# Shell model formalism

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Nuclear structure basics

### 2 Shell model

3 Application to *sd* shell

### Statement of the problem

$$egin{aligned} & \mathcal{H}|\Psi
angle = \mathcal{E}|\Psi
angle \ & \mathcal{H} = \sum_{l_1 l_2} t_{l_1 l_2} c_{l_1}^\dagger c_{l_2} + rac{1}{(2!)^2} \sum_{l_1 l_2 l_3 l_4} ar{v}_{l_1 l_2 l_3 l_4} c_{l_1}^\dagger c_{l_2}^\dagger c_{l_4} c_{l_3} \ & + rac{1}{(3!)^2} \sum_{l_1 l_2 l_3 l_4 l_5 l_6} ar{w}_{l_1 l_2 l_3 l_4 l_5 l_6} c_{l_1}^\dagger c_{l_2}^\dagger c_{l_3}^\dagger c_{l_6} c_{l_5} c_{l_4} + \dots \end{aligned}$$

- Standard eigenvalue problem, but two issues arise for nuclei
- Mesoscopic system
  - 10s-100s of particles
  - Too many permutations to solve computationally
  - System not large enough to treat statistically
- Nuclear Hamiltonian
  - ${\, \bullet \,}$  Nucleons are composite particles (quark/gluon degrees of freedom  ${\, \rightarrow \,} {\sf QCD})$
  - Low-energy nuclear physics is typically insensitive to quark dynamics
  - $\bullet\,$  Chiral effective field theory ( $\chi {\sf EFT})$  provides a low-energy effective approach
  - Nuclear interaction depends on renormalization, is not analytic
  - The nuclear Hamiltonian is scale-dependent
- Exact solution impossible- attempt reasonable approximations

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  - Can determine a two-body interaction from nucleon-nucleon (NN) scattering
  - $\bullet\,$  Can derive a Hamiltonian from  $\chi {\rm EFT}$  based on QCD order by order
  - Coupling constants are parameters, can be fit to experimental scattering data
- Complications
  - Nucleus with mass A has, in principle, A-body Hamiltonian
  - Difficult to implement for structure calculations
  - Even with exact Hamiltonian, calculations are expensive
- Hierarchy in forces ( $NN > NNN > NNN \dots$ )
  - Suggested empirically
  - Confirmed by  $\chi \text{EFT}$
- Limit to three-body forces in nuclear Hamiltonian
- Assume bare microscopic interactions (NN and NNN) are known
  - Underlying approximation to all further results
  - Will not evaluate effect of initial interaction
- Similarities to atomic problem suggest simpler calculational methods

### Nuclear Forces

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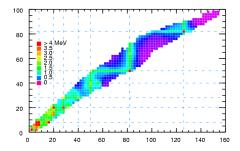
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## Single particle shell structure

- Mean field in the nucleus produced by A nucleons composing it
  - Familiar idea (atoms- low density of electrons and point-like nucleus)
  - Experimental observations: high  $E(2+)^a$ , low B(E2), BE...

 $\rightarrow$  "magic" numbers

- Indicative of single particle shell closures (e.g., group 18 noble gases )
- Collisions within the nucleus are suppressed due to the Pauli principle



# Single particle shell structure

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- Nuclear Hamiltonian:

$$\begin{split} H &= T + V \\ &= \sum_{l_1 l_2} t_{l_1 l_2} c_{l_1}^{\dagger} c_{l_2} + \frac{1}{(2!)^2} \sum_{l_1 l_2 l_3 l_4} \bar{v}_{l_1 l_2 l_3 l_4} c_{l_1}^{\dagger} c_{l_2}^{\dagger} c_{l_4} c_{l_3} + \dots \\ &= [T + V_{mf}] + [V - V_{mf}] \\ &= H_0 + H_1 \end{split}$$

