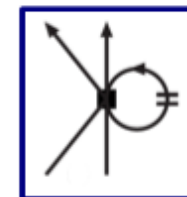
$$c_D(2.1 \text{ fm}^{-1}) = -2.06$$



$$V_{3N}^{(\text{ct})} = \sum_{i \neq j \neq k} \frac{c_E}{2f_\pi^4 \Lambda_\chi} \vec{\tau}_i \cdot \vec{\tau}_j$$

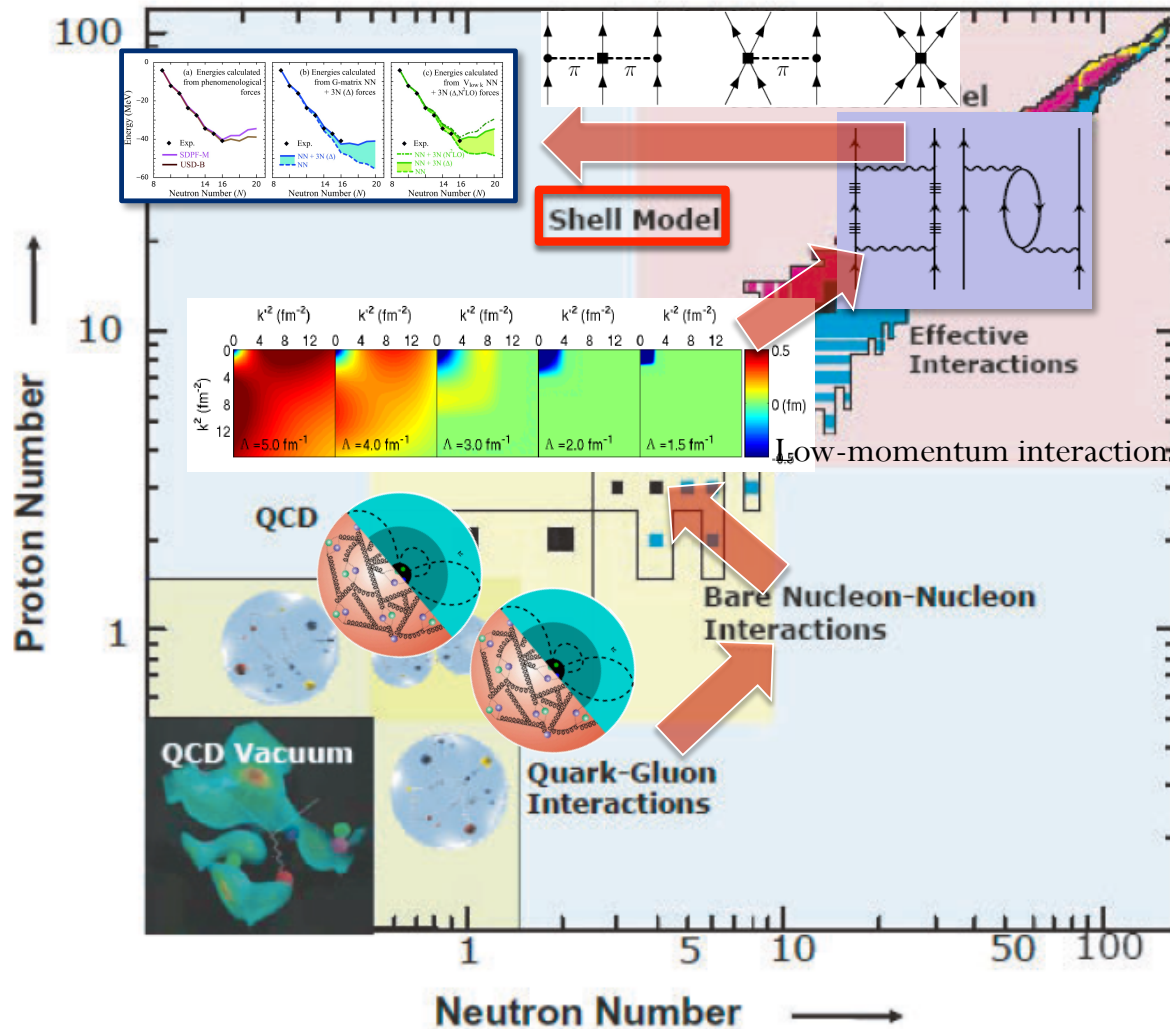
$$c_E(N^3\text{LO}) = -0.205$$

$$c_E(2.1 \text{ fm}^{-1}) = -0.63$$



The Challenge of Microscopic Nuclear Theory

To understand the properties of complex nuclei from elementary interactions



Three-Nucleon Forces

Clear path from symmetries of QCD to shell model
 Ideas of effective field theories
 Renormalization group essential for this progress

Much to do:

How will we approach this problem:

QCD → NN (3N) forces → Renormalize → Solve many-body problem → Predictions

Chiral Effective Field Theory: Philosophy

“Very soft potentials must be excluded because they do not give saturation; they give too much binding and too high density.”

- *H. Bethe*

How might you respond?