Linear momentum restored calculations with the Brink-Boeker interaction.

(Restoration of Galilei invariance in the nuclear many-body problem).

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## Introduction

We consider the nucleus as a closed system of interacting and non relativistic nucleons.

Homogenity of space  $\rightarrow \rightarrow$  conservation of total linear momentum of the system (the nuclear hamiltonian can depend only on relative coordinates and momenta).

a) pependence on total linear momentum is trivial.

One should solve the Schrödinger equation for the internal hamiltonian.

Few-body physics $\rightarrow$  Jacobi coordinates $\rightarrow$  nucleons are fermions $\rightarrow$  explicit antisymmetrization $\rightarrow$  very difficult with increasing number of nucleons.

Many-body physics  $\rightarrow$  expansion of w.f in terms of HF (or generalized HF) determinants  $\rightarrow$  Pauli

principle is fullfilled but there is spontaneous symmetry breaking.  $\rightarrow$  spuriuos

admixtures  $\rightarrow$  **broken Galilei invariance** 

Problem recognized inmediately after the development of the shell model (Elliot and Skyrme, Proc. Roy. Soc. A **232** (1955) 561)

For pure HO configurations and complete  $n\hbar\omega$  spaces there is a solution (Giraud, NPA **71** (1965) 373)



#### Advantages

- a) It works in general model spaces and well as for general (non oscillator) wave functions.
- b) It allows to recover translational (PAV) and even full Galilei (PBV) invariances.

#### Main technical difficulties

a) Any of the bases usually used in nuclear structure is closed under the action of the shift operator  $\rightarrow$  link with states outside of the original single-particle basis  $\rightarrow$  extended Wick's theorems are required.

b) No inert core can be used.

### Linear momentum is a true A-body correlation

Galilei invariance provides important effects on top of the usually assumed 1/A dependence

K.W.Schmid and F. Grümmer, Z. Phys. A 336 (1990) 5; A 337 (1990) 267.

K.W.Schmid and P.-G.Reinhard, NPA 530 (1991) 283.

K.W.Schmid, Eur. Phys. J. A 16 (2003) 475, A 12 (2001) 29; A 13 (2002) 319; A 14 (2002) 413.

### **COM projection before the variation(COM-PBV)**

We used the Brink-Boeker interaction B1 (NPA A 91 (1996) 1)

$$\hat{V}(1,2) = \sum_{i=1,2} e^{-\frac{(\vec{r}_1 - \vec{r}_2)^2}{\mu_i^2}} \left( W_i + B_i \hat{P}^{\sigma} - H_i \hat{P}^{\tau} - M_i \hat{P}^{\sigma} \hat{P}^{\tau} \right)$$

complemented with a short finite range spin-orbit interaction (derived from Gogny-D1S) plus Coulomb.

The projection operator has the Thouless's form (we go to the rest frame !!!)

$$\hat{C}(0) = \int d^3 \vec{a} S(\vec{a}) = \int d^3 \vec{a} e^{i \vec{a} \vec{P}}$$
<sup>(2)</sup>

(1)

$$\hat{H} = \sum_{ir} t_{ir} \hat{c}_i^{\dagger} \hat{c}_i + \frac{1}{4} \sum_{ikrs} \overline{v}_{ikrs} \hat{c}_i^{\dagger} \hat{c}_k^{\dagger} \hat{c}_s \hat{c}_r$$
(3)

**Details** R.Rodriguez-Guzman and K.W.Schmid; Eur. Phys. J. A **19** (2004) 45; A **19** (2004) 61.

$$|D\rangle = \prod_{h=1}^{A} \hat{b}_{h}^{\dagger} |0\rangle \tag{4}$$

$$\hat{b}_{\beta}^{\dagger} = \sum_{i=1}^{M_b} D_{i\beta}^* \hat{c}_i^{\dagger}$$
<sup>(5)</sup>

Trial wave function for symmetry restoration (PBV into the COM rest frame)

$$|D;0\rangle = \frac{\hat{C}(0)|D\rangle}{\langle D|\hat{C}(0)|D\rangle} \tag{6}$$

The variational principle is applied to the projected energy  $\delta E_{pr}=0$ 

$$E_{pr} = \frac{\langle D | \hat{H} \hat{C}(0) | D \rangle}{\langle D | \hat{C}(0) | D \rangle}$$
(7)

We perform COM-PBV for the nuclei  ${}^{4}$ He,  ${}^{12}$ C,  ${}^{16}$ O,  ${}^{28}$ Si,  ${}^{32}$ S and  ${}^{40}$ Ca.









Usual hole-spectroscopic factors

$$S_h^{nor} = \sum_{\sigma} \int d^3 \vec{k} \, |f_{h\tau\sigma}^{nor}(\vec{k}\,)|^2 = \delta_{\tau\tau_h}.$$
(11)

**Galilei invariant hole-spectroscopic factors**  $\rightarrow$  **Details** EPJ A **19** (2004) 61.

$$S_{\tilde{h}}^{proj} \equiv \sum_{\sigma} \int d^3\vec{k} \, |f_{\tilde{h}\tau\sigma}^{proj}(\vec{k}\,)|^2 = \delta_{\tau\tau_h} \int_{0}^{\infty} dk \, k^2 \, g_{\tau_h \tilde{h} \, l_h j_h}^{proj}(k)^2 \tag{12}$$

sum rule violation for Galilei invariant hole-spectroscopic factors

$$\sum_{\tilde{h}} S_{\tilde{h}}^{proj} = A - \epsilon \tag{13}$$

 $\epsilon/A$  varies only between 0.12 and 0.35 percent for the considered cases.



**Density in momentum representation** 

$$\hat{\rho}_n \equiv \sum_{\tau} f_\tau(Q^2) \sum_{i=1}^{N_\tau} \exp\{i\vec{q}\cdot\vec{r}_i\}$$
(14)

Normal elastic charge form factor

$$F_{ch}^n(Q^2) = \langle D|\hat{\rho}_n|D\rangle \tag{15}$$

Density in momentum representation+Gartenhaus-Schwartz operator

$$\hat{\rho}_{inv} \equiv \hat{\rho}_n \exp\{-i\vec{q} \cdot \vec{R}\}$$
(16)

Galilei invariant charge form factor

$$F_{ch}^{pr}(Q^2) = \frac{\langle D|\hat{\rho}_{inv}\hat{C}(0)|D\rangle}{\langle D|\hat{C}(0)|D\rangle}$$
(17)

$$F_{ch}^{dy}(Q^2) = F_{ch}^n(Q^2) \exp\left\{\frac{3}{8} \frac{\vec{q}^2}{\langle D | \hat{P}^2 | D \rangle}\right\}$$
(18)









# Conclusions







3) In the long range this symmetry restoration should also be included in shell model, QMC diagonalization method as well as in the VAMPIR family.



This symmetry restoration is also considered at the moment in electronic structure studies



**4-a)** 1D repulsive Hubbard model (3D spin projection+ 1D linear momentum projection)  $\rightarrow$ 

K.W.Schmid et al., Phys. Rev. B 72 (2005) 085116



**4-b)** 2D repulsive Hubbard model (3D spin projection+ 2D linear momentum projection)  $\rightarrow$ 

R.Rodriguez-Guzman et al., in progress



A lot remains to be done !!!!!!)