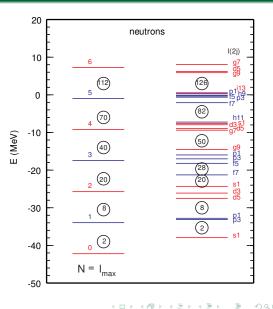
• Analytic solutions to one-body  $H_0 = \sum_i \epsilon_i a_i^{\dagger} a_i$  provide typical single particle bases

• A-body (in principle) residual interaction treated approximately

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# Harmonic oscillator (HO) single particle bases

- Neutron single particle orbits
- Determine  $\hbar\omega$  empirically
- For  $^{132}{\rm Sn},~\hbar\omega\approx 8~{\rm MeV}$



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Nuclear structure basics





### Motivation

- Need to simplify the nuclear many-body problem
- Harmonic oscillator potential (with spin-orbit) reproduces magic numbers
- Results in large energy gaps between bunches of single particle orbits
- Fundamental assumptions
  - Interested in low-energy nuclear properties
  - 2) Strongly bound single particle orbits are rarely excited ightarrow core
  - Properties outside core can be represented by few orbits
- Physical energy scale limited approximately by single particle energy gaps

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## Independent particle model

- Simplest approximation: easy to solve, limited applicability
- Find exact solution to basis  $H_0 |\phi_i\rangle = \epsilon_i |\phi_i\rangle$
- Core of mass  $A_c$  with  $Z_c$  protons and  $N_c$  neutrons ( $Z_c$ ,  $N_c$  closed subshells)
- Single Slater determinant describes wavefunctions

$$|\Phi^{A_c}\rangle \equiv |\phi_1\phi_2\dots\phi_{Z_c}\rangle \otimes |\phi_1\phi_2\dots\phi_{N_c}\rangle$$

Energy of core

$$E_c = \sum_{i=1}^{Z_c} \epsilon_i + \sum_{i=1}^{N_c} \epsilon_i$$

- Nuclei with  $A_c + 1$  are given by  $|\Phi^{A+1}\rangle = a_i^{\dagger} |\Phi^A\rangle$  with  $i = Z_c + 1, N_c + 1$
- Energy relative to core

$$E_{(c+p)} - E_c = \epsilon_{(Z_c+1)} \qquad \qquad E_{(c+n)} - E_c = \epsilon_{(N_c+1)}$$

- Slater determinant is tailored to reproduce experimental data
  - Correlations implicitly included ightarrow experiments measure many-body system
  - Dominant states are reproduced, but fragmentation occurs
  - Distinct from other Slater determinant methods like EDF methods

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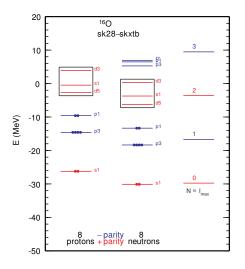
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# <sup>16</sup>O in independent particle model



- Single Slater determinant
- $BE(^{17}O) BE(^{16}O) = 4.14$  MeV
- $BE(^{18}O) BE(^{16}O) = 12.19 \text{ MeV}$
- Too simple model already!

# Connection to Hartree-Fock (HF)

Schrödinger equation  $H|\Psi
angle=E|\Psi
angle$ 

- Independent particle model:
  - Approximate H by  $H_0 = T + V_{mf}$
  - Select reasonable V<sub>mf</sub>
  - Solve  $H_0 |\phi_i\rangle = \epsilon_i |\phi_i\rangle$ •  $E_0 = \sum_{i=1}^{A} \epsilon_i$  for A nucleons
  - No correlations
  - Only appropriate for closed shells  $\pm \ 1$

• (symmetry-restricted) Hartree-Fock:

• Approximate 
$$|\Psi
angle$$
 by  $|\Phi
angle = \prod_{i=1}^{A} a_{i}^{\dagger} |0
angle$ 

• Variational principle (minimizes energy)

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- Determine single particle basis  $\{|\phi\rangle\}$
- Limited correlations
- Best near closed shells

- Independent particle model is not rich enough
- Single-reference Hartree-Fock is better but suffers from same limitations
- Need to include correlations without solving full many-body problem
- Select core with A<sub>c</sub> as before (closed subshells)
  - Treat core as vacuum based on large energy gap in single particle orbits
  - Fewer nucleons treated explicitly (valence particles  $A_{val} = A A_c$ )
- Limit to "valence orbits" up to another large energy gap
- Reduction in model space (and number of particles)
  - Schrödinger equation can be solved completely
  - Separation into long-range and short-range correlations
  - Long-range correlations are completely included
- Can other effects be included?
  - Polarization of the core by valence particles
  - ② Virtual scattering of valence particles into higher-lying orbits
  - In Full short-range correlations

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### Effective interaction

- Incorporate effects from outside of the model space into the Hamiltonian
- Still, cannot account for everything (e.g., short-range correlations)
- Reduction in degrees of freedom = reduction in possibilities
- Production of interactions is extremely important
  - Multiple lectures will focus on various procedures and mindsets
  - Tutorial sessions devoted to derivation and implementation
  - Will proceed currently with a general Hamiltonian

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## Effective interaction

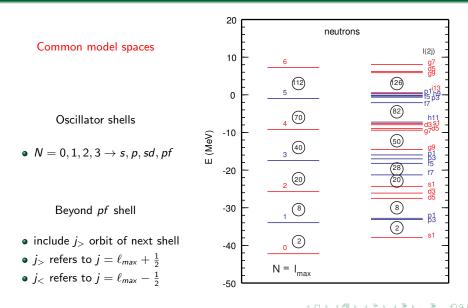
- Representation of effective interaction in reduced model space
  - **(**) Single particle energies (SPE):  $\epsilon_i$  where *i* refers to valence orbits
  - Two-body matrix elements (TBME):

$$\langle (ab)_{JT} | V_{ms} | (cd)_{JT} \rangle$$

- Model space orbits a, b, c, d
- Angular momentum and isospin J and T, respectively
- $V_{ms}$  is the effective interaction in the reduced model space
- Finite number of TBME for a given model space determined by J and T coupling
- $V_{ms}$  distinct from original Hamiltonian
- Lowest order (monopole) of multipole expansion of the interaction

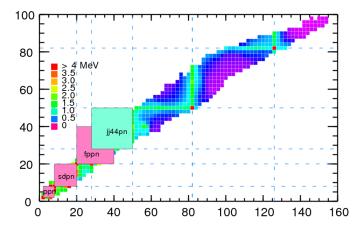
$$\bar{V}_{ab}^{T} = \frac{\sum_{J} (2J+1) \langle (ab)_{JT} | V_{ms} | (ab)_{JT} \rangle [1-(-1)^{J+T} \delta_{ab}]}{\sum_{J} (2J+1) [1-(-1)^{J+T} \delta_{ab}]}$$

## Model Spaces



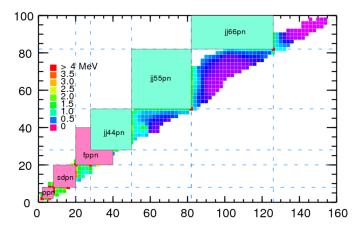
## Model Spaces

• Most common model spaces (courtesy of Alex Brown)



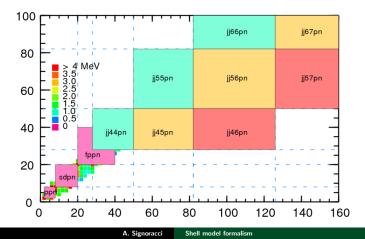
## Model Spaces

• Inclusion of  $j_{>}$  orbit (courtesy of Alex Brown)



### Model Spaces

- Extension to heavier nuclei with N > Z (courtesy of Alex Brown)
- Based on stable magic numbers  $\rightarrow$  island of inversion region?



### Summary

### • Cannot solve full Schrödinger equation beyond lightest nuclei

### • Search for approximate techniques

- Because nuclei display shell structure via "magic" numbers
  - Break Hamiltonian into mean field and residual interaction
  - Solve mean field Hamiltonian for single particle basis
  - Utilize large energy gaps in basis to isolate few valence orbits
  - Treat core of model space as vacuum
  - Solve Schrödinger equation exactly in reduced model space
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- Falls under more general category of configuration interaction (CI) theory

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### Outline

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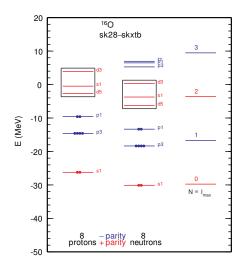
### 2 Shell model

3 Application to *sd* shell

### Opening remarks

- Hamiltonian operating only in the reduced model space is required
- $\bullet\,$  For now, given empirically by USDB interaction ^1
- SPE and TBME parameterized (66 parameters in all)
- Fit to experimental energies in *sd* shell
  - Iterative procedure of CI calculations for states accessible in the model space
  - See Lecture VII for details
- Use sd shell as example, same principles apply throughout nuclear chart

# <sup>16</sup>O as core

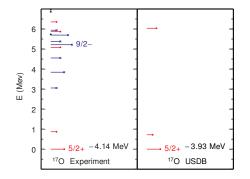


- sd model space outlined
- Separation from *p* and *pf* shells
- N = 0, 1 oscillator shells filled

• 
$$|^{16}\mathrm{O} \rangle \equiv |0
angle$$

• 
$$E(^{16}O) = 0$$

### Simplest system- <sup>17</sup>O



- One valence neutron
- Three orbits to occupy
- <sup>(12)</sup>
   <sub>1</sub> = 3 possibilities to add nucleons
- Three possible states<sup>a</sup>

<sup>a</sup>See Lecture II for counting procedure

- Experiment displays richer behavior
  - Intruder states (only positive-parity many-body states from sd orbits)
  - Is Fragmentation of single particle strength due to correlations
- Calculated  $BE_{USDB}(^{17}O) = 3.93$  MeV, whereas  $BE_{exp}(^{17}O) = 131.76$  MeV
- USDB selects  $^{16}$ O as vacuum with E = 0 MeV

• 
$$BE_{exp}(^{17}O) - BE_{exp}(^{16}O) = 4.14 \text{ MeV}$$

### Spectroscopic factors

- Basis-independent, but not observable
- Spectroscopic probability matrices

$$\mathcal{S}^{+
ho q}_{\mu}\equiv \langle \Psi^A_0|a_{
ho}|\Psi^{A+1}_{\mu}
angle \langle \Psi^{A+1}_{\mu}|a^{\dagger}_q|\Psi^A_0
angle$$

and

$$\mathcal{S}_{
u}^{-
ho q}\equiv \langle \Psi_{0}^{A}|a_{q}^{\dagger}|\Psi_{
u}^{A-1}
angle \langle \Psi_{
u}^{A-1}|a_{
ho}|\Psi_{0}^{A}
angle$$

- Spectroscopic factors (SF) found from tracing spectroscopic probability matrices
- In reduced model space, recover typical "definitions"

$$SF^+_\mu \equiv |\langle \Psi^{A+1}_\mu | a^\dagger_q | \Psi^A_0 
angle|^2$$

and

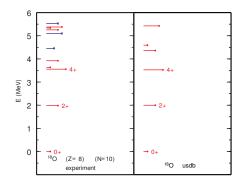
$$SF_{
u}^{-}\equiv |\langle \Psi_{
u}^{A-1}|a_{
ho}|\Psi_{0}^{A}
angle |^{2}$$

### Spectroscopic factors

- For  ${}^{17}\text{O}$  in *sd* shell, each state carries full spectroscopic strength
  - Only one valence particle
- "Experimental" SF
  - SF = 0.81 for  $\frac{5}{2}^+$  ground state of <sup>17</sup>O <sup>a</sup>
  - SF = 0.67 for  $\frac{3}{2}^+$  "single particle peak" at 5.09 MeV<sup>b</sup>
  - SF = 0.06 for  $\frac{3}{2}^+$  state at 5.87 MeV<sup>c</sup>
- Problematic for calculations? Is <sup>16</sup>O a good enough core?

<sup>a</sup>J. Lee, M.B. Tsang, and W.G. Lynch, Phys. Rev. C **75**, 064320 (2007)
 <sup>b</sup>M. Yasue et al., Phys. Rev. C **46**, 1242 (1992)
 <sup>c</sup>M. Yasue et al., Phys. Rev. C **46**, 1242 (1992)

### Next simplest system- <sup>18</sup>O

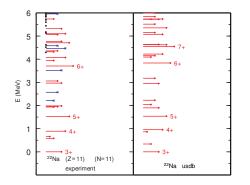


- Two valence neutrons
- Three orbits to occupy
- $\binom{12}{2} = 66$  possibilites to add nucleons
- Only 14 possible states<sup>a</sup>

<sup>a</sup>See Lecture II for counting procedure

- Experiment displays richer behavior, including negative-parity intruder states
- Calculated  $BE_{USDB}(^{18}O) = 11.93$  MeV, whereas  $BE_{exp}(^{18}O) = 139.81$  MeV
- $BE_{exp}(^{18}\text{O}) BE_{exp}(^{16}\text{O}) = 12.19 \text{ MeV}$

## More complicated-<sup>22</sup>Na



- Odd-odd nucleus
- Three neutrons, three protons
- $\binom{12}{3}^2 = 48400$  possibilities
- 3266 possible states<sup>a</sup>

<sup>a</sup>See Lecture II for counting procedure

- USDB interaction is isospin-symmetric (no Coulomb force!)
- $E_{USDB}(^{22}Na) = -58.44 \text{ MeV}$
- $E_{USDB}(^{22}Na) = -46.71 \text{ MeV}$  (Coulomb corrected)
- $BE_{exp}(^{22}Na) BE_{exp}(^{16}O) = 46.53 \text{ MeV}$

- All sd nuclei have such level schemes available<sup>2</sup>
- $\bullet\,$  Overall, root-mean-square (rms) deviation of  $\approx 170$  keV to low-energy states
- Only considering states accessible in the model space
- USDB has been used in hundreds of calculations
- Only energies thus far, but good agreement for other nuclear properties
  - To be discussed in more detail later
- Slater determinant of N = 0, 1 HO orbits not accurate description of <sup>16</sup>O
  - Seen from Hartree-Fock or realistic calculation (e.g., coupled cluster)
  - Correlations contribute multiple MeV to ground state
  - Excited states exist- experimentally  $E(0_2^+) = 6.05$  MeV
  - Single particle strength fragmented in <sup>17</sup>C
- Still reproduce results in *sd* shell well!
- Long-range correlations cause low-energy behavior in sd shell nuclei

 $^{2} http://www.nscl.msu.edu/{\sim}brown/resources/resources.html$ 

T. Duguet and G. Hagen, Phys. Rev. C 85, 034330 (2012

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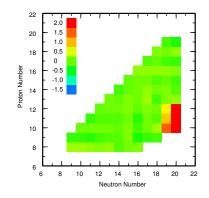
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Shell model

#### Problems remain

• Even with correct interaction, can fail to reproduce low-energy states



 $BE_{exp}(Z, N) - BE_{th}(Z, N)$ 

• Wrong degrees of freedom (missing necessary valence orbits)

- Invalidates assumption that shell gap excludes pf orbits
- Referred to as the island of inversion region